

**Arc maps for the He, Ne, Hg, ThAr, and Cd
lamps when used with WYFFOS/Red+4**

Version 1.0

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Introduction

The following document is a compilation of the arc lamp maps for AF2+WYFFOS, the multi-object, wide-field, fiber spectrograph at the Prime focus of the 4.2m William Herschel Telescope (WHT).

The document contains a table with the information of the representative setups. The data is ordered from lowest to highest resolution and from the bluest to the reddest wavelength. After the table the document continues with the arc maps for all setups in the order of the table. Appendix 1 contains the list with all lines labeled in the plots. In Appendix 2 and 3 the source code of the two scripts used for creating the plots: First the *idl* script to create the plots and second the *python* script, which processes the output of the WYFFOS reduction pipeline¹ for the *idl* script. Additionally, the latter one creates the tables shown in this document.

The data was obtained in May, October, and November 2015 with the new CCD (Red+4) after the upgrades in the spectrograph, intended to enhance the throughput at blue spectra ranges. For more information about the Red+4 CCD check the AF2 webpage². For all setups a CCD binning of 2x2 was used. The reduction and wavelength calibration was done with the instrument pipeline version 3.

The entries in Table 1 show the information of all used setups. The first three columns give the grating (1), the central wavelength (2) to which WYFFOS was set, and the wavelength range (3). Column (4) gives the dispersion of the grating for the used wavelength. The dispersion values were either copied directly from the AF2+WYFFOS webpage³ or calculated from the values available on the webpage. The following two columns show the used arc lamp (5) and exposure time (6). Column (7) gives the number of the usable lines. The first number shows all non-blended lines which are visible in the subplots and have enough signal-to-noise to be used for wavelength calibration. The second number shows only these lines which have at least 10% of the flux of the brightest arc-line. If a neutral density filter was used, the value is given in the last column (8).

In the atlas plots the median flux of all fibers in [e^-/px] (pixel-step-in-wavelength) against the wavelength in [\AA] is plotted. As each fiber has a different throughput the fibers were scaled and a 2σ clipping was applied before combining them. To determine the scaling factor, the scaling ratio against the median flux for each pixel was determined. These factors were combined by using the weighted average with the median flux as weights. In the plots the flux is normalized to one second exposure time and collapsed to one pixel. λ_{cen} gives the central wavelength as set up in WYFFOS. The plotted wavelength range is the same as in Table 1. Each of the four subplots shows approximately one quarter of the wavelength range of the top plot.

The gratings used in WYFFOS are the same as used in ISIS⁴, except for the H1800V grating, which was designed for IDS⁵. The H1800V grating is smaller in size than WYFFOS/ISIS gratings, hence some vignetting in the spectral range occurs.

NOTES:

- The Hg and Cd lamps show strong warming effects. During the first minutes the relative intensity of the lines may change drastically, after ~ 2 to ~ 5 min it will have stabilized. The spectra in this document were obtained at least 2 to 3 min after the Hg or Cd lamp was switched on but still for some spectra the lamps might not have been stabilized.
- The Hg and Cd lamps contain some Ar gas, hence their arc map contains Ar lines together with the Cd or Hg.

¹<http://www.ing.iac.es/astrometry/instruments/af2/reduction.html>

²<http://www.ing.iac.es/astrometry/instruments/af2/detector.html>

³<http://www.ing.iac.es/astrometry/instruments/af2>

⁴<http://www.ing.iac.es/astrometry/instruments/isis/>

⁵<http://www.ing.iac.es/astrometry/instruments/ids/>

- The plots without sufficient arc lines are plotted for homogeneity and to assist the decision for optimum arc lamp. See ⁶ for further notice.

Detailed lamp properties

He lamp from SpectroLamps

http://www.spectrolamps.com.au/spectral_lamps.html

The Helium lamp is a stable, high quality source of discrete spectral lines. It consists of a simple borosilicate discharge tube. It uses a high-current electrode and a thin coil of tungsten wire is wound around a pellet of thoria which imparts excellent electron-emissive properties to this electrode, minimising its operating temperature and allowing it to handle a higher current density.

Ne lamp from the Oriel brand, now sold by Newport

http://search.newport.com/?q=*%&x2=sku&q2=6032

Model 6032

The 6032 spectral calibration lamp produces narrow, intense lines from the excitation of Neon gas. It is a pencil lamp, called like this because of its size and shape. It is made of double bore quartz tubing with two electrodes at one end sealed into a phenolic handle.

Hg(Ar) lamp from the Oriel brand, now sold by Newport

http://search.newport.com/?q=*%&x2=sku&q2=6035

Model 6035

The 6035 spectral calibration lamp produces narrow, intense lines from the excitation of Argon gas and Mercury vapor. The 6035 Hg(Ar) Lamp is insensitive to temperature. It requires a two-minute warm-up for the mercury vapor to dominate the discharge, then 30 minutes for complete stabilization. The average intensity is remarkably constant and reproducible after the thermal conditions stabilize. It is a pencil lamp, called like this because of its size and shape. It is made of double bore quartz tubing with two electrodes at one end sealed into a phenolic handle.

ThAr lamp from Heraeus

http://www.heraeus-noblelight.com/en/productsandsolutions/opticalanalysis/hollow_cathode_lamps.aspx

The Thorium-Argon lamp is a hollow cathode lamp. It consists of a cathode made from the element of interest, Thorium in this case, an anode and an inert filler gas contained in a glass envelope, e.g. Argon. In addition various mica discs, ceramic sheaths and glass shields assist in alignment and insulation.

Cd lamp from UVP

<http://uvp.com/zinccadmium.html>

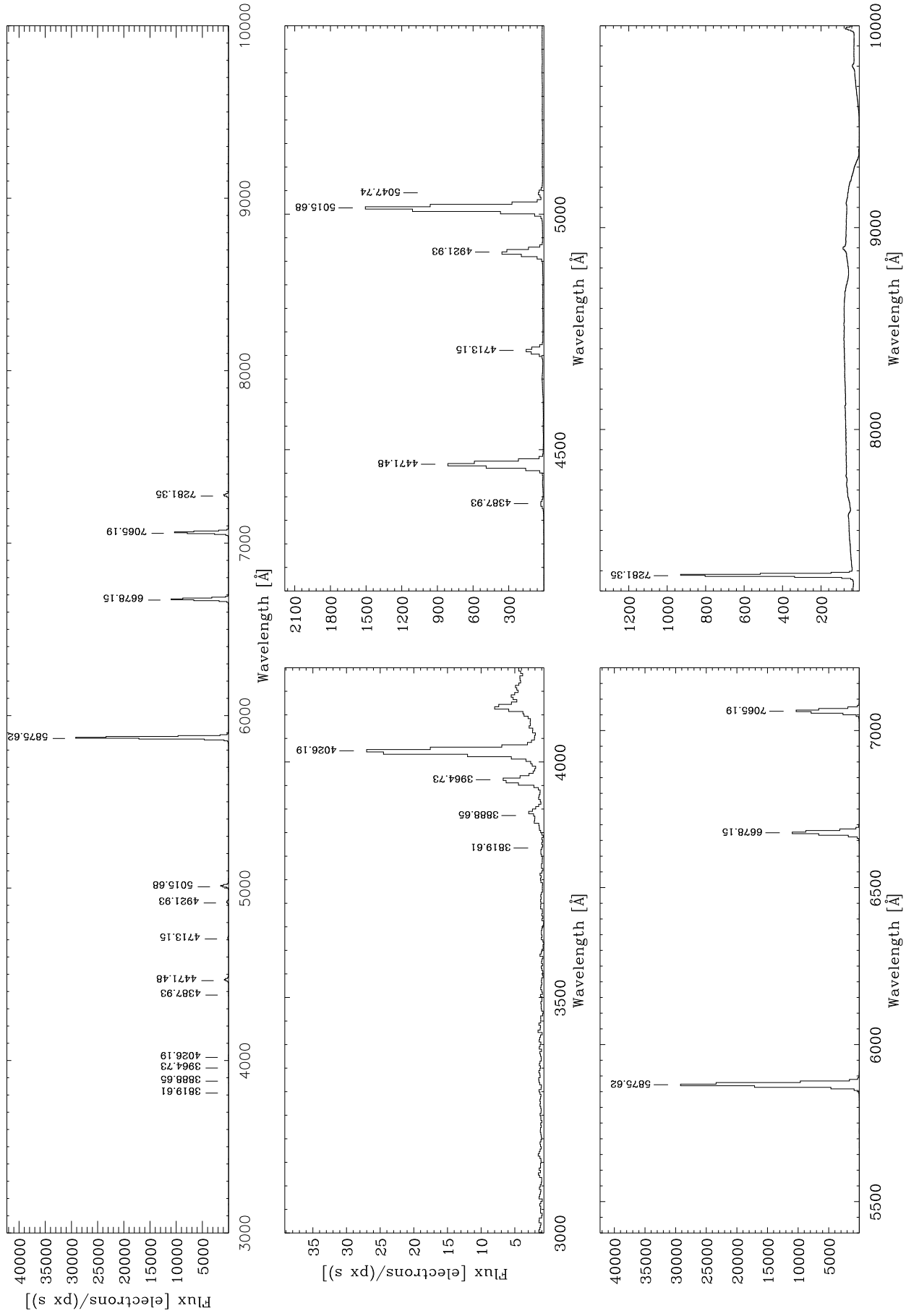
The Cadmium is a Pen-Ray Lamp, which are low pressure, cold cathode UV lamps made of double bore quartz tubing. This design allows the electrodes to be positioned at one end of the lamp providing easy access into small apertures.

⁶<http://www.ing.iac.es/astrometry/instruments/af2/calibration.html>

Grating	λ_{cen} Å	$\lambda_{\text{min}} - \lambda_{\text{max}}$ Å	Dispersion Å/px	Lamp	Exp.time s	Lines	ND filter
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)
R158B	4500	3000 – 10000	6.4	He	3	9 / 8	
				Ne	2	31 / 25	
				Hg	1	18 / 15	
R158R	7500	3000 – 10000	6.4	He	2	8 / 7	
				Ne	1	30 / 23	ND.5
				Hg	1	23 / 17	ND.5
R300B	5300	3000 – 8730	3.6	He	2	9 / 8	
				Ne	2	46 / 34	
				Hg	1	21 / 14	
R316R	7500	4418 – 10000	3.4	He	2	5 / 4	
				Ne	1	47 / 32	ND.5
				Hg	1	17 / 12	ND.5
R600B	5000	3351 – 6682	1.8	He	5	9 / 6	
				Ne	2	25 / 21	
				Hg	1	7 / 6	ND.5
				ThAr	60	72 / 60	
				Cd	10	4 / 4	
R600R	6500	4834 – 8177	1.8	He	3	6 / 6	
				Ne	1	32 / 27	ND.5
				Hg	1	12 / 12	ND.5
				ThAr	5	39 / 29	
				Cd	1	12 / 11	
	8000	6371 – 9668	1.8	He	3	3 / 3	
				Ne	1	36 / 23	ND.5
				Hg	3	19 / 13	
				ThAr	2	30 / 17	
				Cd	20	21 / 14	
R1200B	4000	3177 – 4815	0.86	He	120	7 / 5	
				Ne	120	2 / 2	
				Hg	2	5 / 3	
				ThAr	120	47 / 43	
				Cd	30	15 / 13	
	4500	3678 – 5298	0.84	He	60	9 / 6	
				Ne	120	11 / 10	
				Hg	1	4 / 3	
				ThAr	60	66 / 48	
				Cd	30	3 / 3	
	5000	4193 – 5804	0.81	He	30	6 / 6	
				Ne	60	28 / 23	
				Hg	1	4 / 4	
				ThAr	120	67 / 54	
				Cd	30	20 / 18	
R1200R	5500	4665 – 6292	0.79	He	15	5 / 4	
				Ne	2	21 / 18	
				Hg	3	4 / 4	
				ThAr	120	64 / 47	
				Cd	10	3 / 3	
6000	5169 – 6746	0.76	He	15	2 / 2	ND.5	
			Ne	2	25 / 23	ND.5	
			Hg	3	5 / 5		

Grating	λ_{cen} Å	$\lambda_{\text{min}} - \lambda_{\text{max}}$ Å	Dispersion Å/px	Lamp	Exp.time s	Lines	ND filter
				ThAr	10	51 /43	
				Cd	30	26 /23	
	6500	5681 – 7266	0.74	He	5	3 /3	
				Ne	1	28 /23	
				Hg	10	13 /10	
				ThAr	30	57 /39	
				Cd	30	4 /3	
	7000	6195 – 7748	0.71	He	10	3 /3	
				Ne	1	21 /19	
				Hg	10	18 /10	
				ThAr	5	29 /19	
				Cd	10	11 /10	
	7500	6711 – 8229	0.69	He	5	2 /2	
				Ne	1	15 /13	ND.5
				Hg	10	18 /14	
				ThAr	5	24 /13	
				Cd	10	17 /14	
	8000	7208 – 8708	0.66	He	40	1 /1	
				Ne	1	24 /16	
				Hg	5	17 /16	
				ThAr	3	18 /15	
				Cd	5	16 /15	
	8500	7724 – 9184	0.64	He	80	0 /0	
				Ne	1	31 /14	
				Hg	1	11 /10	
				ThAr	5	11 /11	
				Cd	5	12 /11	
	9350	8601 – 10000	0.59	He	120	0 /0	
				Ne	5	20 /14	
				Hg	2	3 /3	
				ThAr	10	8 /5	
				Cd	30	7 /6	
H1800V	5120	4604 – 5579	0.44	He	60	4 /4	
				Ne	60	30 /23	
				ThAr	30	56 /52	
H2400B	4000	3645 – 4343	0.38	He	120	3 /2	
				Ne	120	0 /0	
				Hg	4	5 /5	
				ThAr	300	35 /35	
				Cd	300	9 /9	
	4500	4144 – 4833	0.38	He	60	2 /2	
				Ne	120	6 /5	
				Hg	1	0 /0	
				ThAr	120	32 /28	
				Cd	60	14 /13	

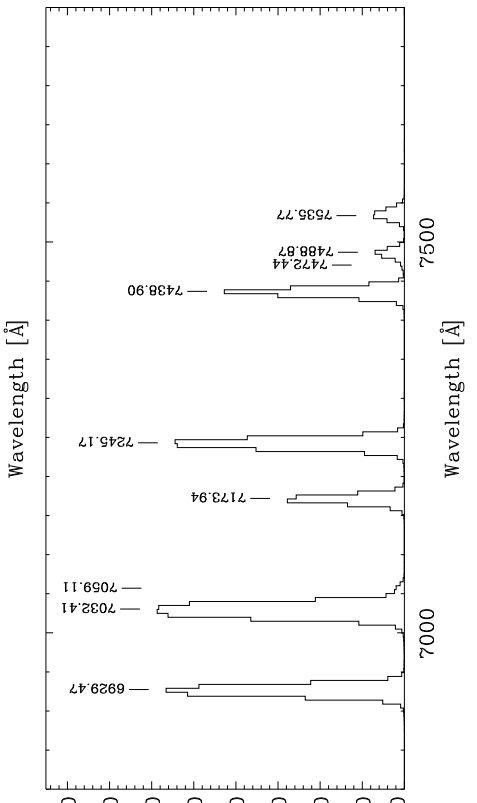
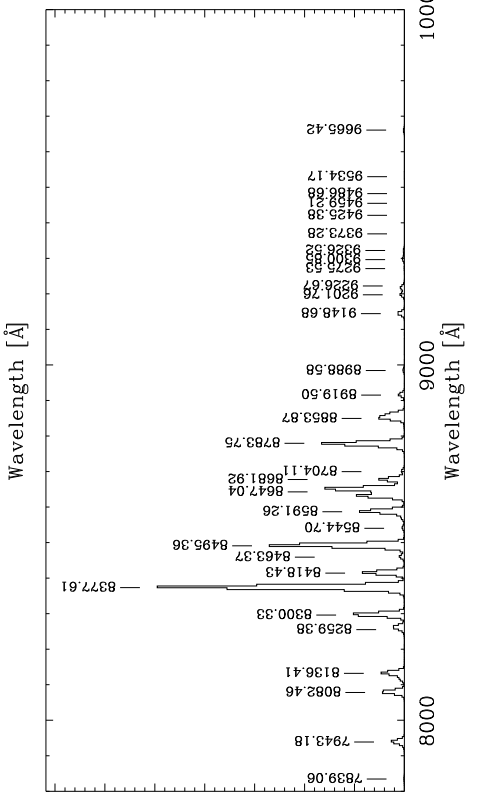
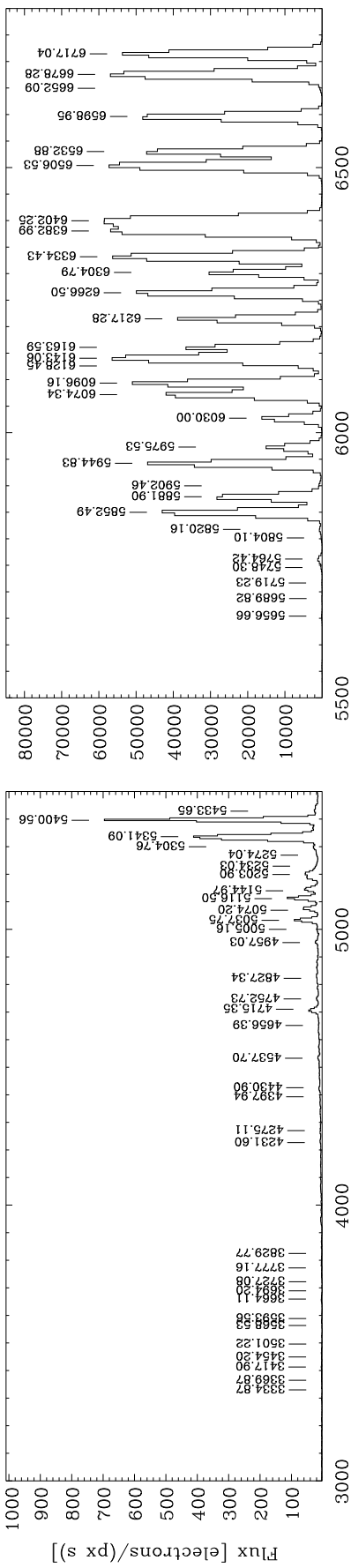
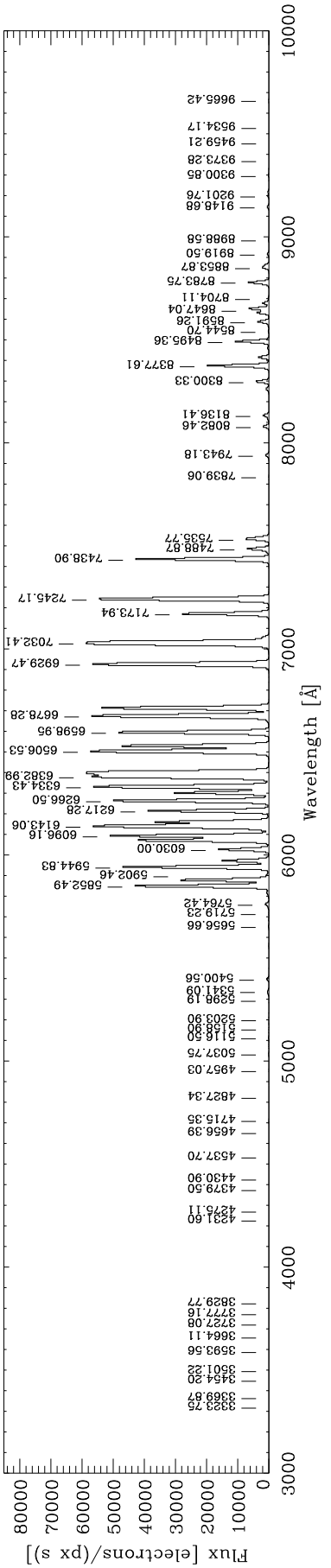
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R158B

$\lambda_{\text{cen}} = 4500\text{\AA}$

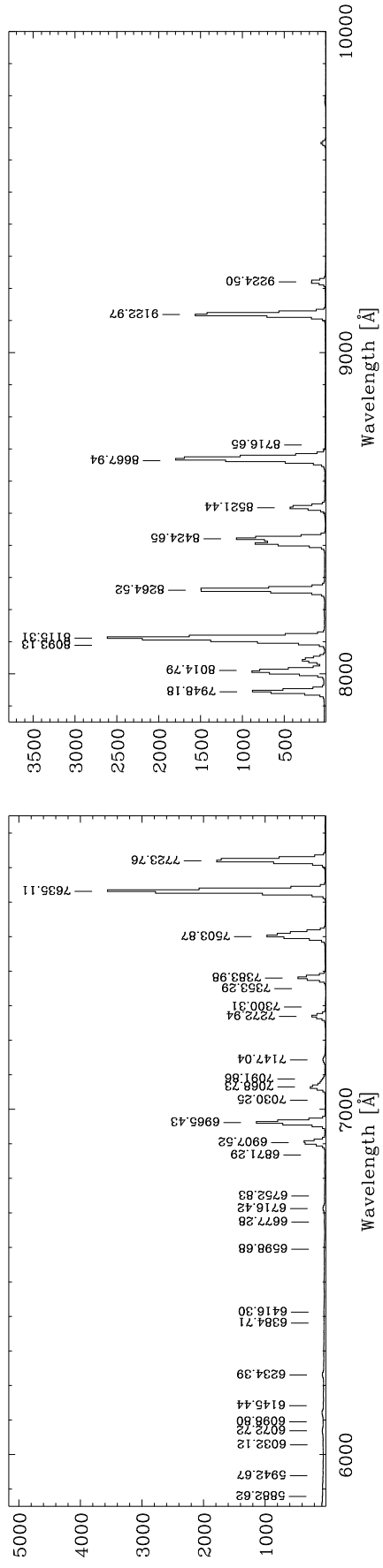
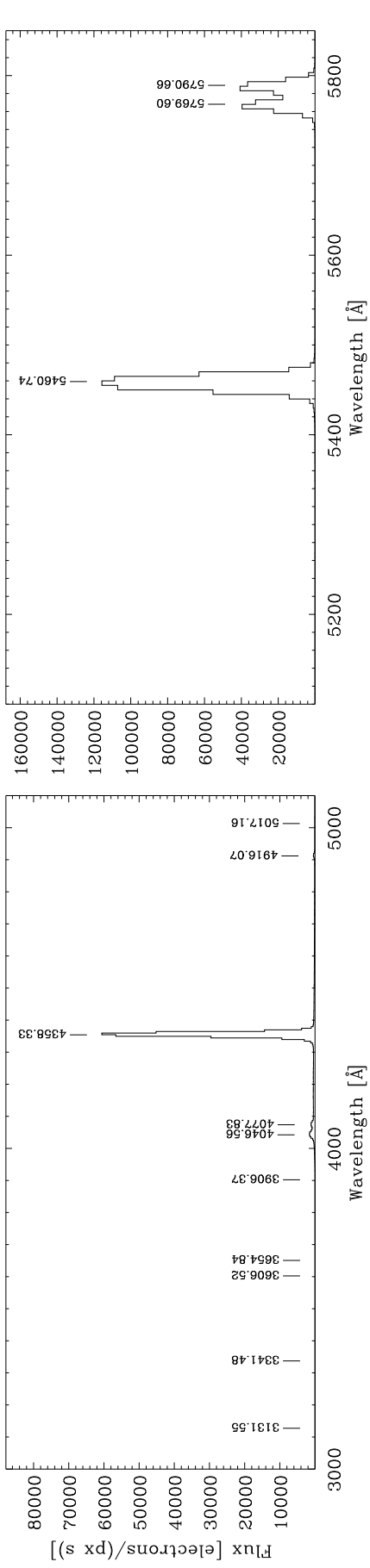
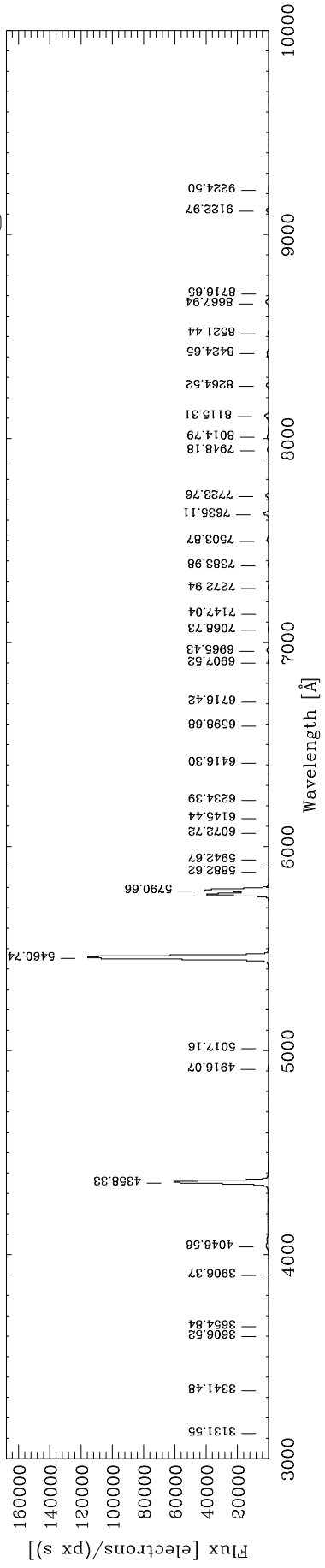
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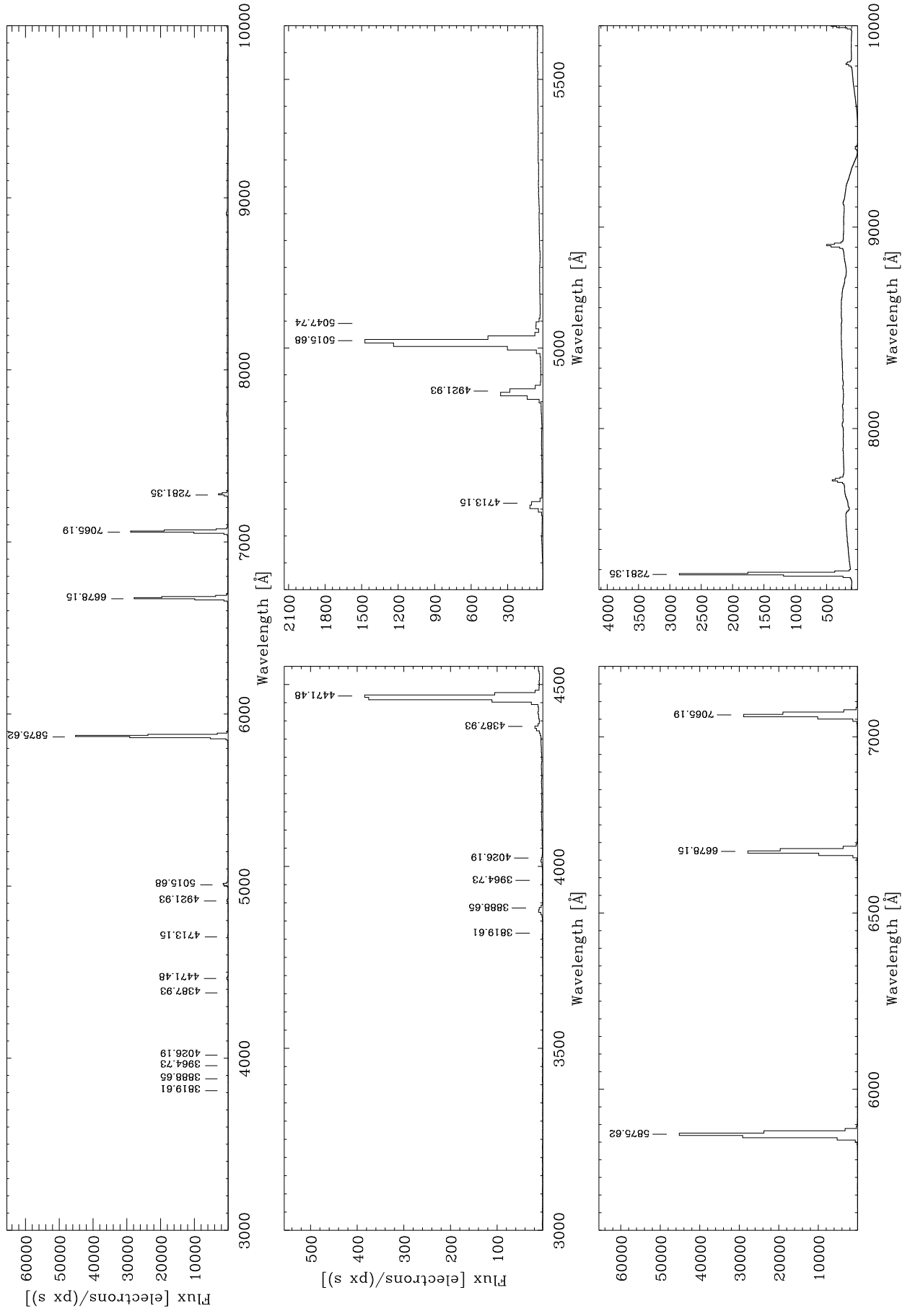
R158B

$\lambda_{\text{cen}} = 4500\text{\AA}$

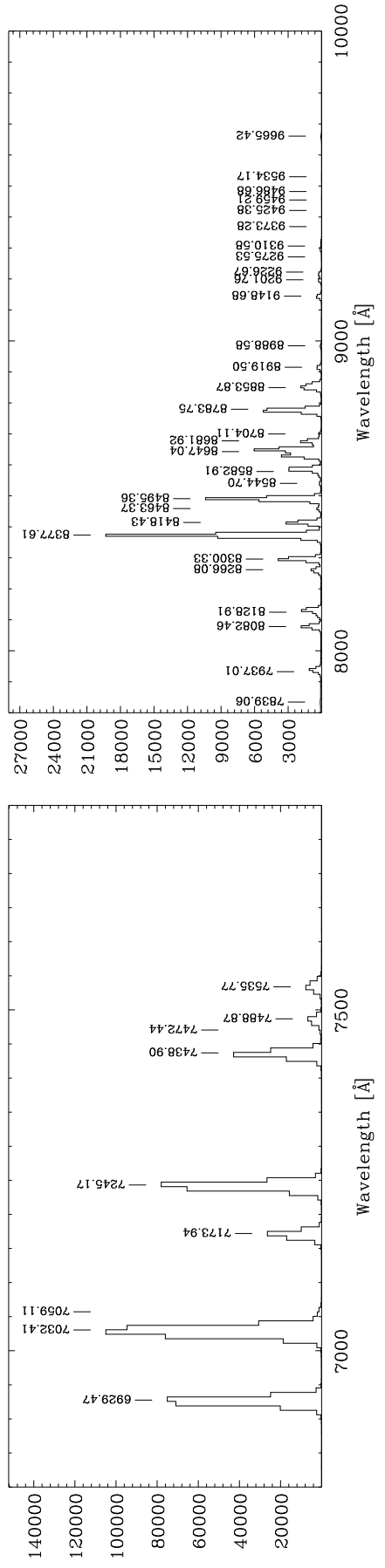
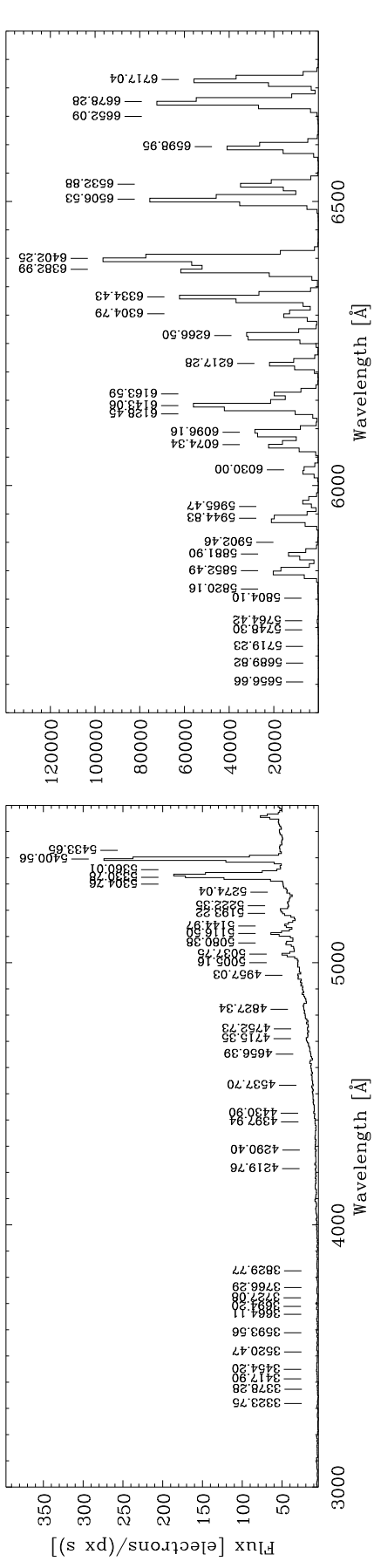
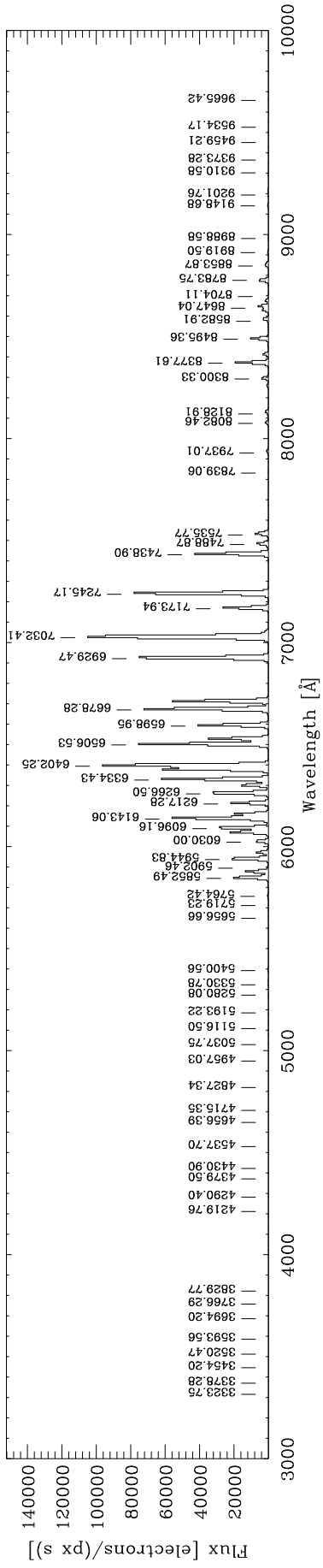
H γ



R158R $\lambda_{\text{cen}} = 7500\text{\AA}$ He



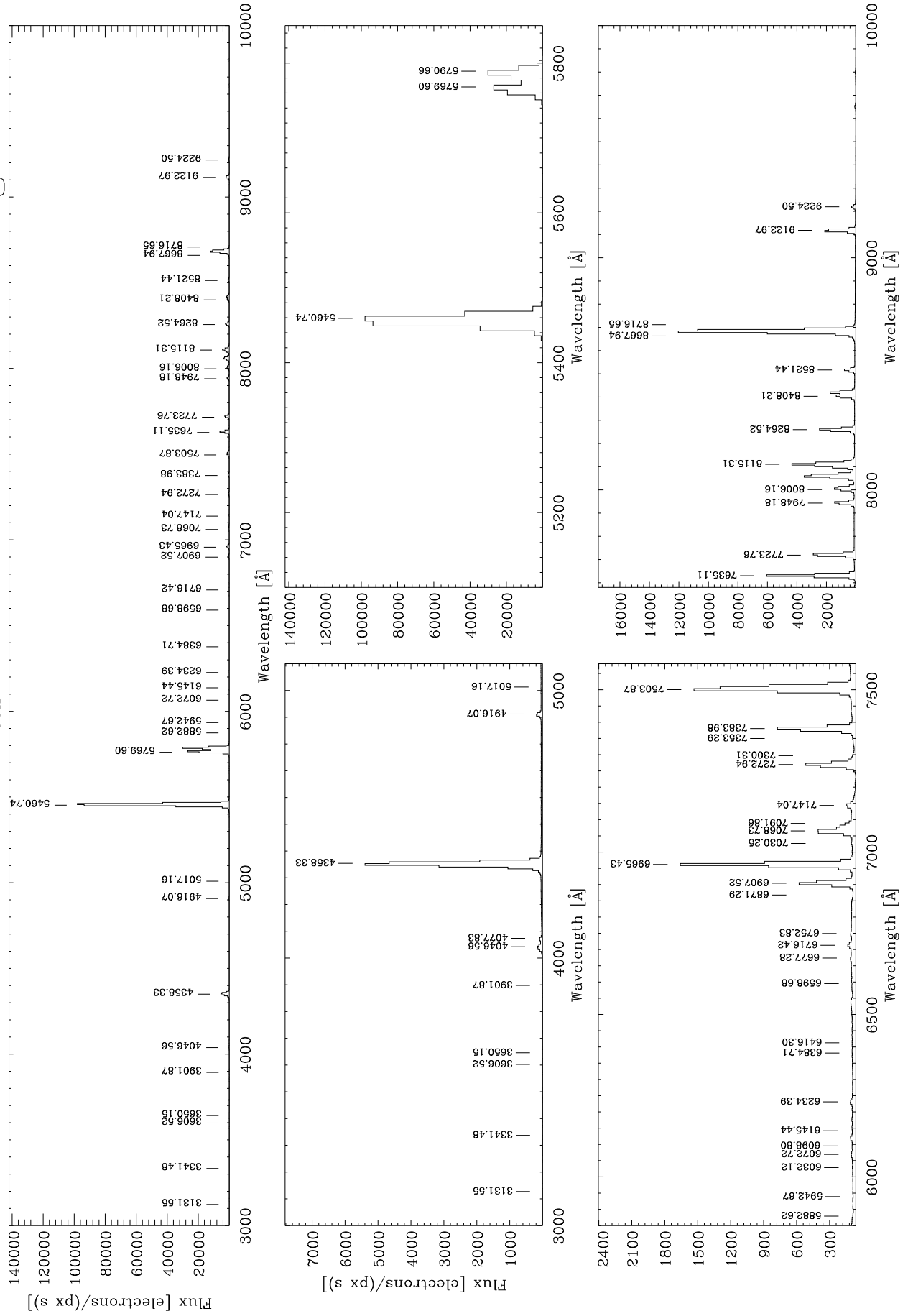
R158R $\lambda_{\text{cen}} = 7500\text{\AA}$ Ne



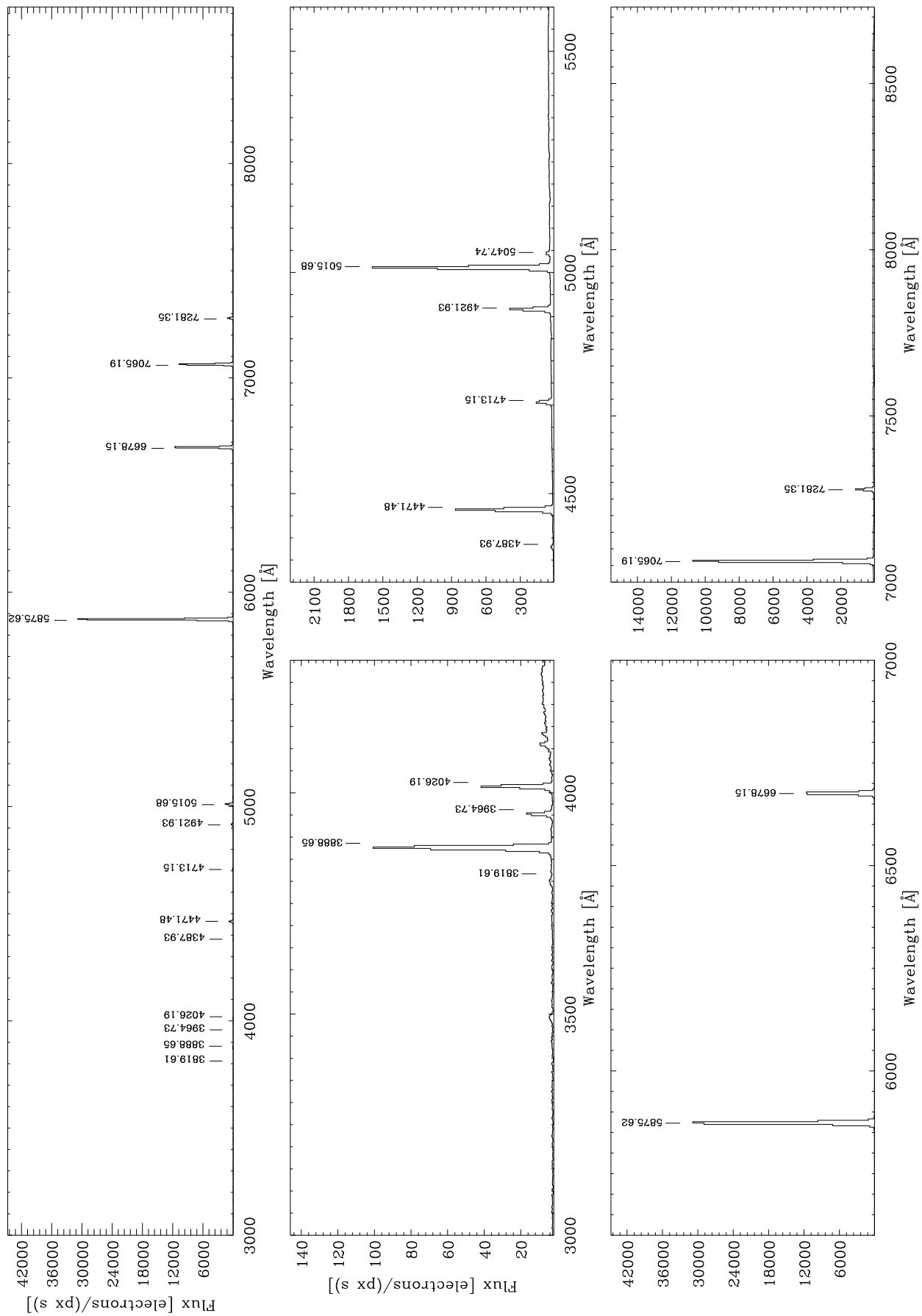
R158R

$\lambda_{\text{cen}} = 7500\text{\AA}$

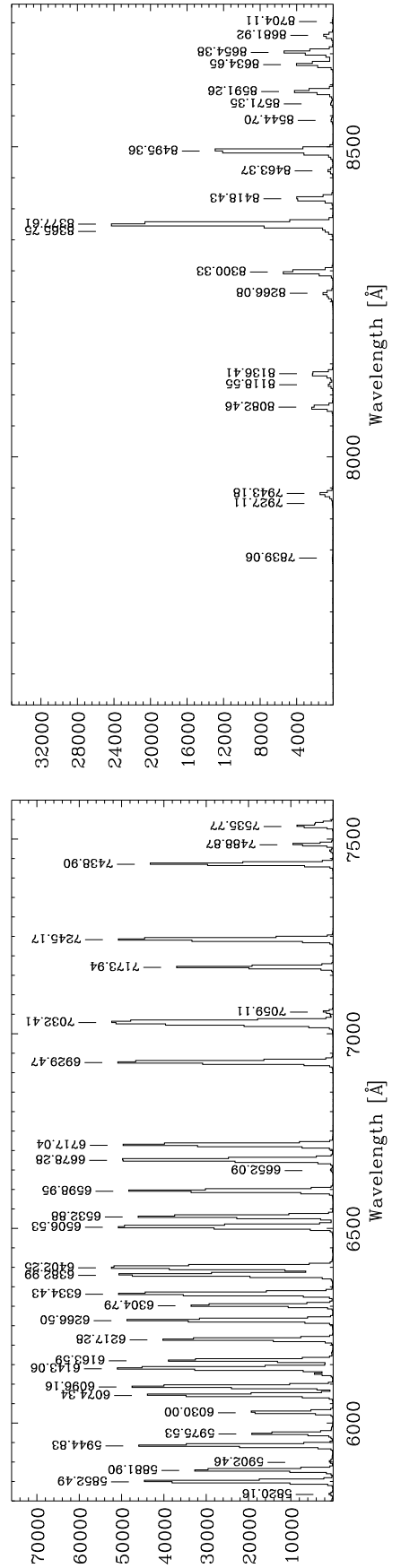
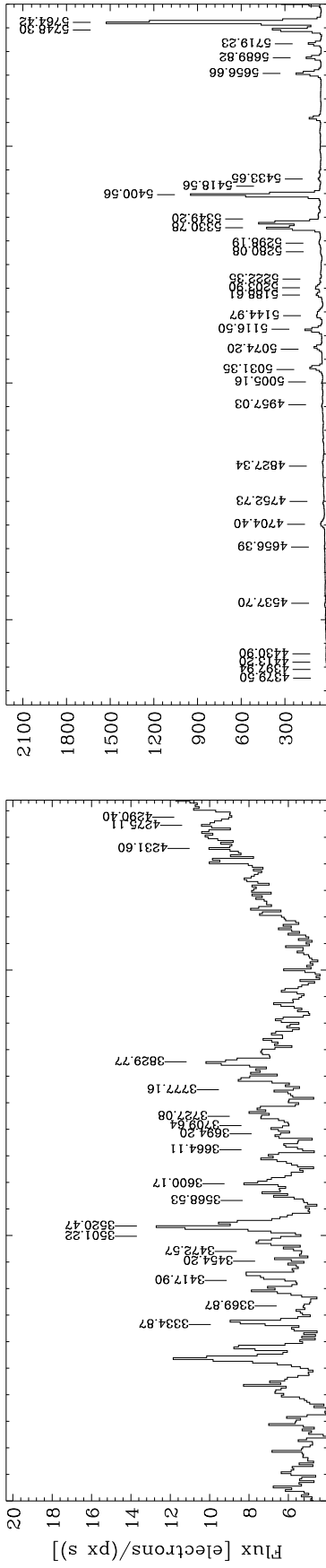
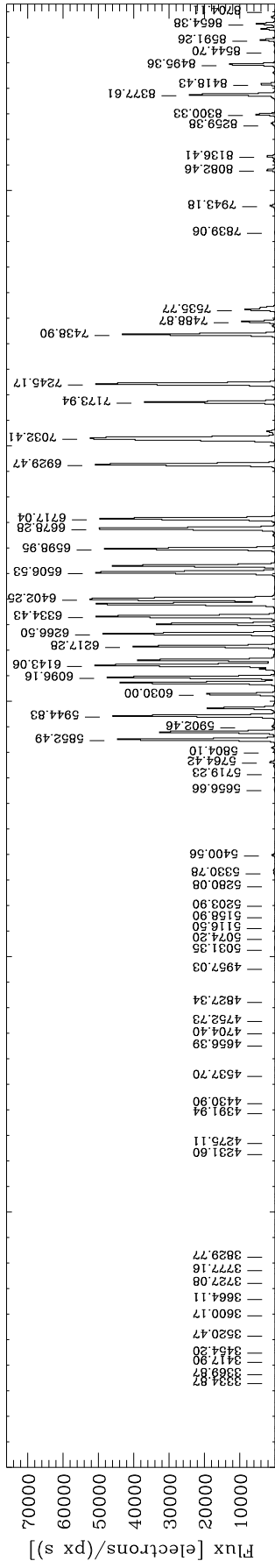
Hg



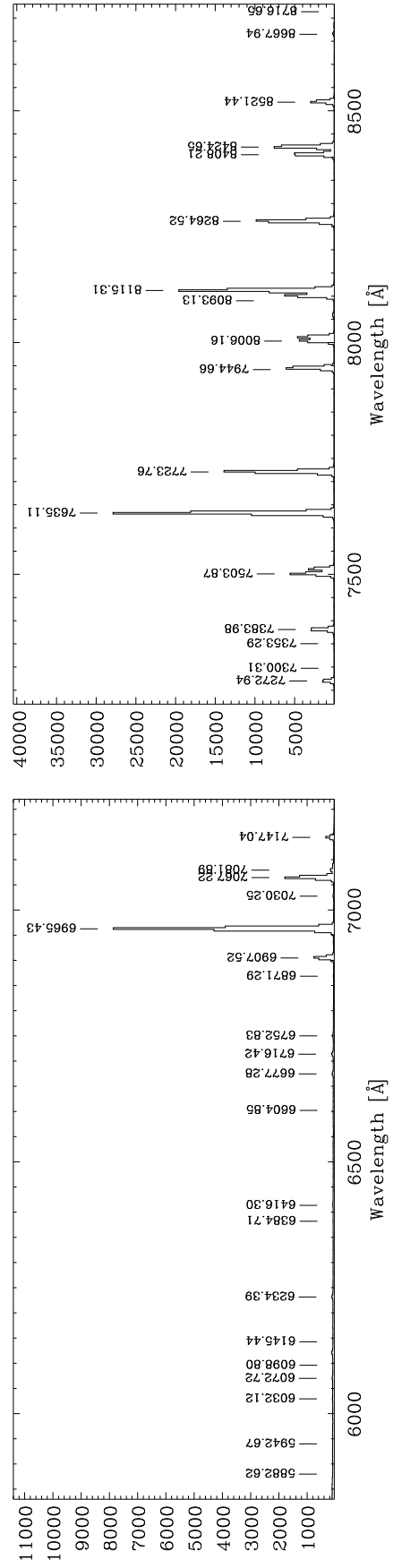
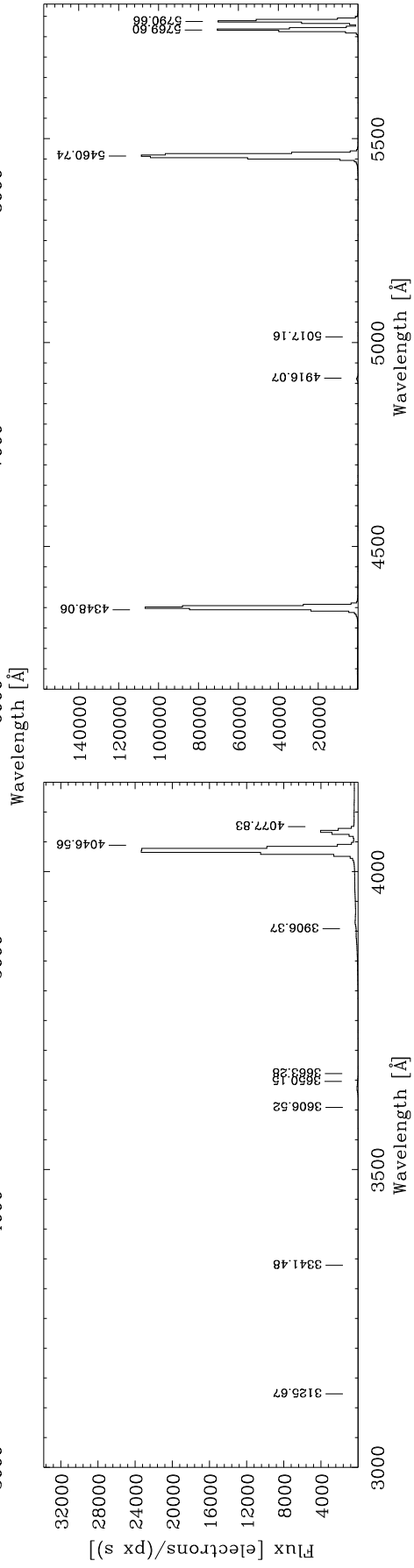
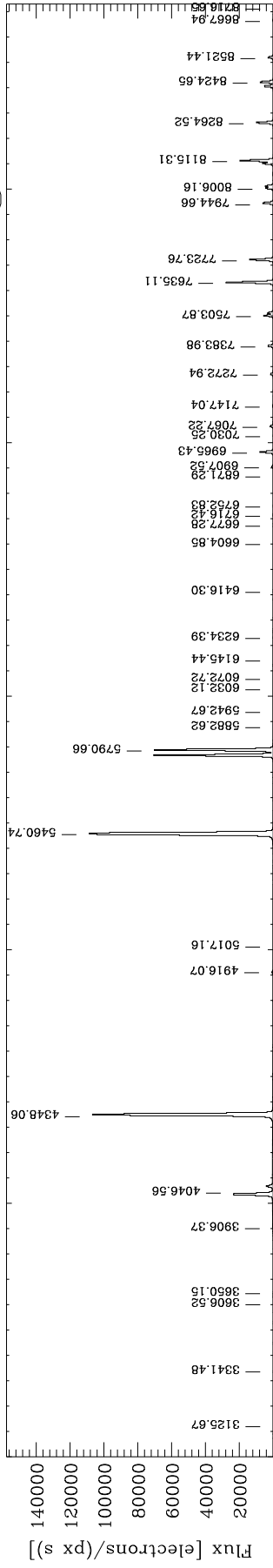
R300B $\lambda_{\text{cen}} = 5300\text{\AA}$ He



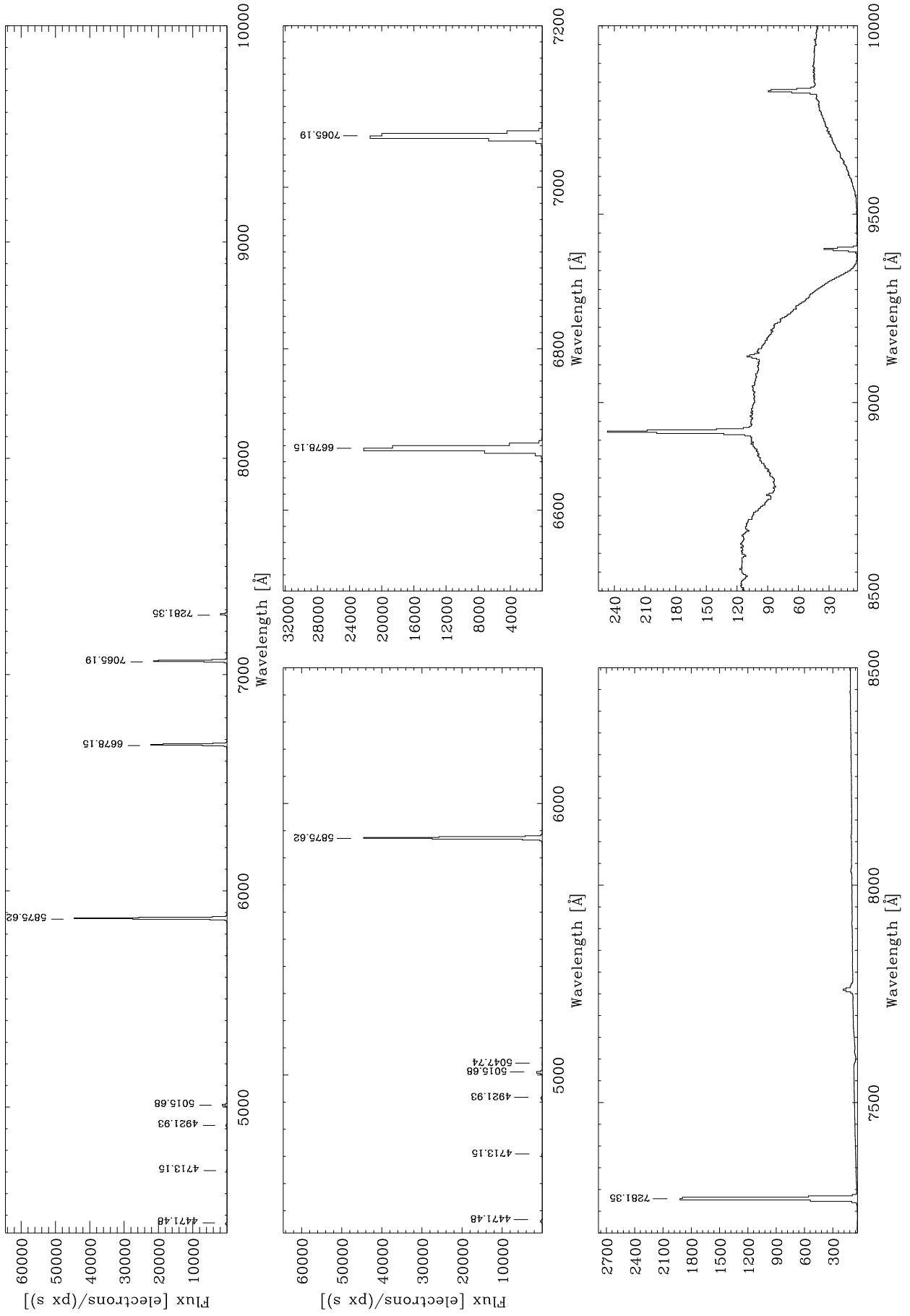
R300B $\lambda_{\text{cen}} = 5300\text{\AA}$ Ne



R300B $\lambda_{\text{cen}} = 5300\text{\AA}$ Hg



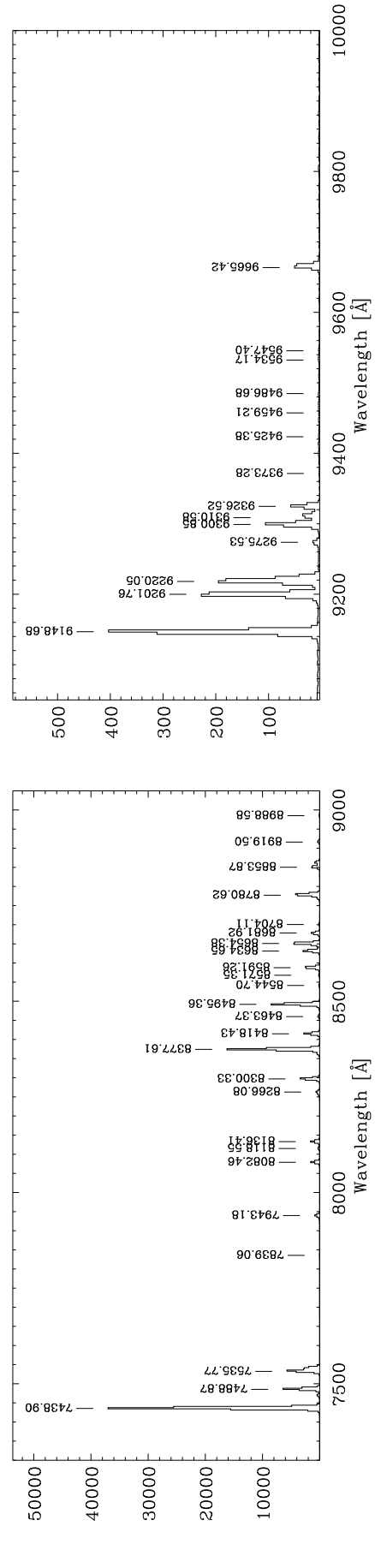
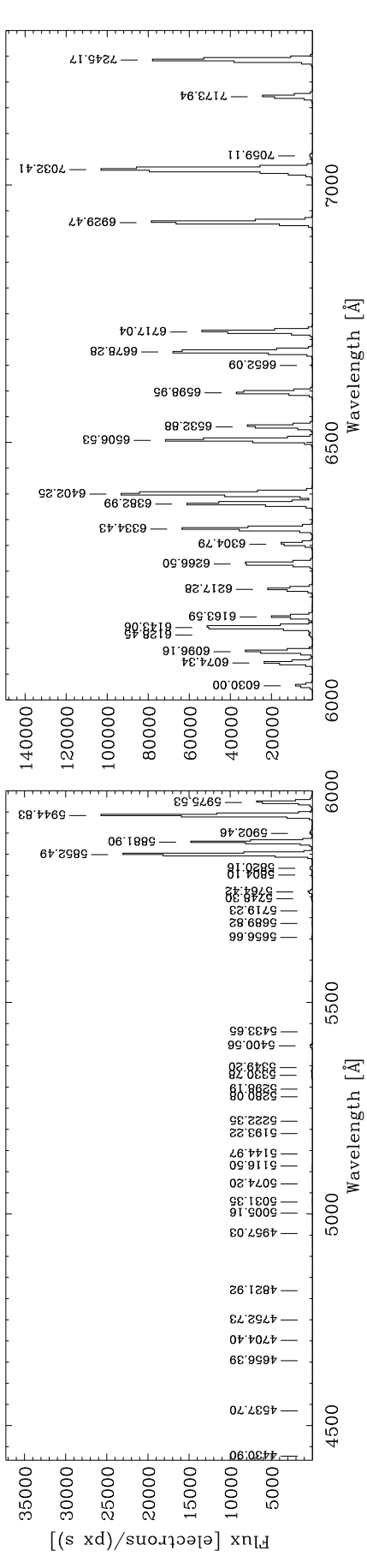
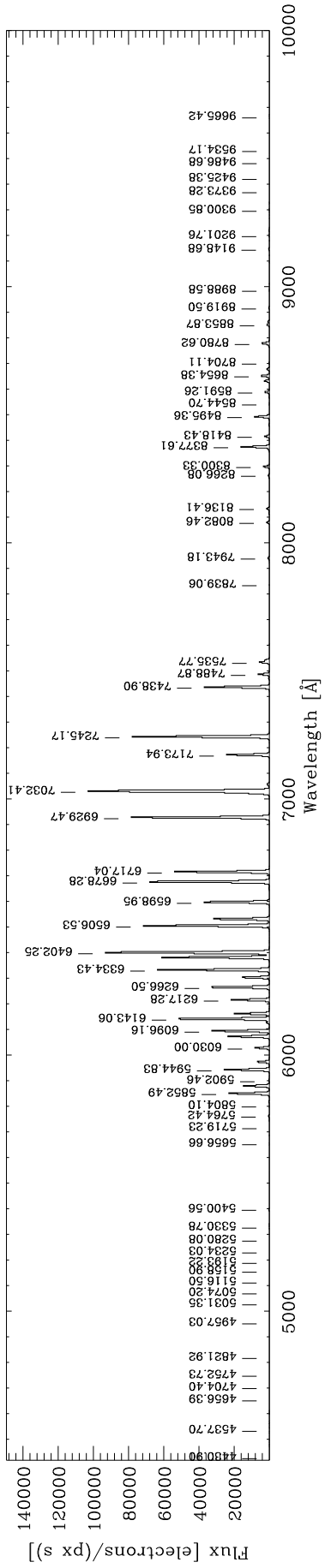
R316R $\lambda_{\text{cen}} = 7500\text{\AA}$ He



R316R

$\lambda_{\text{cen}} = 7500\text{\AA}$

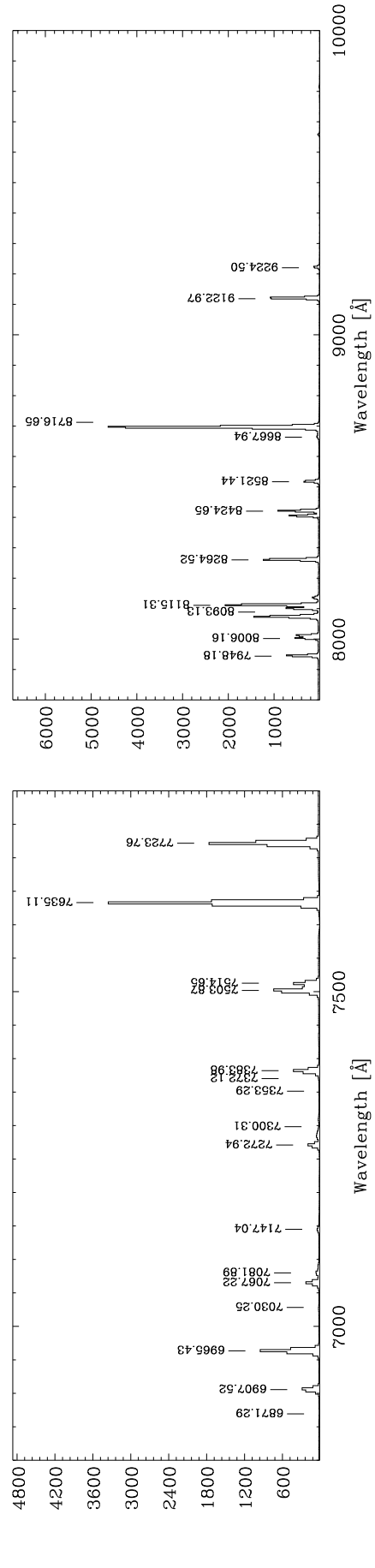
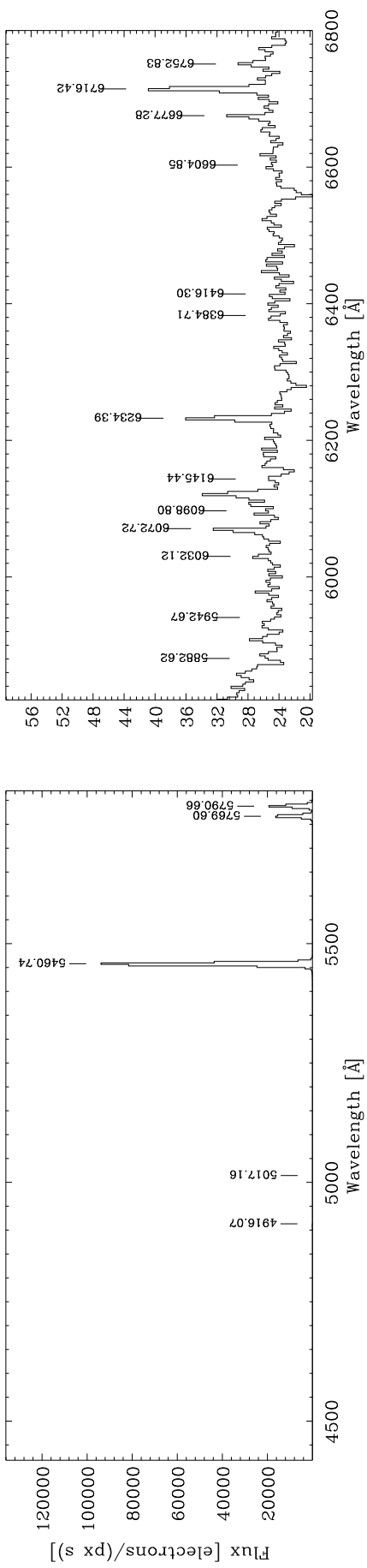
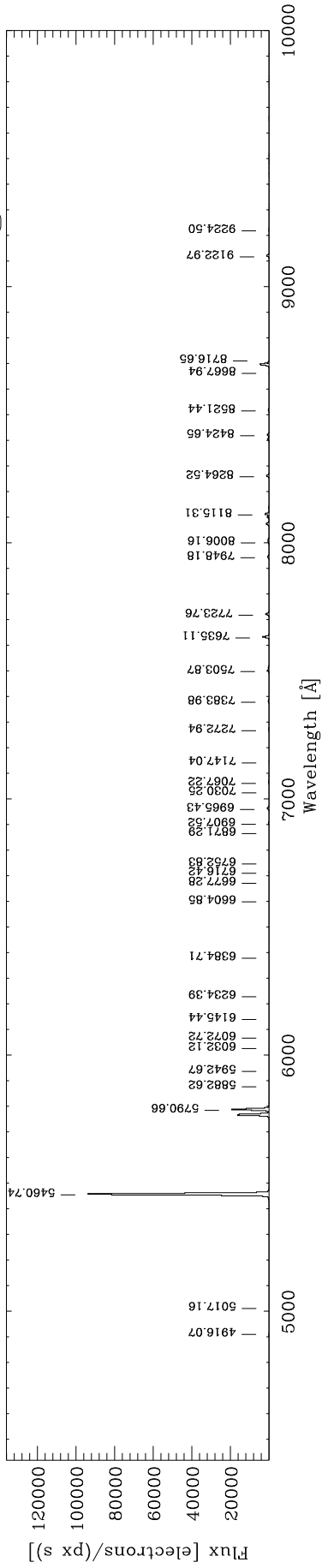
Ne



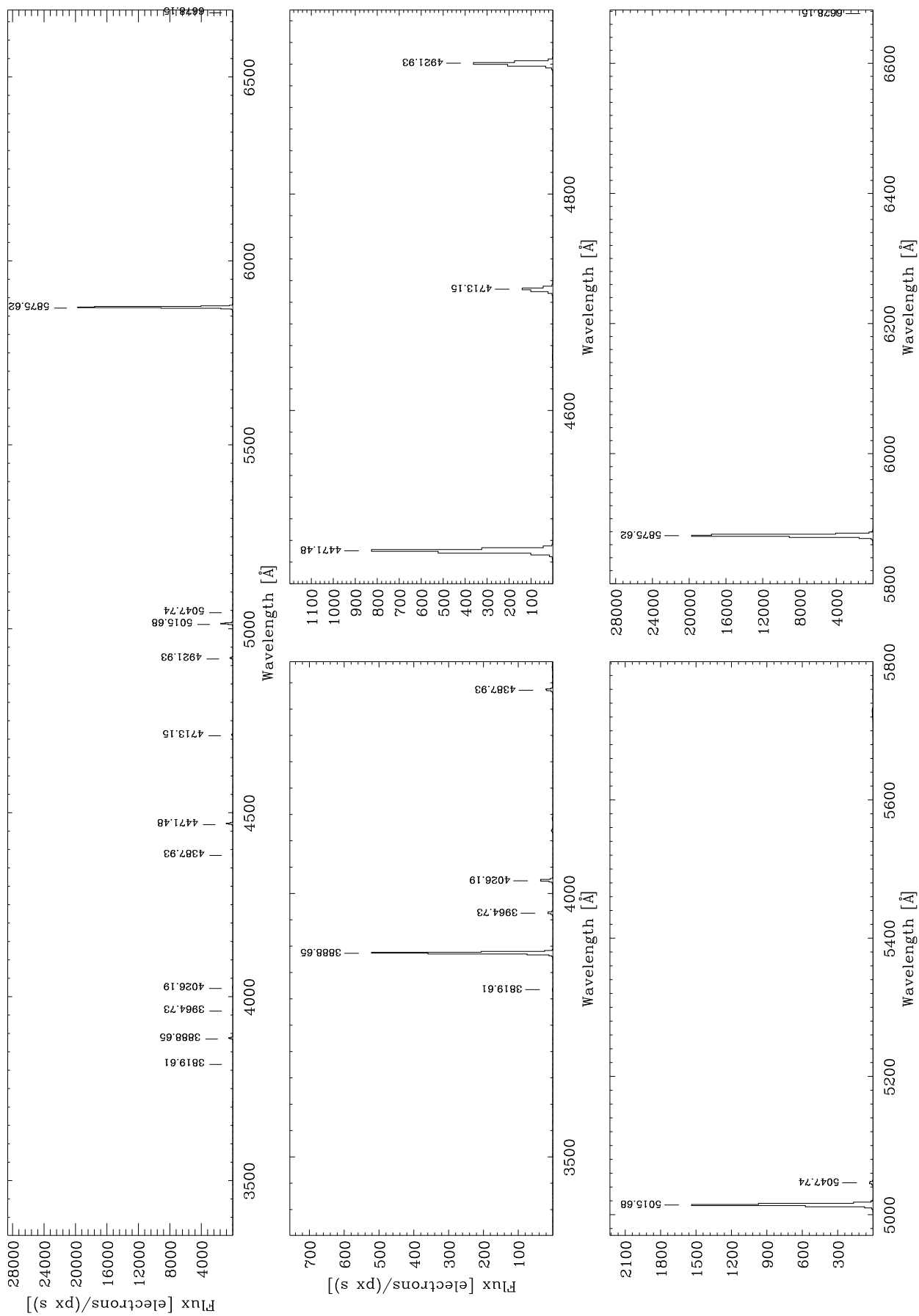
R316R

$\lambda_{\text{cen}} = 7500\text{\AA}$

Hg



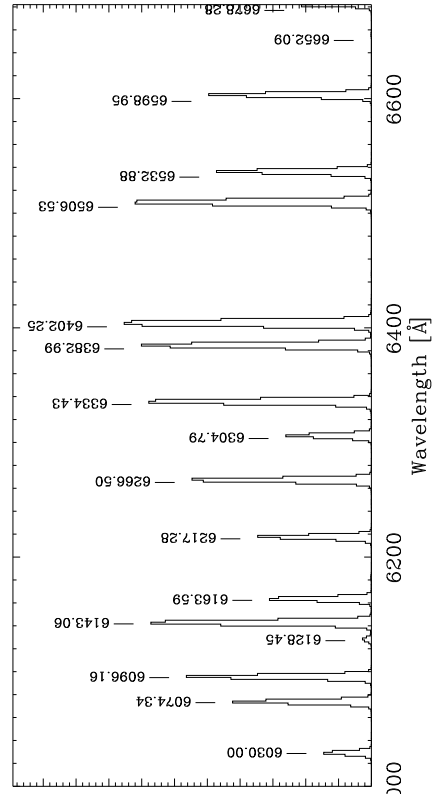
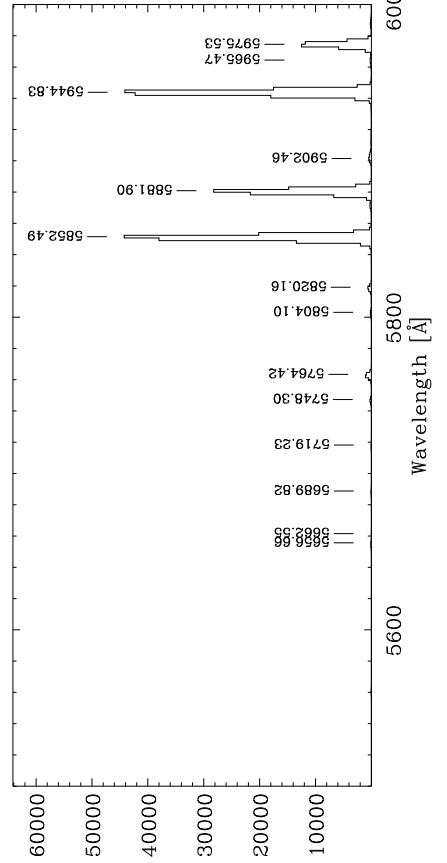
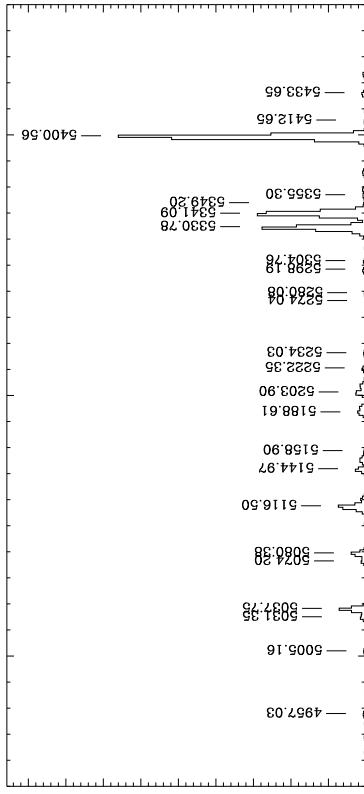
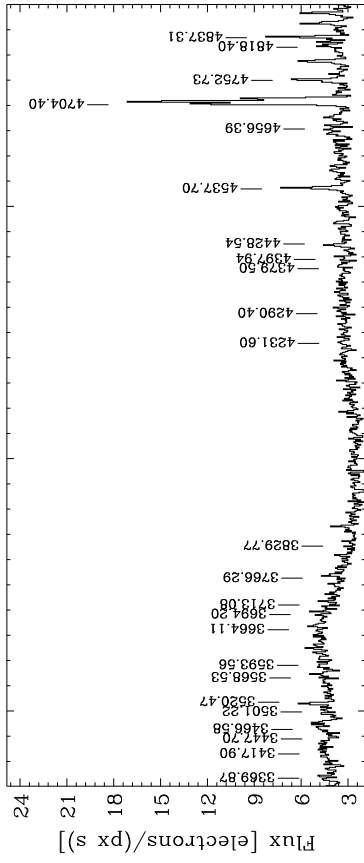
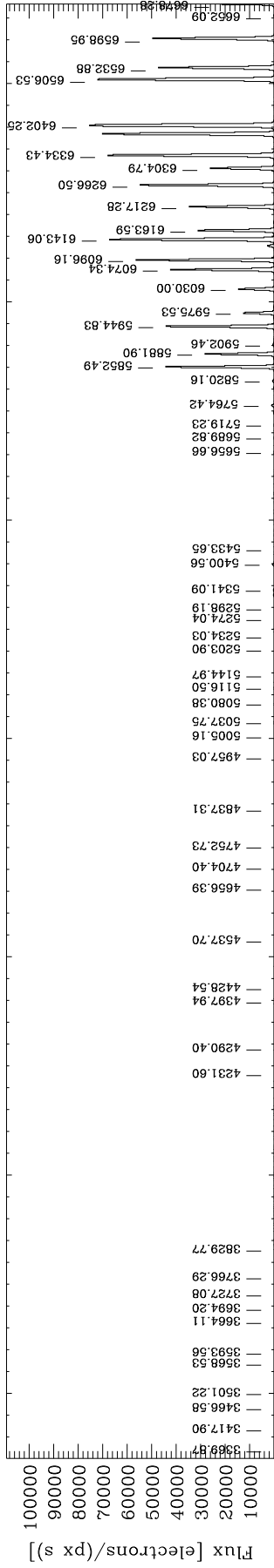
R600B $\lambda_{\text{cen}} = 5000\text{\AA}$ He



R600B

$\lambda_{\text{cen}} = 5000\text{\AA}$

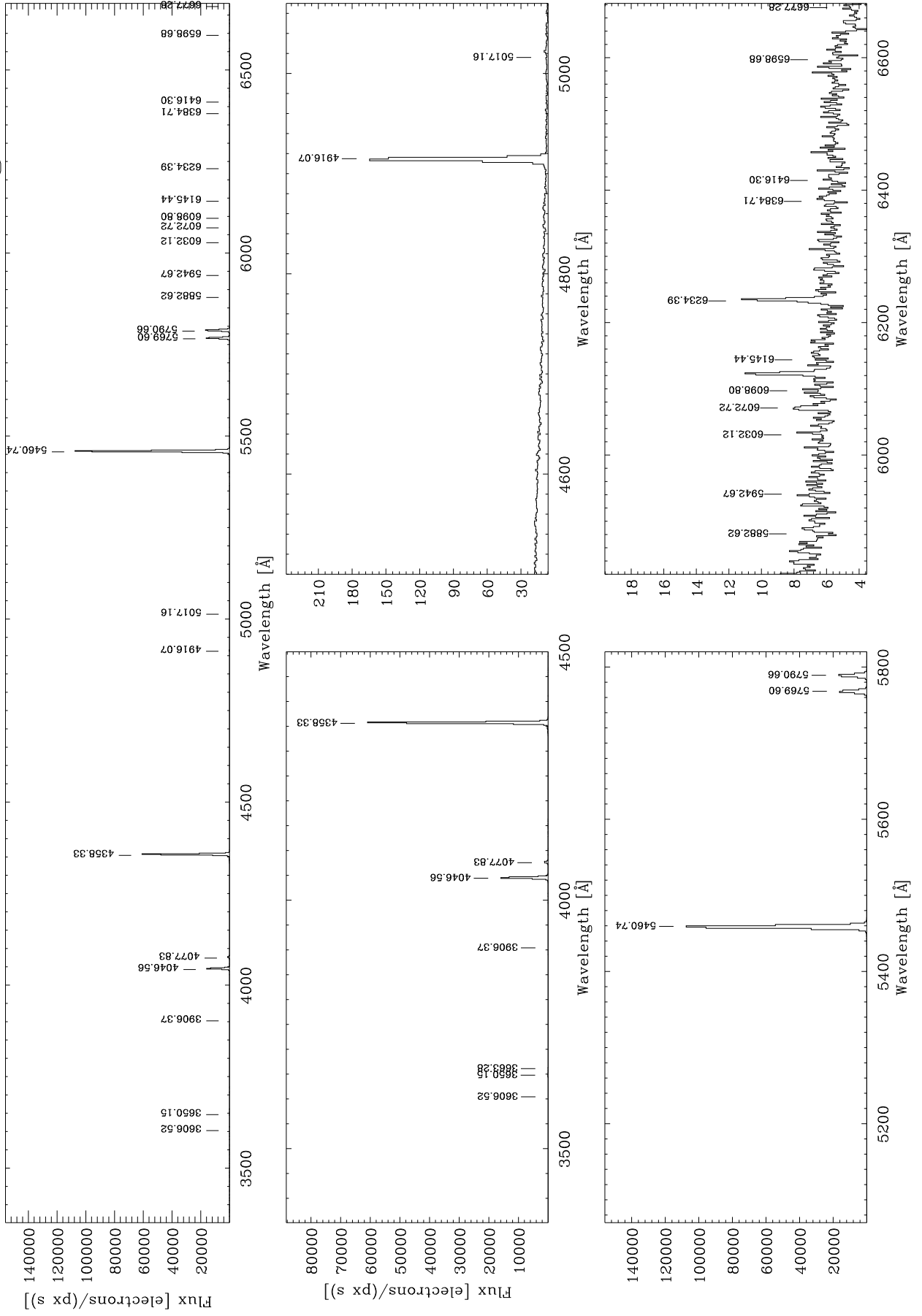
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R600B

$\lambda_{\text{cen}} = 5000\text{\AA}$

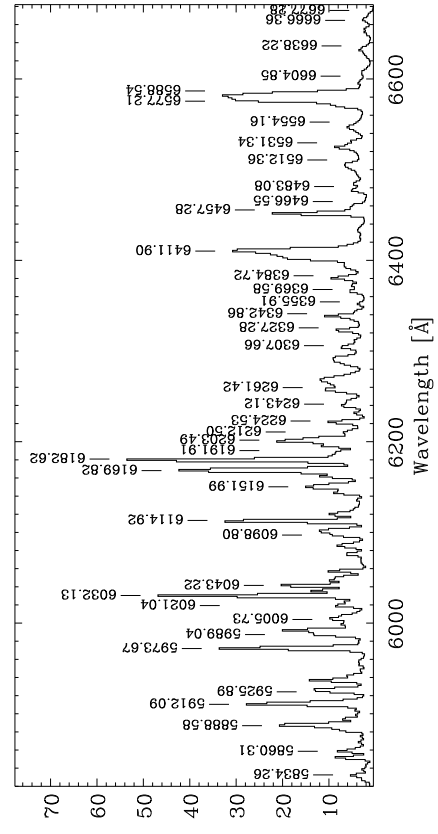
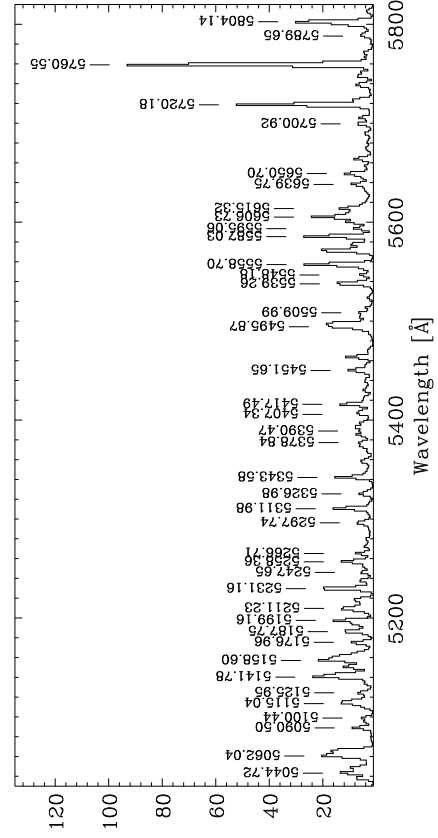
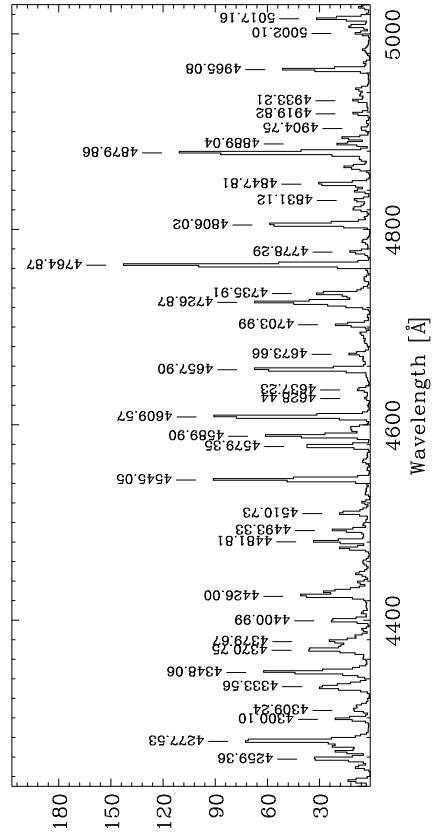
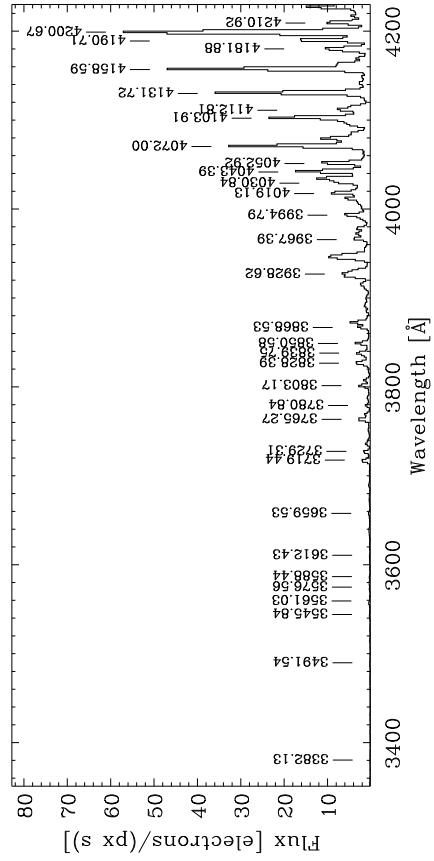
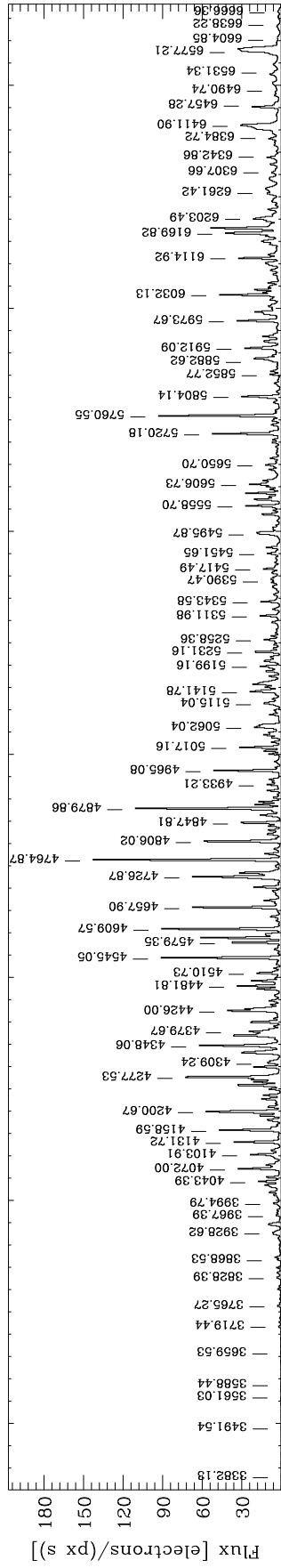
H γ



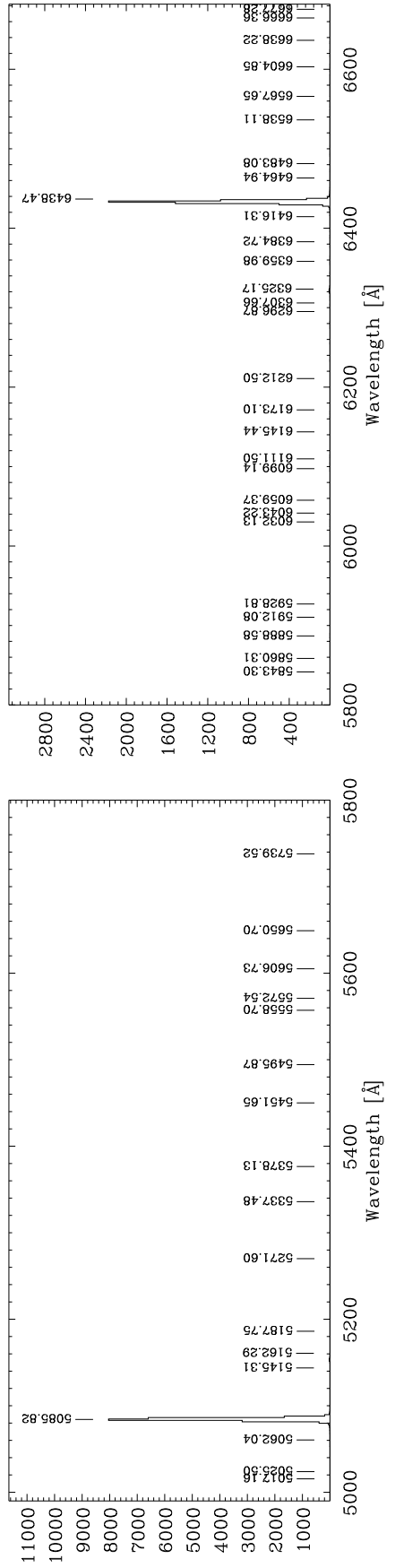
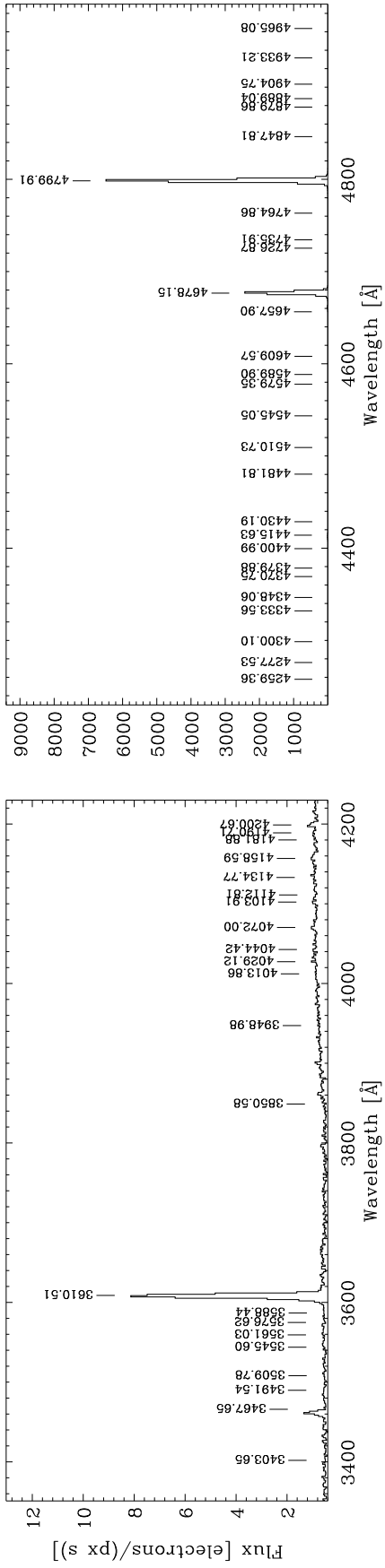
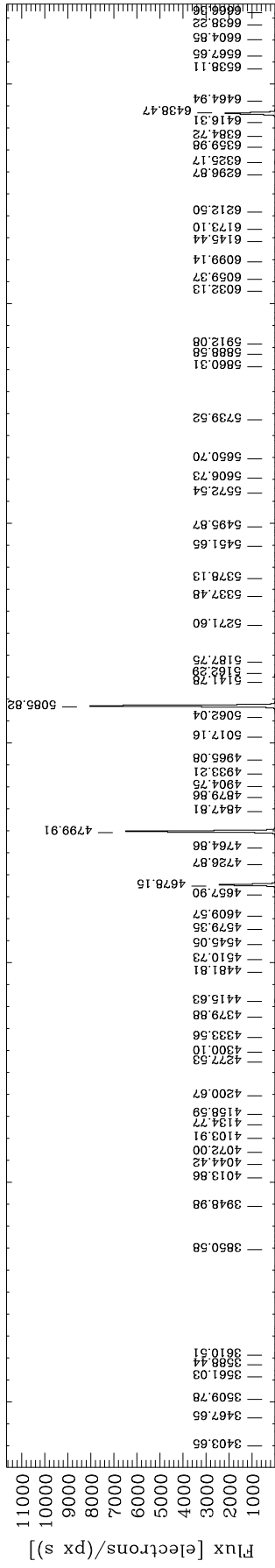
R600B

$\lambda_{\text{cen}} = 5000\text{\AA}$

ThAr



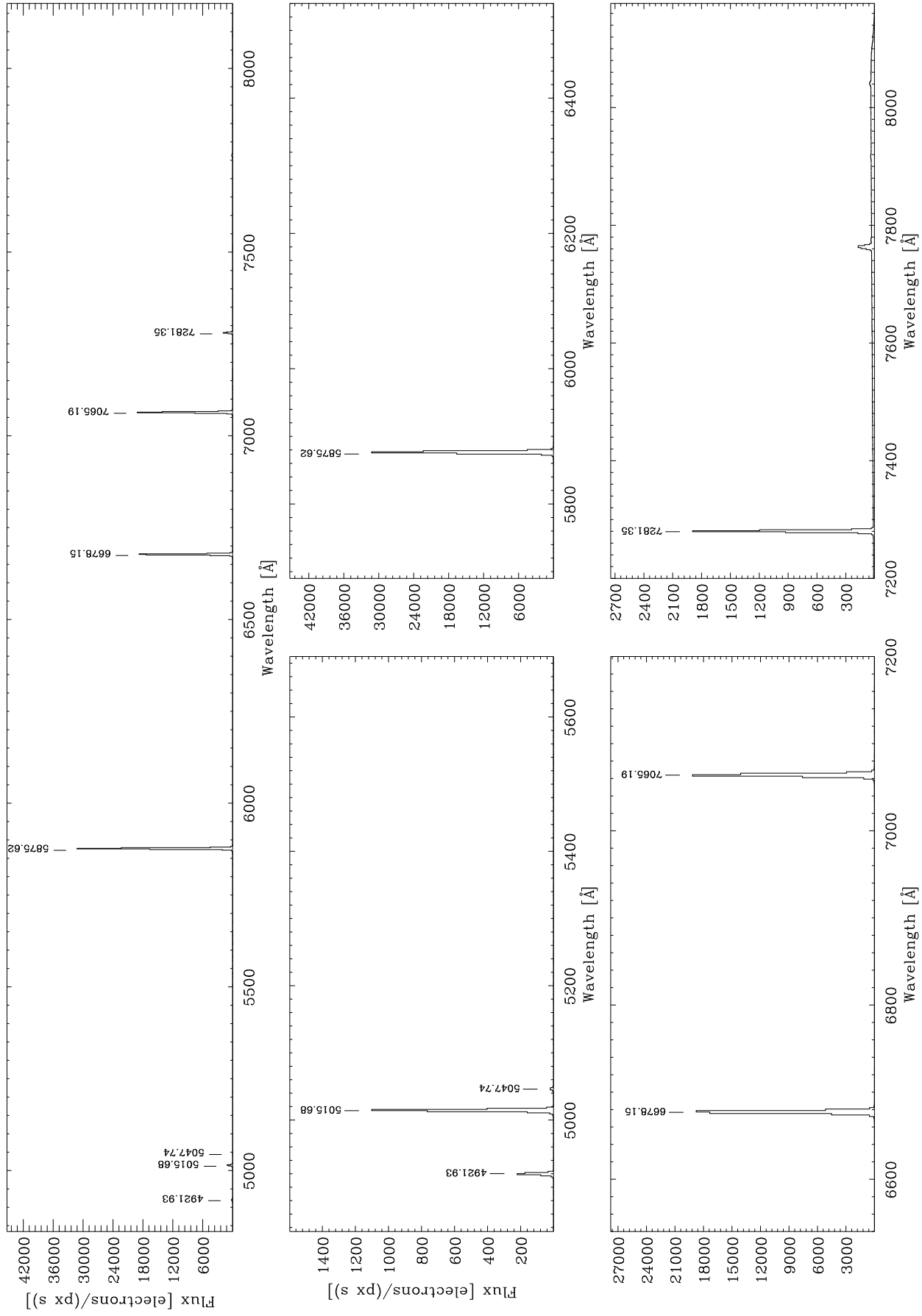
R600B $\lambda_{\text{cen}} = 5000\text{\AA}$ Cd



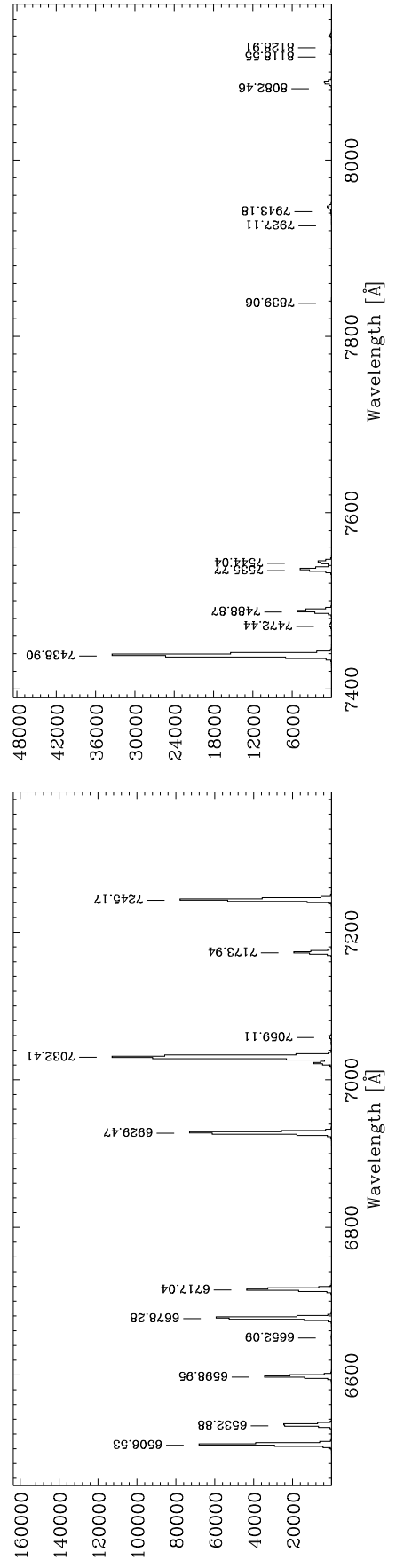
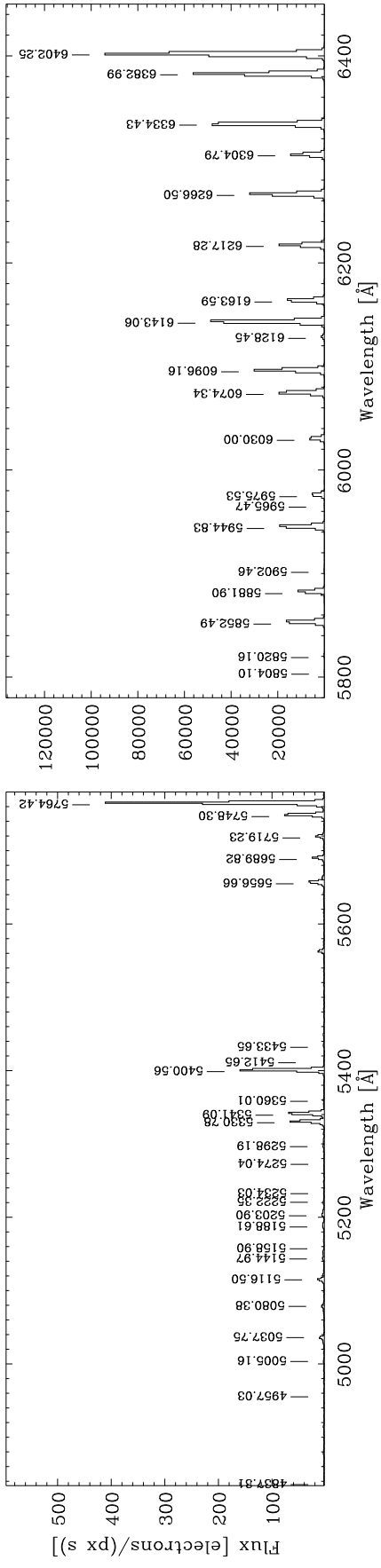
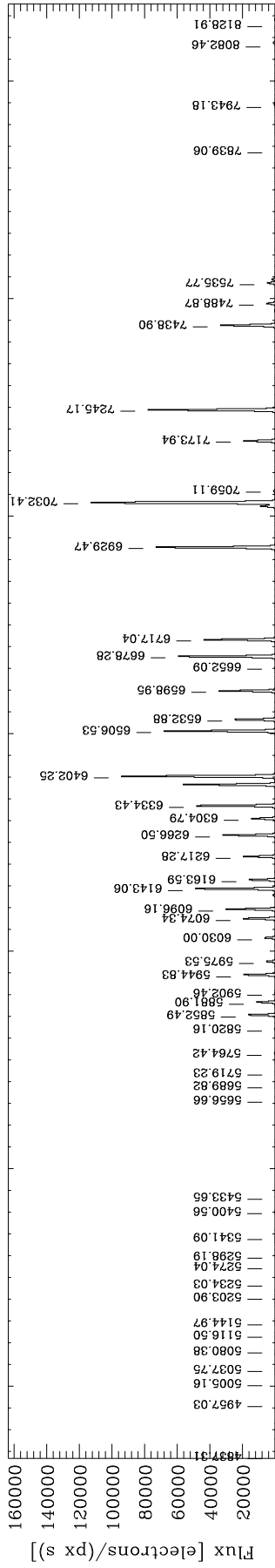
R600R

$\lambda_{\text{cen}} = 6500\text{\AA}$

He



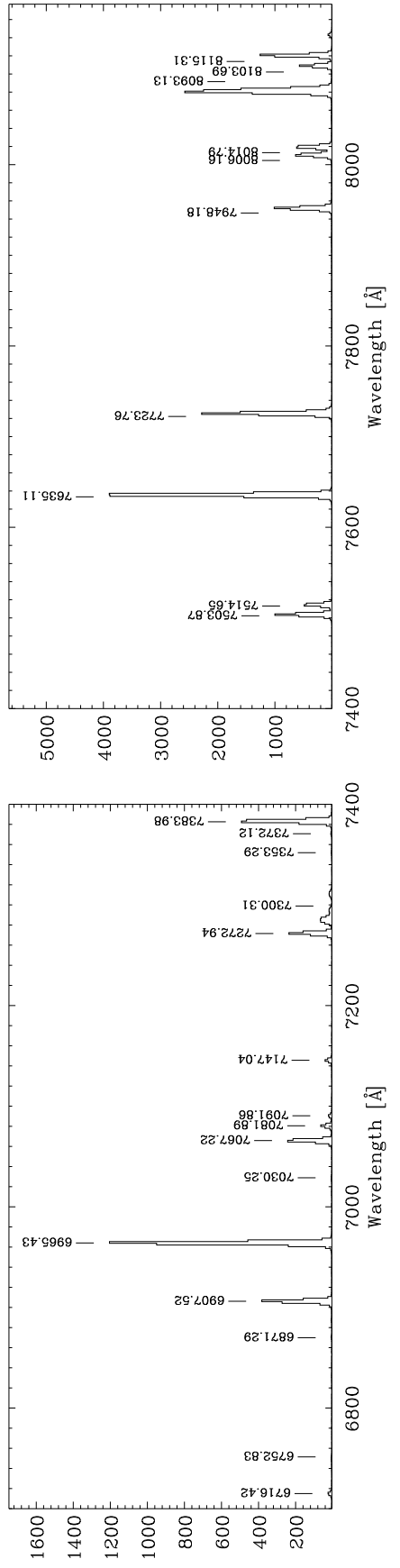
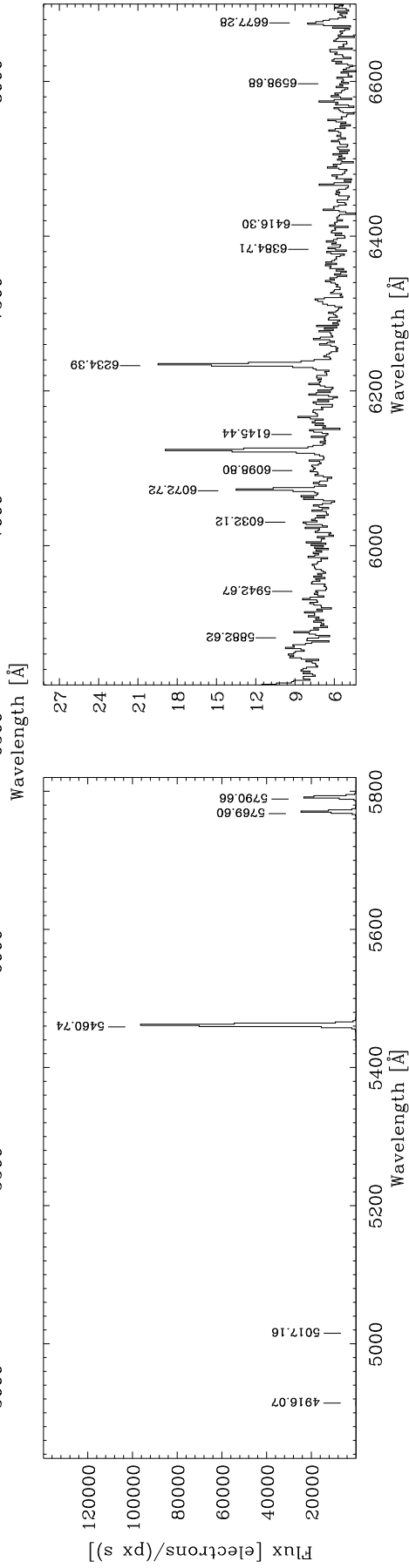
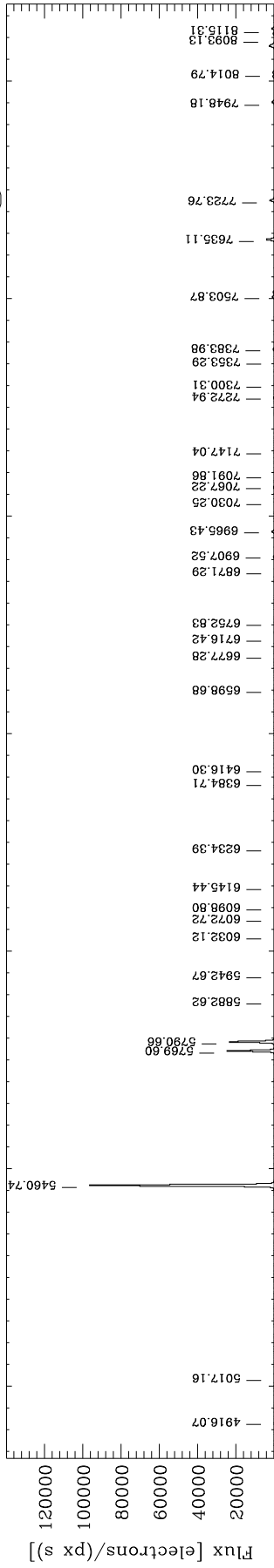
R600R $\lambda_{\text{cen}} = 6500\text{\AA}$ Ne



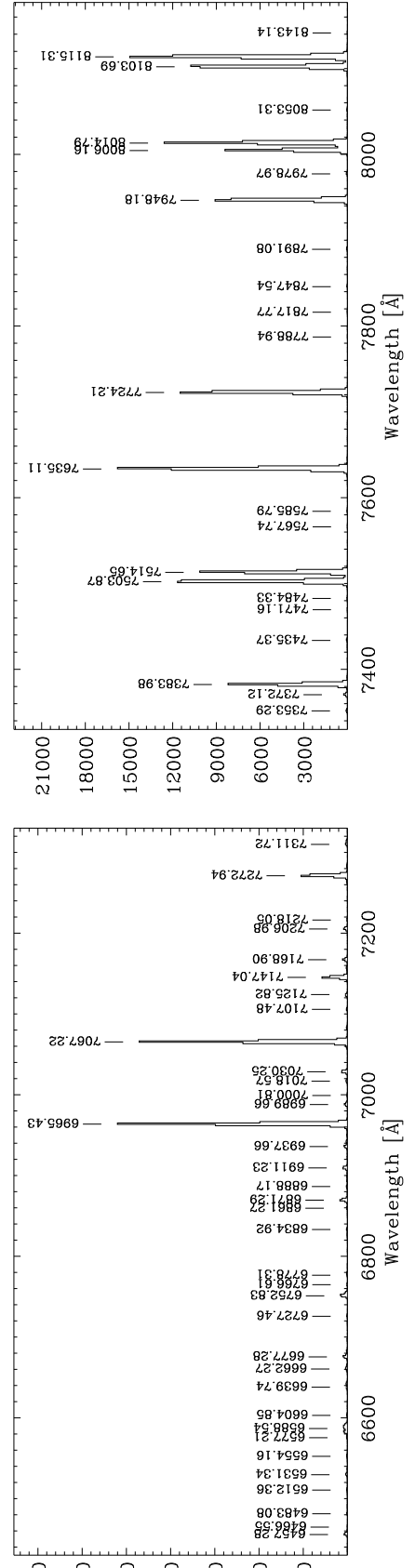
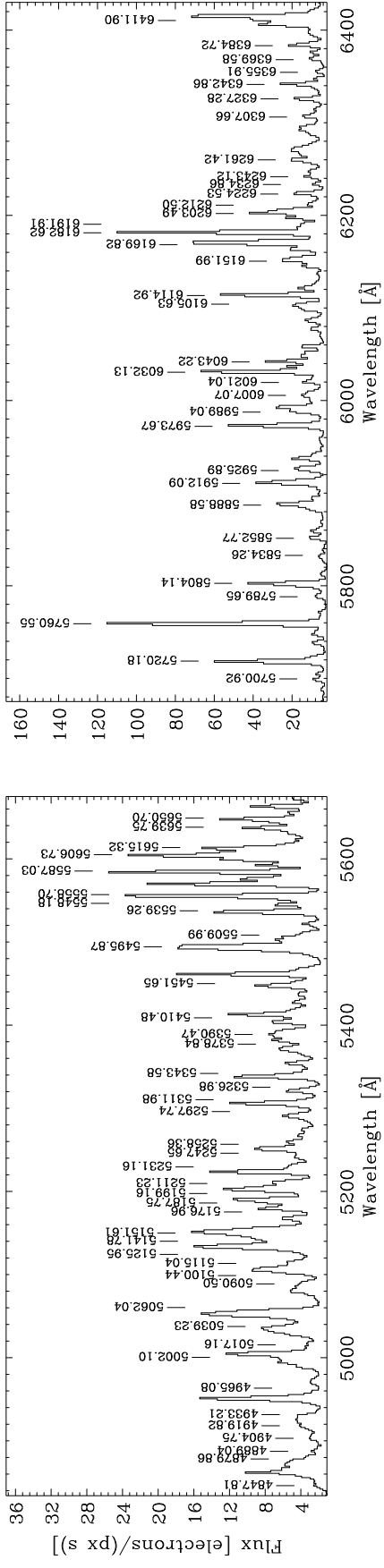
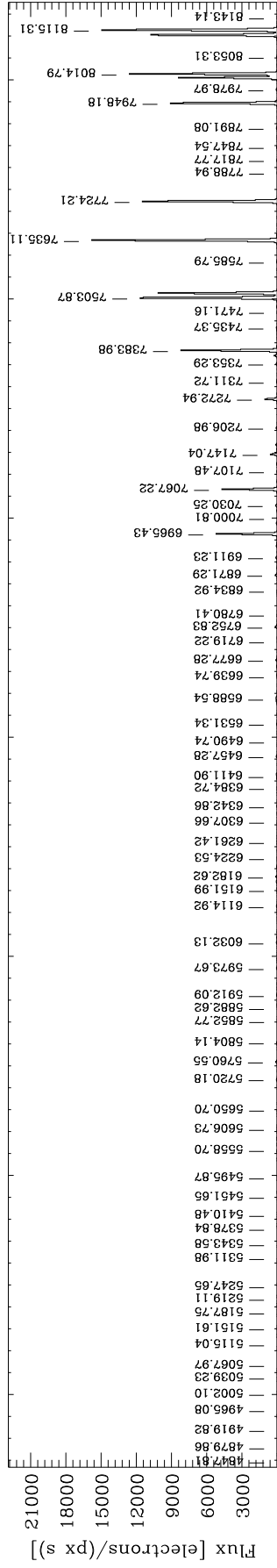
R600R

$\lambda_{\text{cen}} = 6500\text{\AA}$

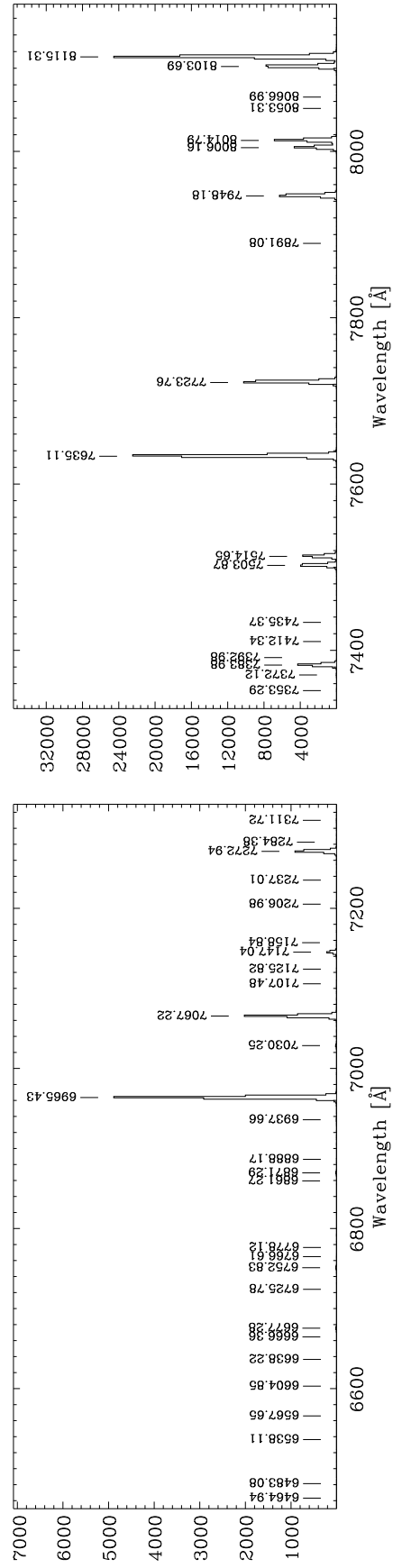
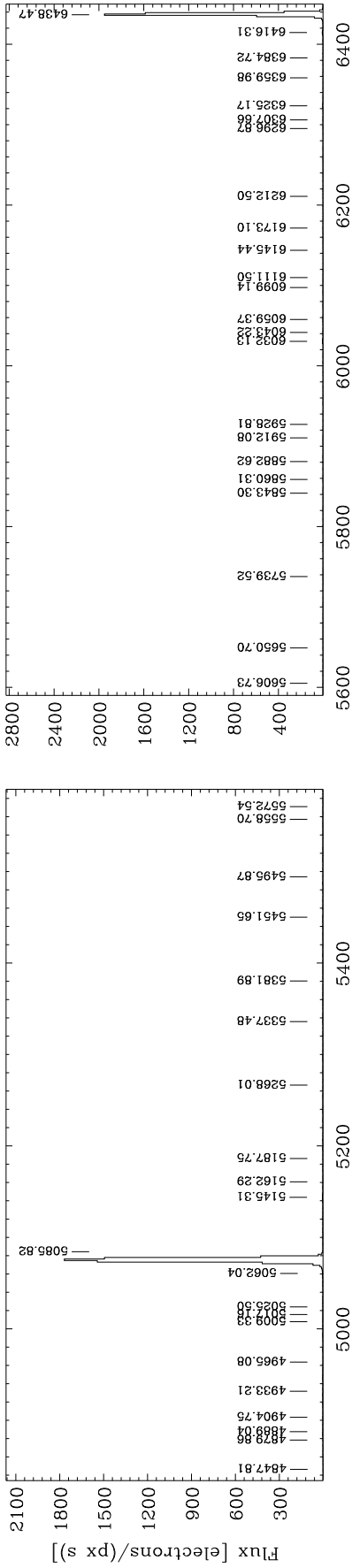
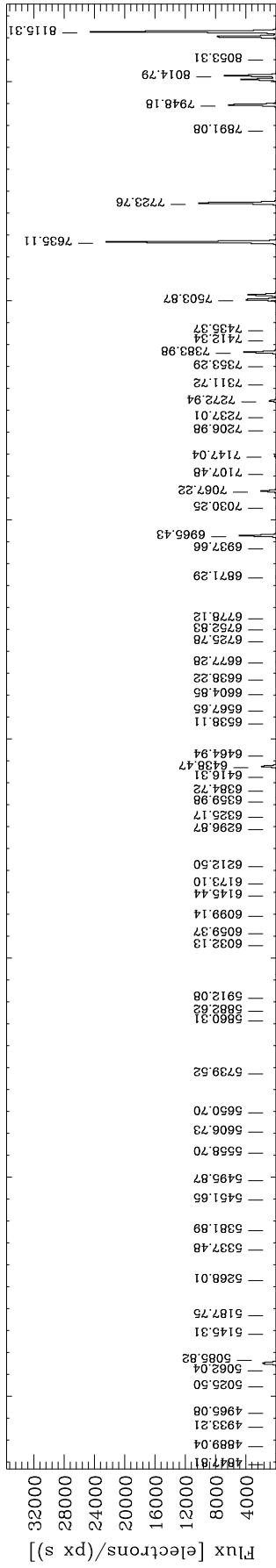
H β



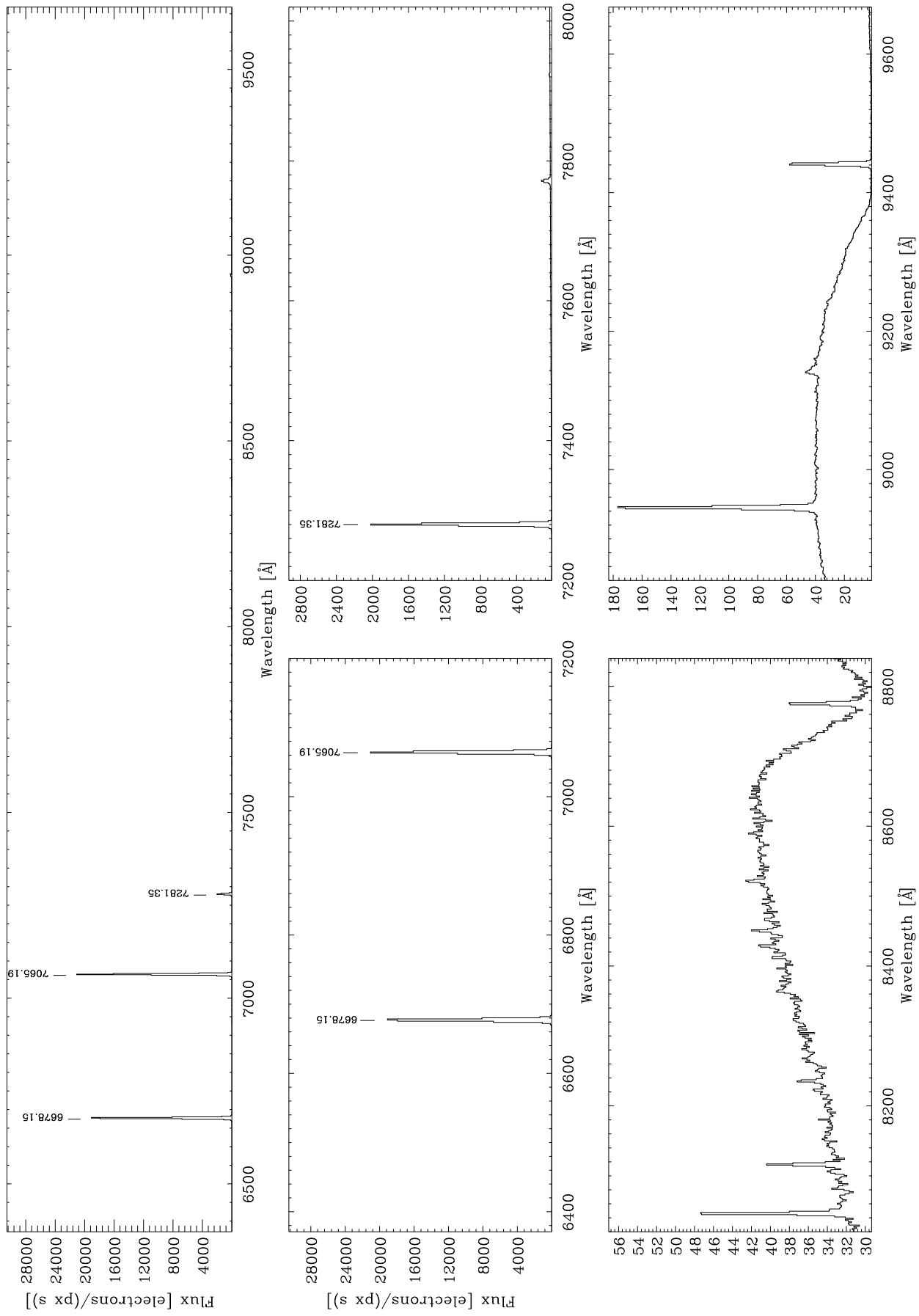
R600R $\lambda_{\text{cen}} = 6500\text{\AA}$ ThAr



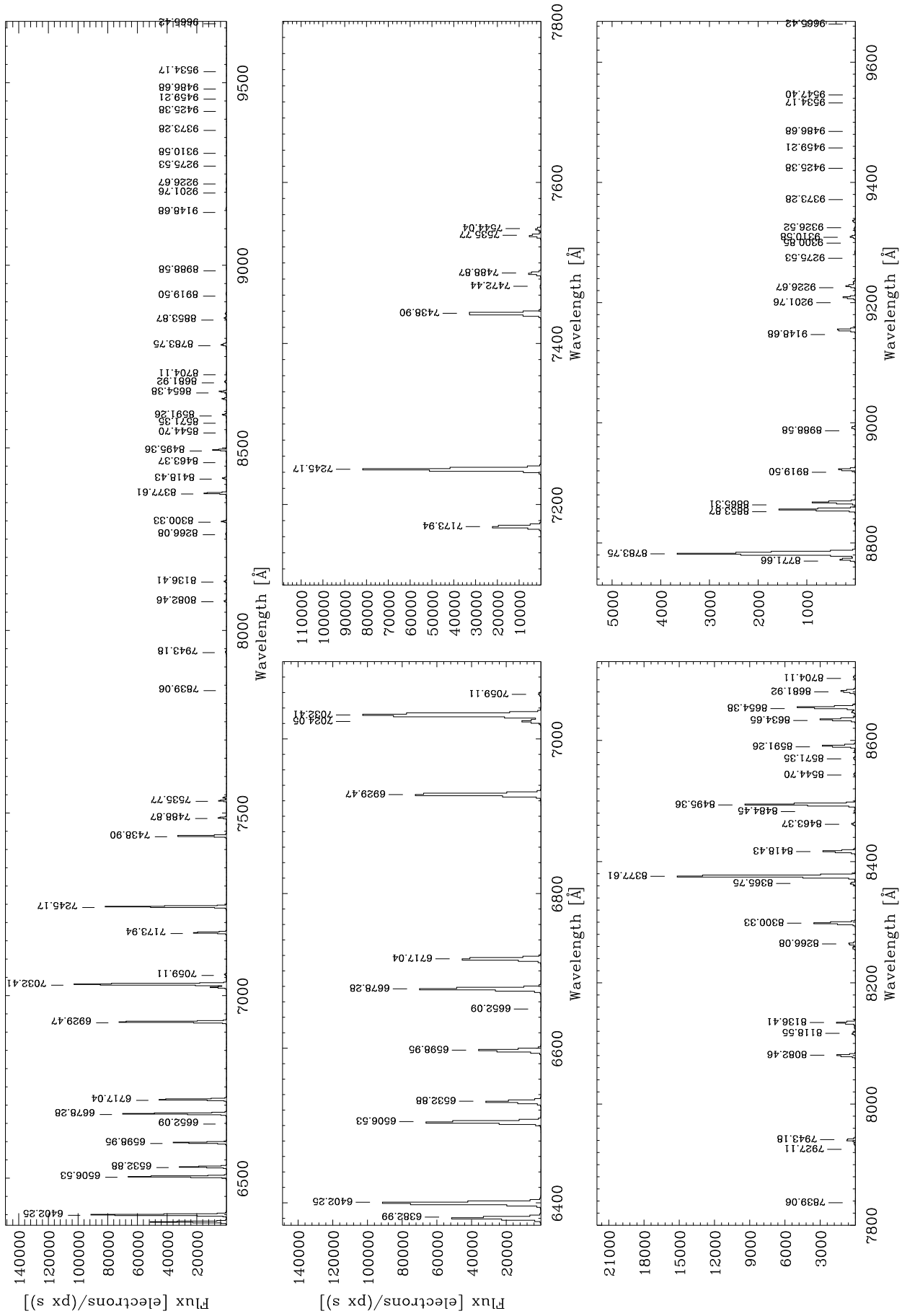
R600R $\lambda_{\text{cen}} = 6500\text{\AA}$ Cd



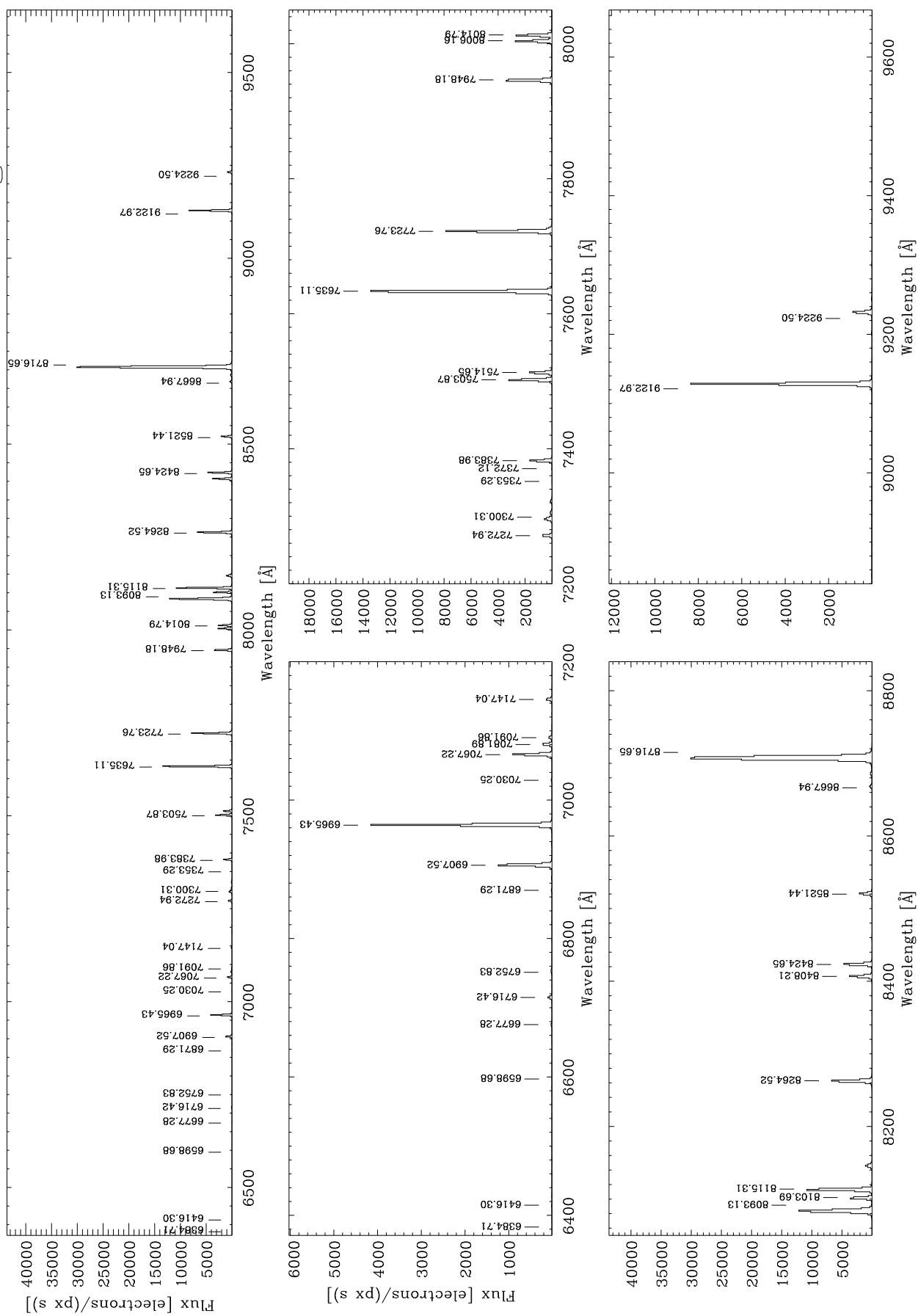
R600R $\lambda_{\text{cen}} = 8000\text{\AA}$ He



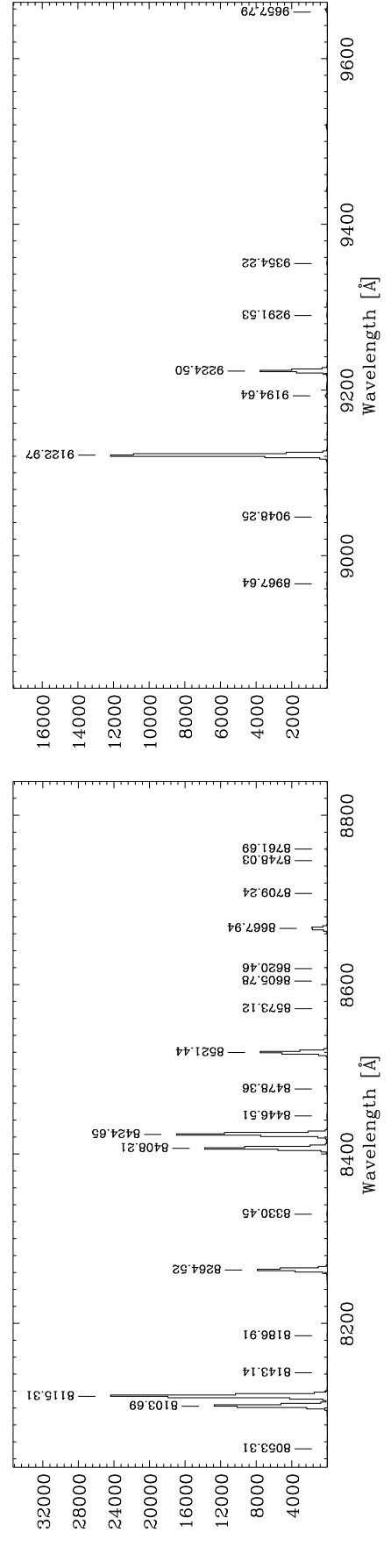
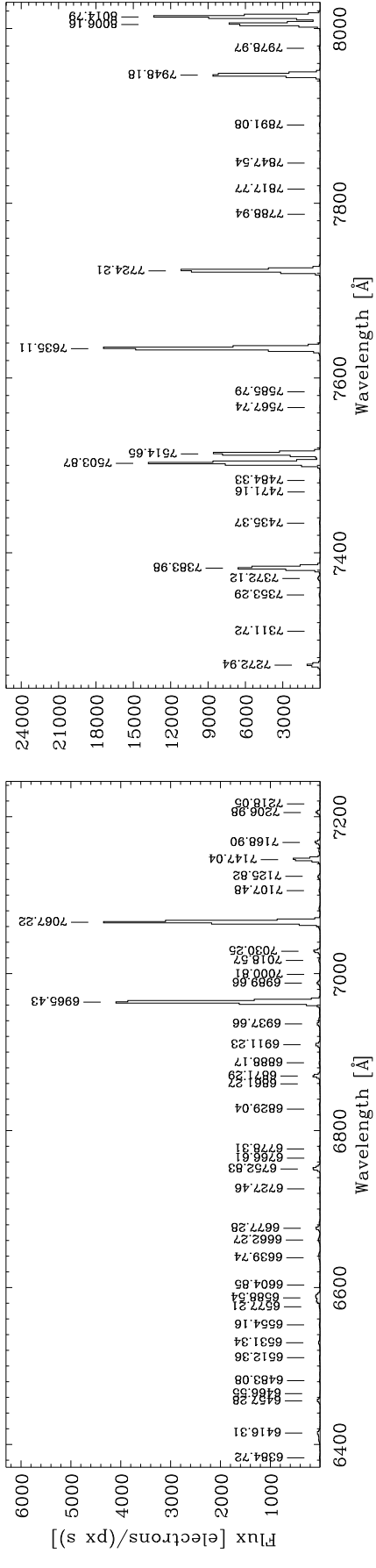
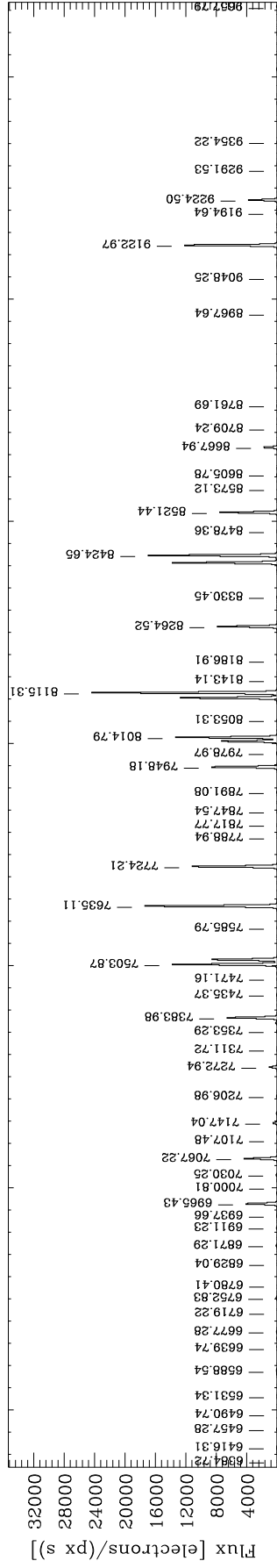
R600R $\lambda_{\text{cen}} = 8000\text{\AA}$ Ne



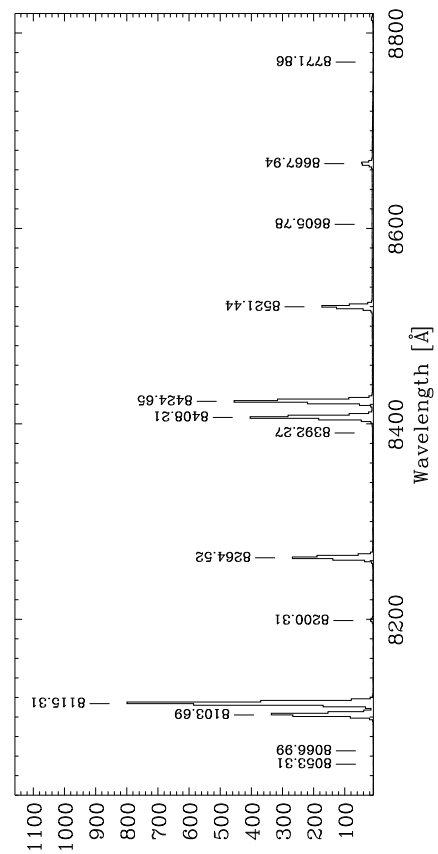
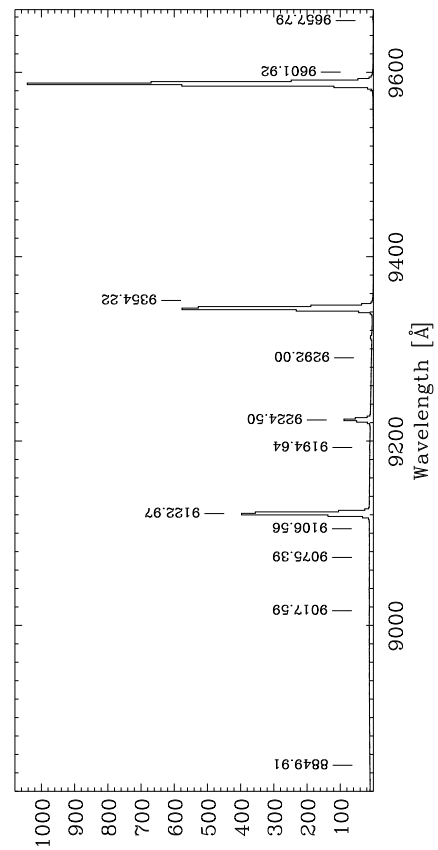
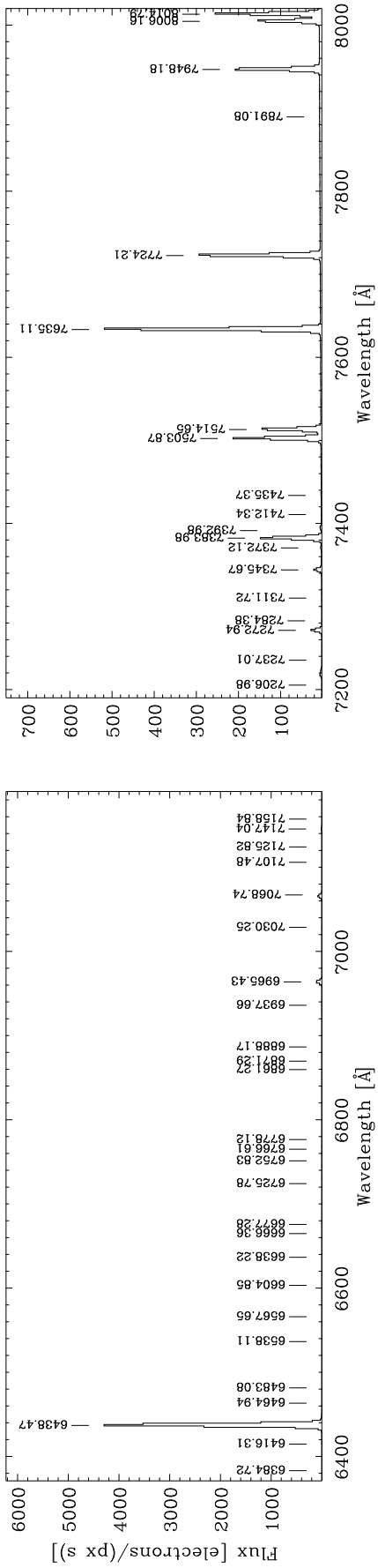
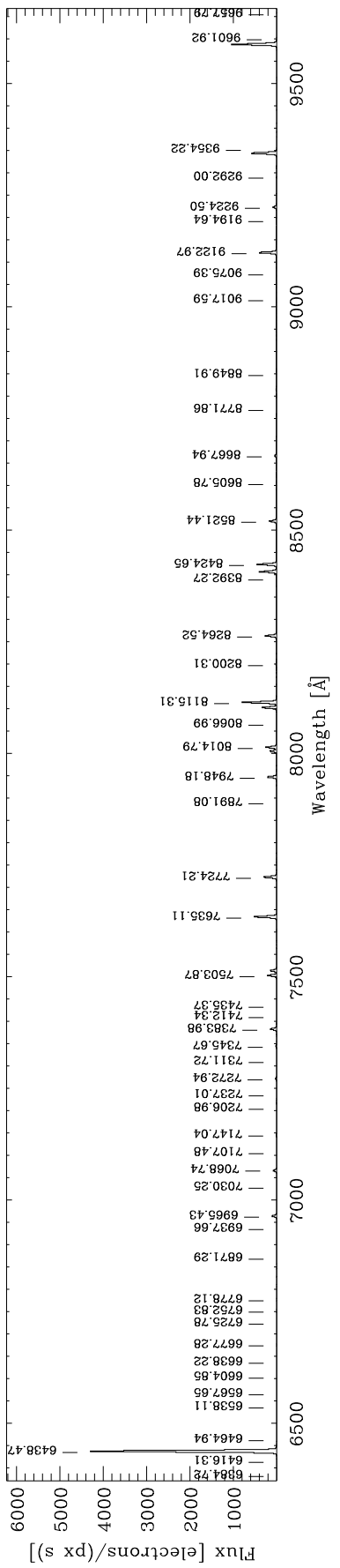
R600R $\lambda_{\text{cen}} = 8000\text{\AA}$ Hg



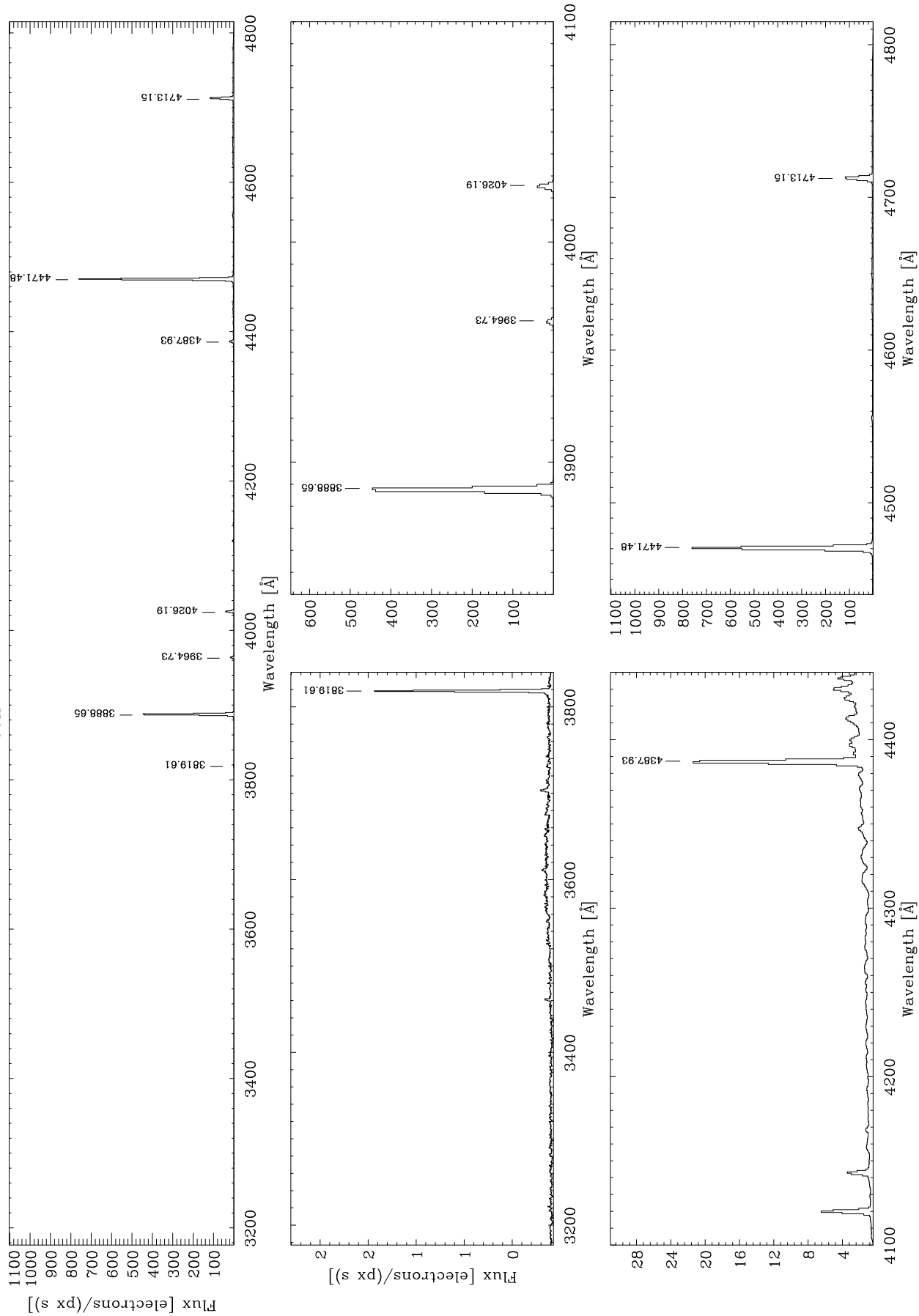
R600R $\lambda_{\text{cen}} = 8000\text{\AA}$ ThAr



R600R $\lambda_{\text{cen}} = 8000\text{\AA}$ Cd



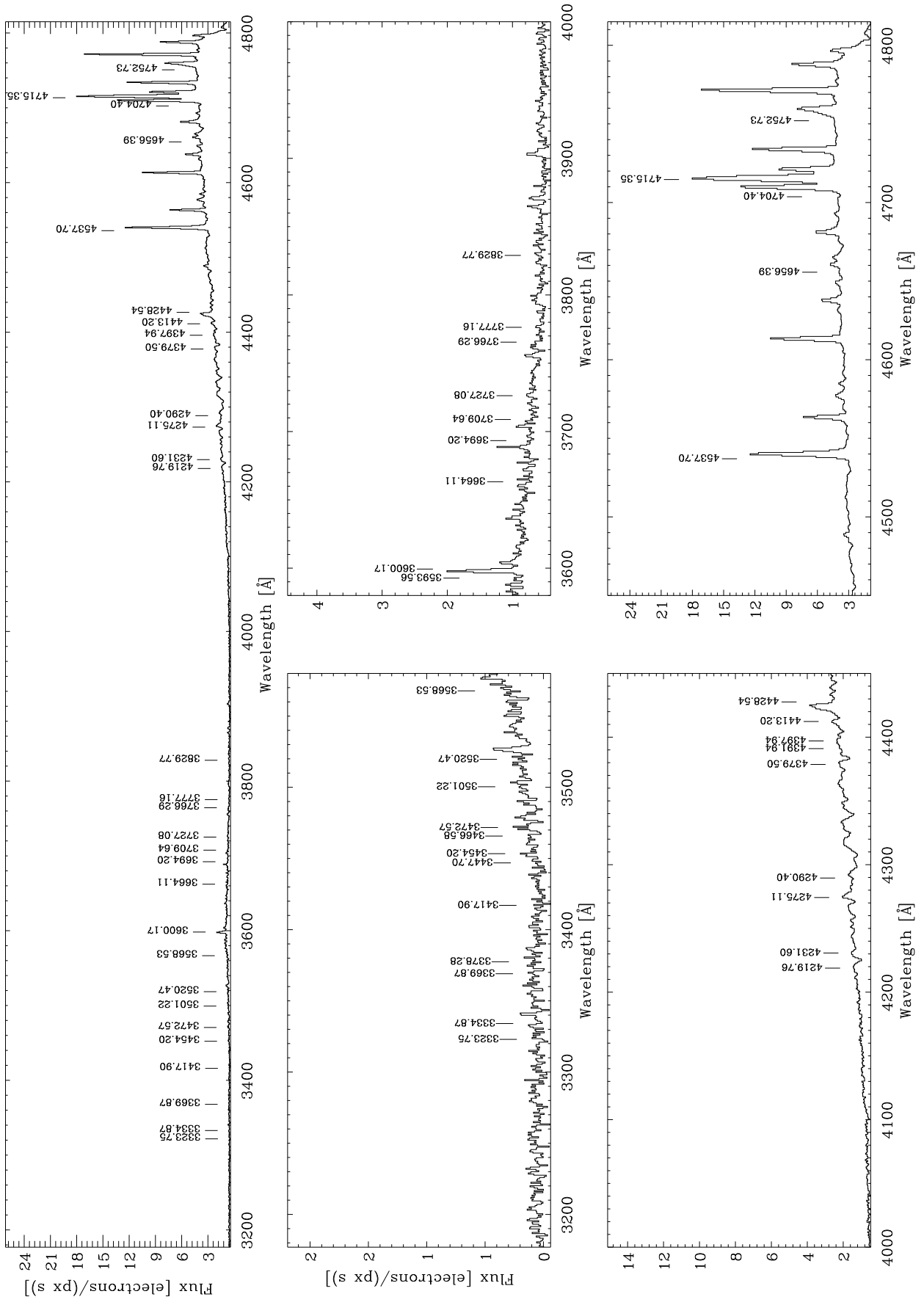
R1200B $\lambda_{\text{cen}} = 4000\text{\AA}$ He



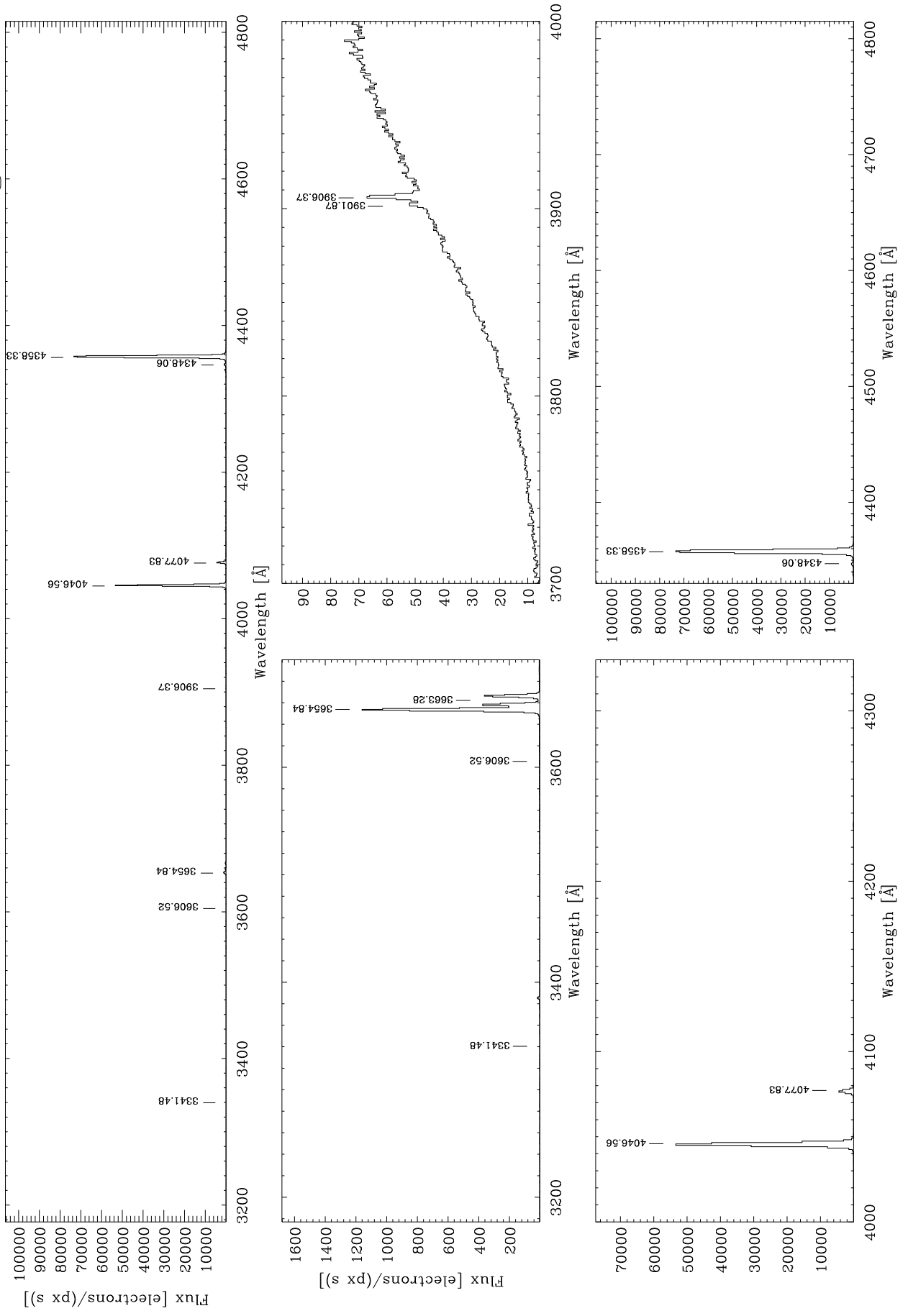
R1200B

$\lambda_{\text{cen}} = 4000\text{\AA}$

Ne



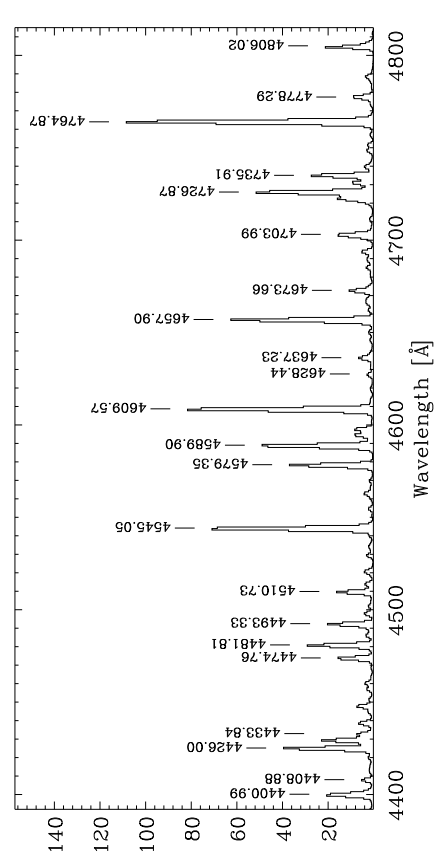
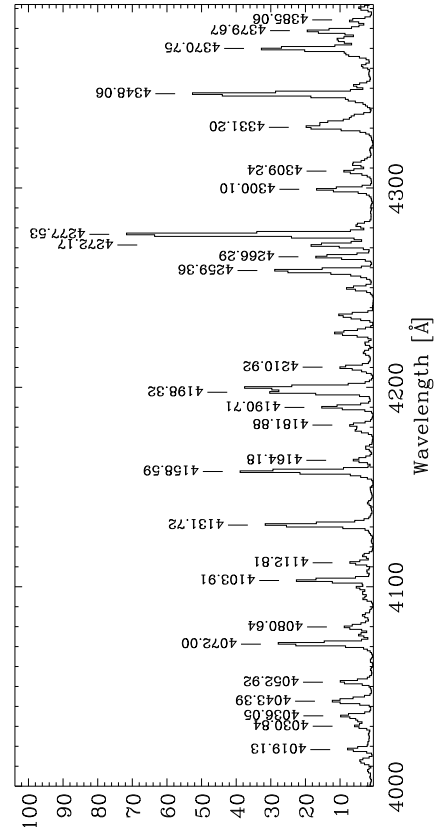
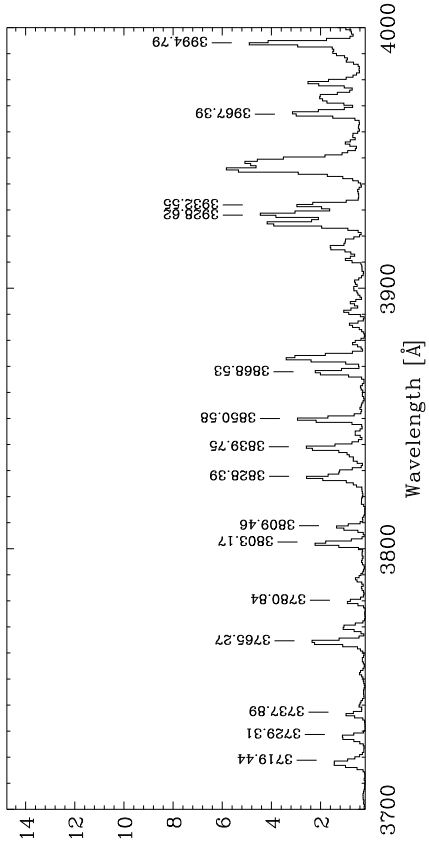
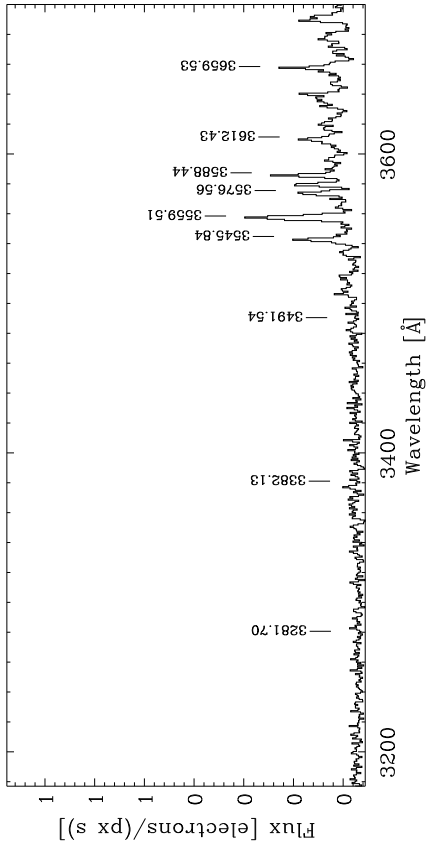
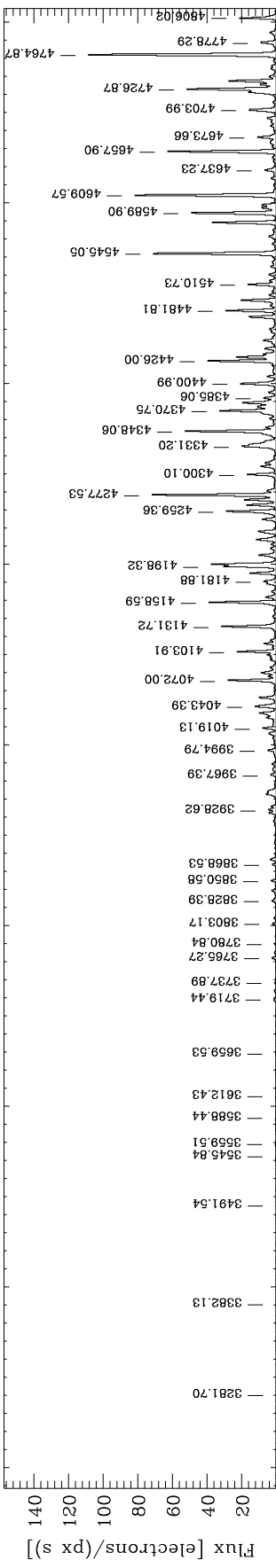
R1200B $\lambda_{\text{cen}} = 4000\text{\AA}$ H γ



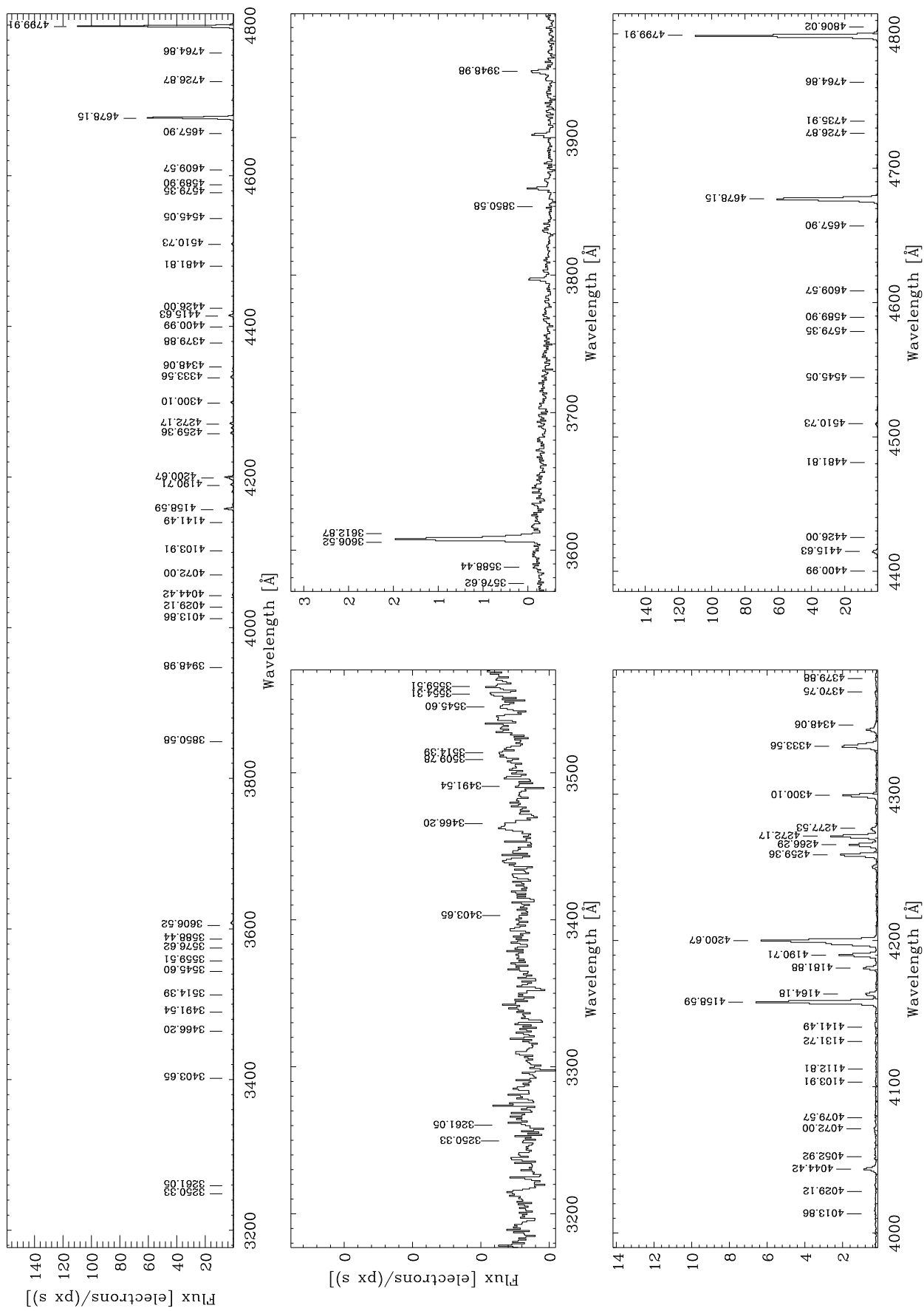
R1200B

$\lambda_{\text{cen}} = 4000\text{\AA}$

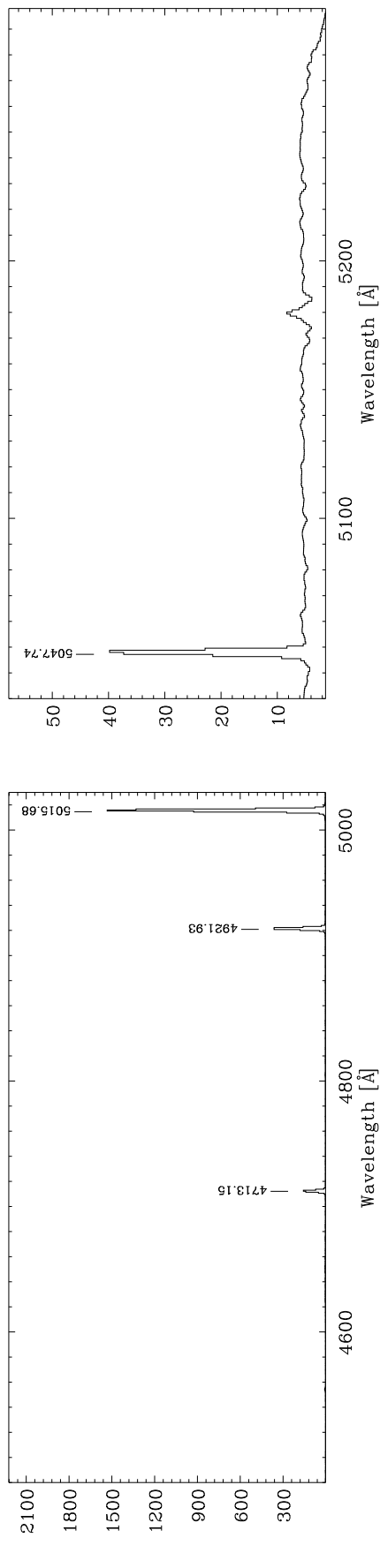
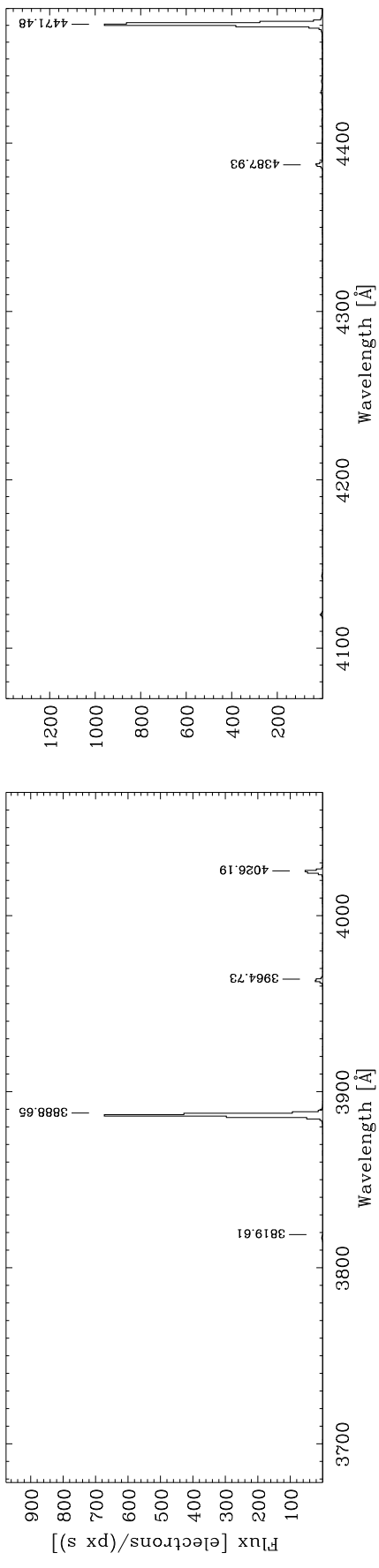
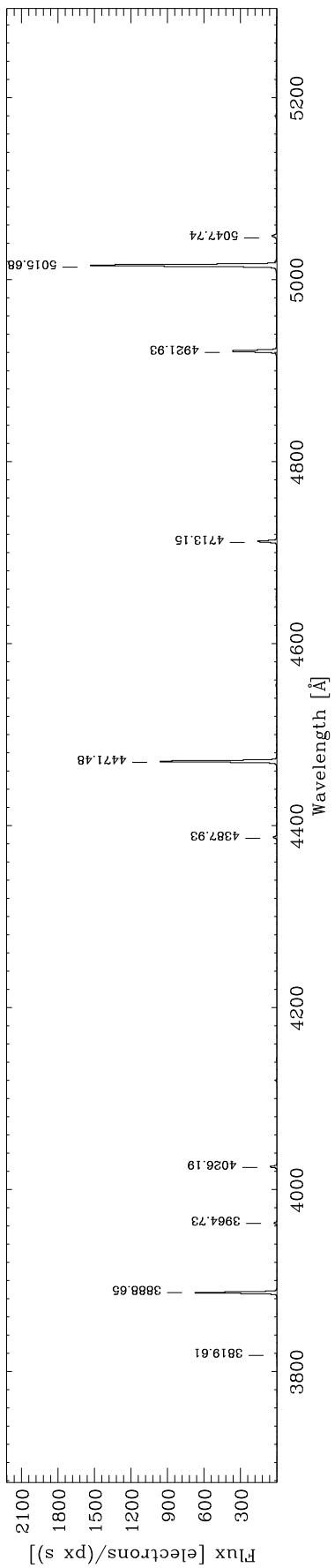
ThAr



R1200B $\lambda_{\text{cen}} = 4000\text{\AA}$ Cd



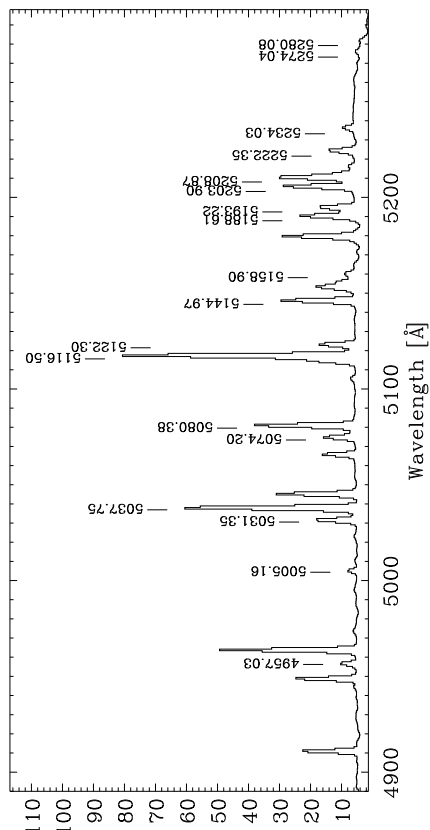
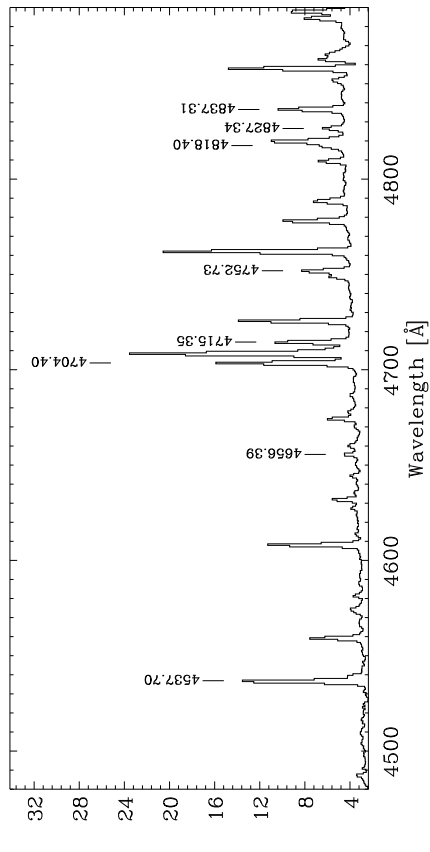
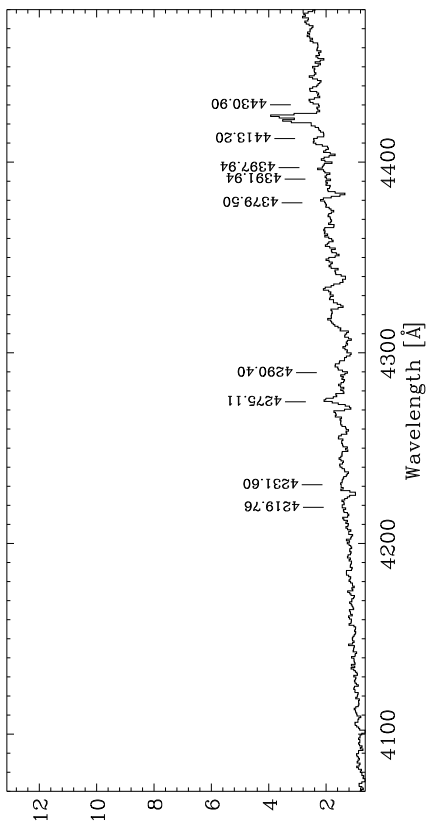
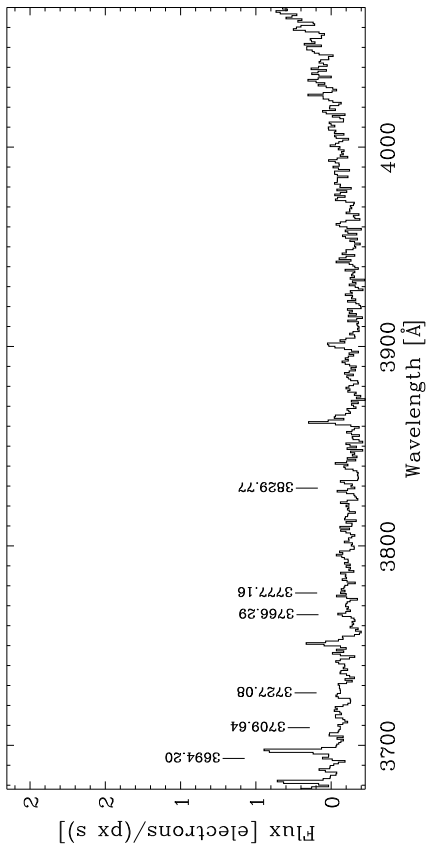
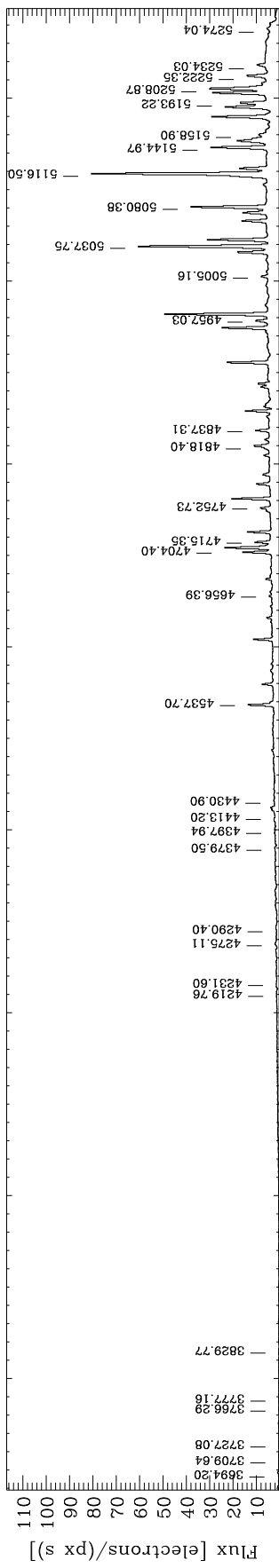
R1200B $\lambda_{\text{cen}} = 4500\text{\AA}$ He



R1200B

$\lambda_{\text{cen}} = 4500\text{\AA}$

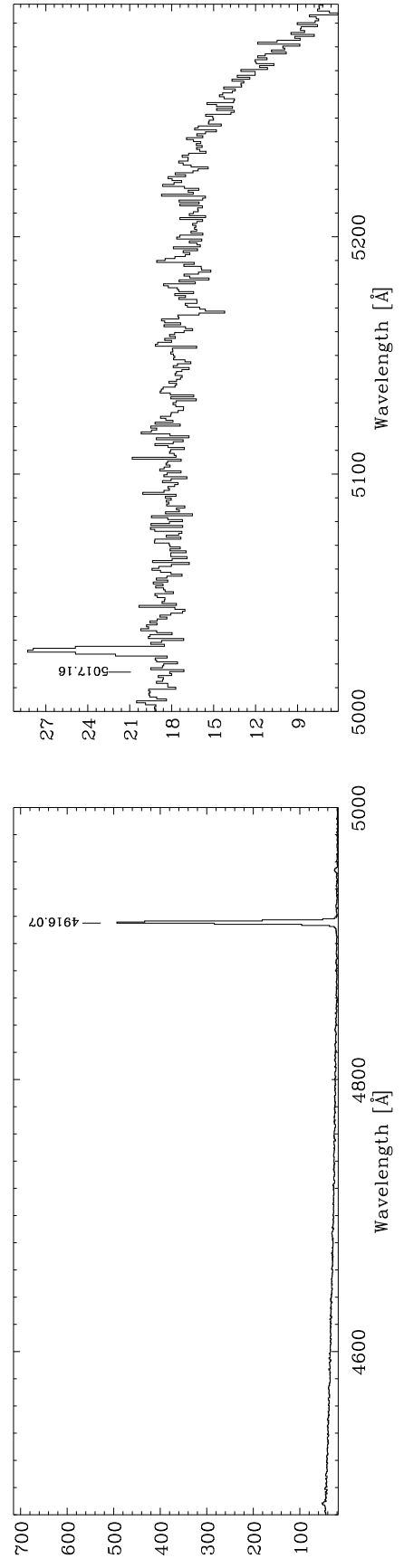
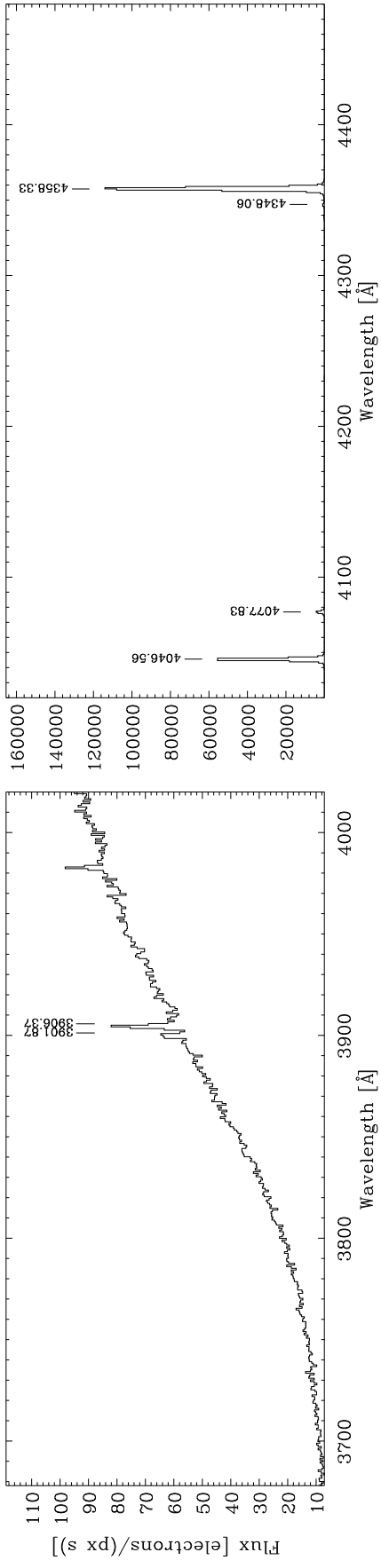
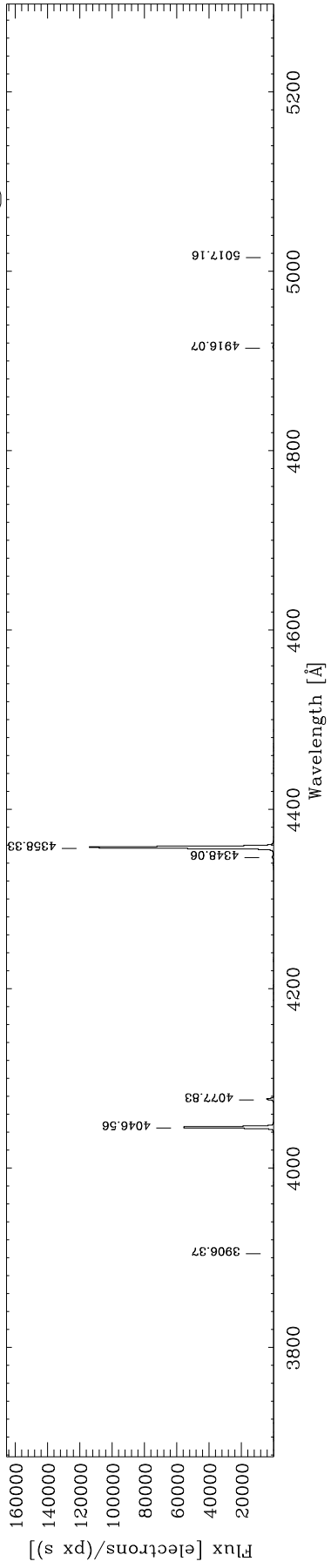
Ne



R1200B

$\lambda_{\text{cen}} = 4500\text{\AA}$

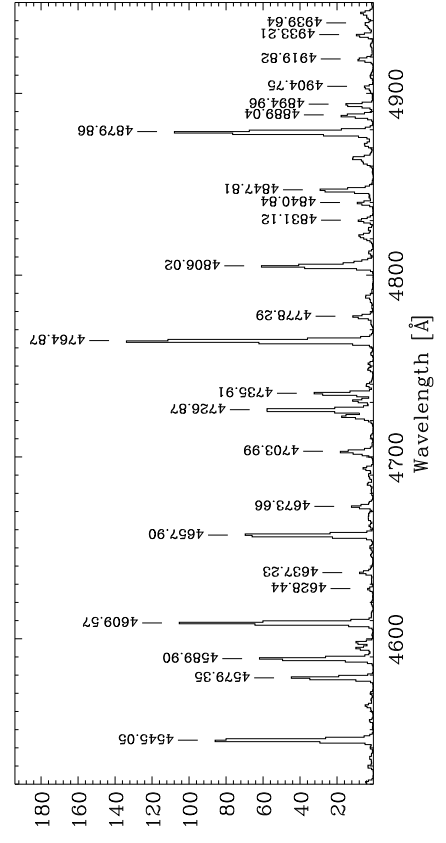
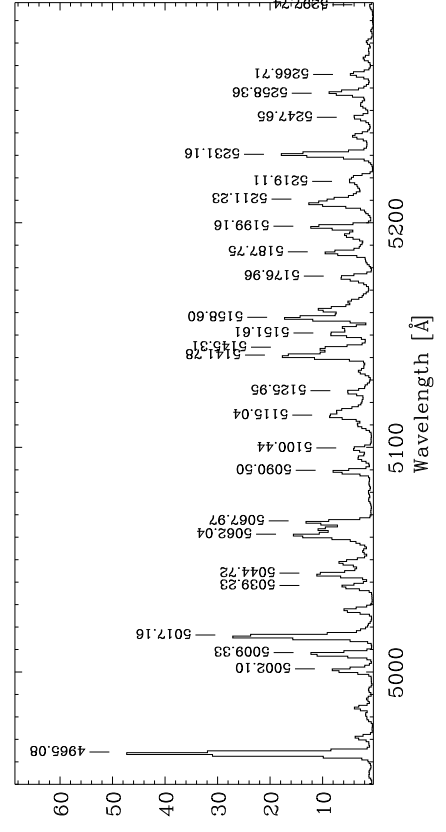
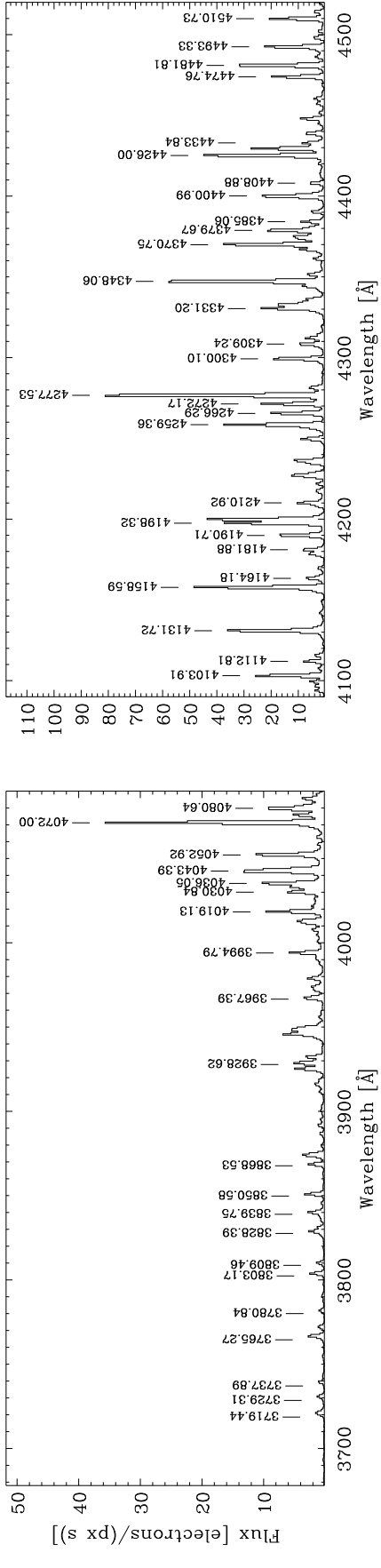
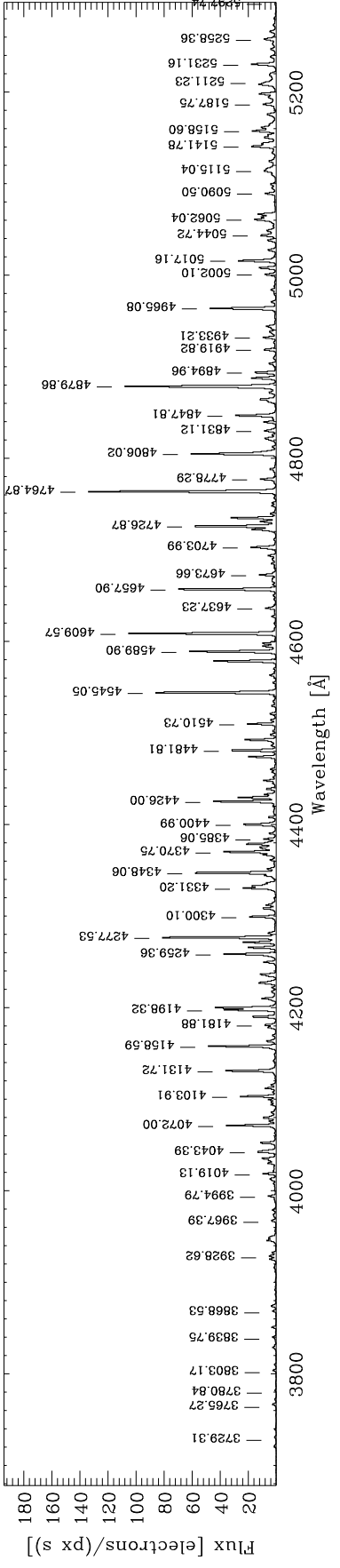
Hg



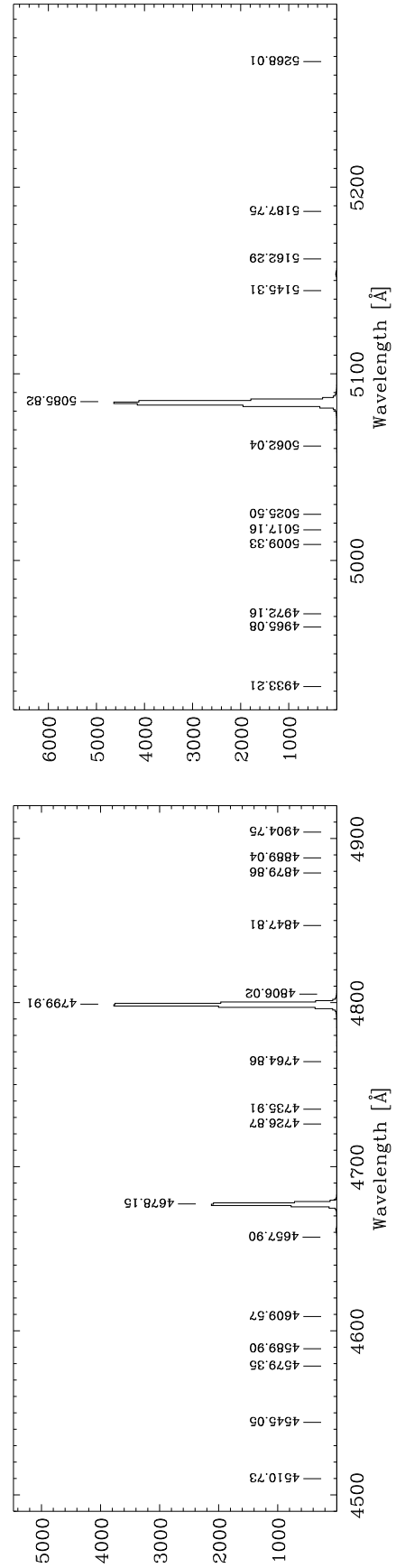
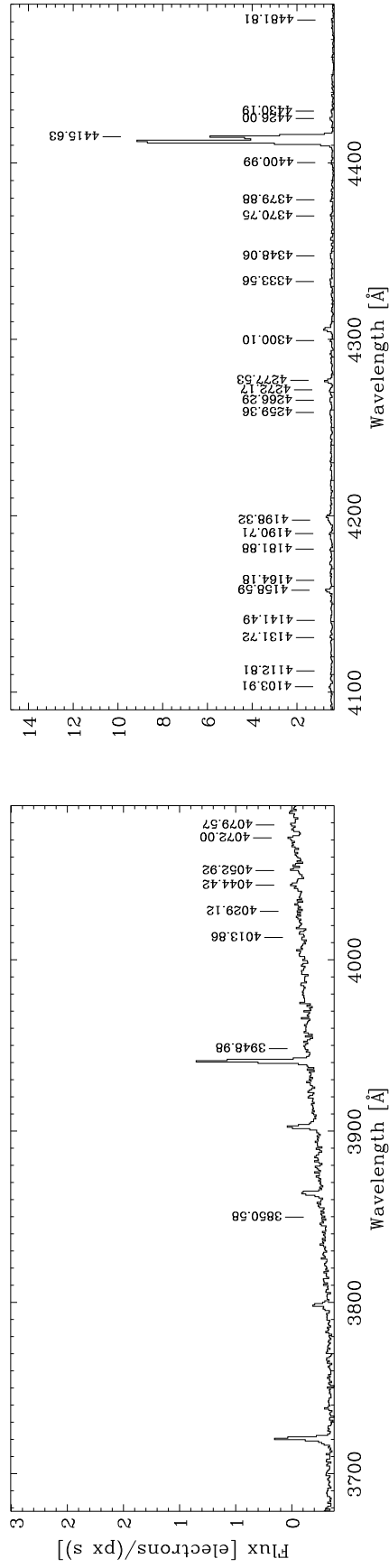
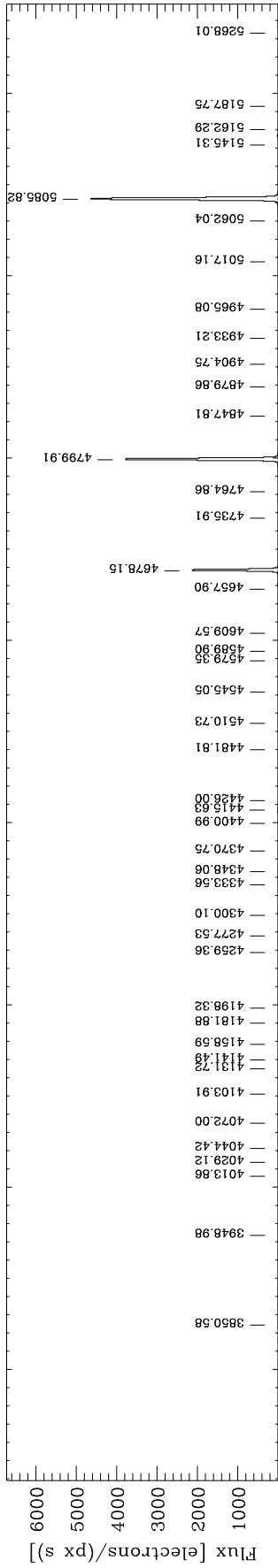
R1200B

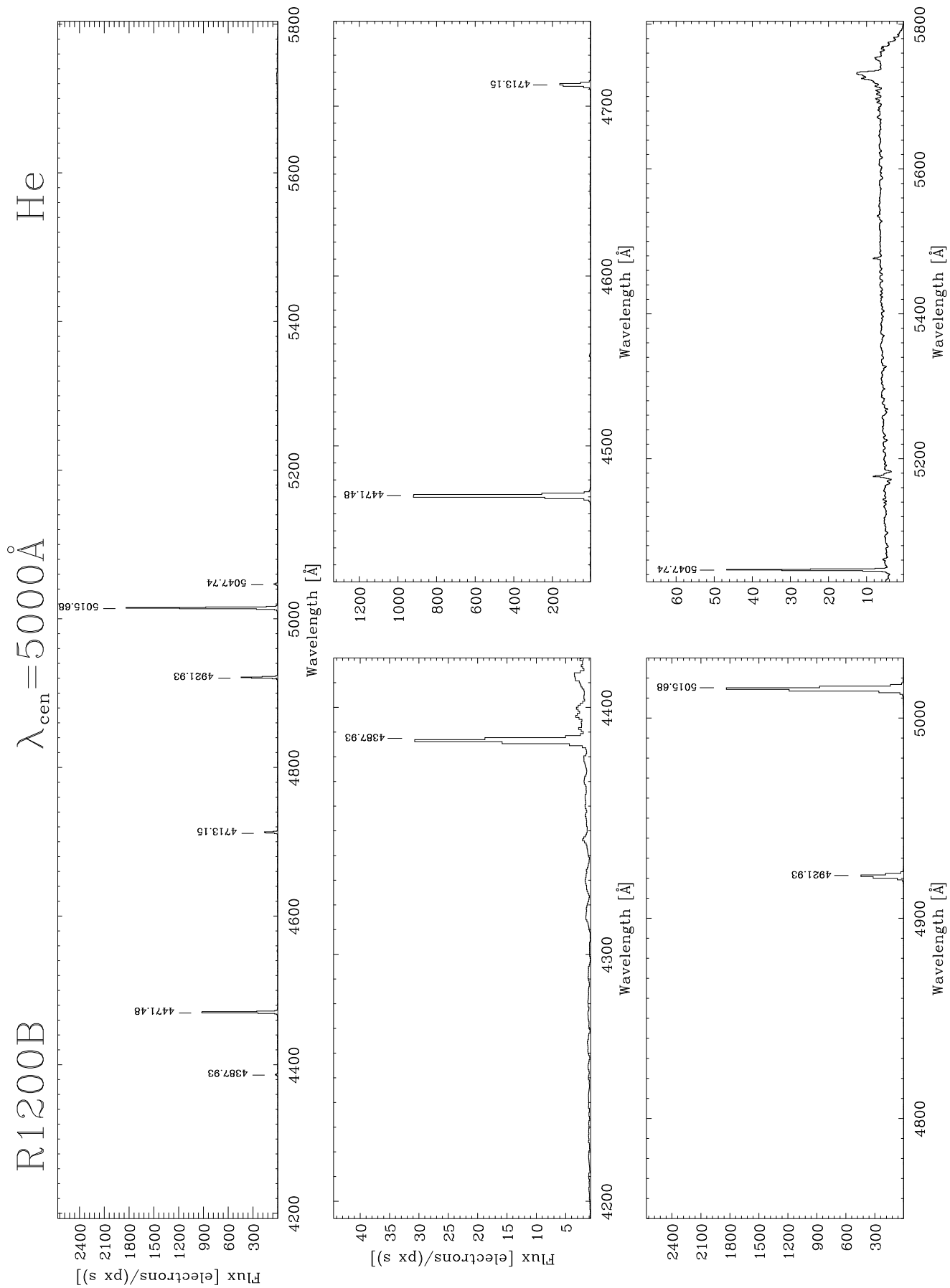
$\lambda_{\text{cen}} = 4500\text{\AA}$

ThAr



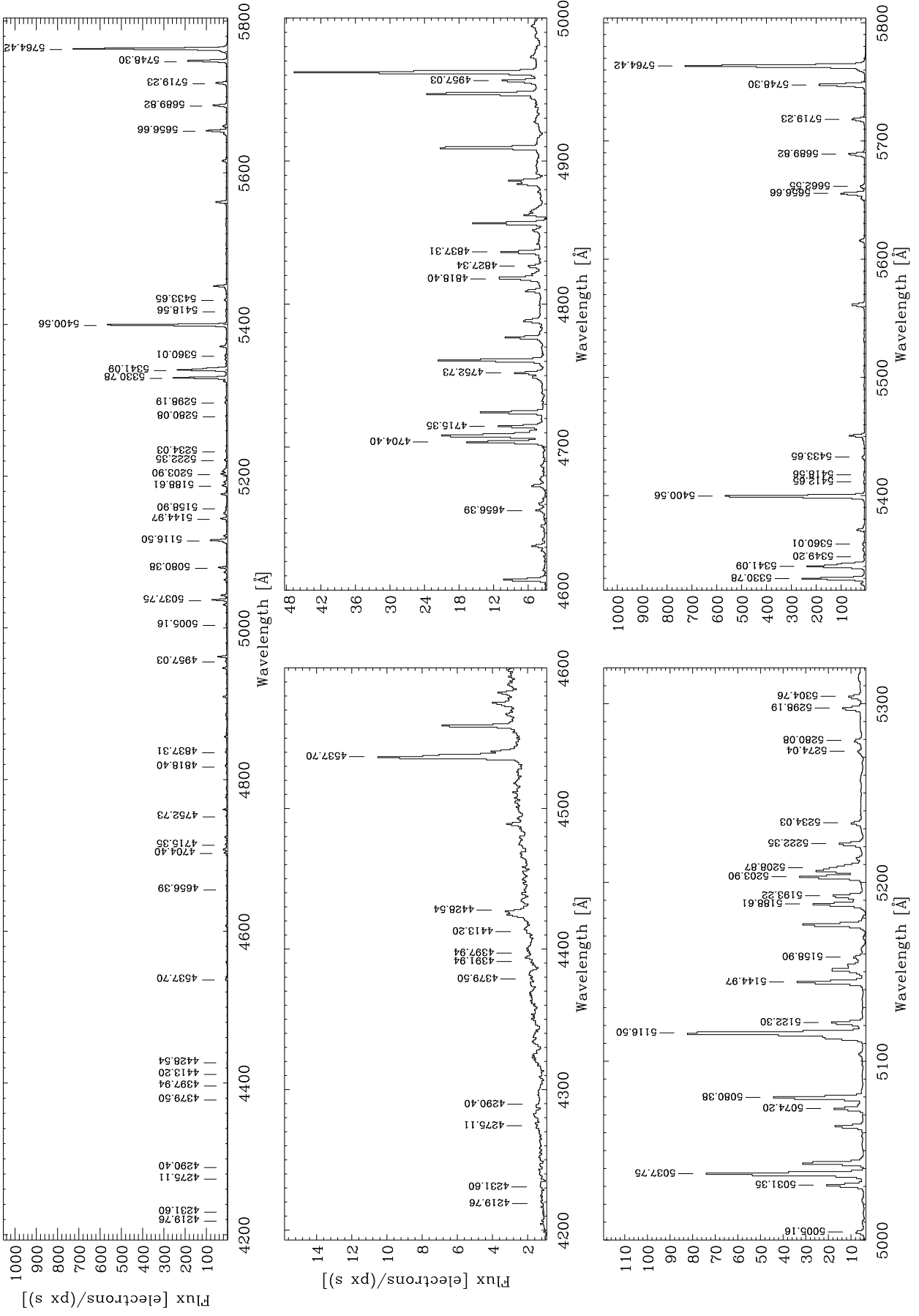
R1200B $\lambda_{\text{cen}} = 4500\text{\AA}$ Cd





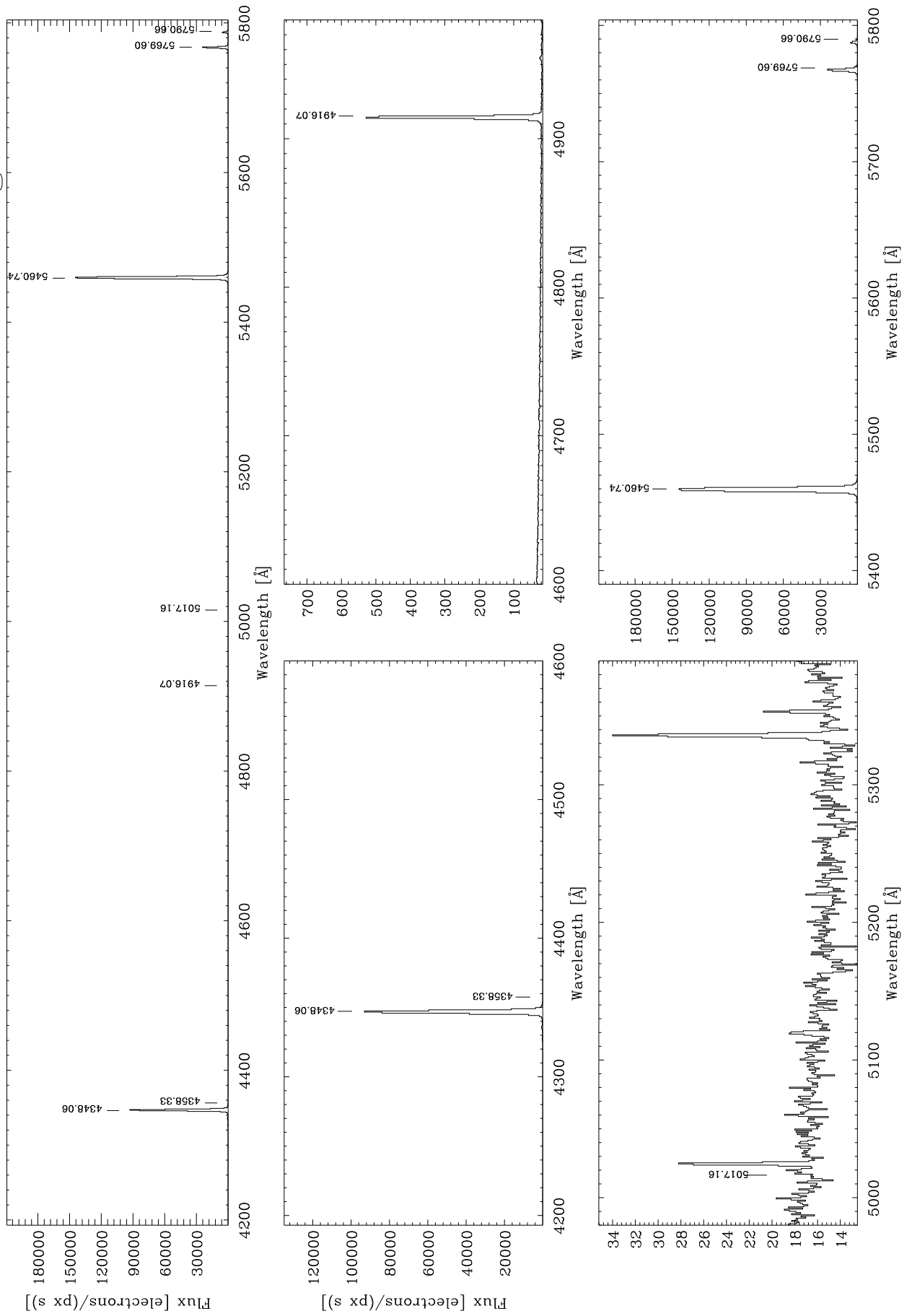
R1200B $\lambda_{\text{cen}} = 5000\text{\AA}$

Ne



R1200B $\lambda_{\text{cen}} = 5000\text{\AA}$

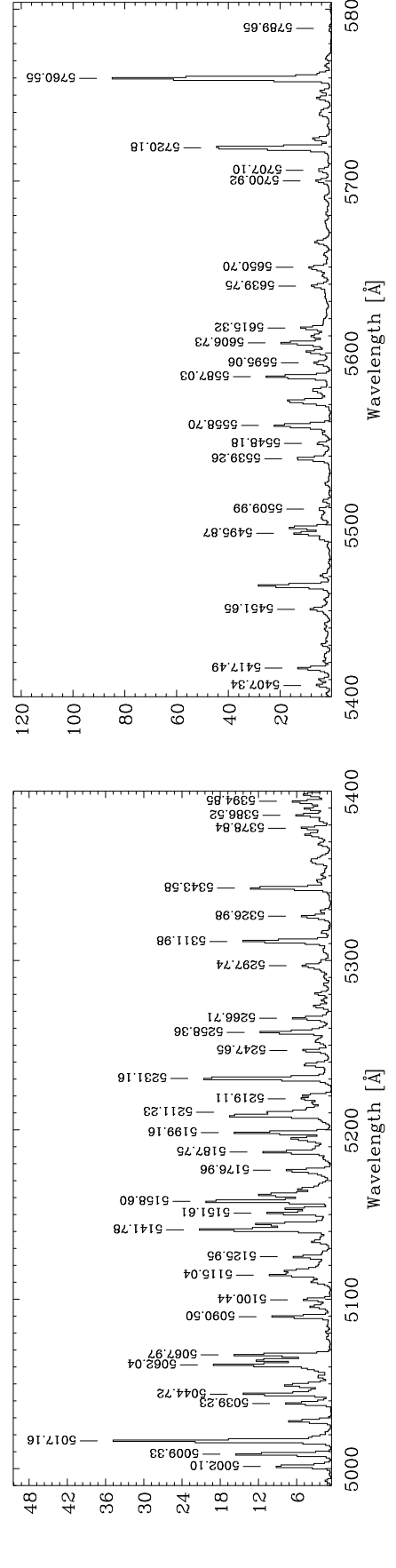
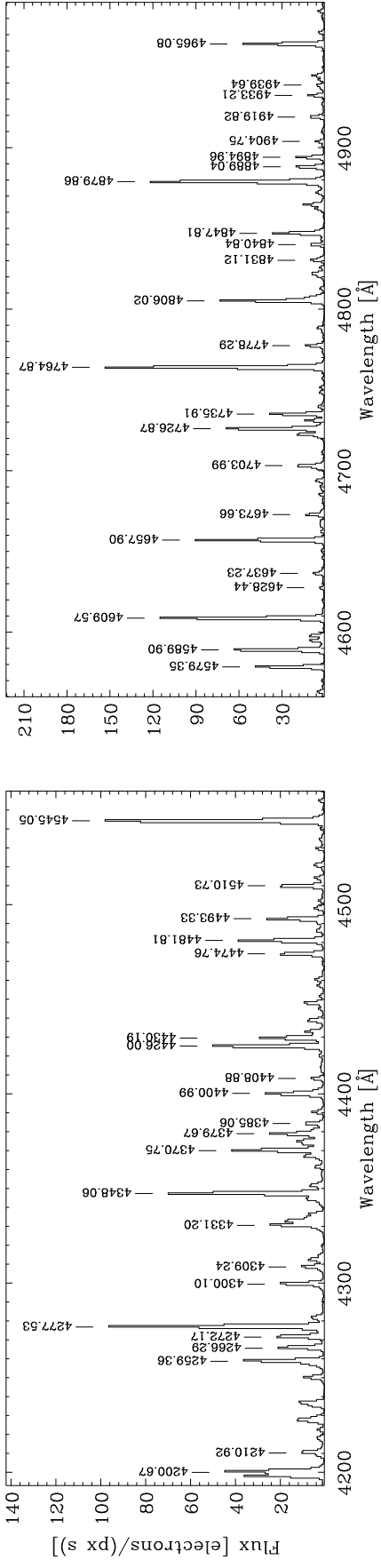
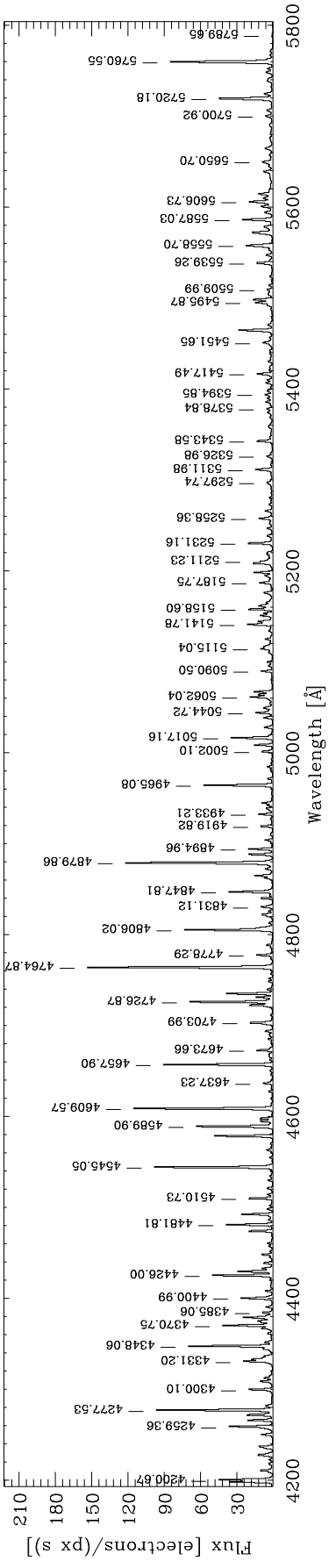
Hg



R1200B

$\lambda_{\text{cen}} = 5000\text{\AA}$

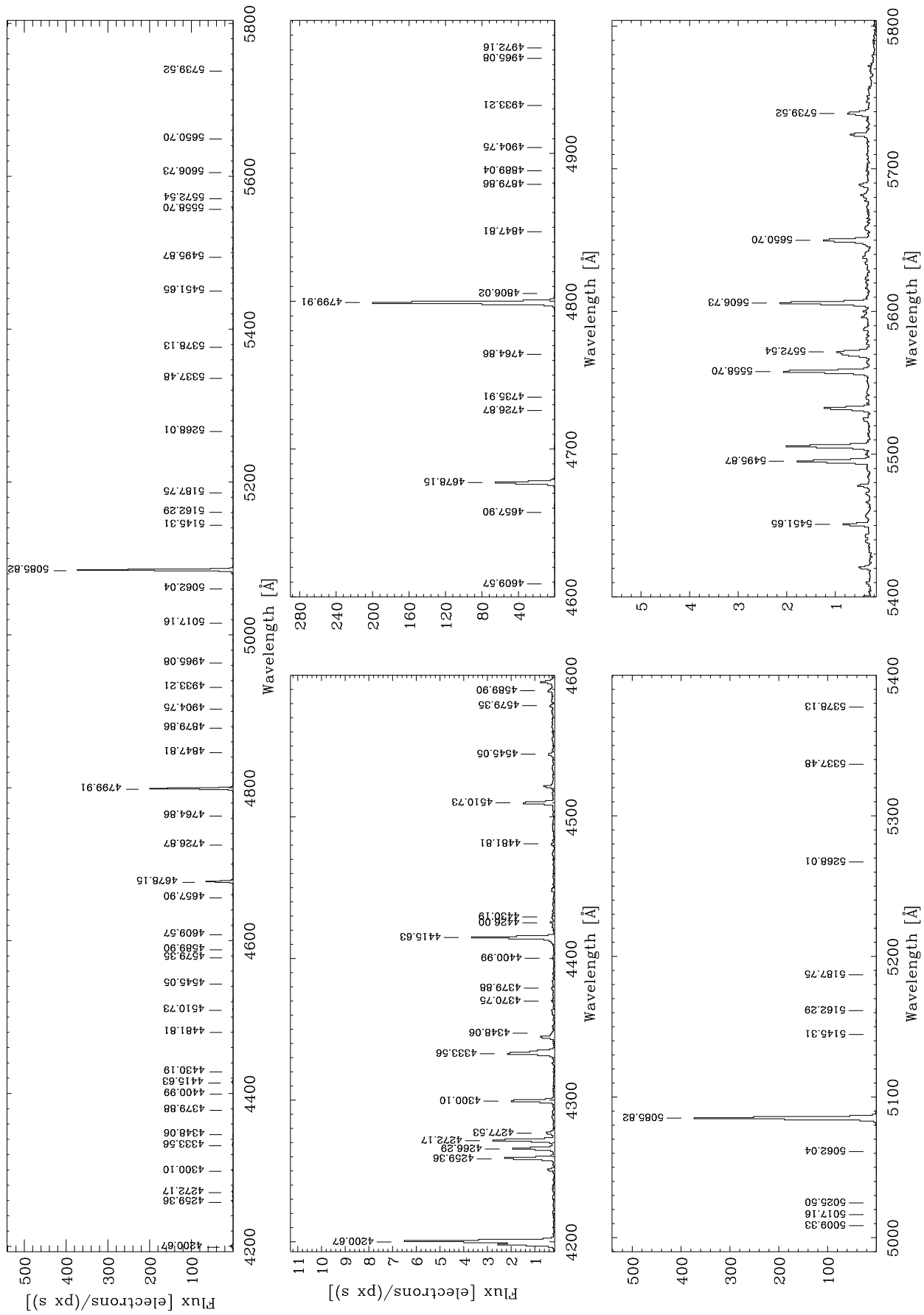
ThAr



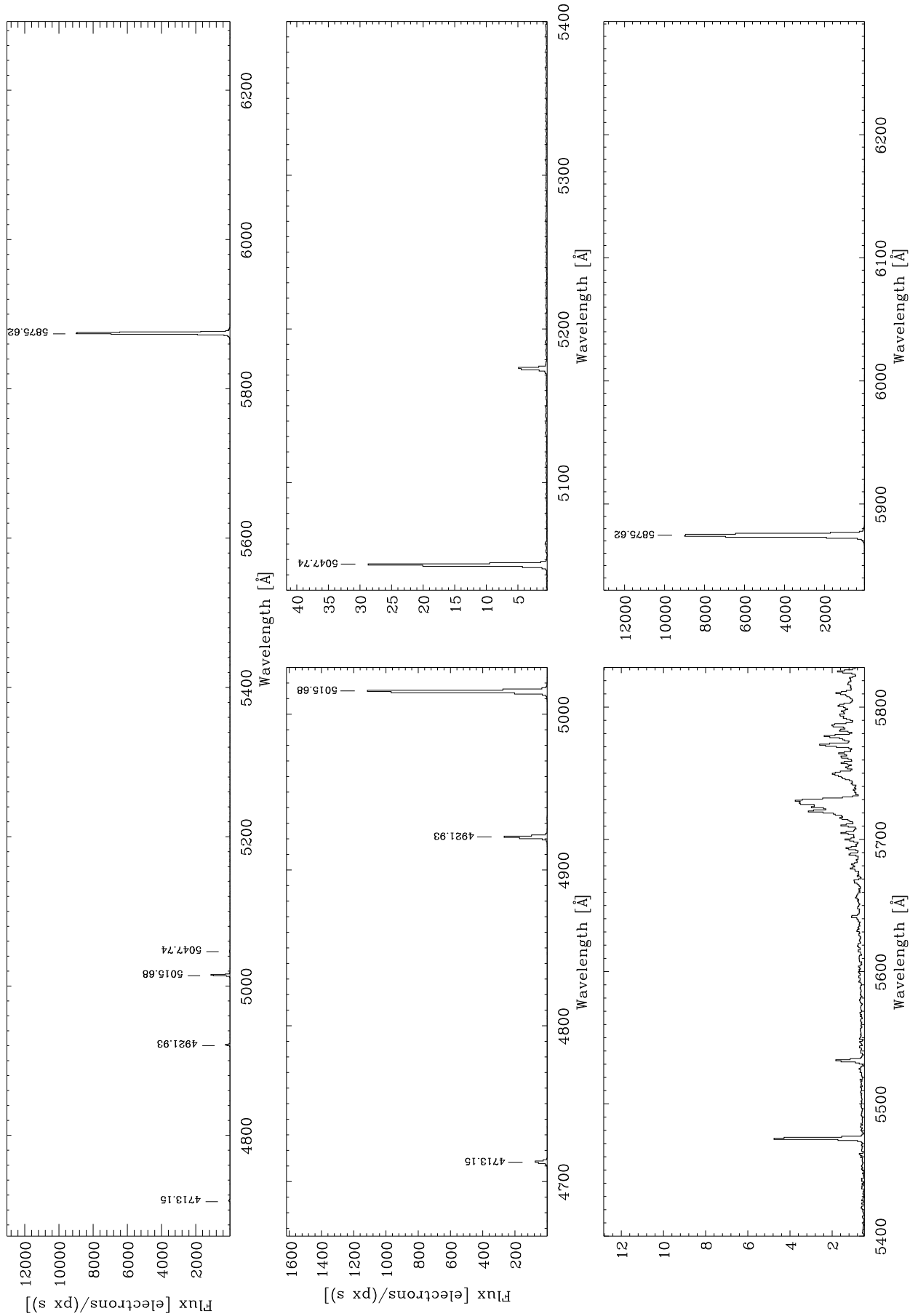
R1200B

$\lambda_{\text{cen}} = 5000\text{\AA}$

Cd



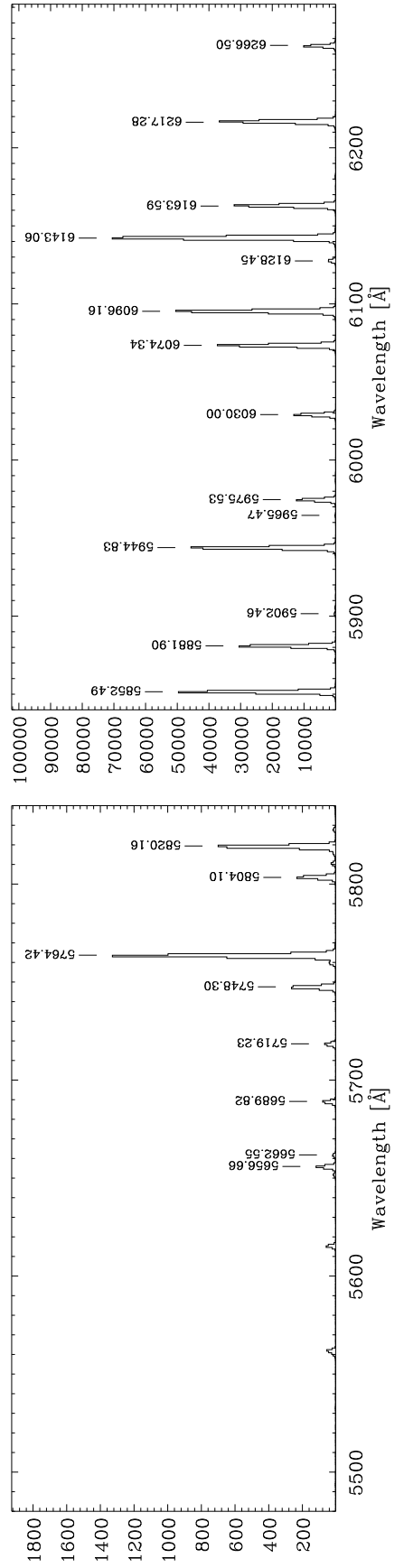
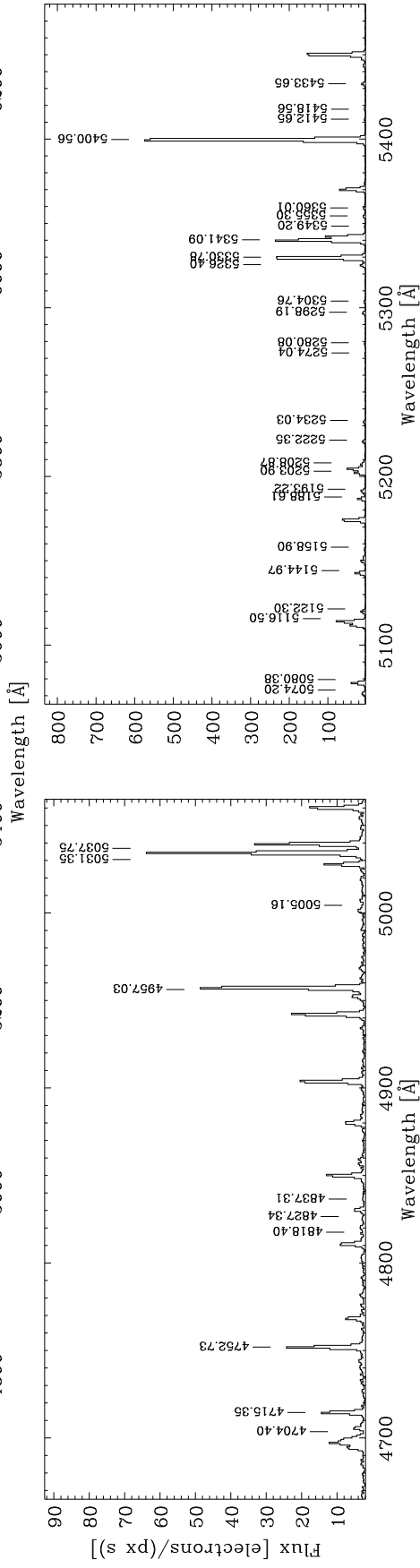
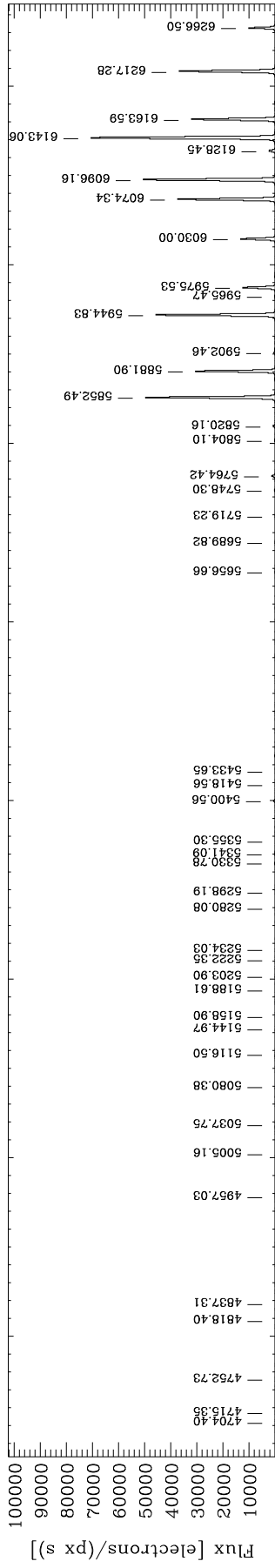
R1200R $\lambda_{\text{cen}} = 5500\text{\AA}$ He



R1200R

$\lambda_{\text{cen}} = 5500\text{\AA}$

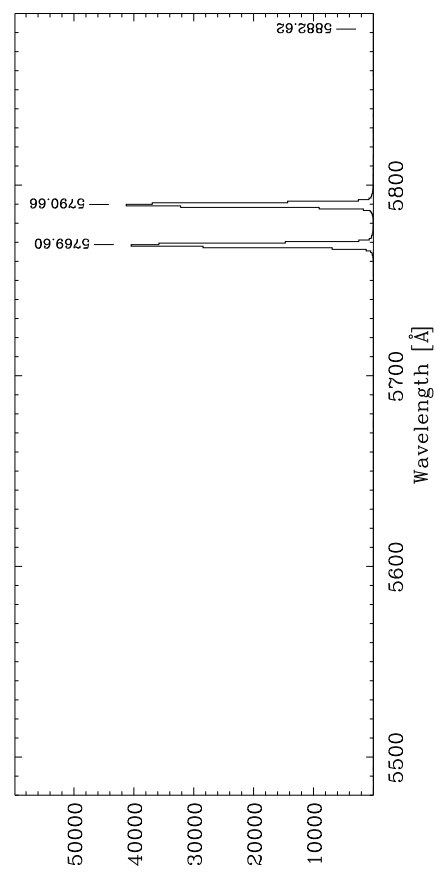
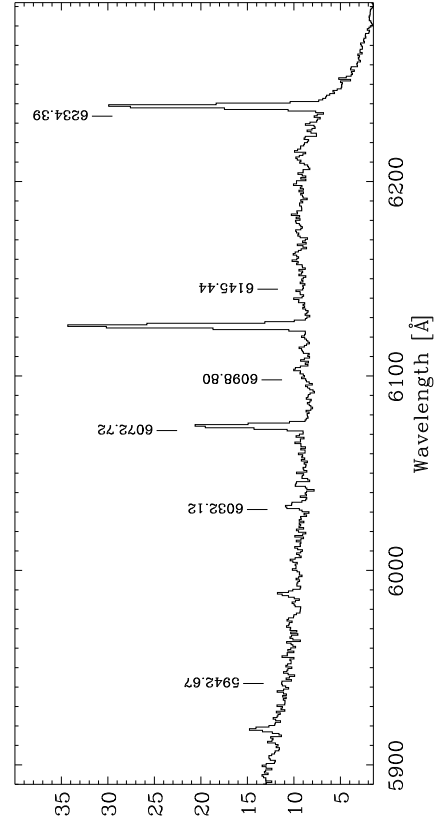
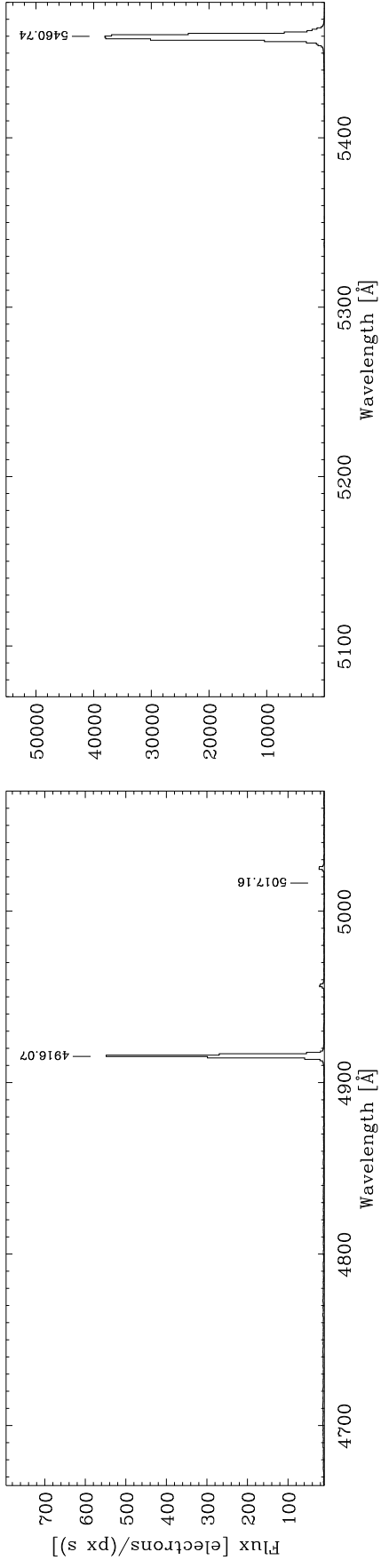
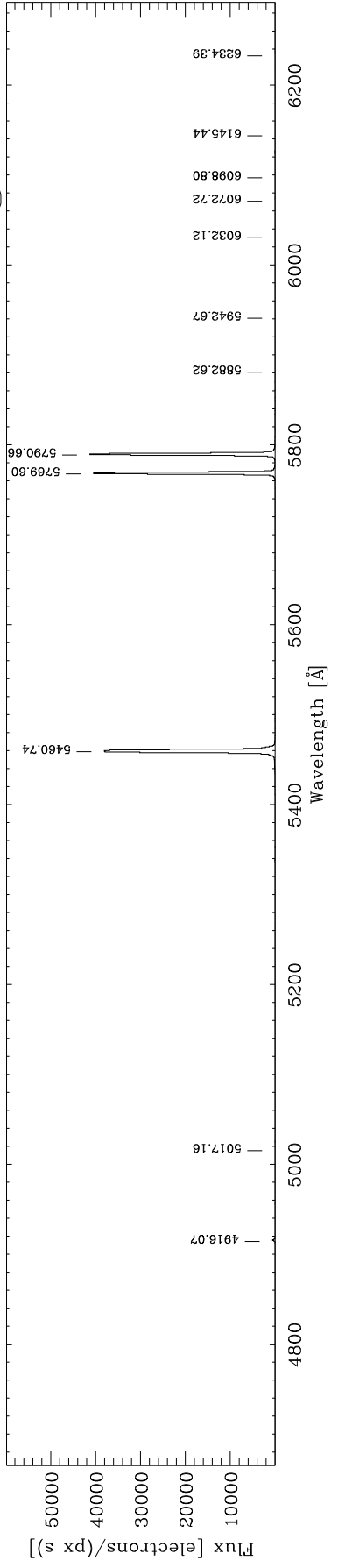
Ne



R1200R

$\lambda_{\text{cen}} = 5500\text{\AA}$

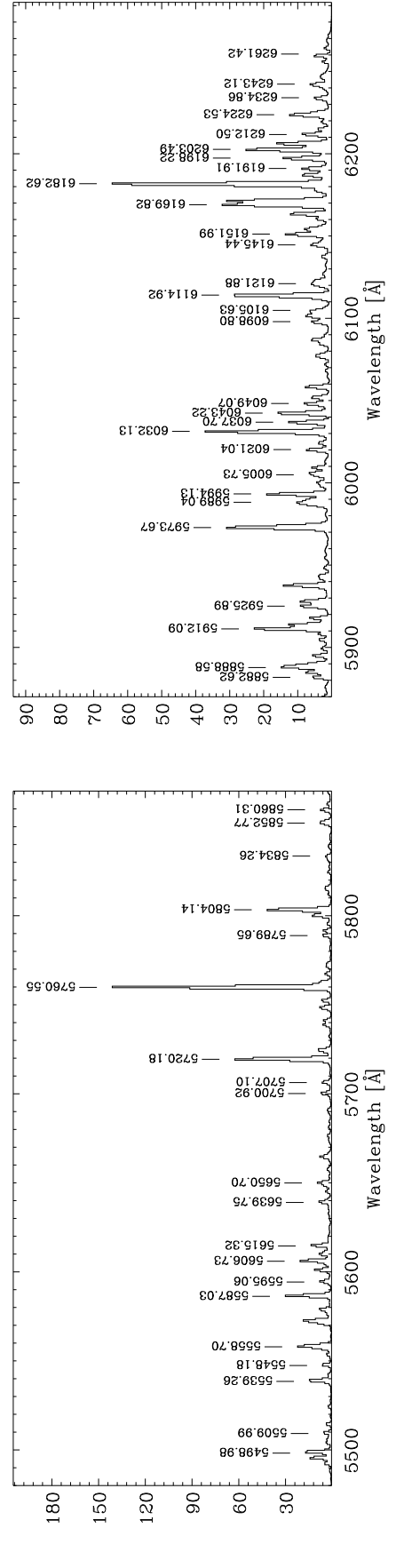
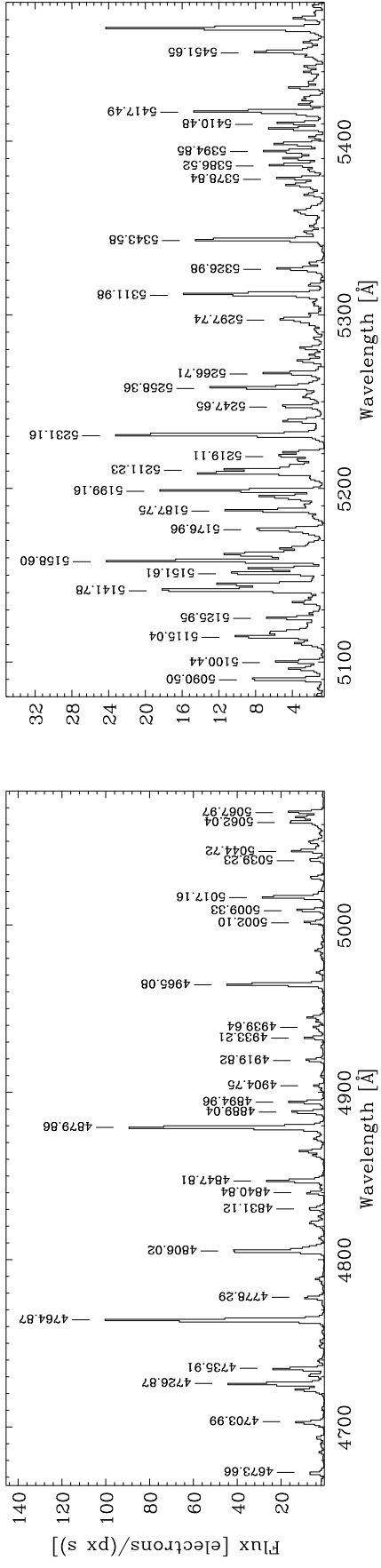
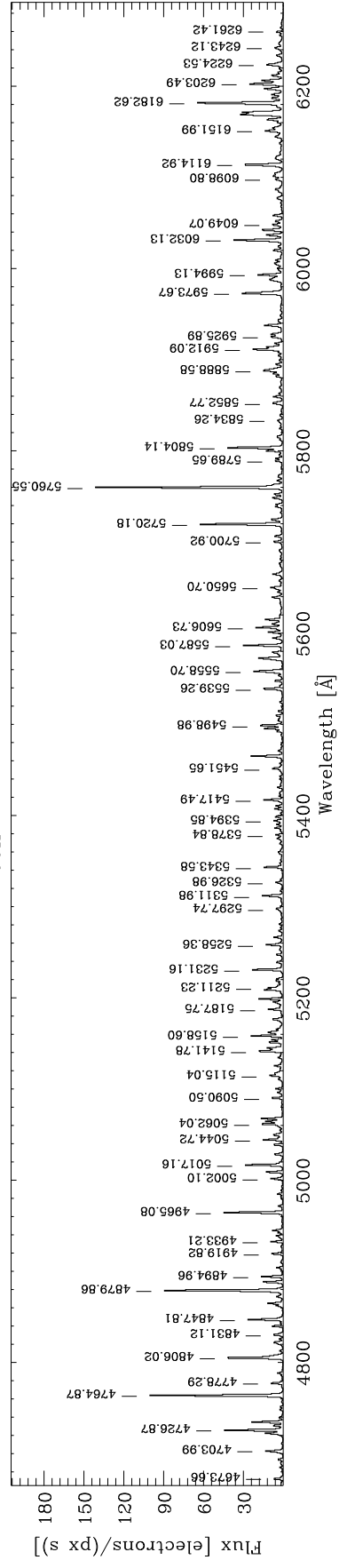
Hg



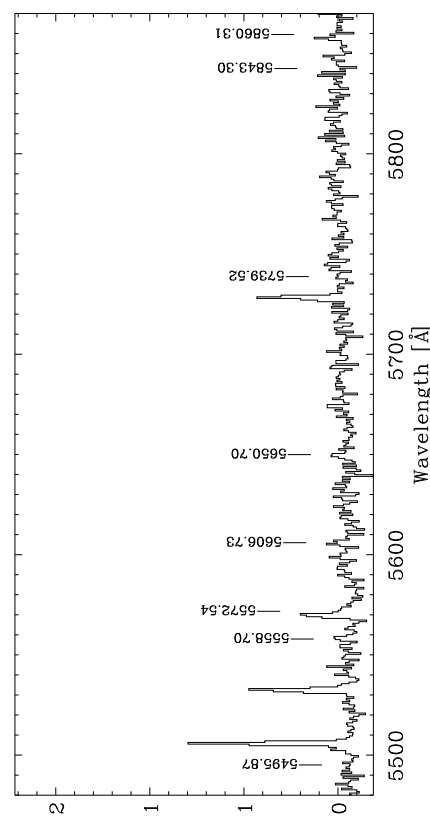
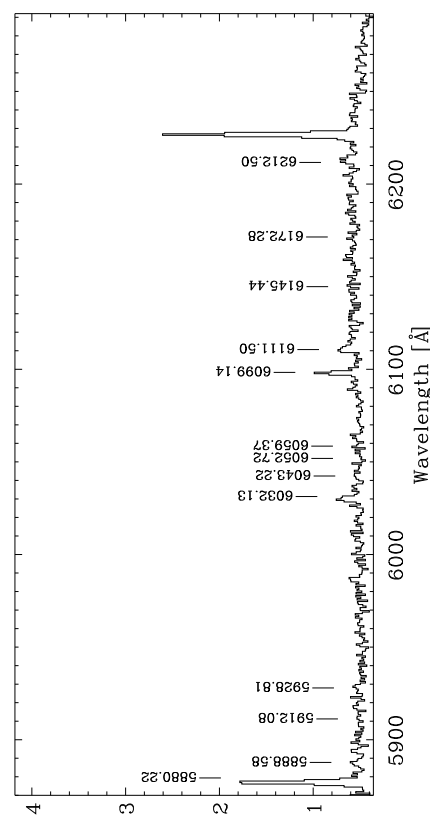
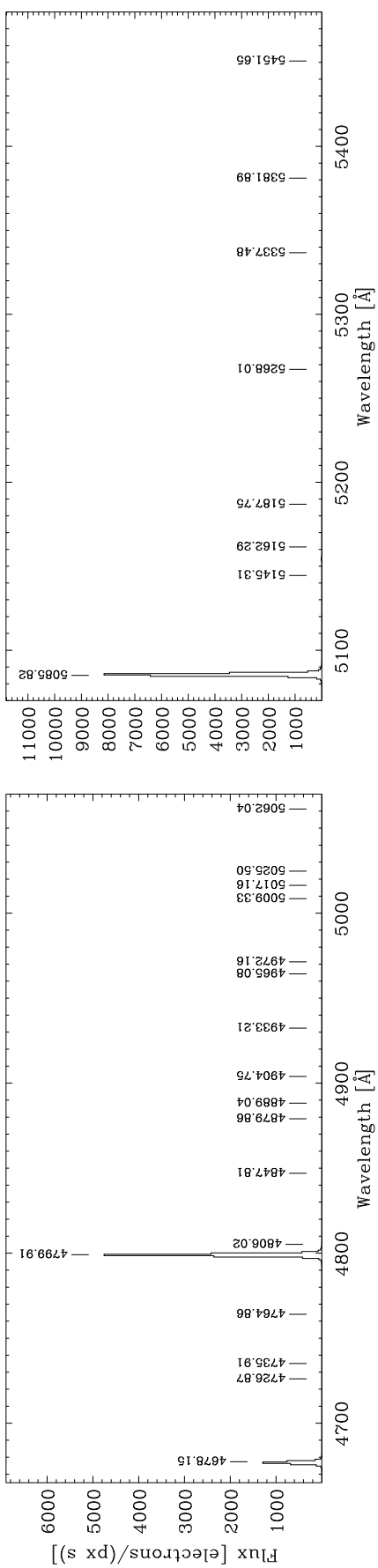
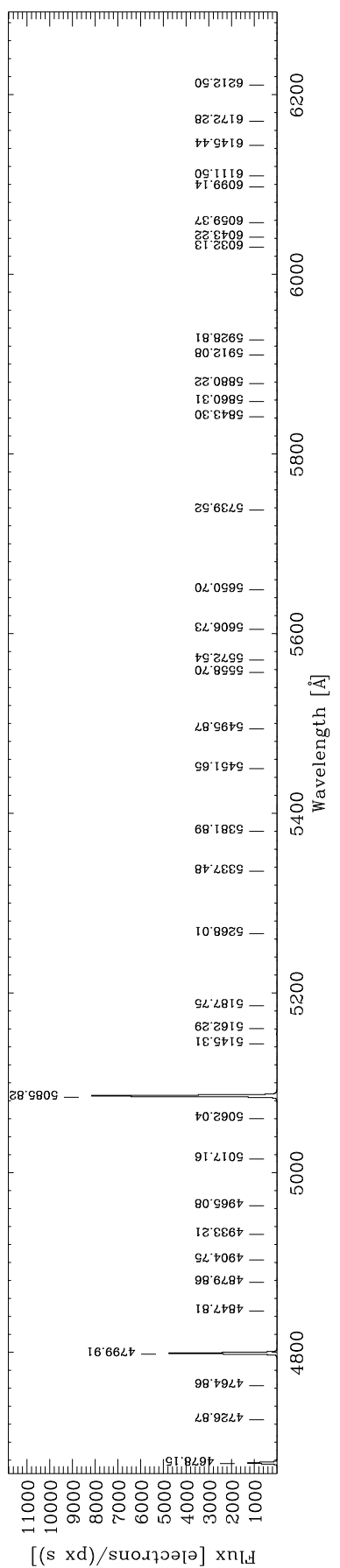
ThAr

$\lambda_{\text{cen}} = 5500\text{\AA}$

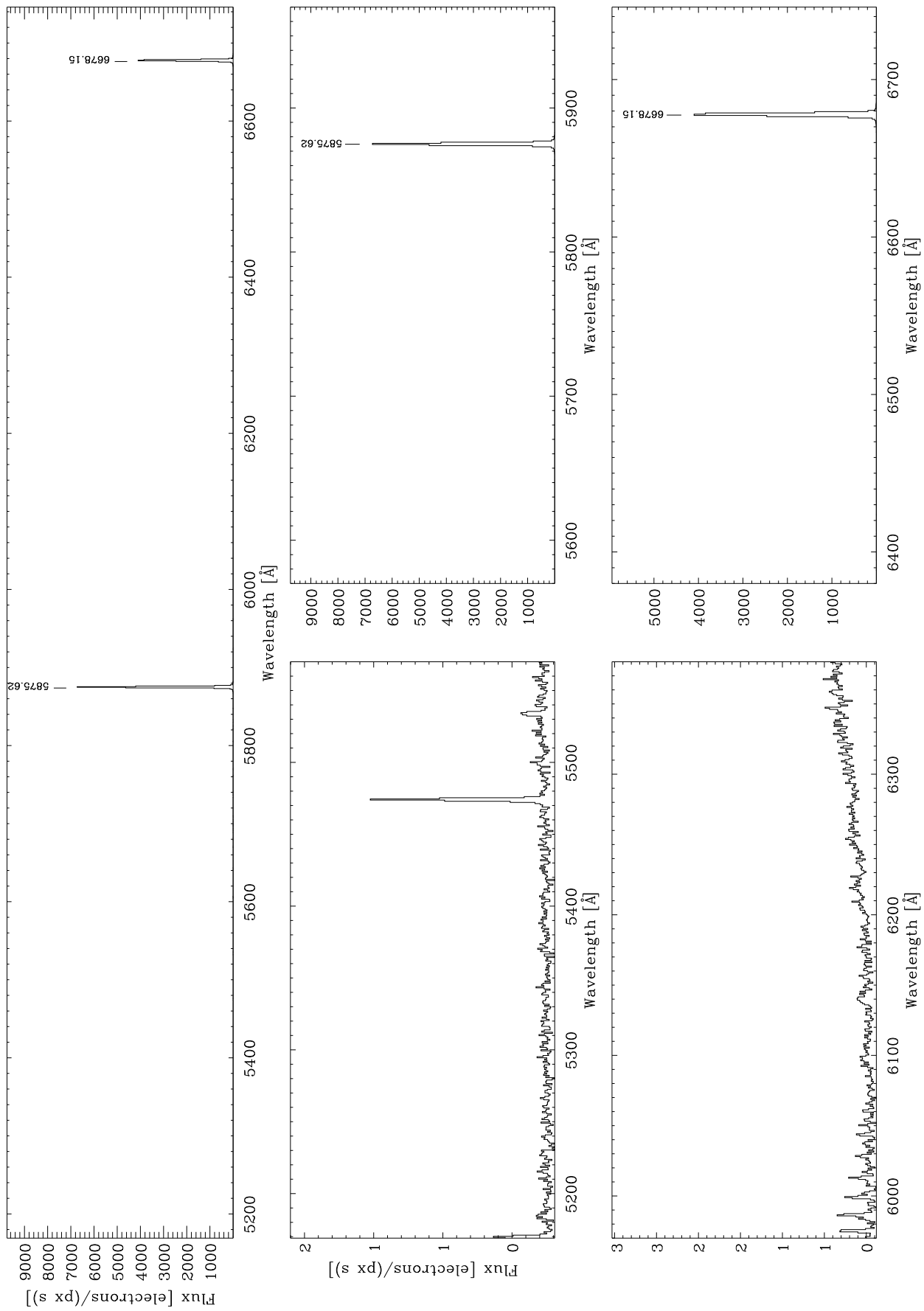
R1200R



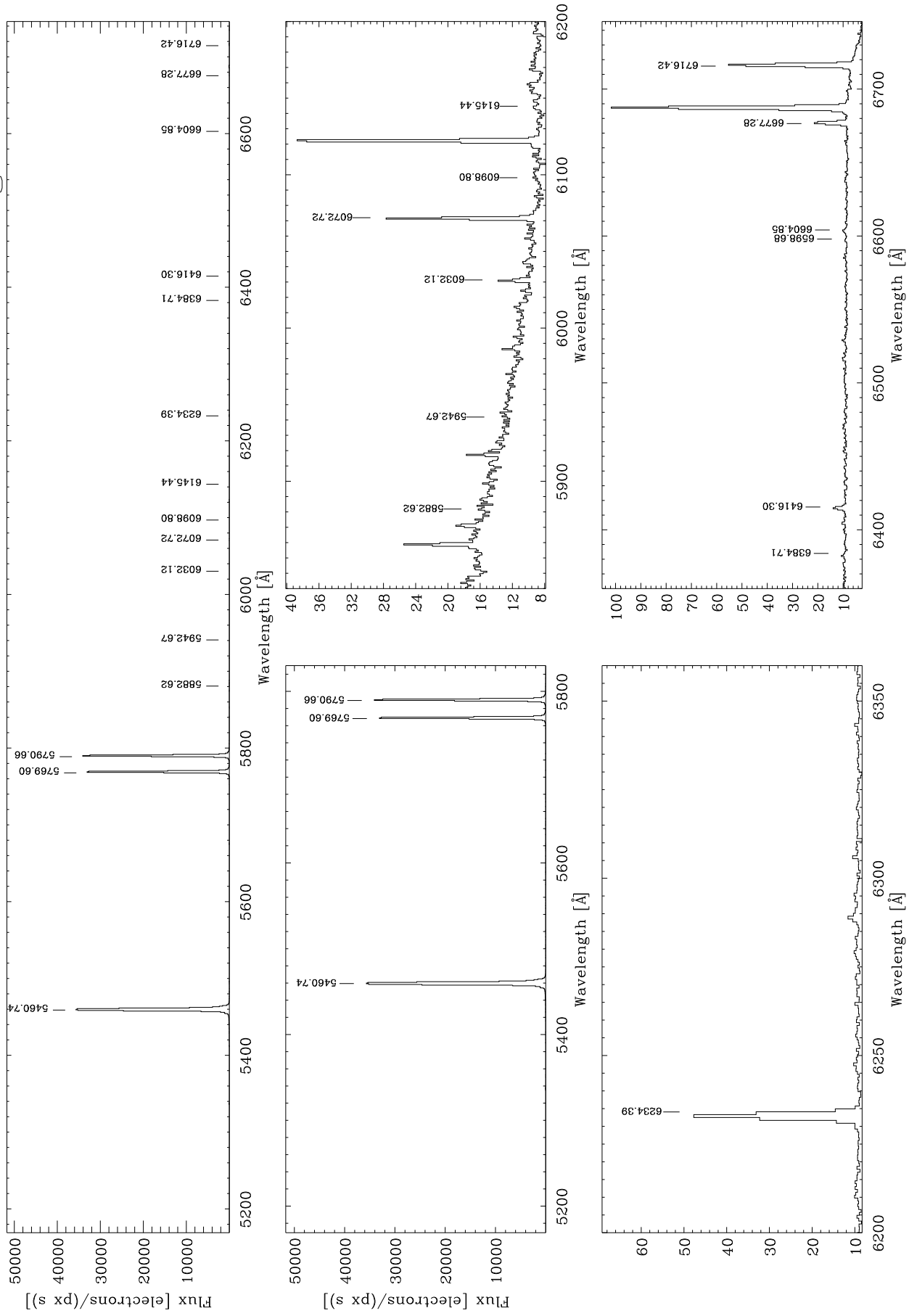
R1200R $\lambda_{\text{cen}} = 5500\text{\AA}$ Cd



R1200R $\lambda_{\text{cen}} = 6000\text{\AA}$ He



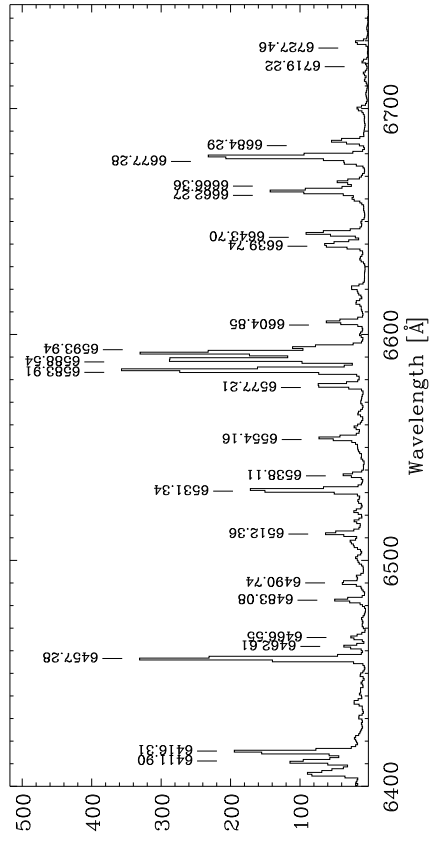
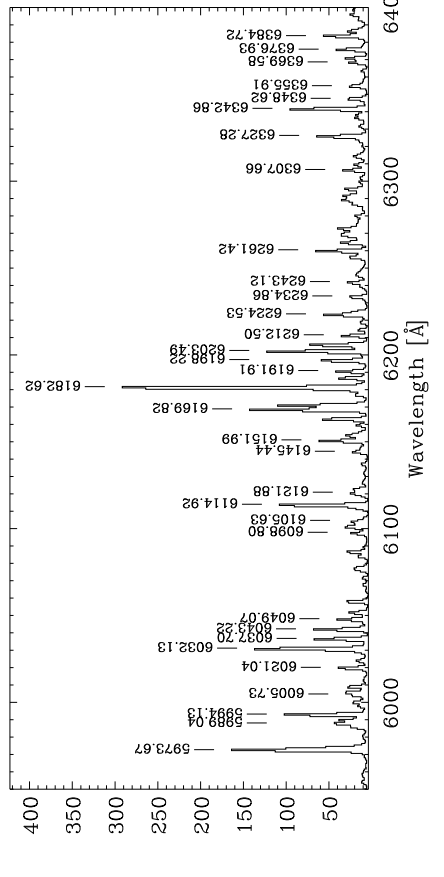
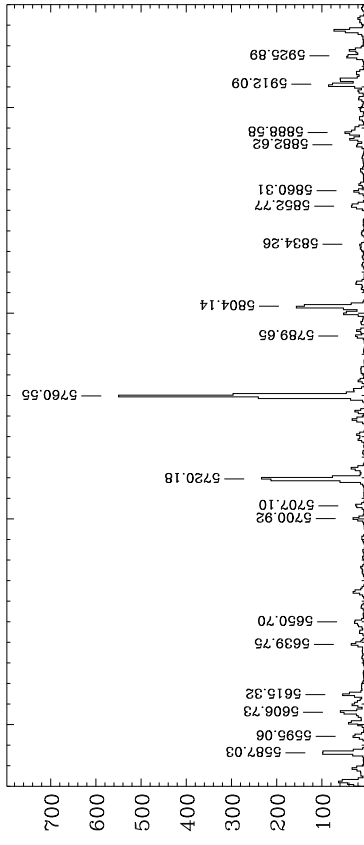
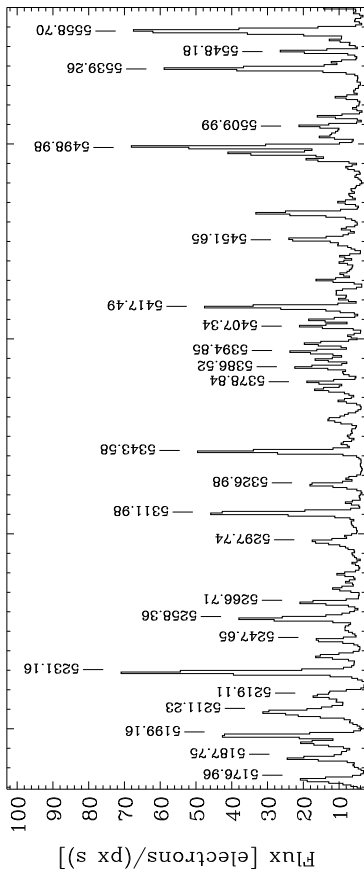
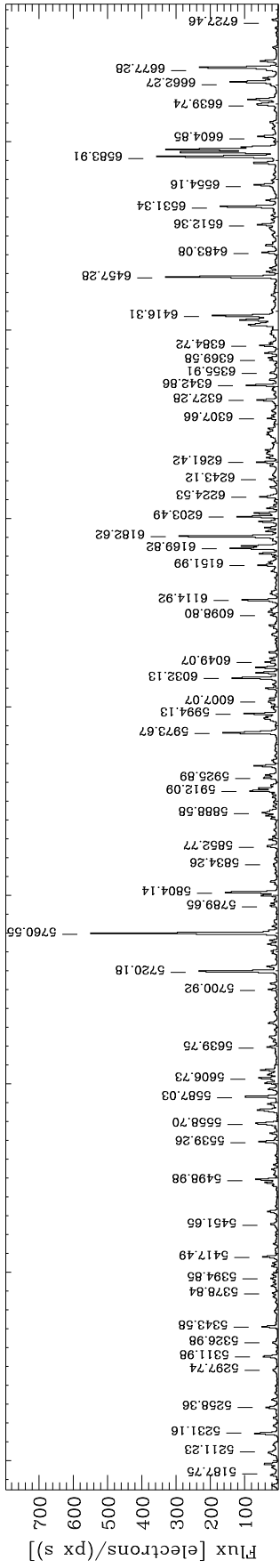
R1200R Hg $\lambda_{\text{cen}} = 6000\text{\AA}$



R1200R

$\lambda_{\text{cen}} = 6000\text{\AA}$

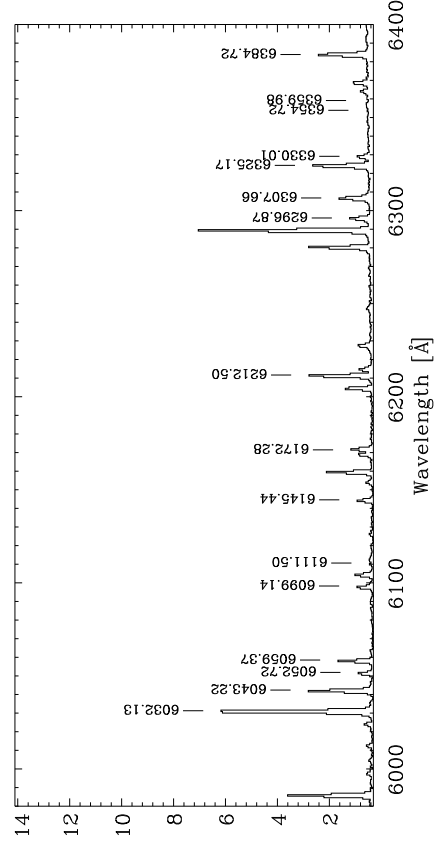
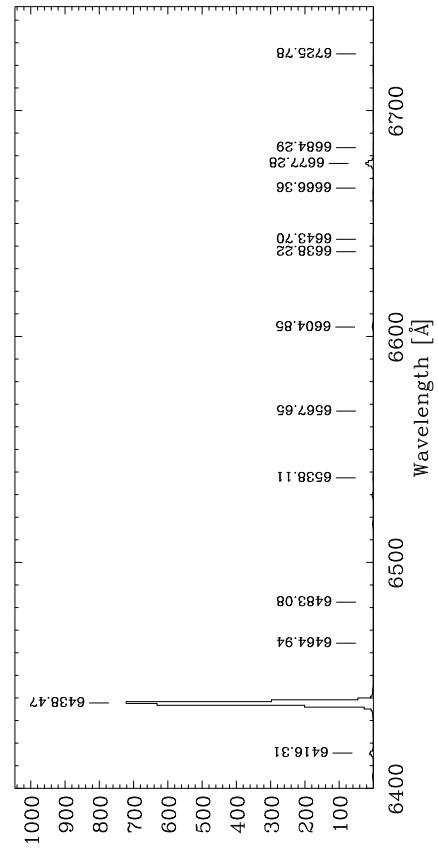
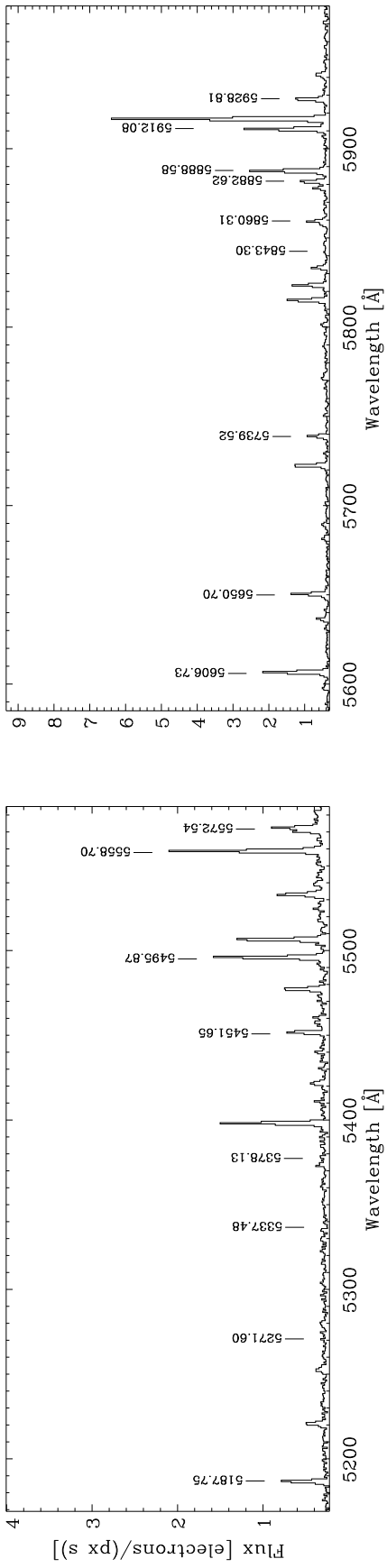
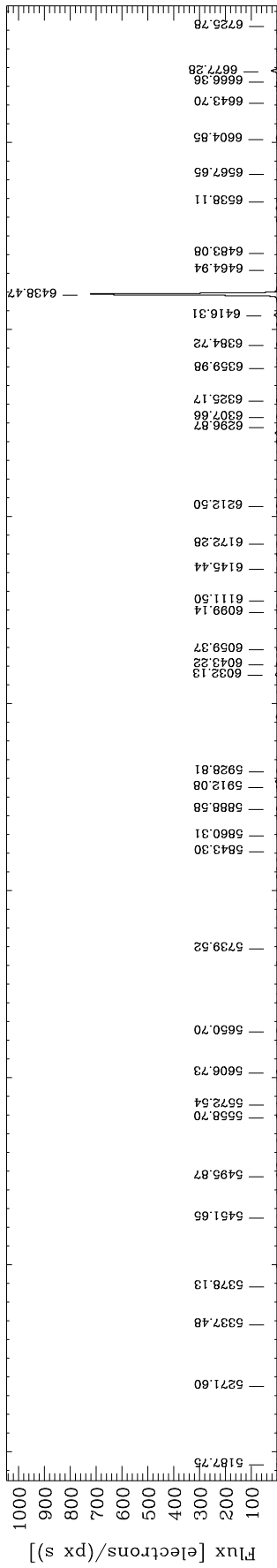
ThAr



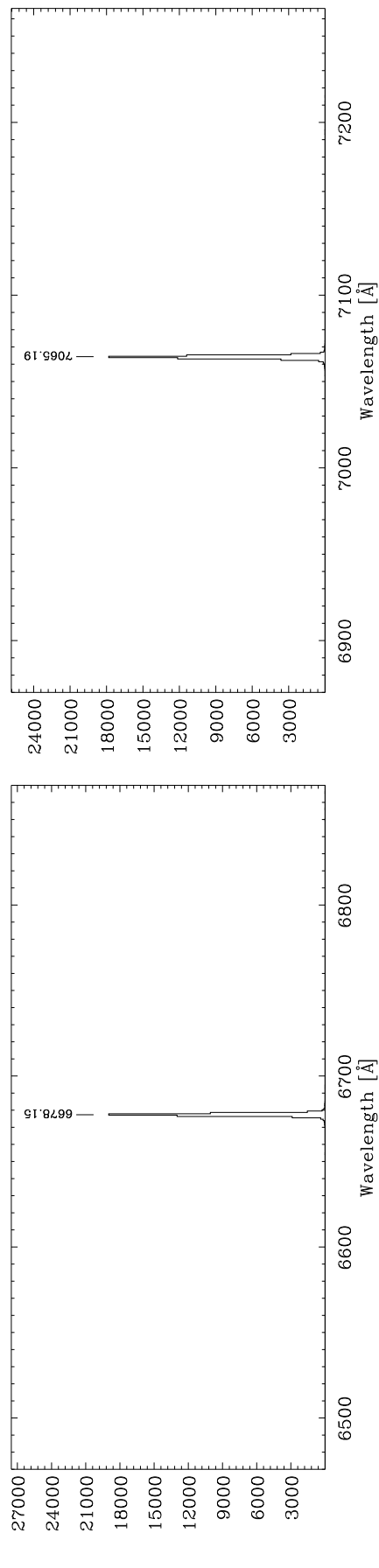
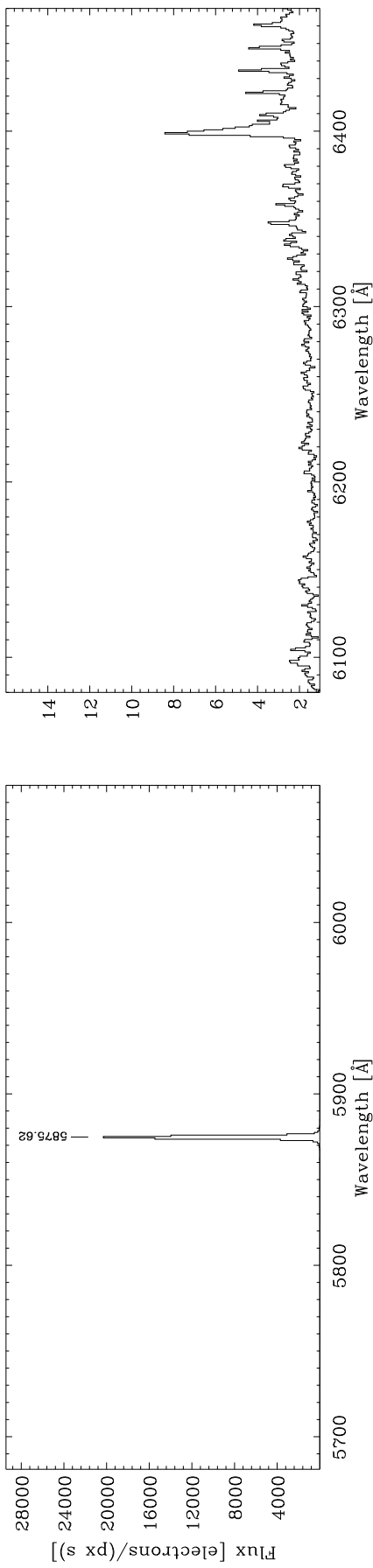
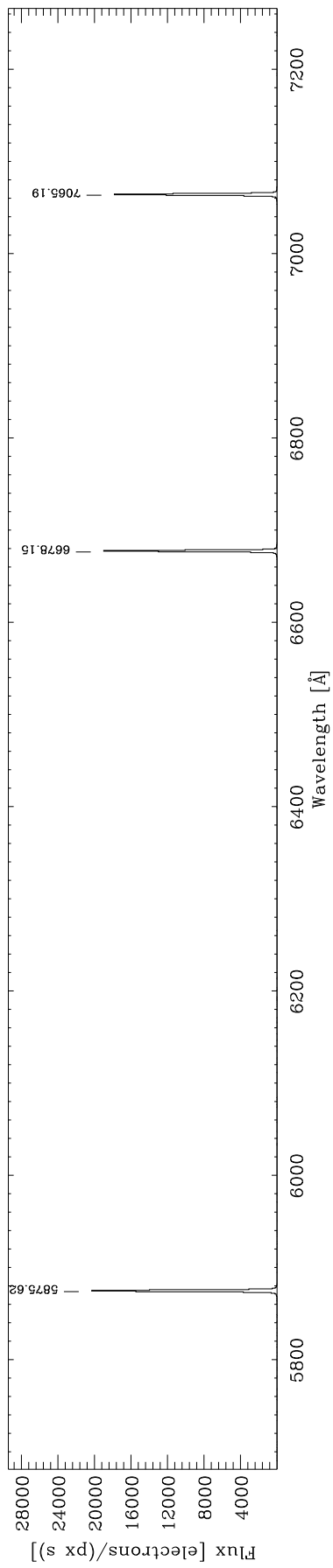
R1200R

$\lambda_{\text{cen}} = 6000\text{\AA}$

Cd



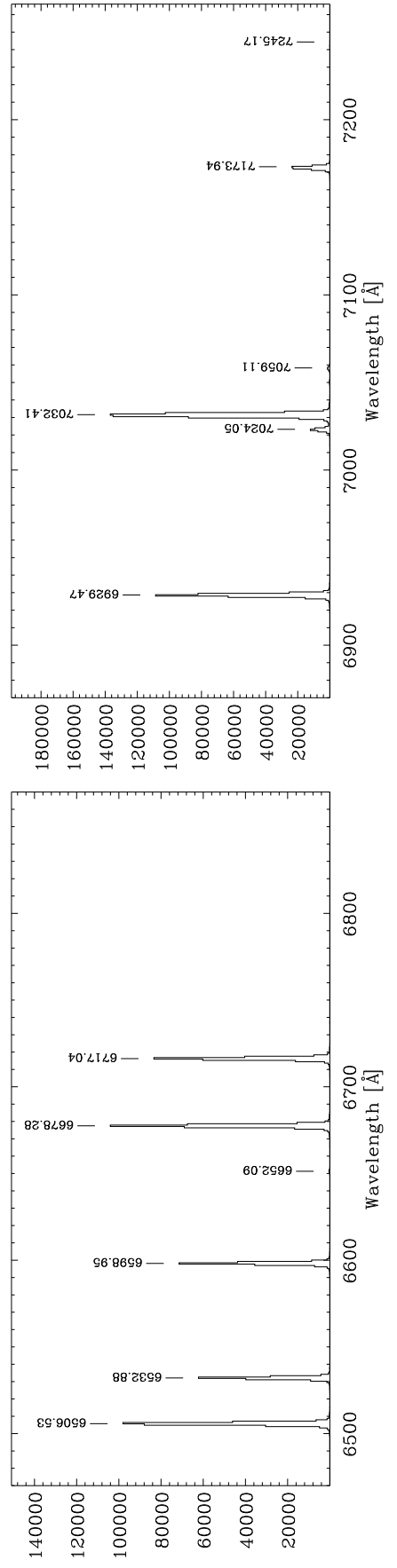
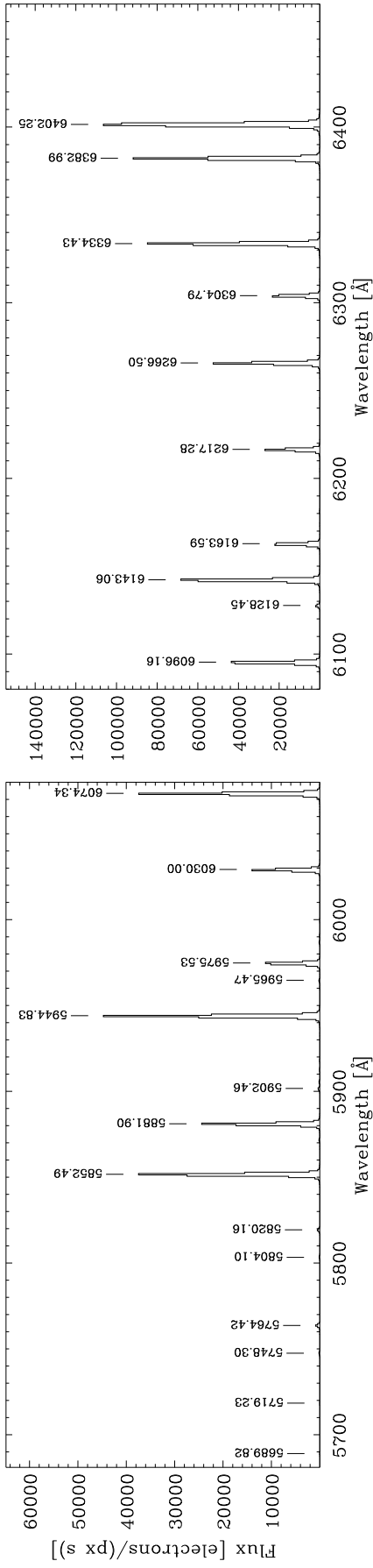
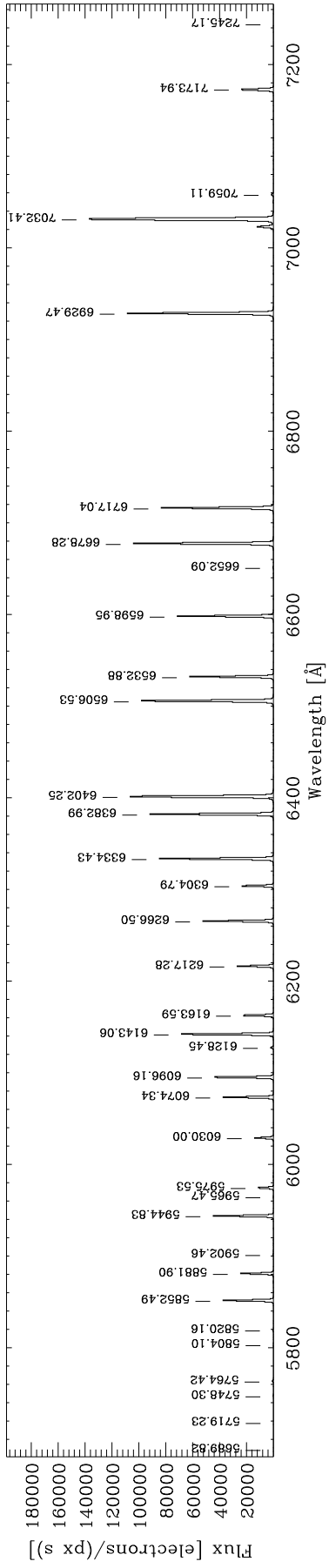
R1200R $\lambda_{\text{cen}} = 6500\text{\AA}$ He



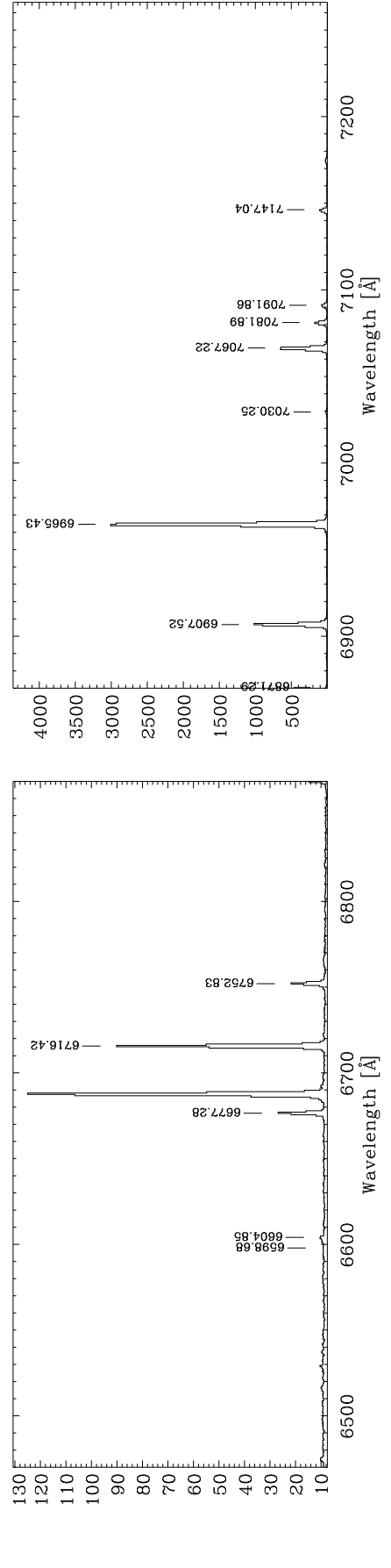
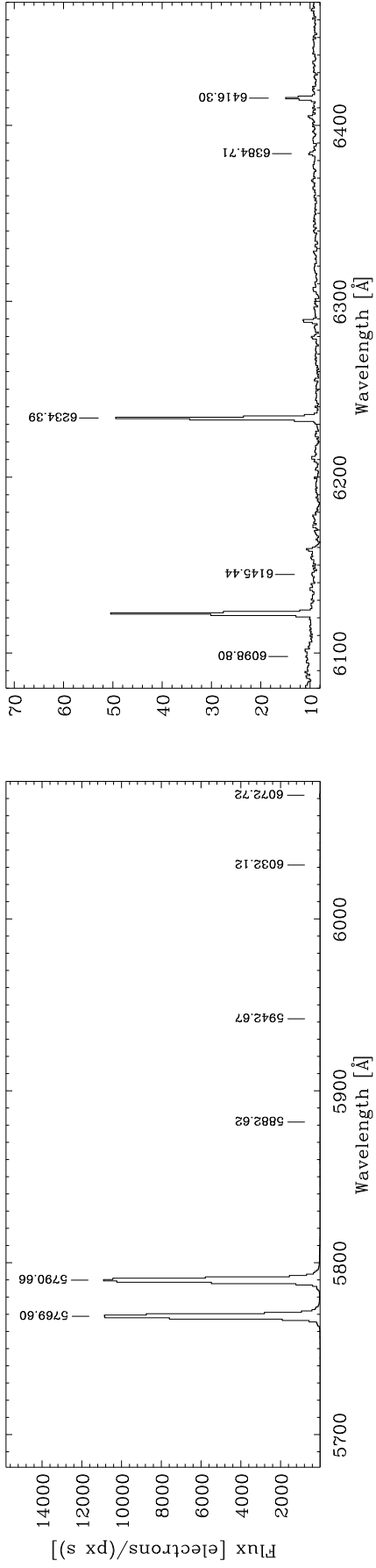
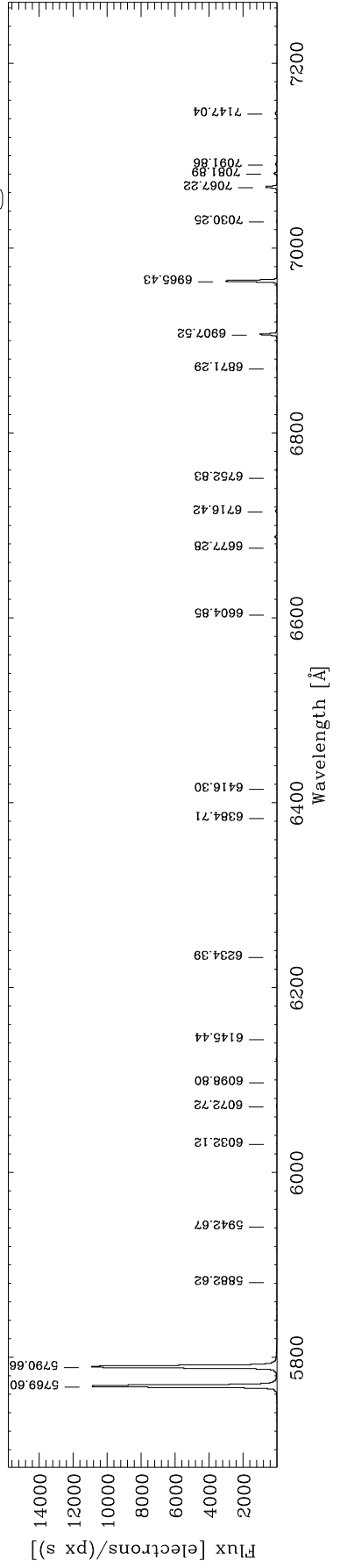
R1200R

$\lambda_{\text{cen}} = 6500\text{\AA}$

Ne



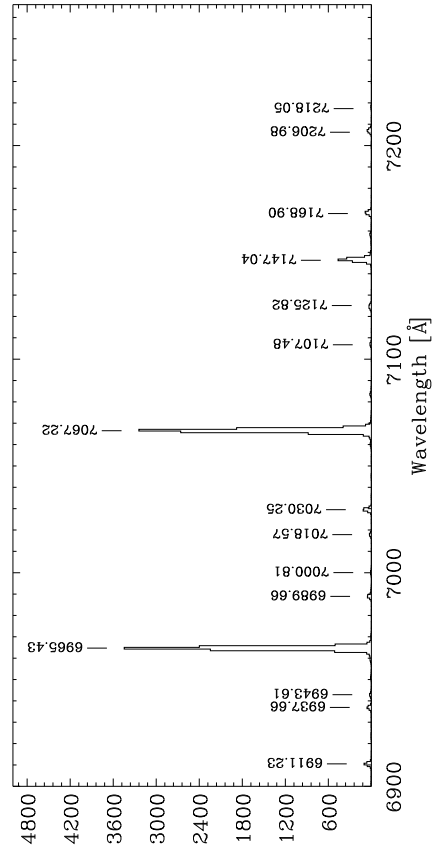
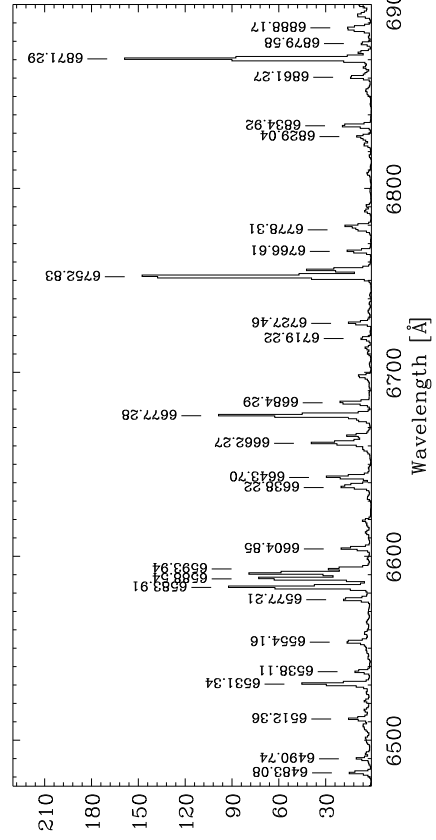
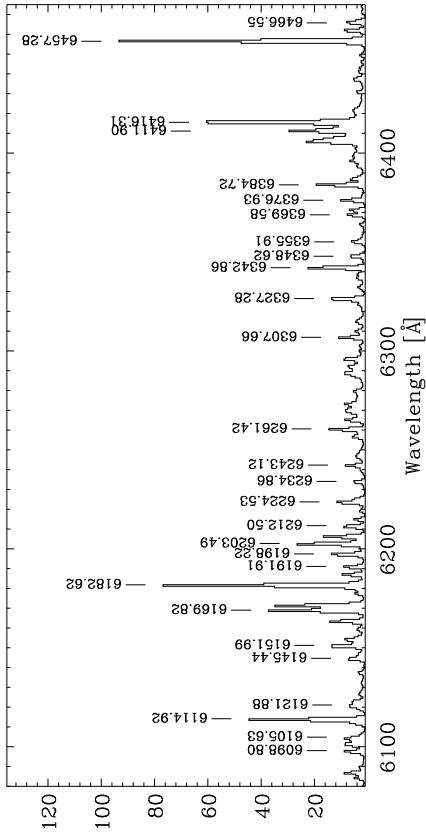
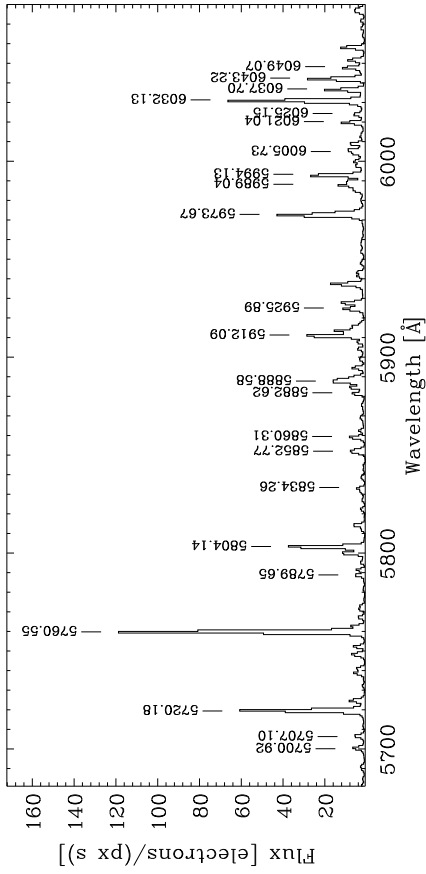
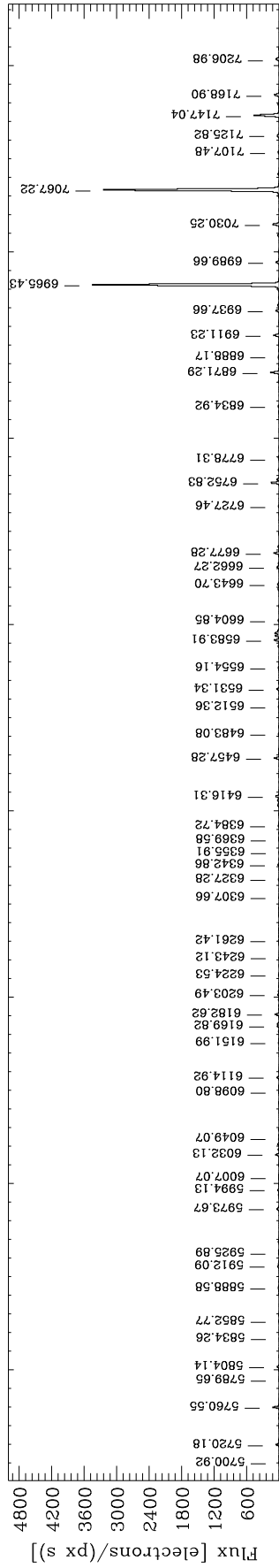
R1200R $\lambda_{\text{cen}} = 6500\text{\AA}$ Hg



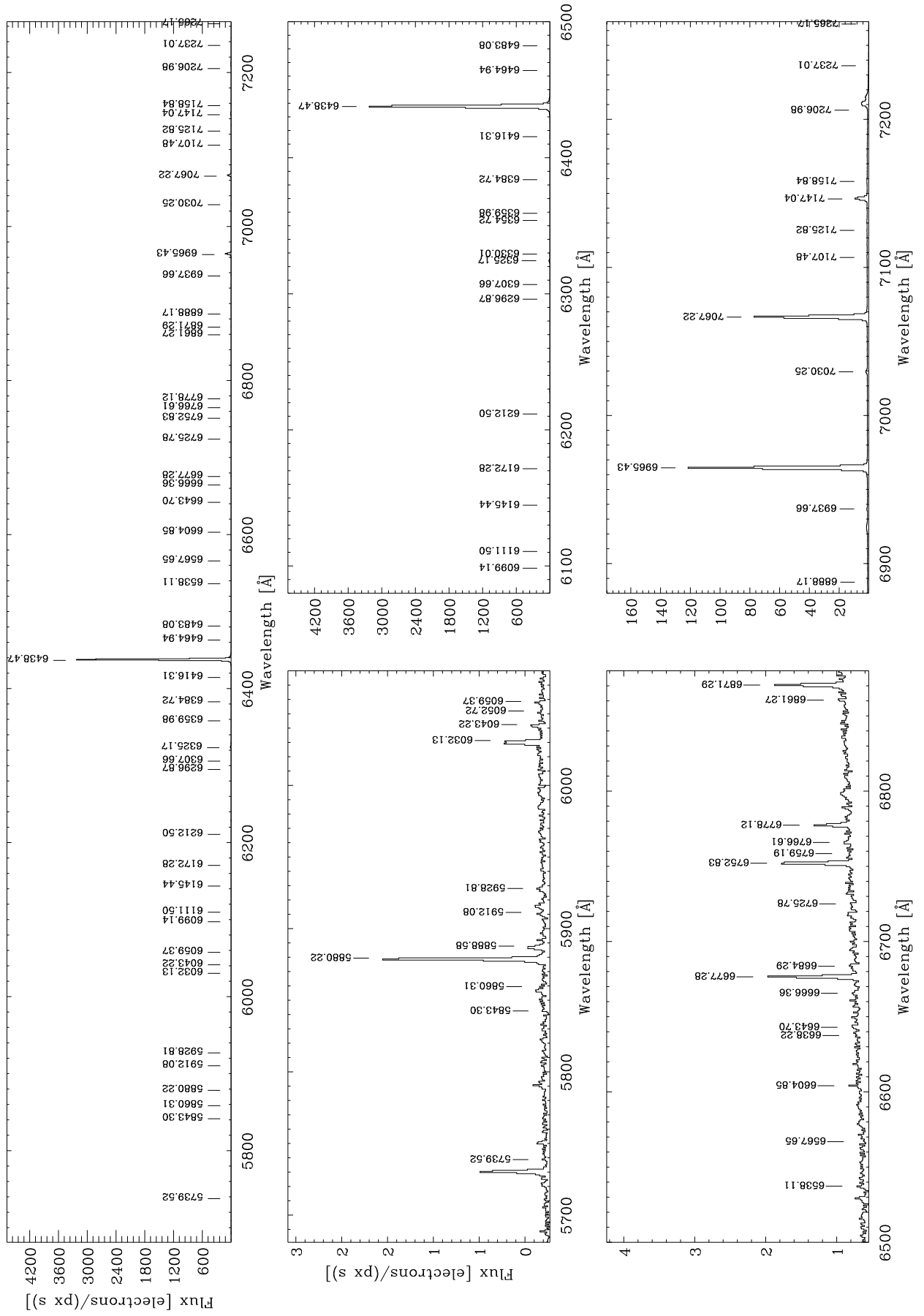
R1200R

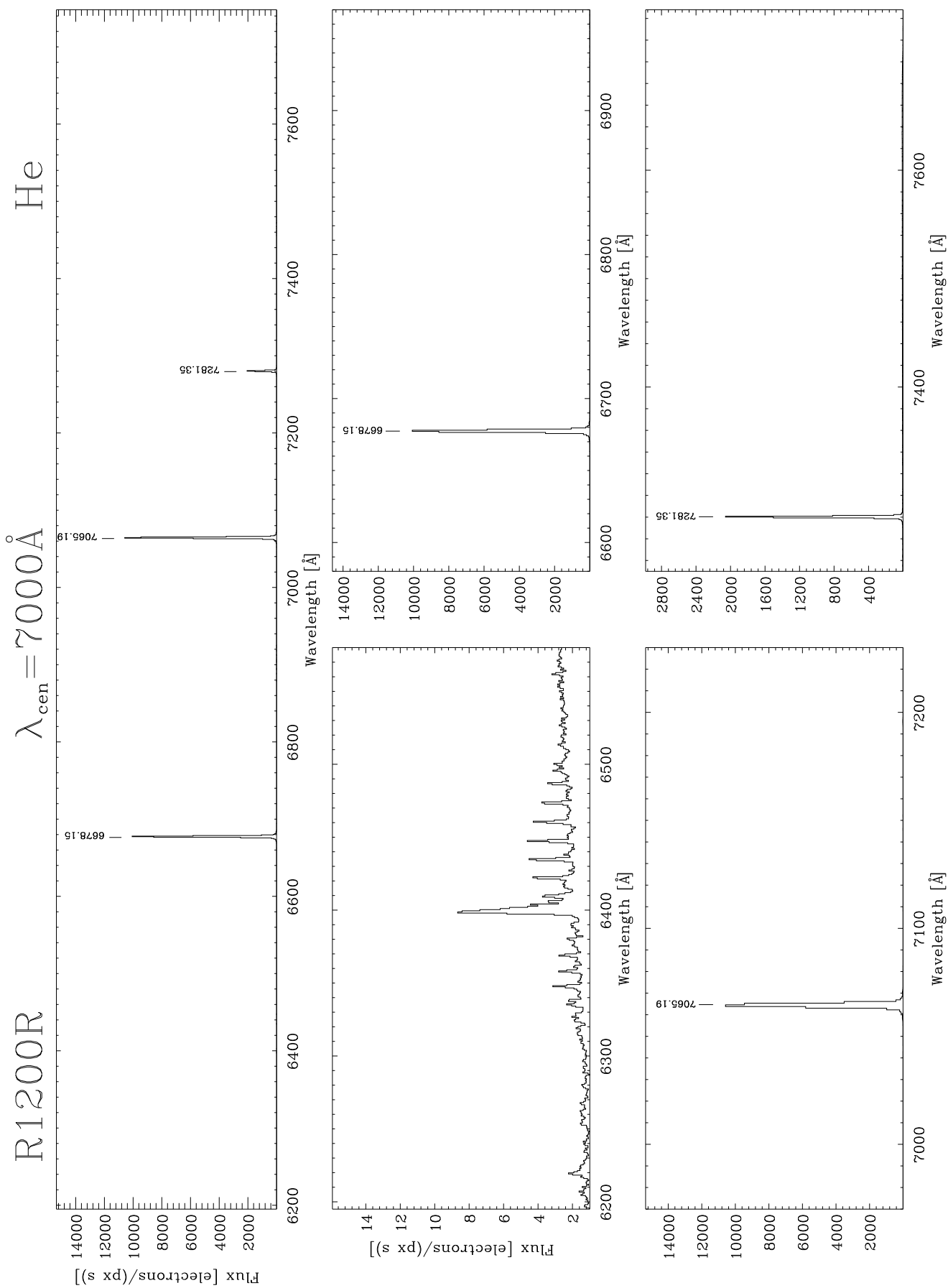
$\lambda_{\text{cen}} = 6500\text{\AA}$

ThAr

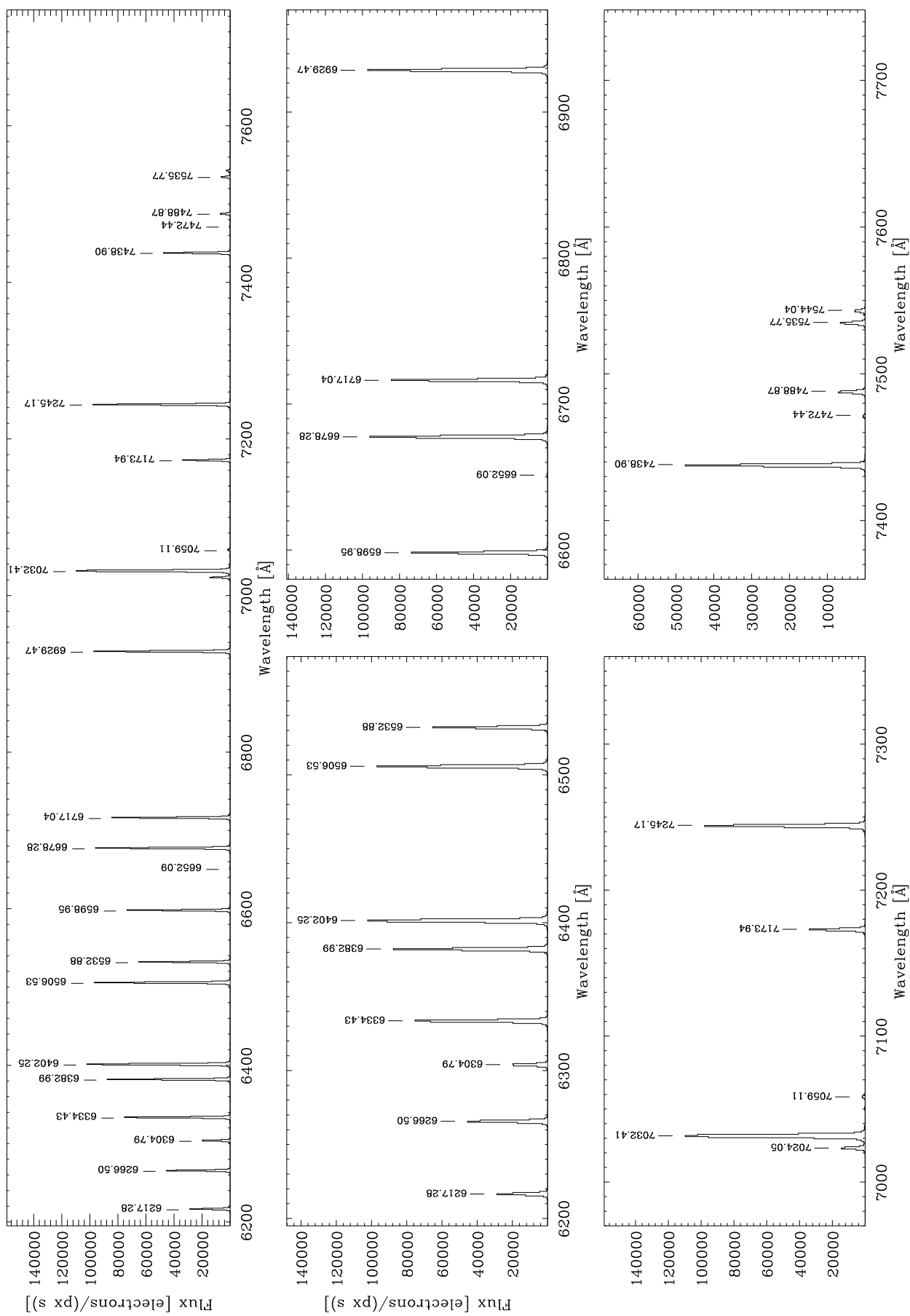


R1200R $\lambda_{\text{cen}} = 6500\text{\AA}$ Cd

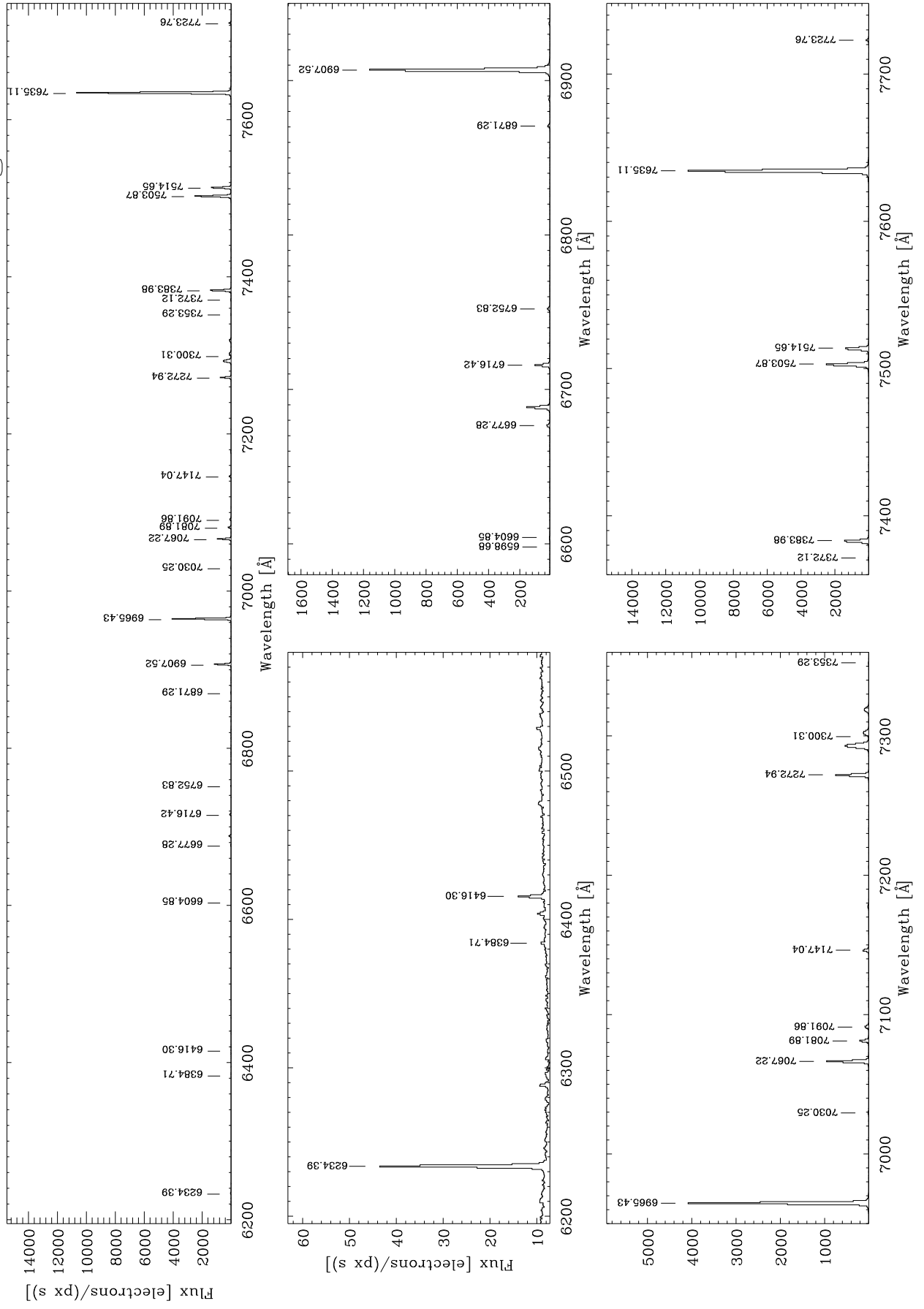




R1200R $\lambda_{\text{cen}} = 7000\text{\AA}$ Ne



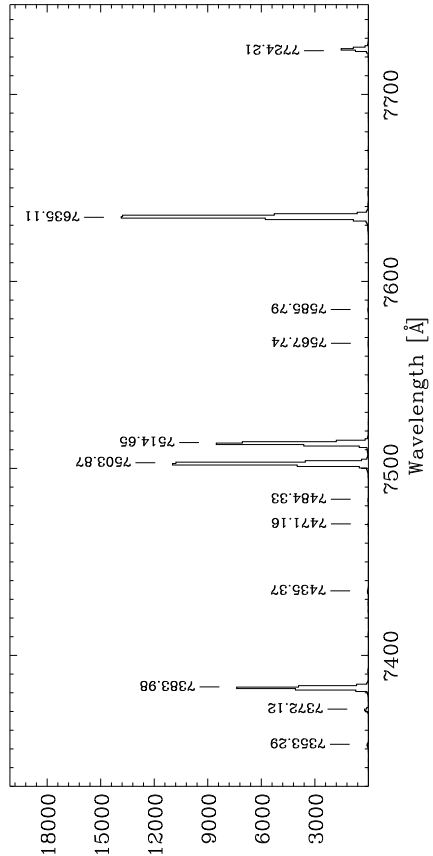
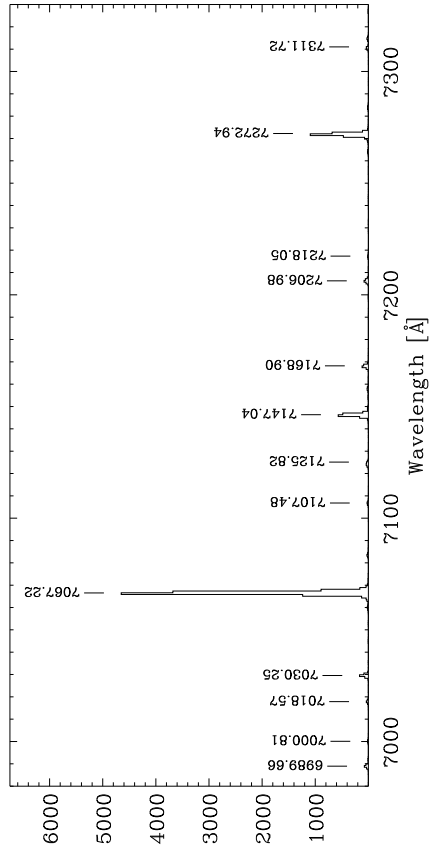
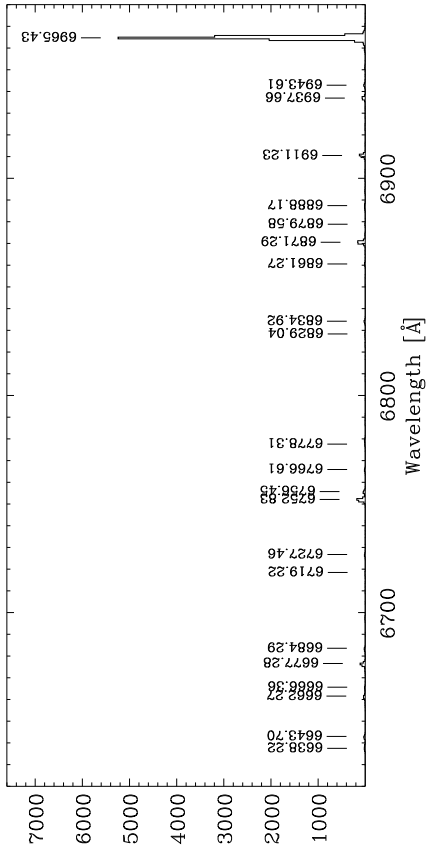
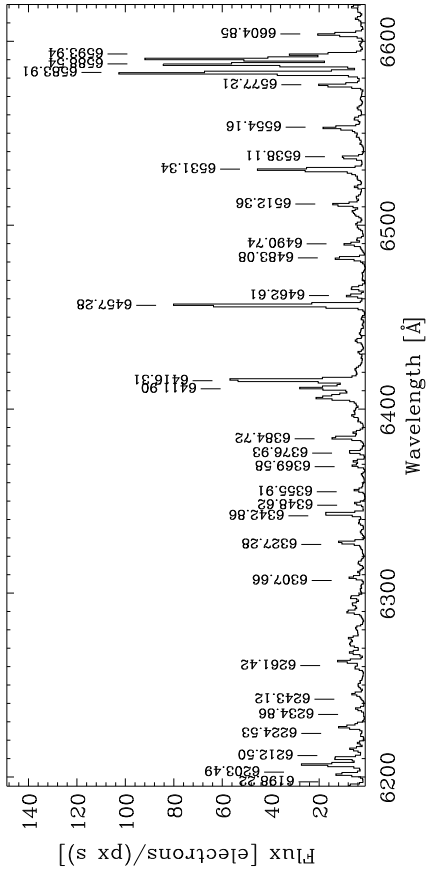
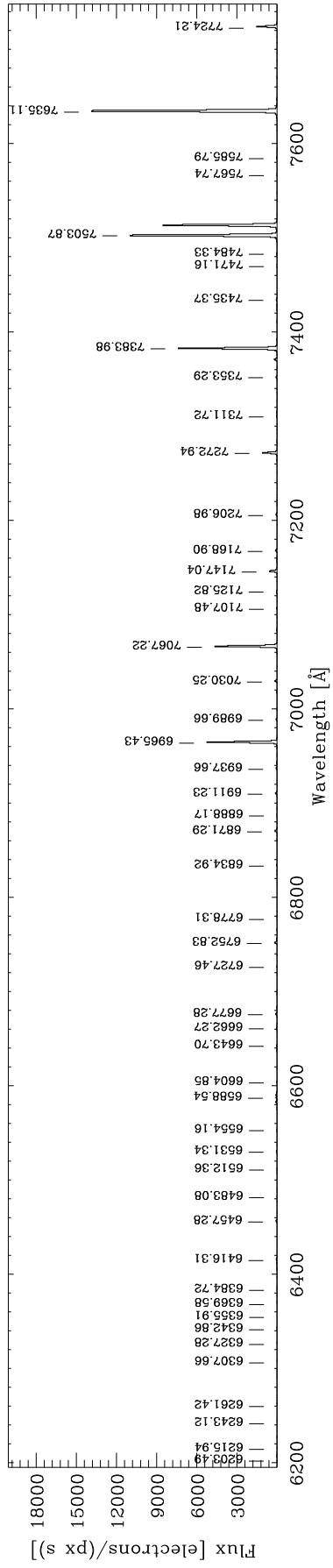
R1200R $\lambda_{\text{cen}} = 7000\text{\AA}$ Hg



R1200R

$\lambda_{\text{cen}} = 7000\text{\AA}$

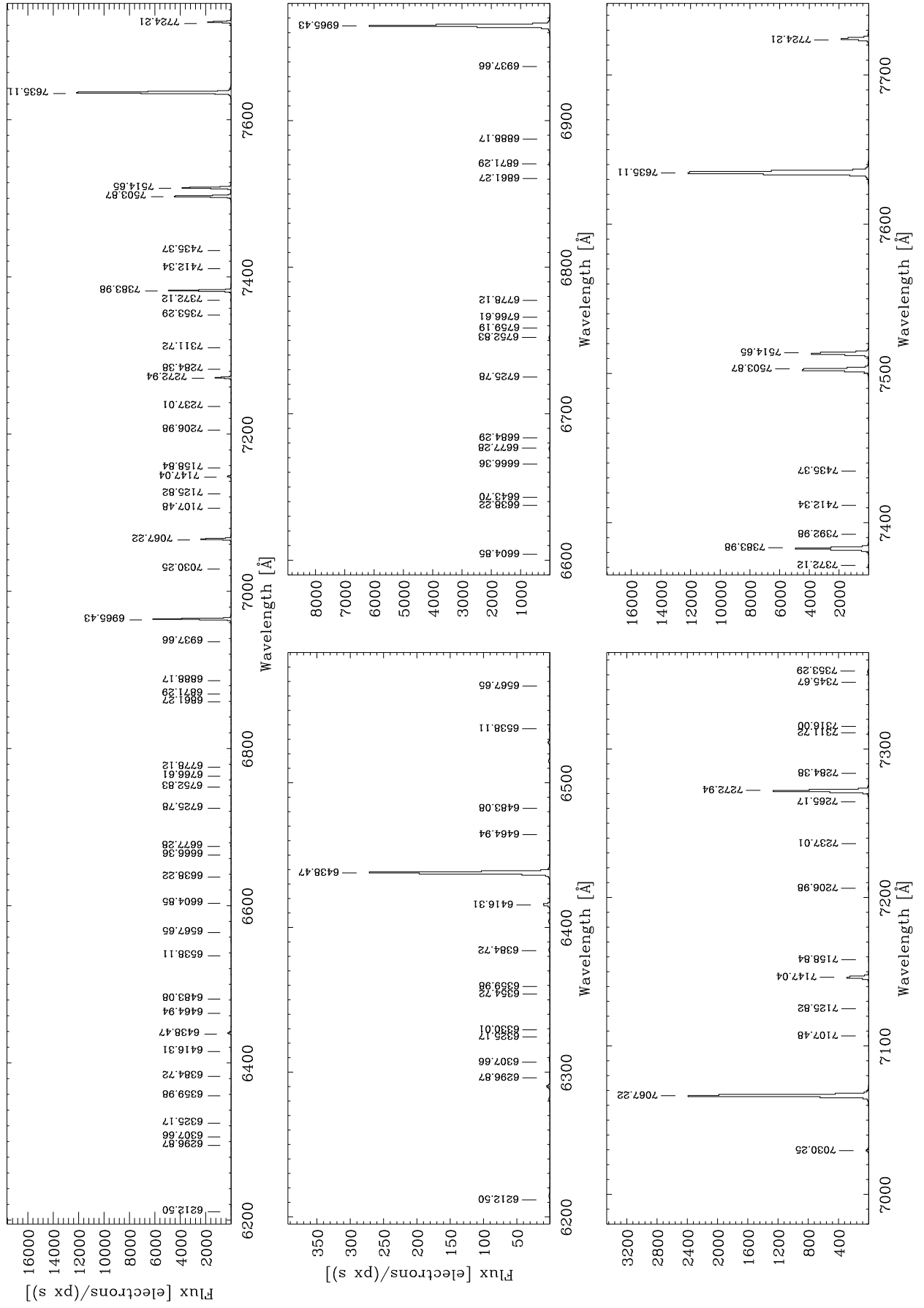
ThAr



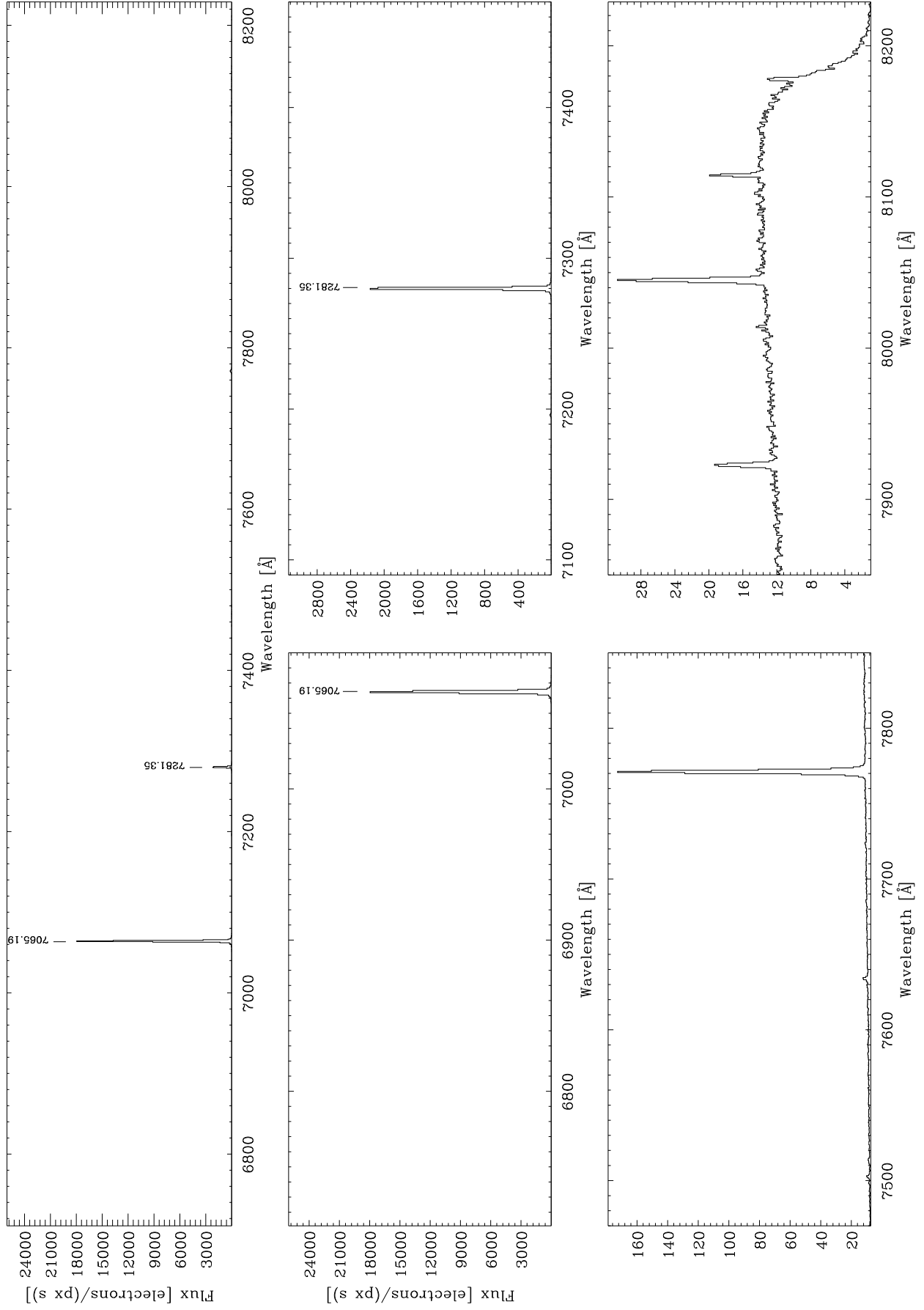
R1200R

$\lambda_{\text{cen}} = 7000\text{\AA}$

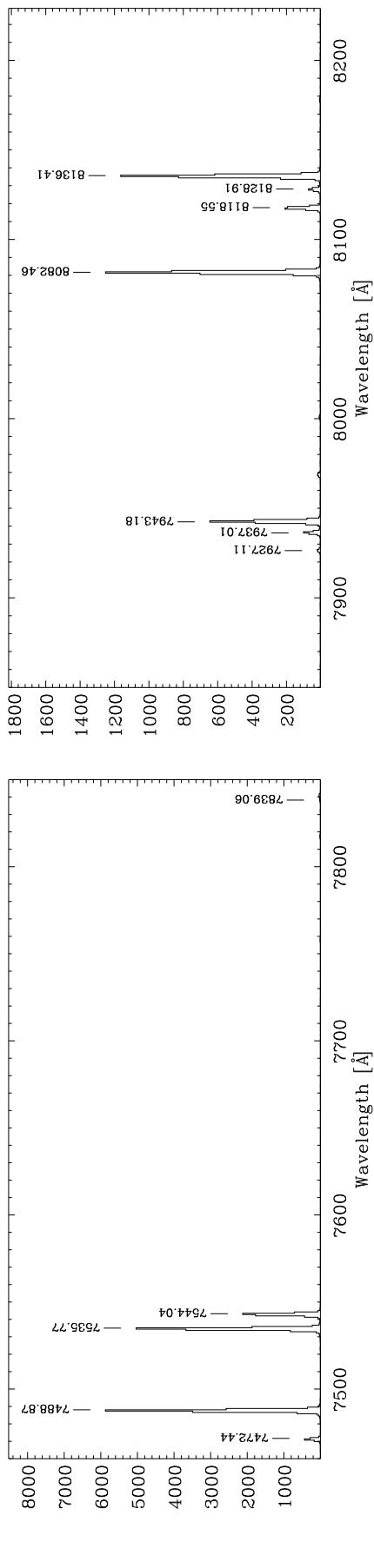
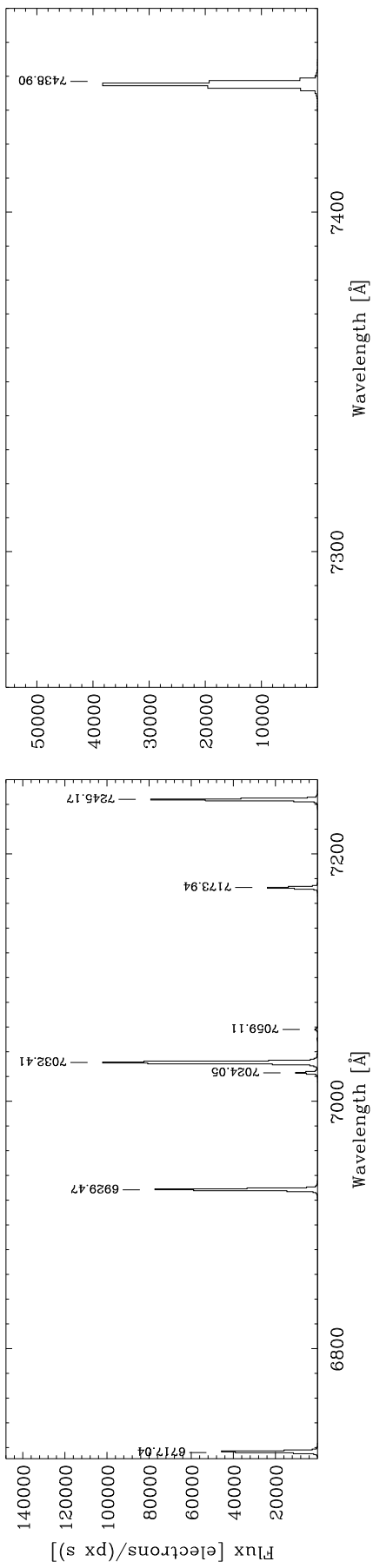
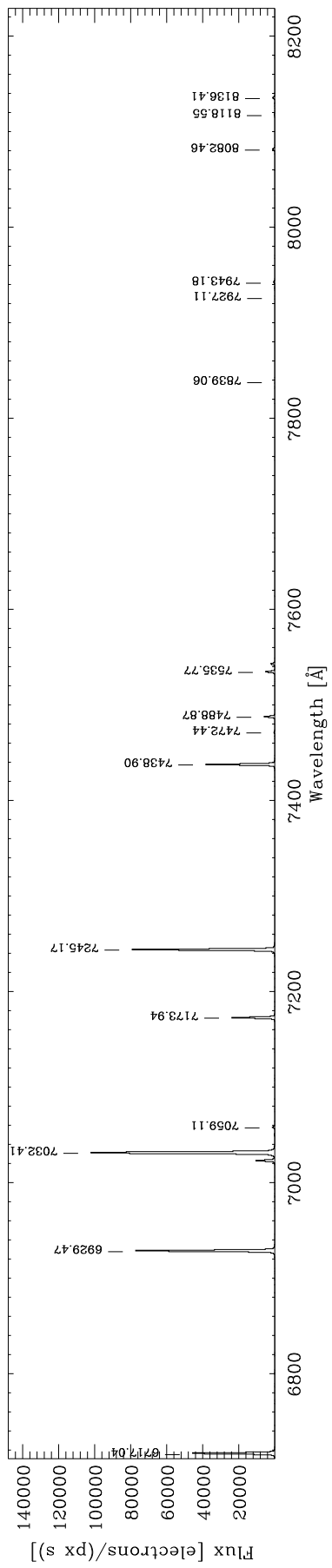
Cd



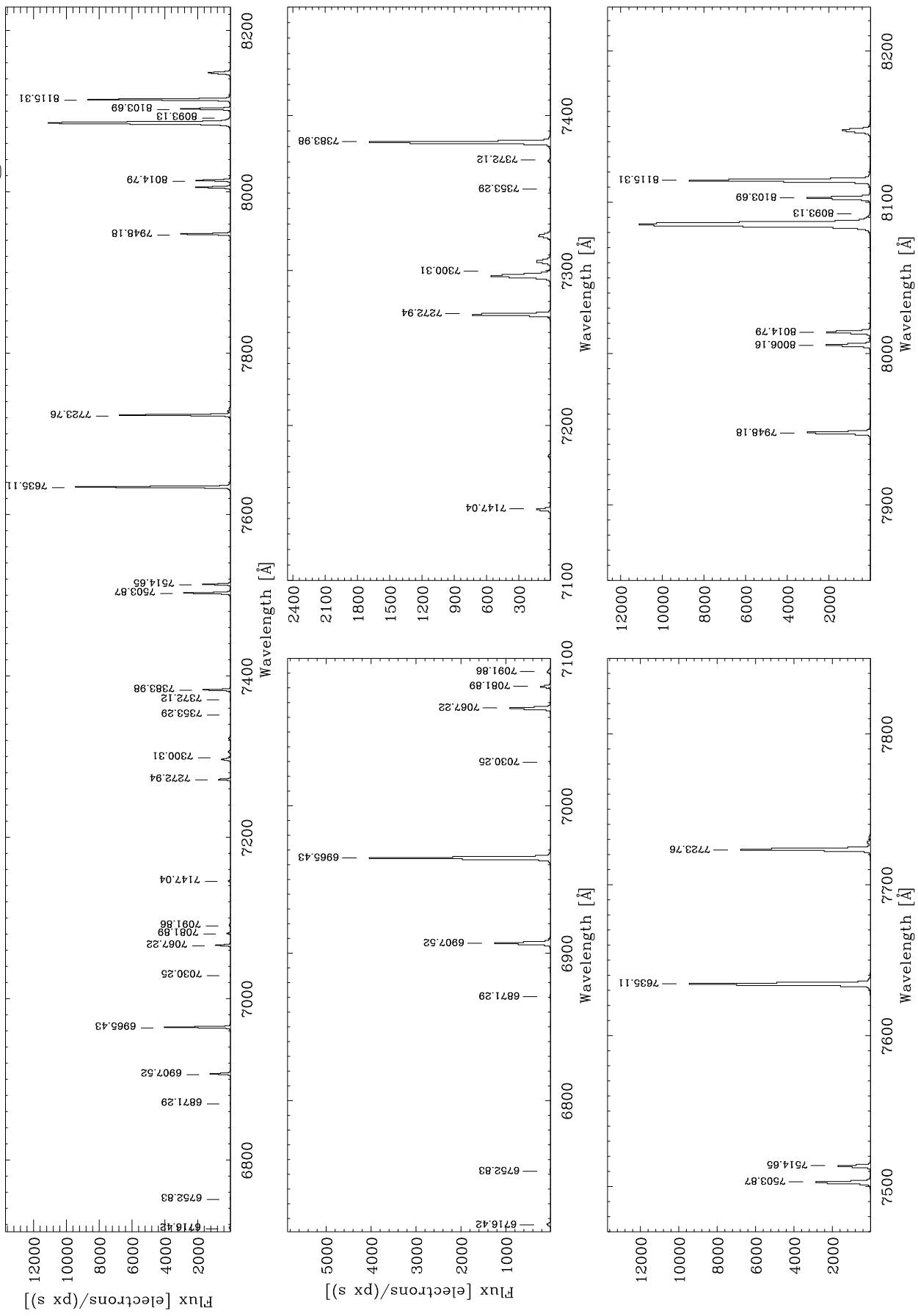
R1200R $\lambda_{\text{cen}} = 7500\text{\AA}$ He



R1200R $\lambda_{\text{cen}} = 7500\text{\AA}$ Ne



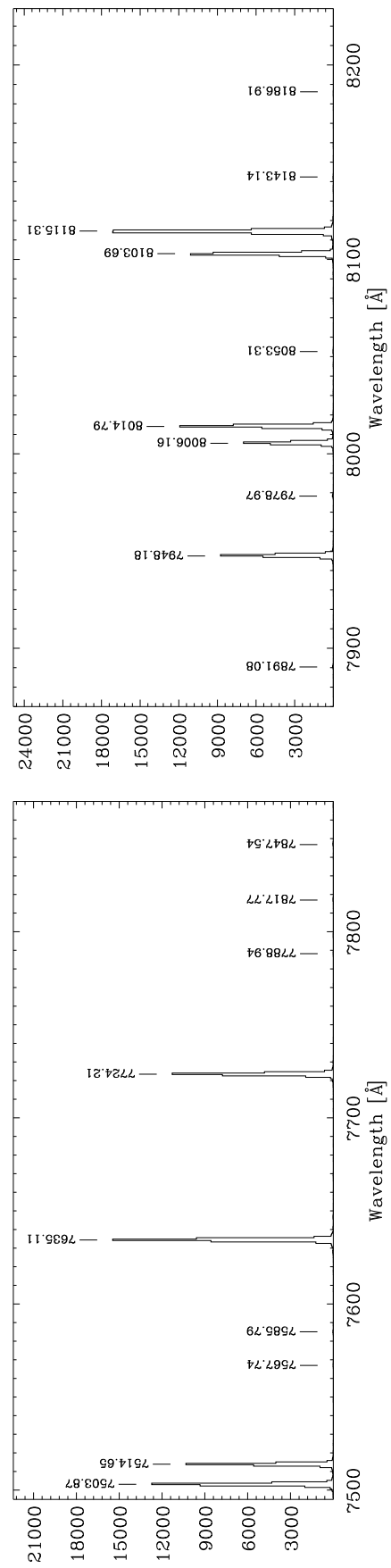
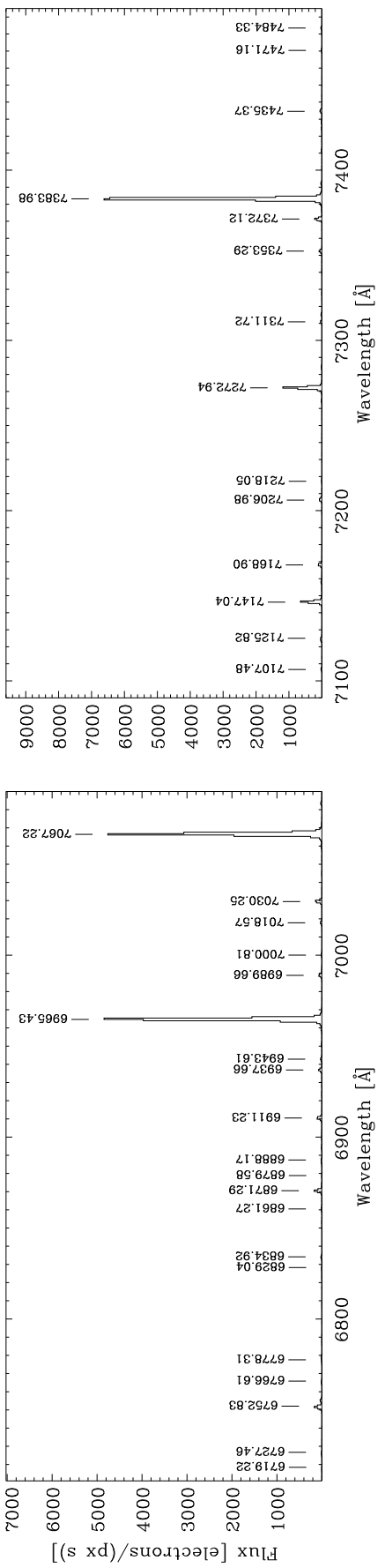
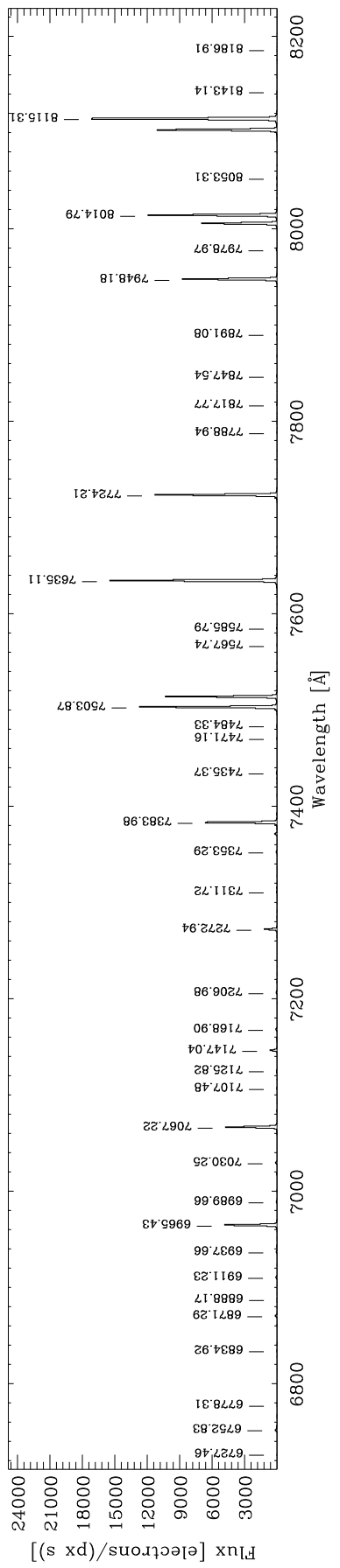
R1200R $\lambda_{\text{cen}} = 7500\text{\AA}$ Hg



R1200R

$\lambda_{\text{cen}} = 7500\text{\AA}$

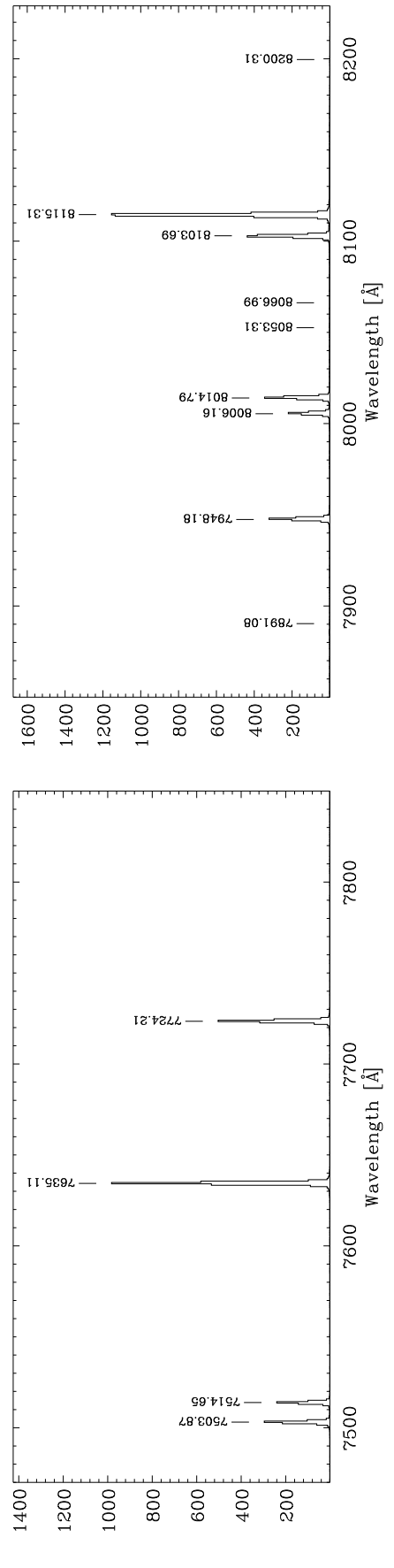
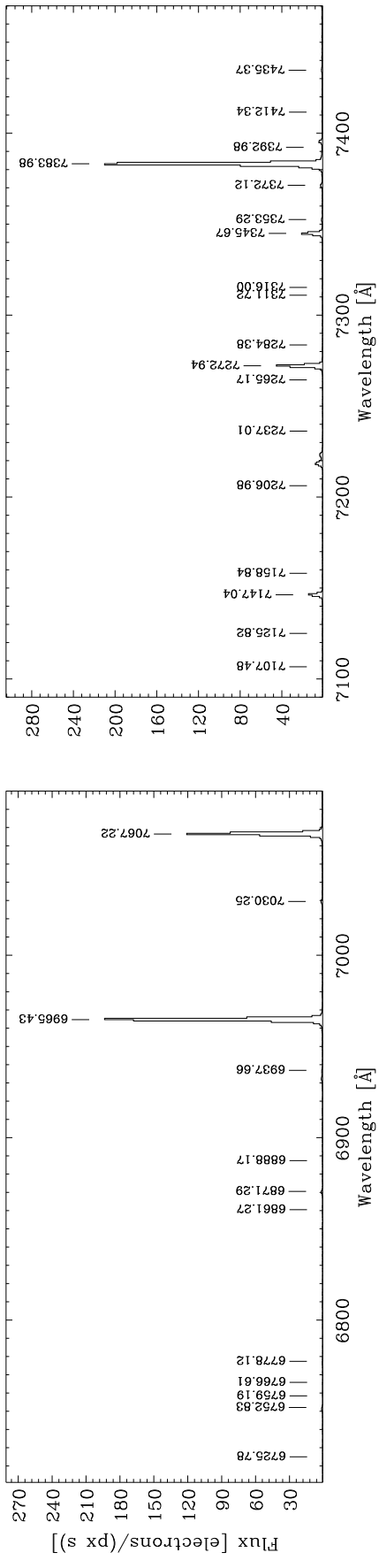
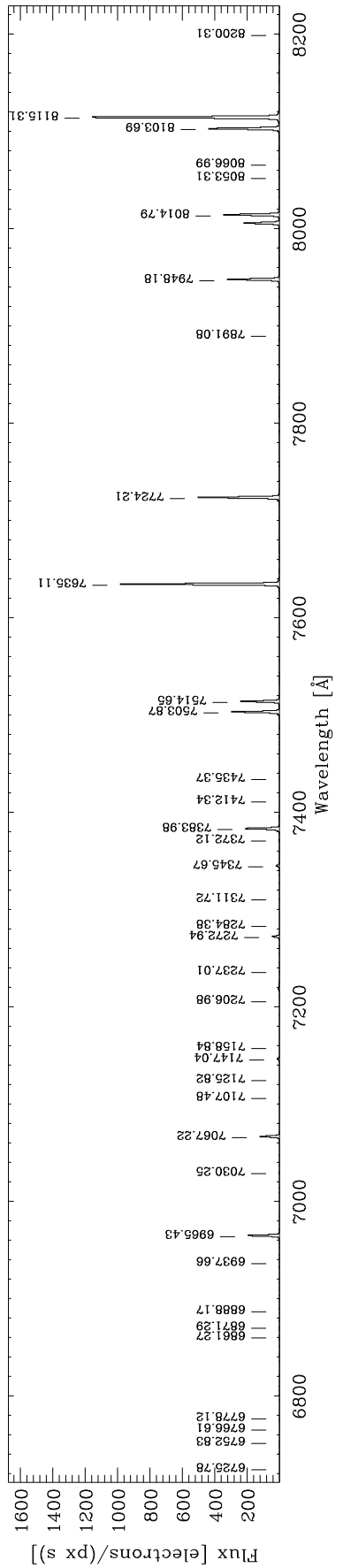
ThAr



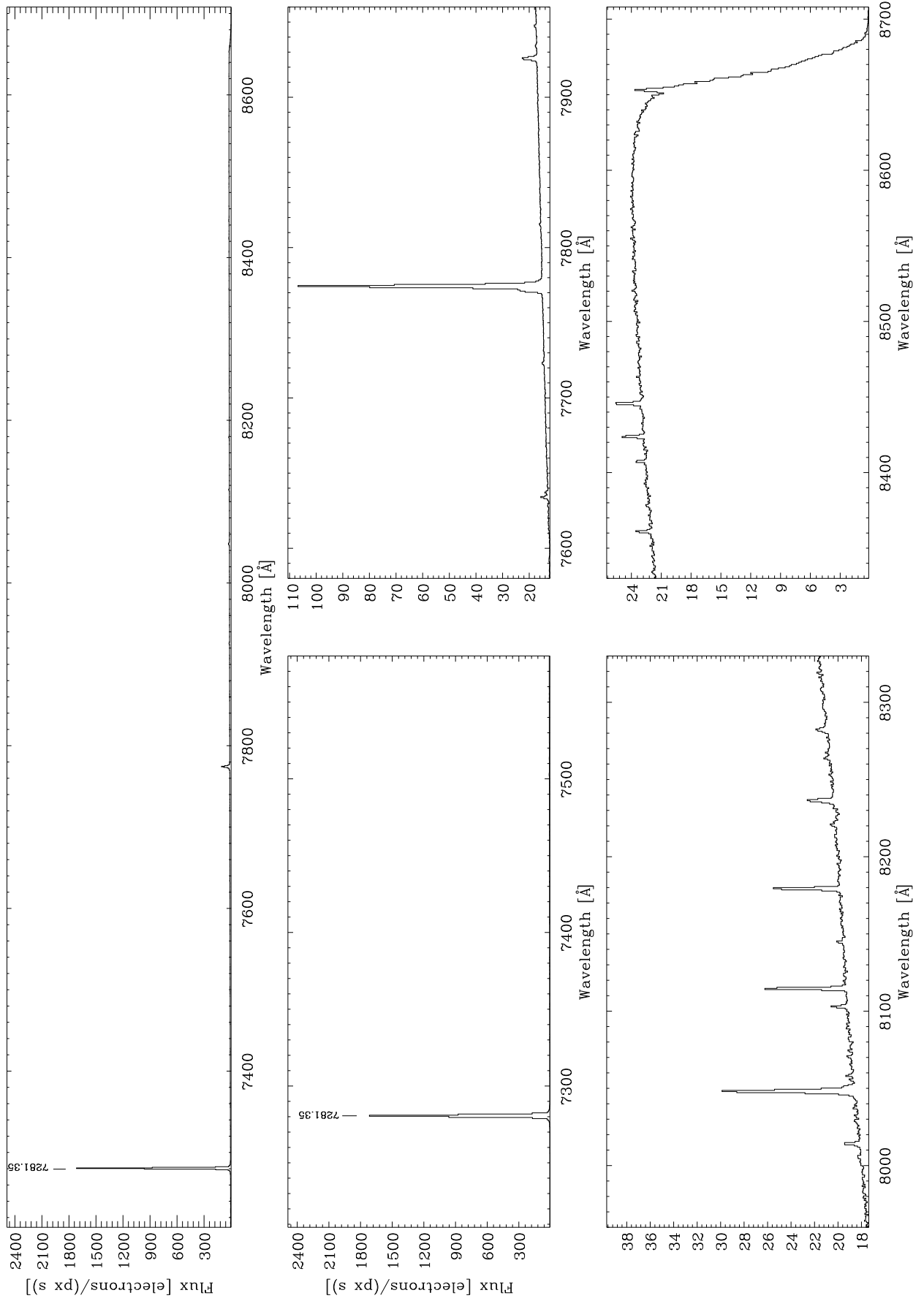
R1200R

$\lambda_{\text{cen}} = 7500\text{\AA}$

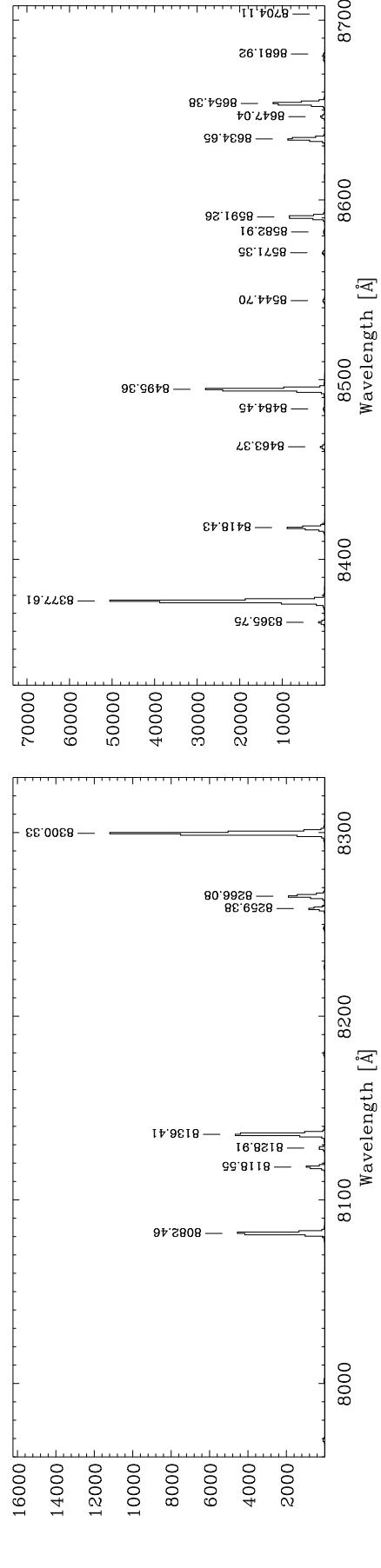
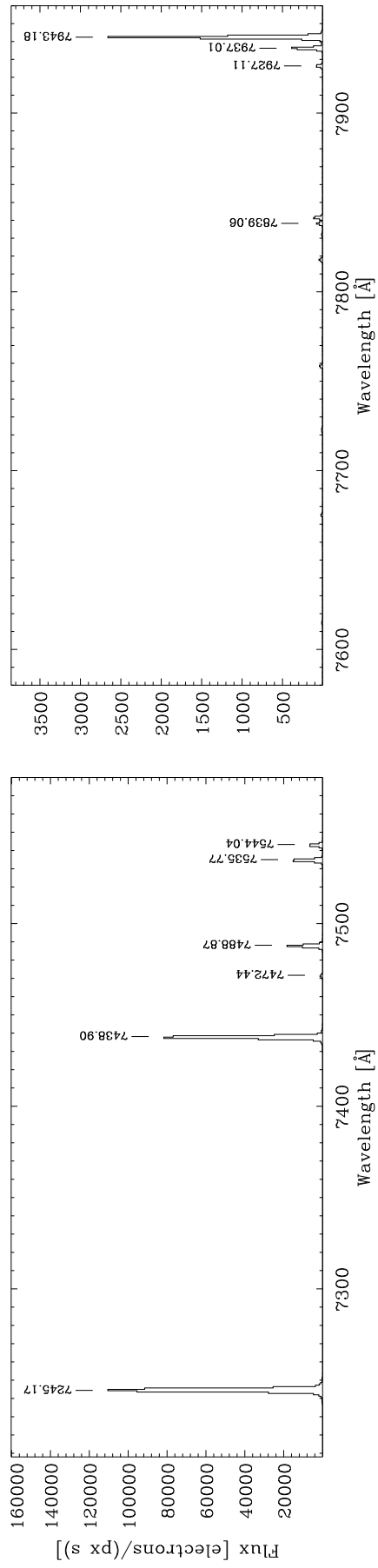
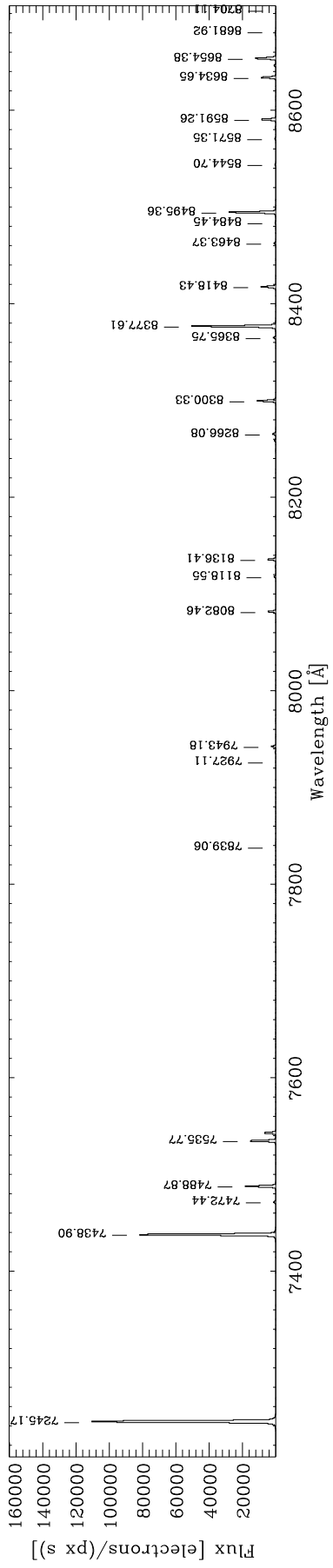
Cd



R1200R $\lambda_{\text{cen}} = 8000\text{\AA}$ He



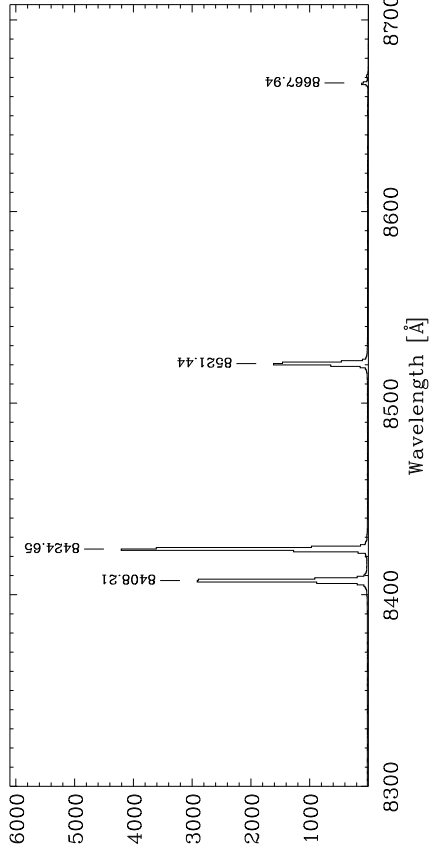
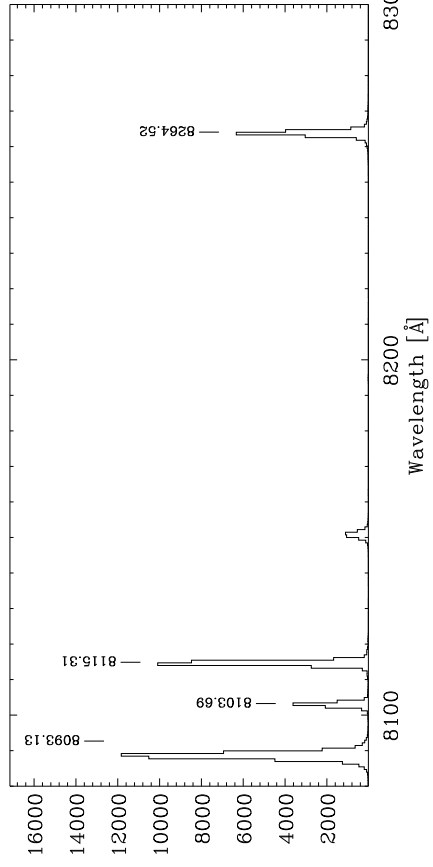
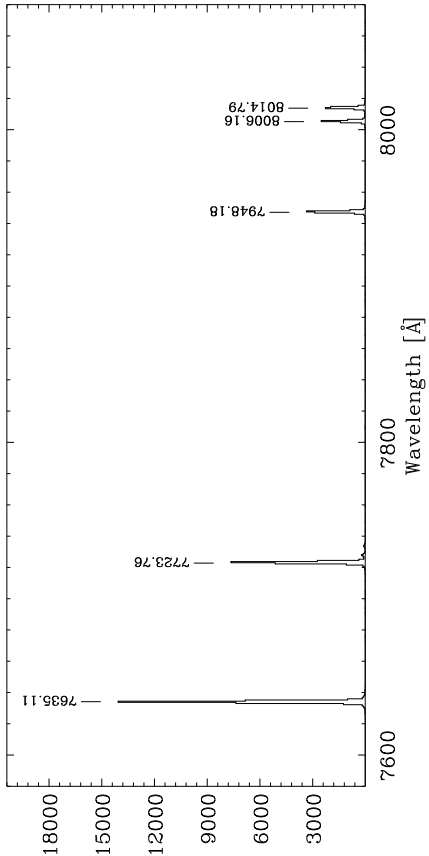
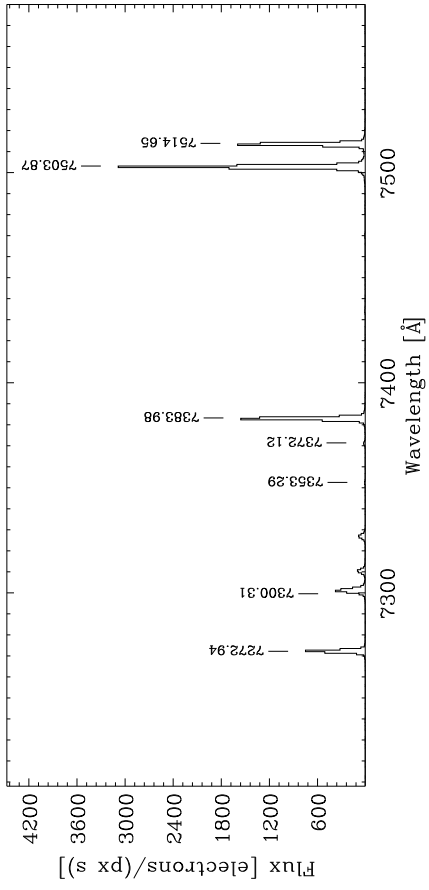
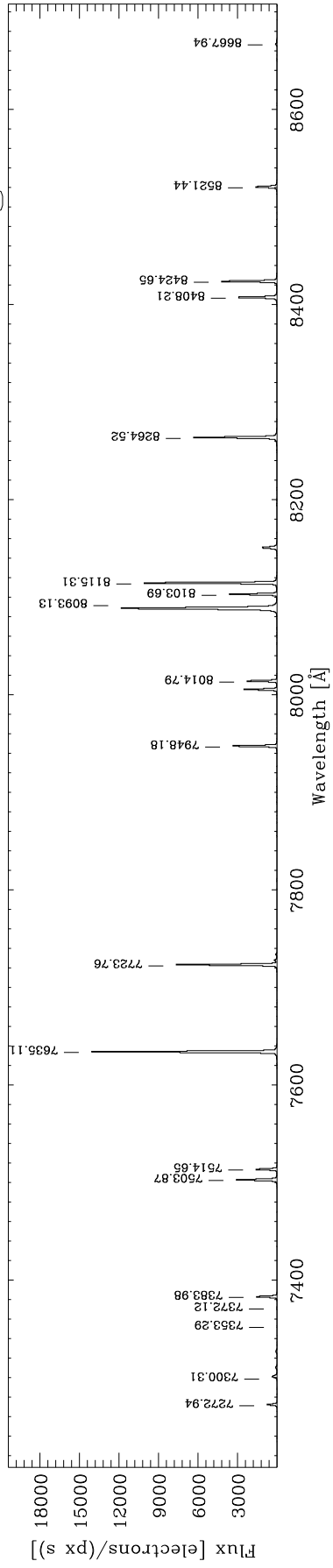
R1200R $\lambda_{\text{cen}} = 8000\text{\AA}$ Ne



R1200R

$\lambda_{\text{cen}} = 8000\text{\AA}$

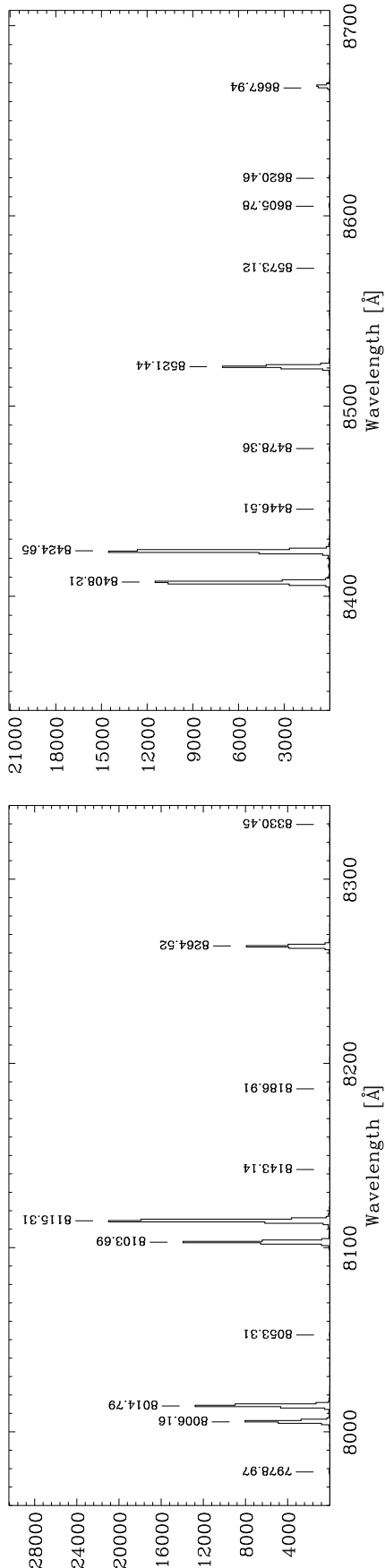
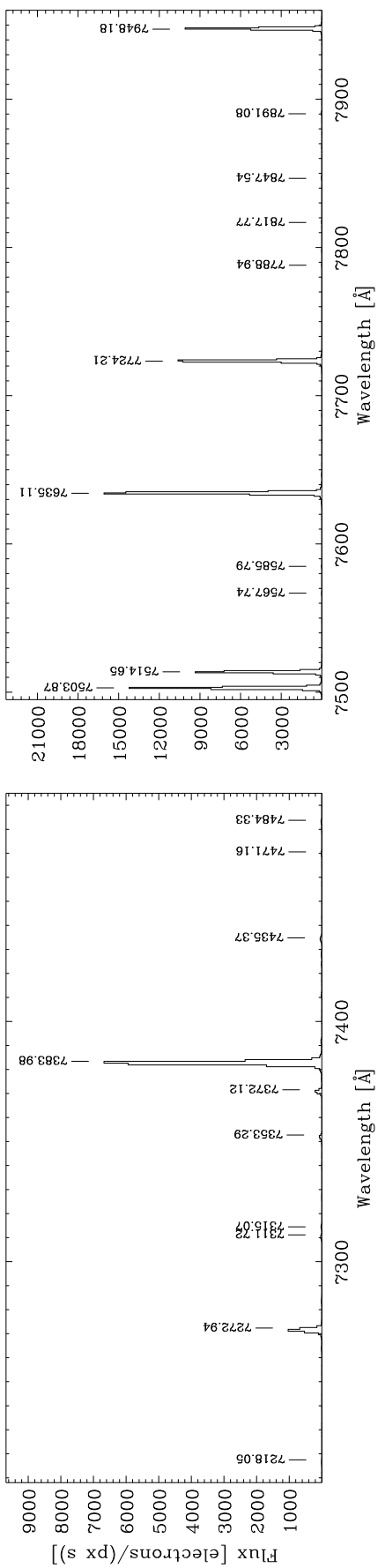
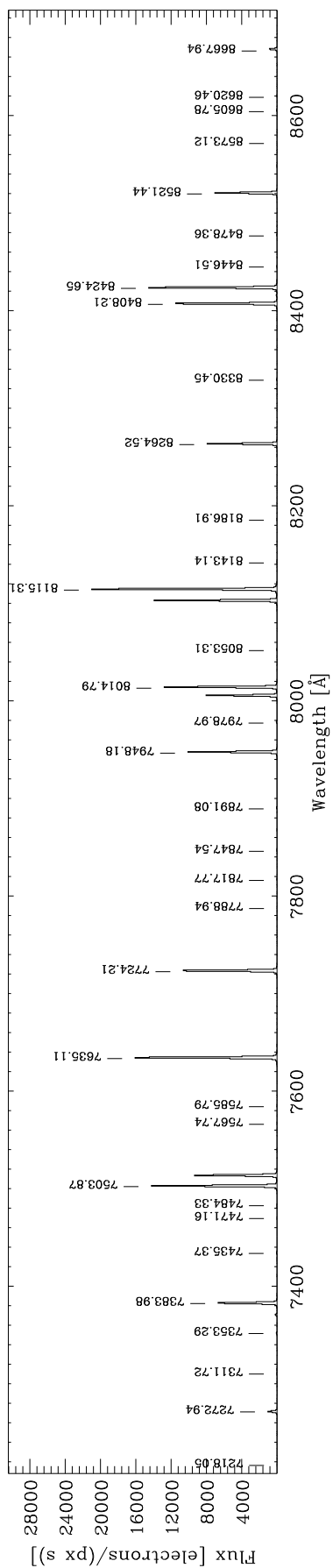
Hg



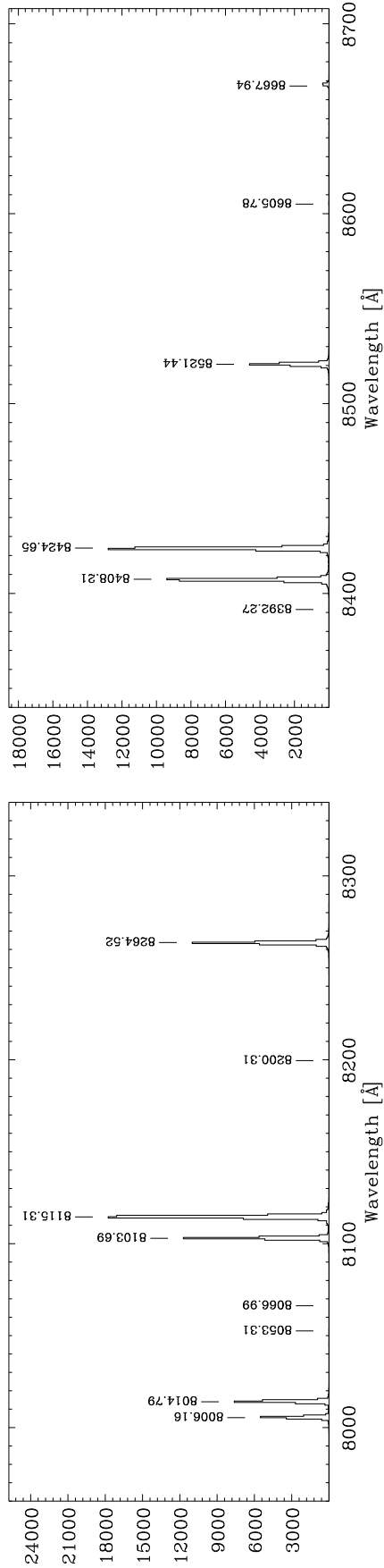
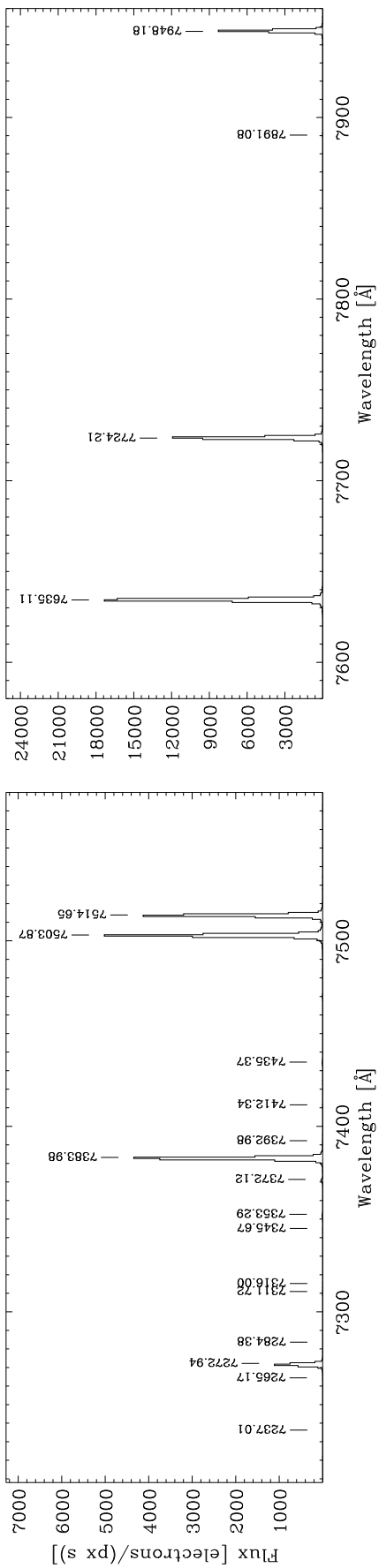
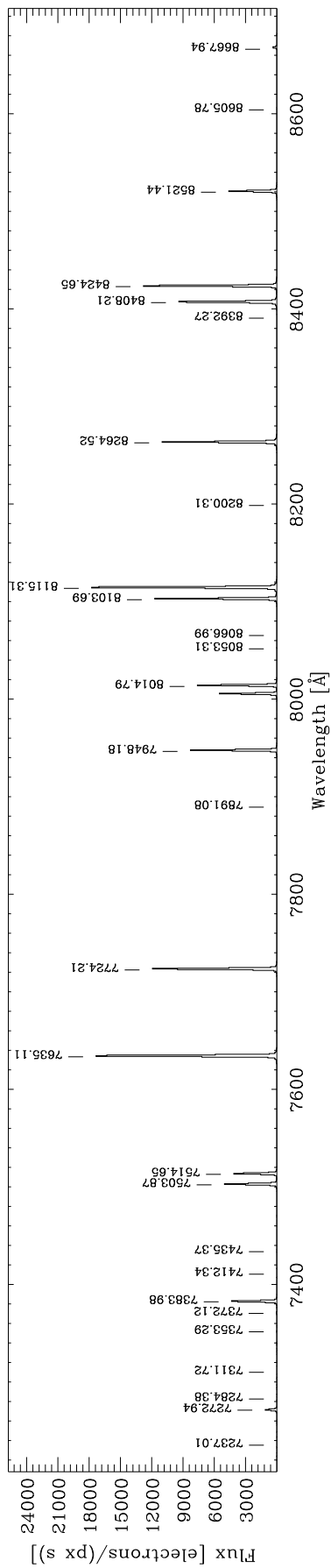
ThAr

$\lambda_{\text{cen}} = 8000\text{\AA}$

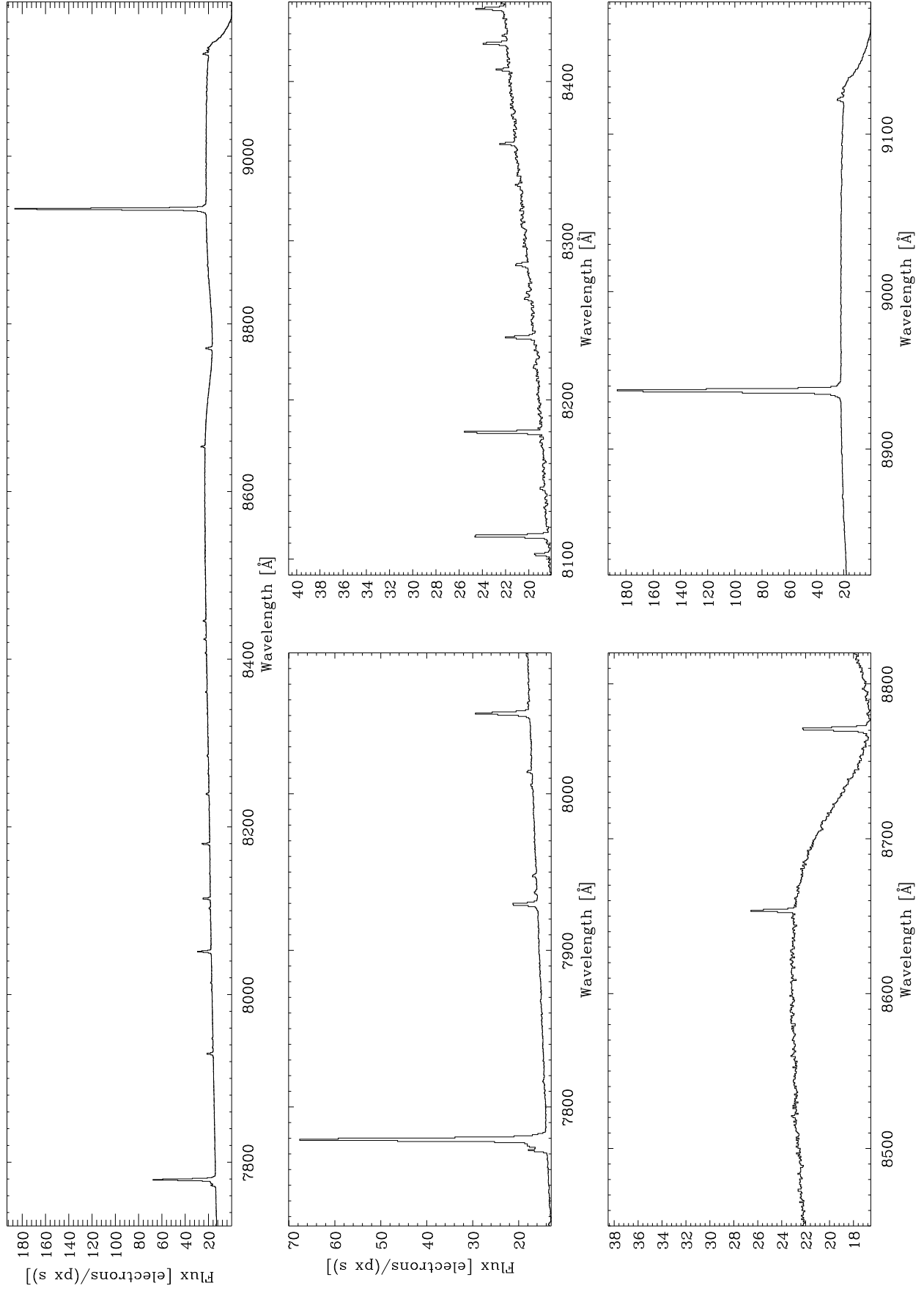
R1200R



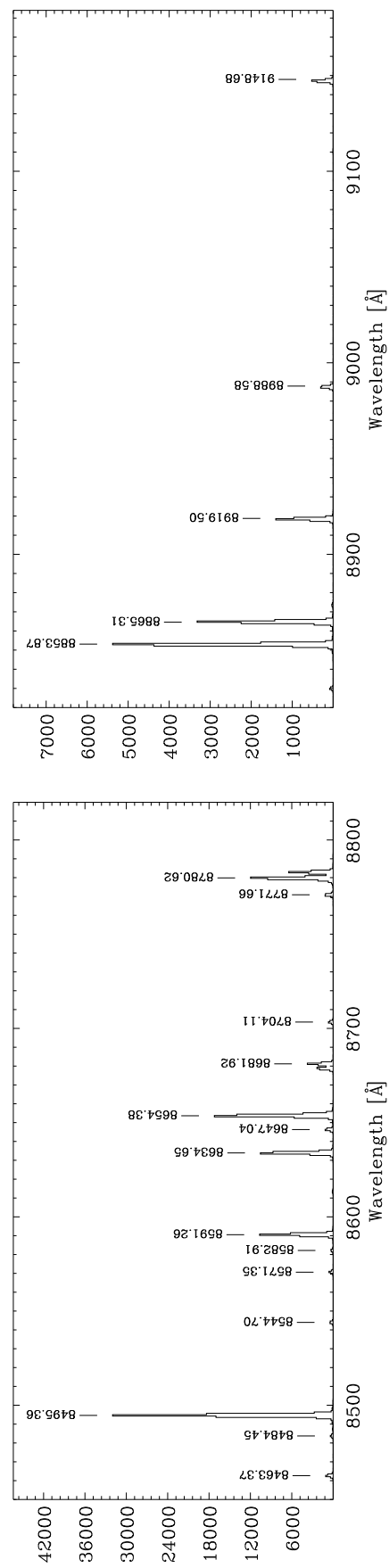
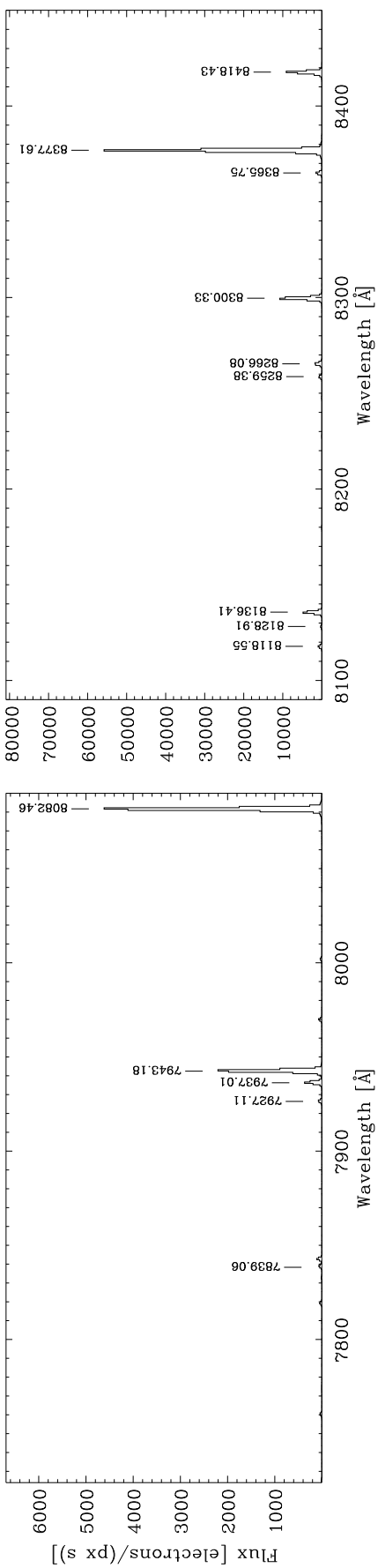
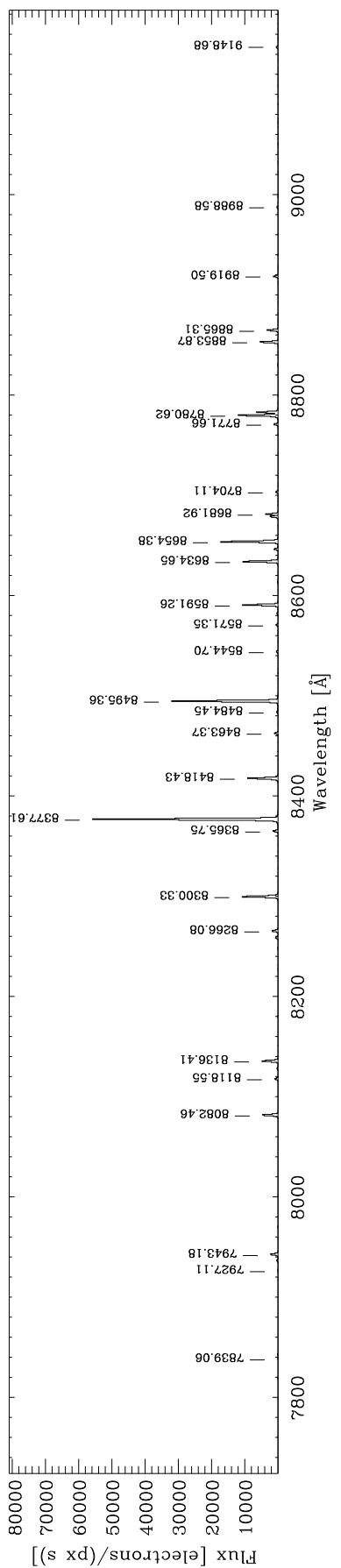
R1200R $\lambda_{\text{cen}} = 8000\text{\AA}$ Cd



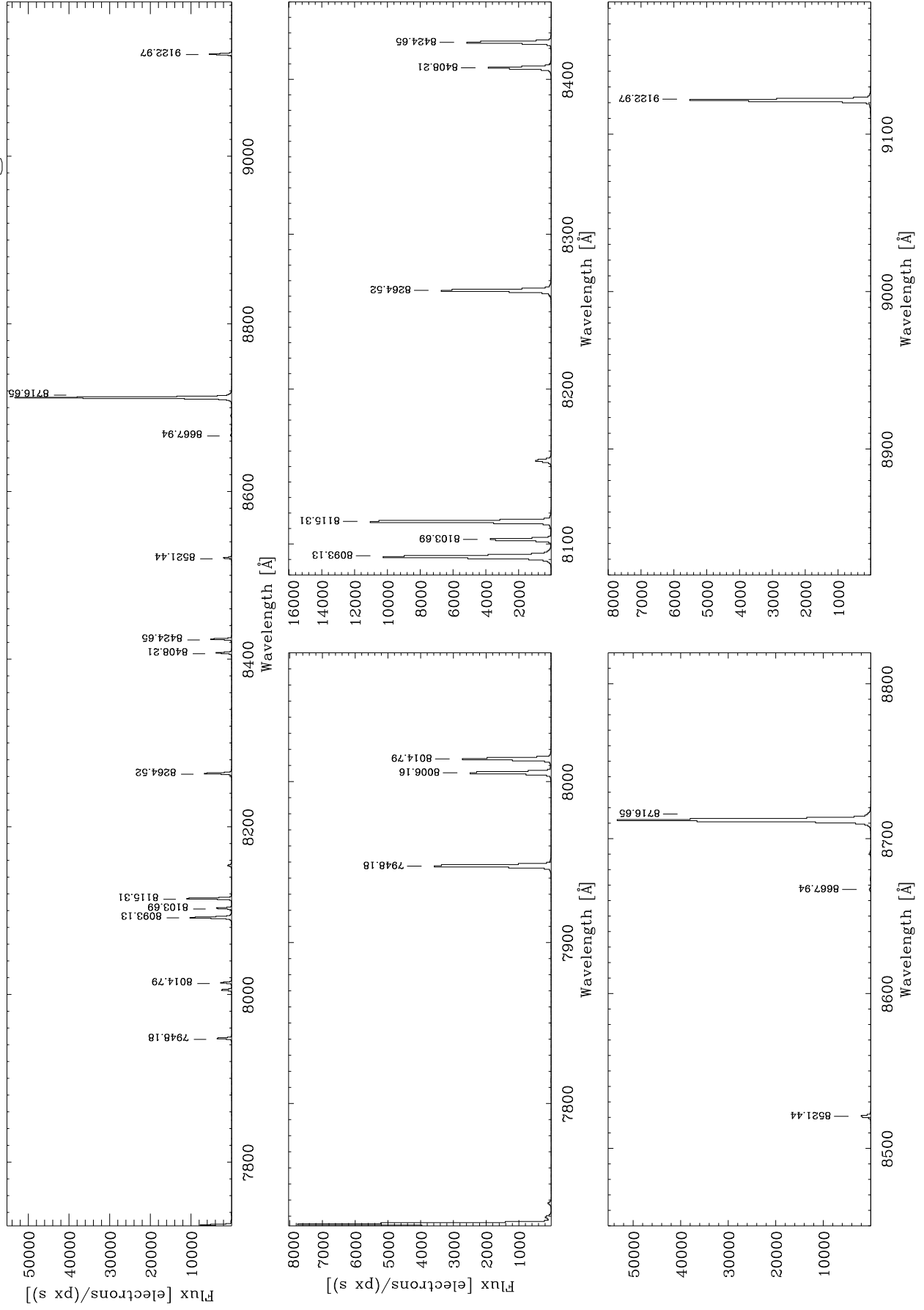
R1200R $\lambda_{\text{cen}} = 8500\text{\AA}$ He



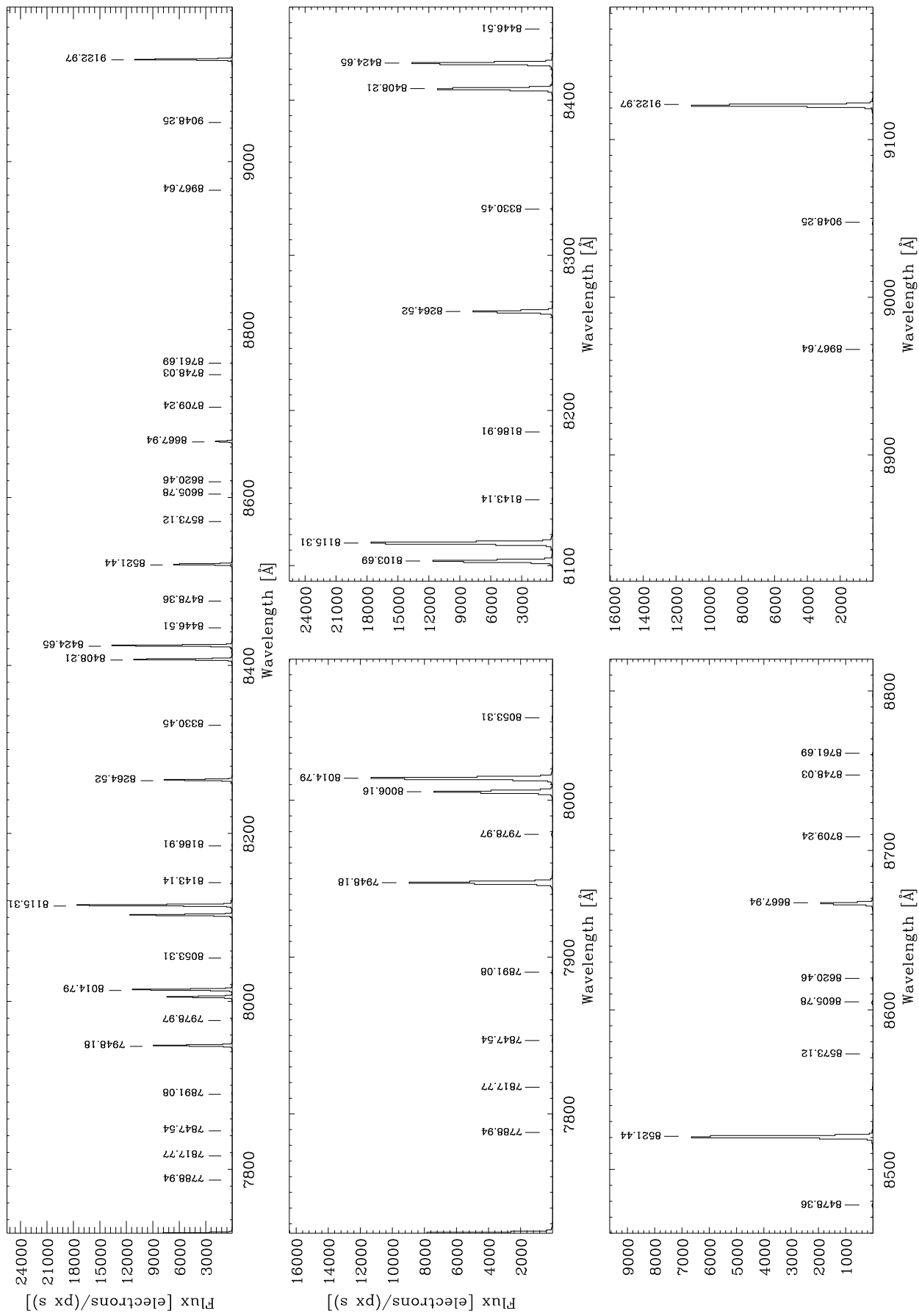
R1200R $\lambda_{\text{cen}} = 8500\text{\AA}$ Ne



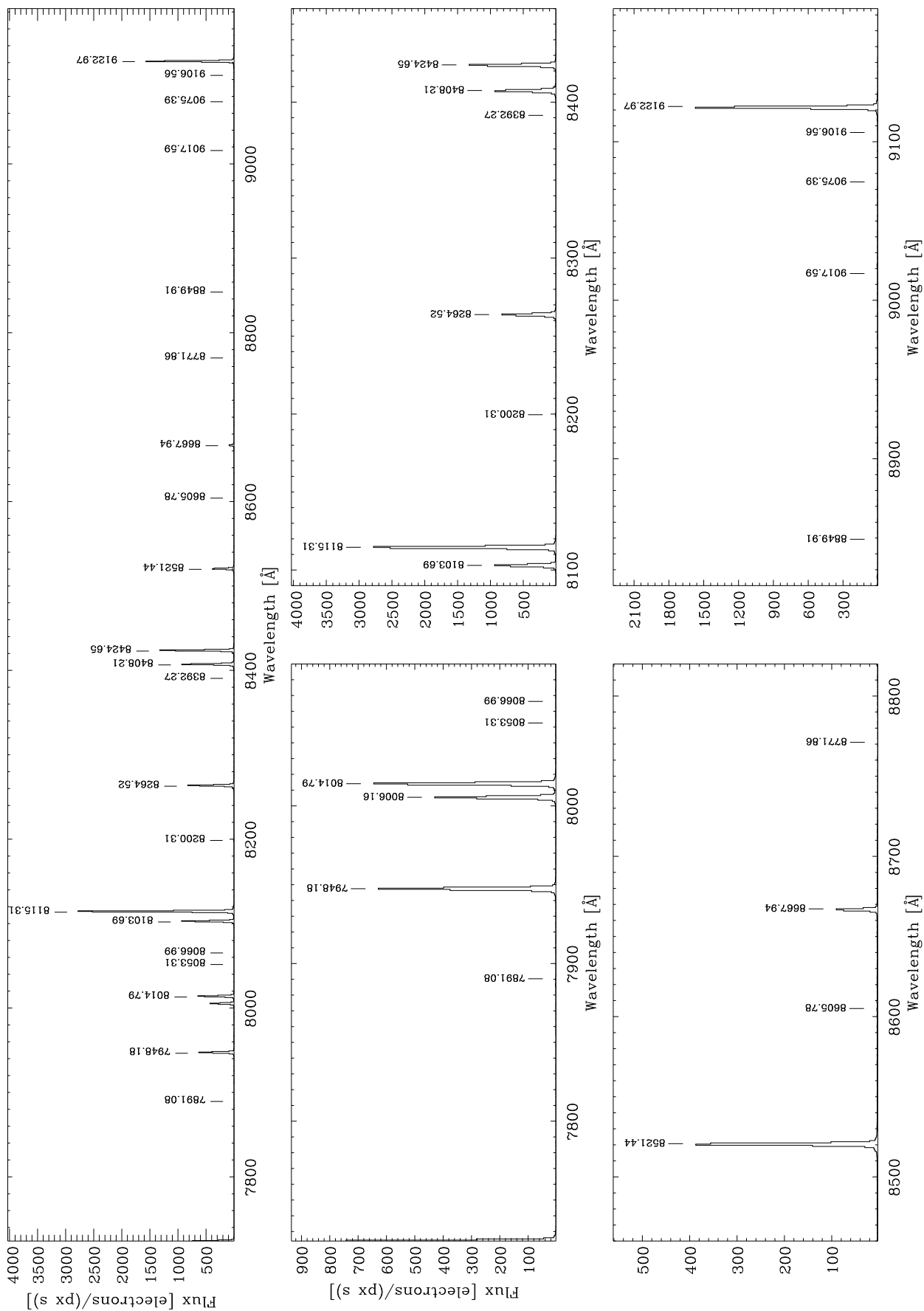
R1200R $\lambda_{\text{cen}} = 8500\text{\AA}$ Hg



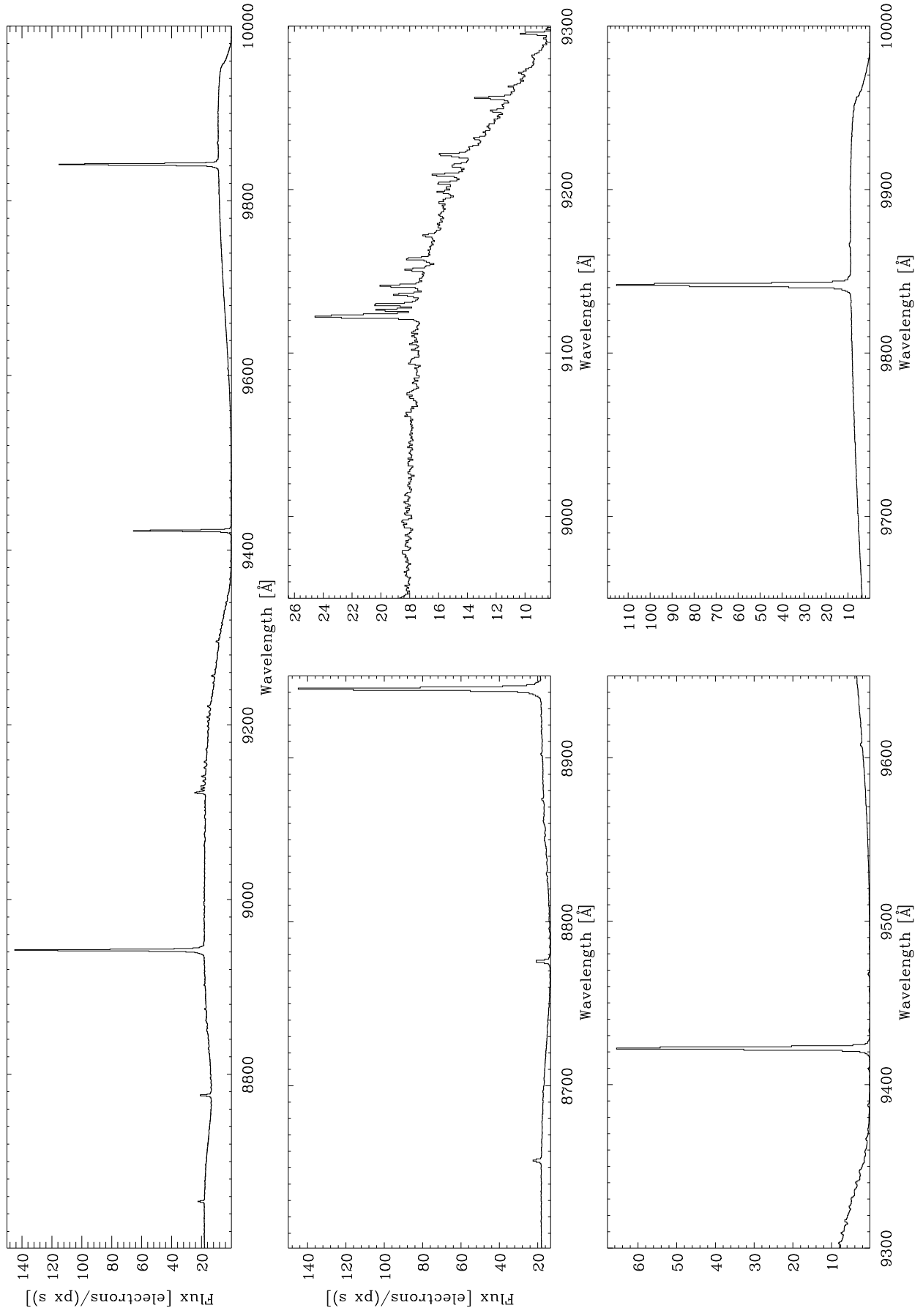
R1200R $\lambda_{\text{cen}} = 8500\text{\AA}$ ThAr



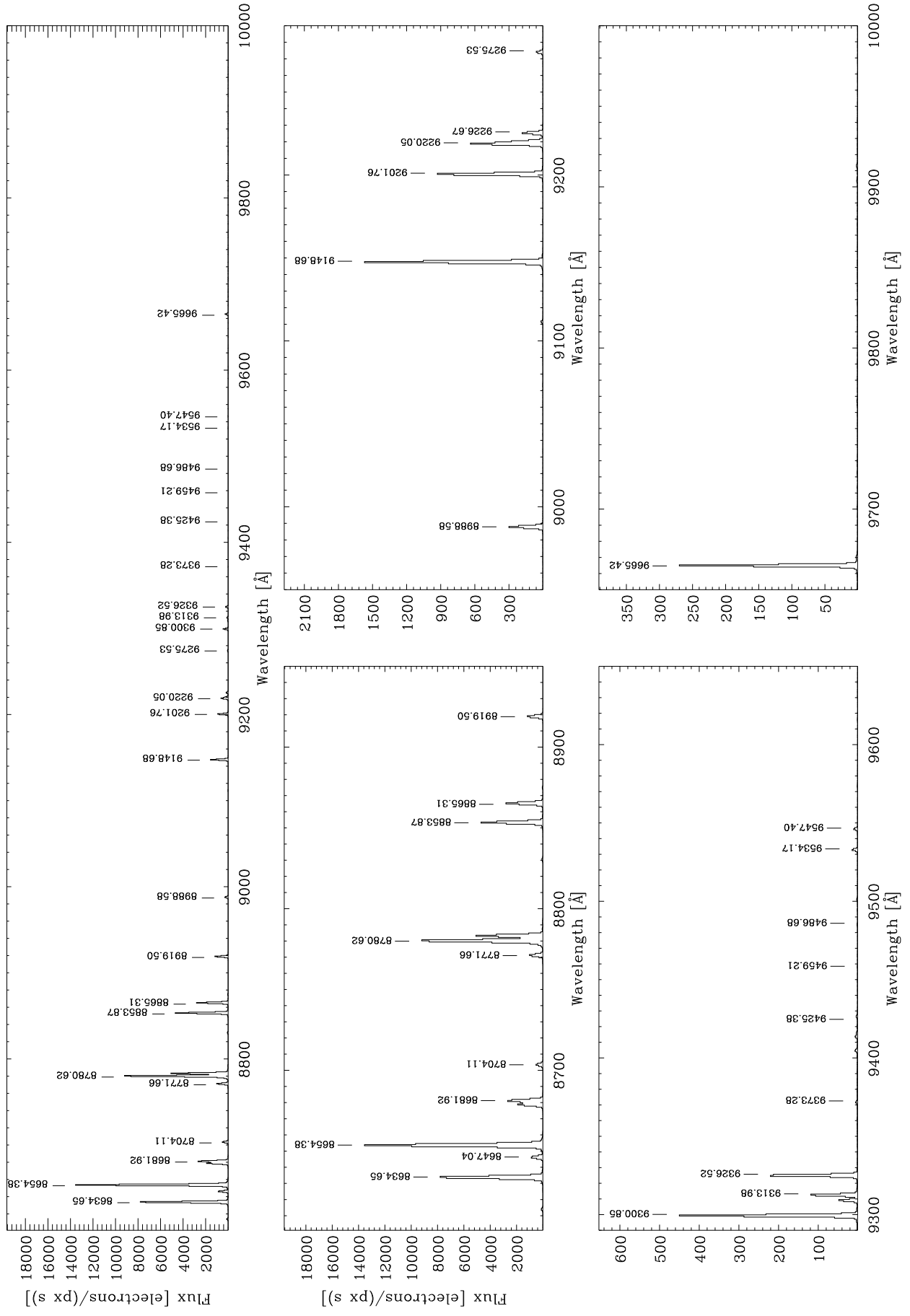
R1200R $\lambda_{\text{cen}} = 8500\text{\AA}$ Cd



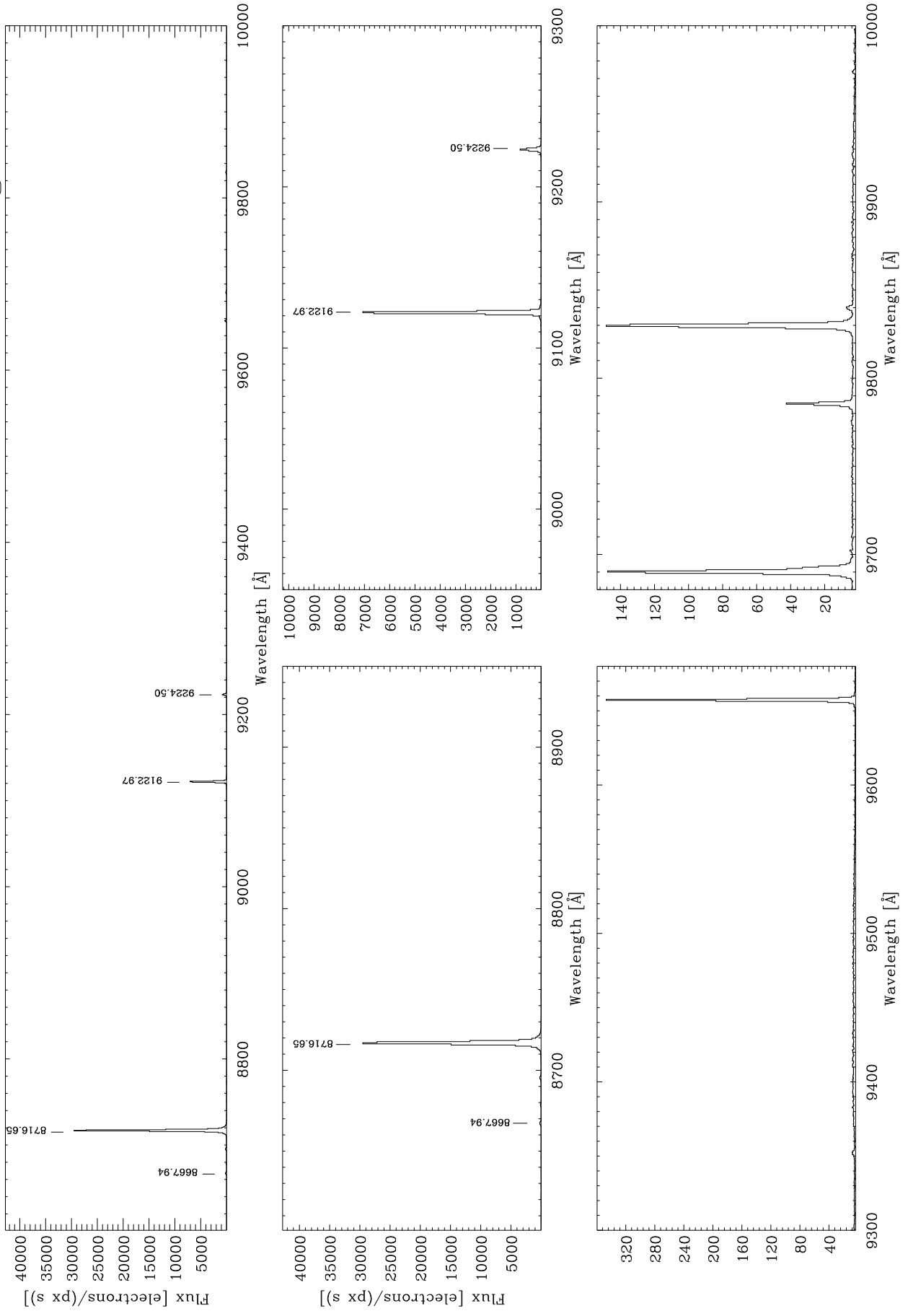
R1200R $\lambda_{\text{cen}} = 9350\text{\AA}$ He



R1200R $\lambda_{\text{cen}} = 9350\text{\AA}$ Ne



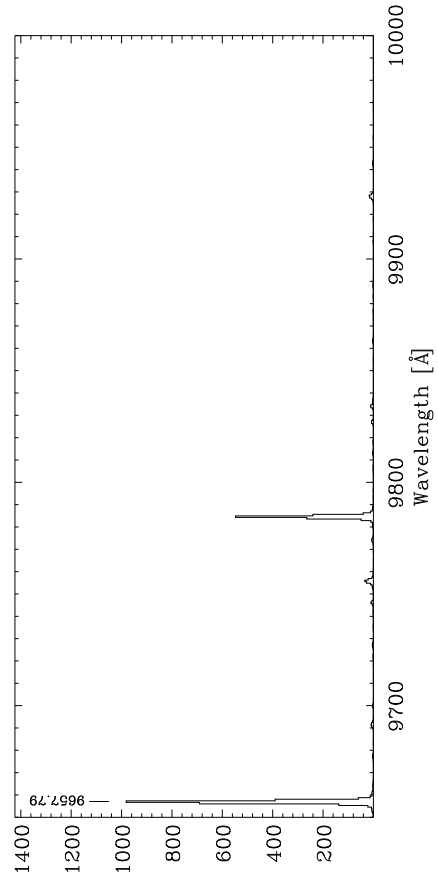
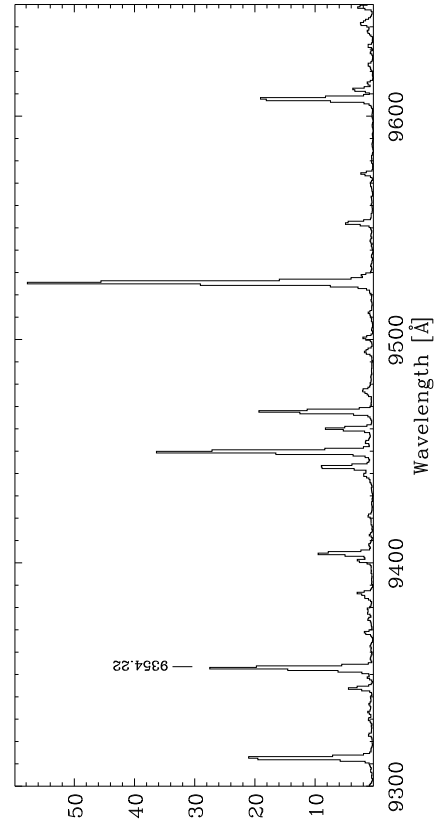
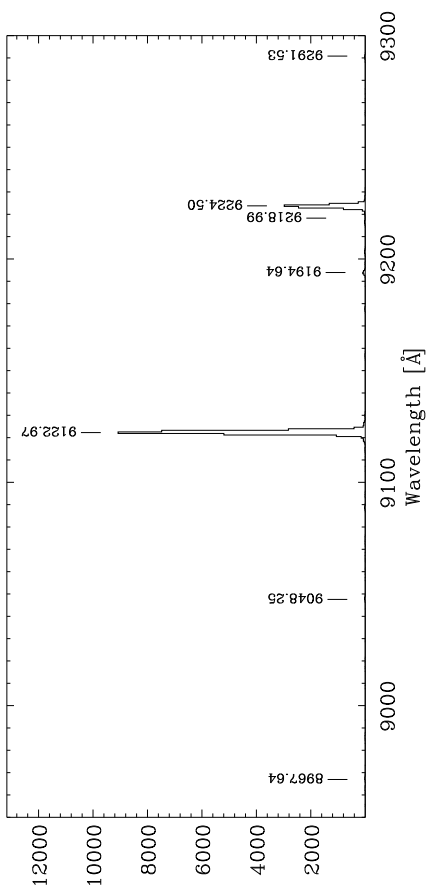
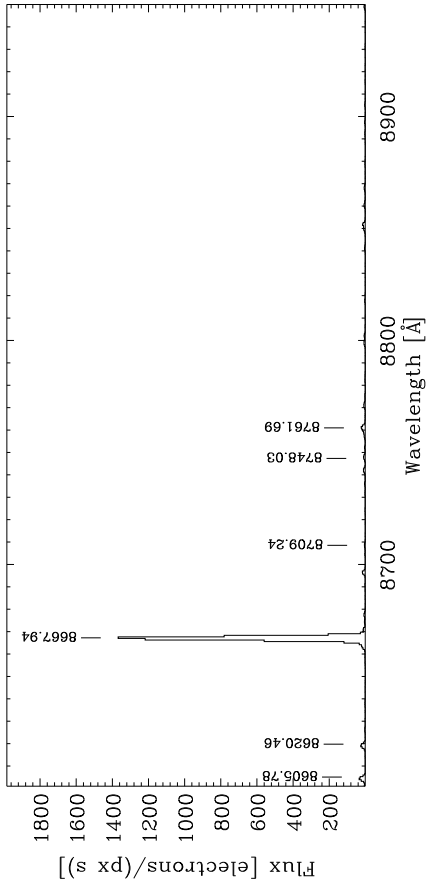
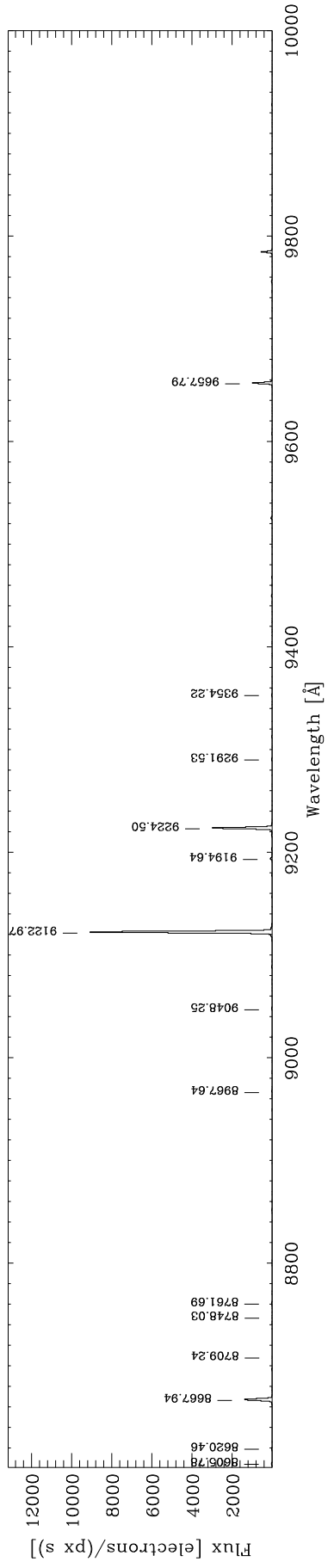
R1200R $\lambda_{\text{cen}} = 9350\text{\AA}$ Hg



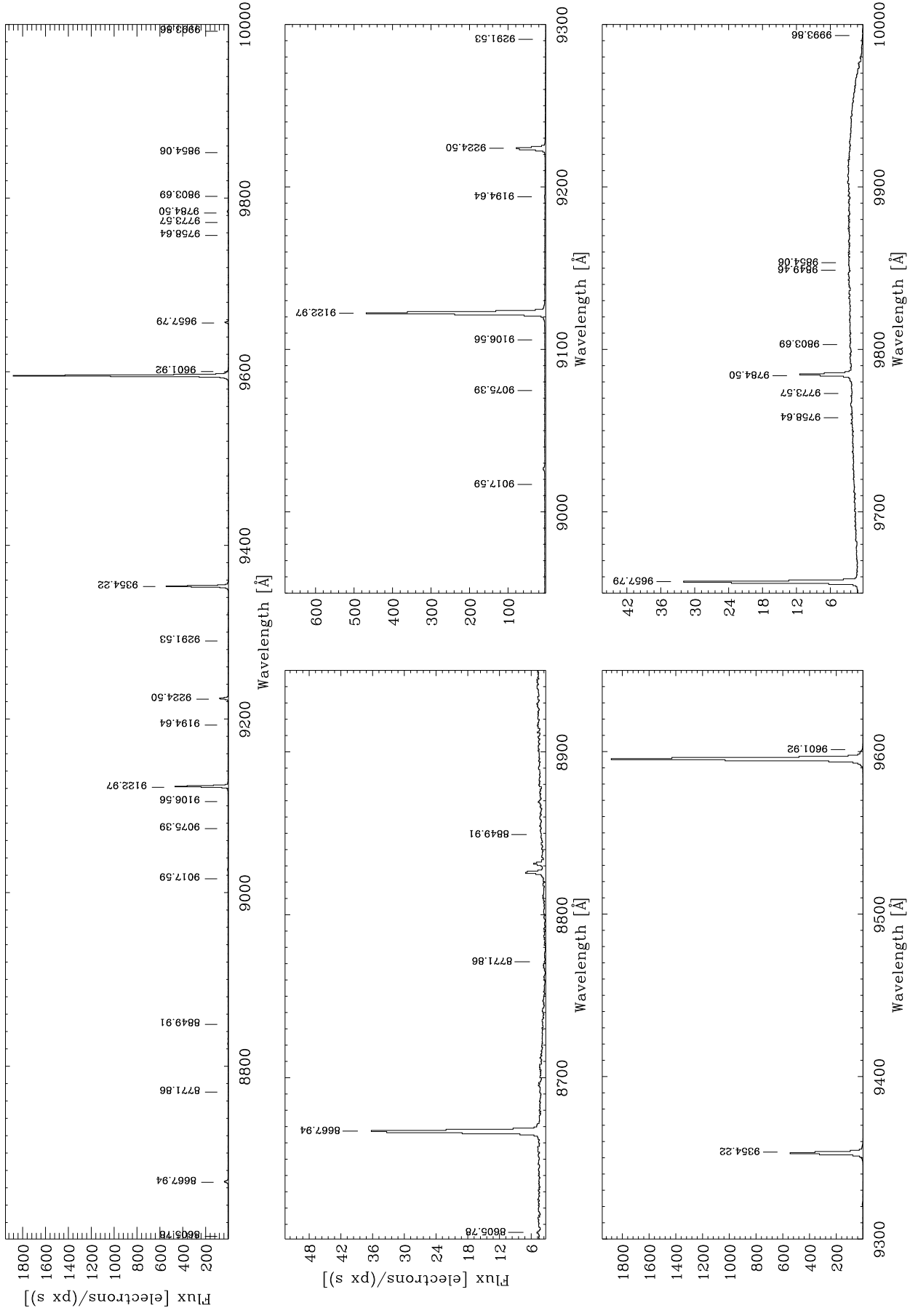
R1200R

$\lambda_{\text{cen}} = 9350\text{\AA}$

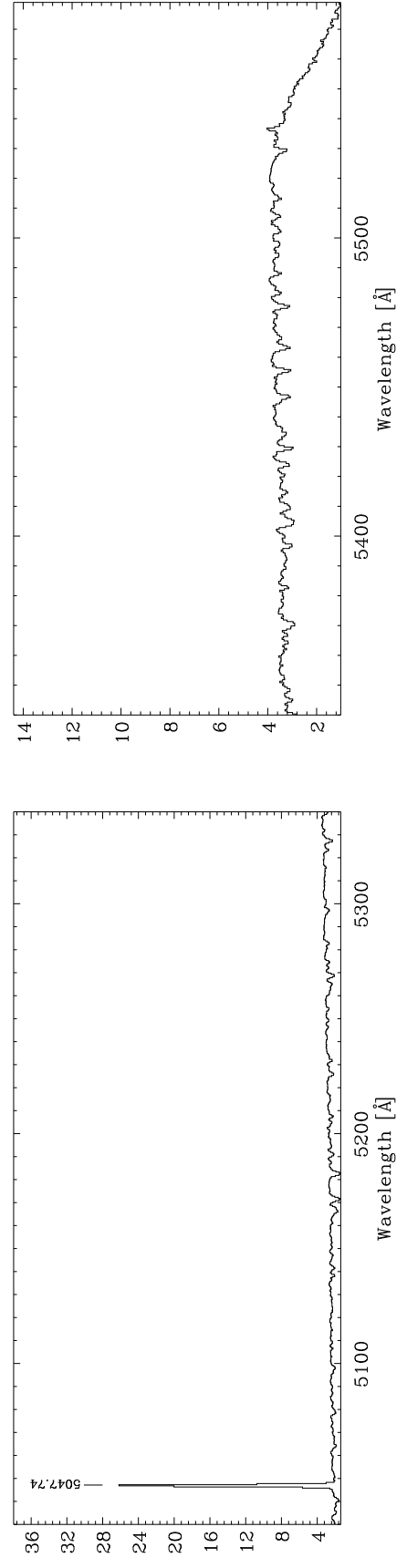
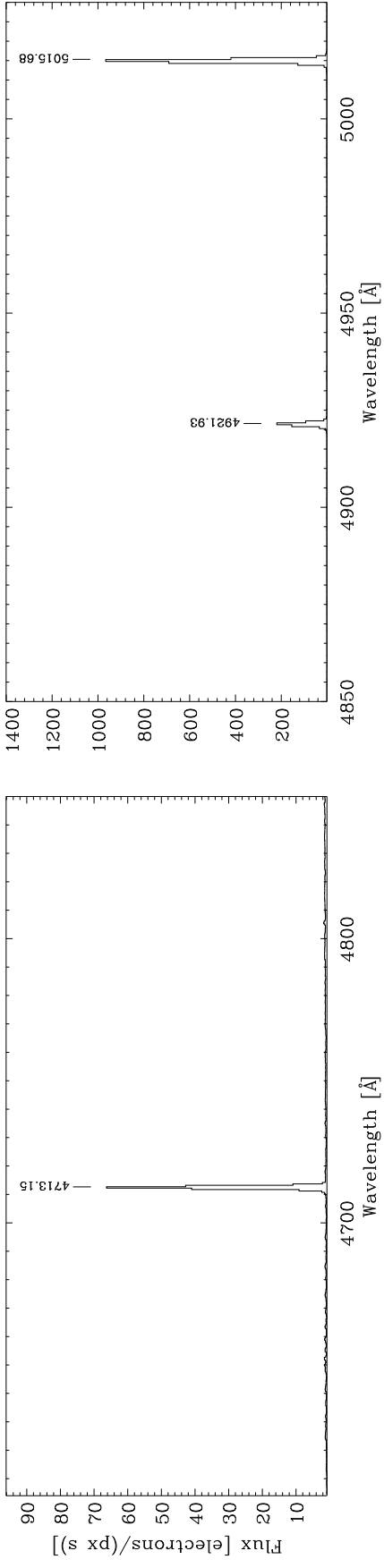
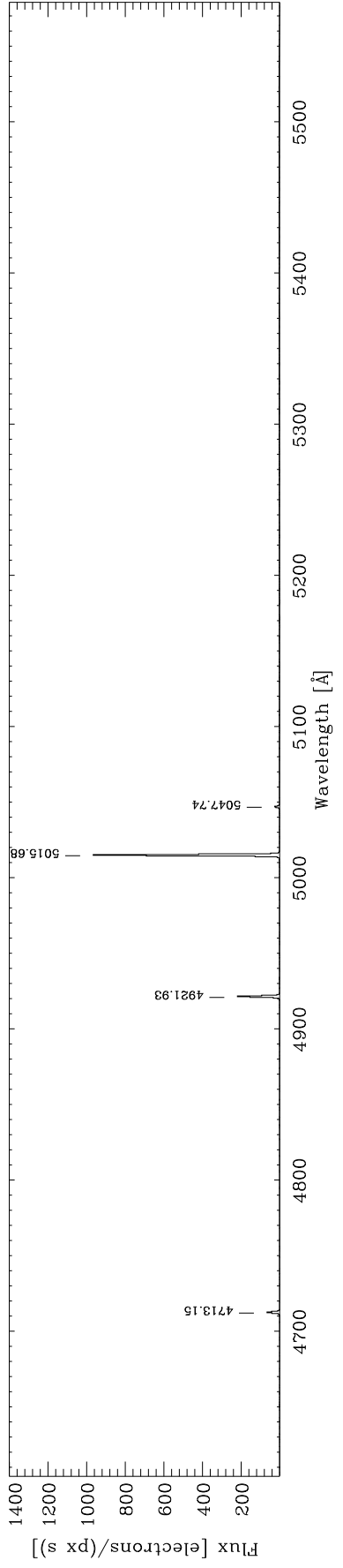
ThAr



R1200R $\lambda_{\text{cen}} = 9350\text{\AA}$ Cd



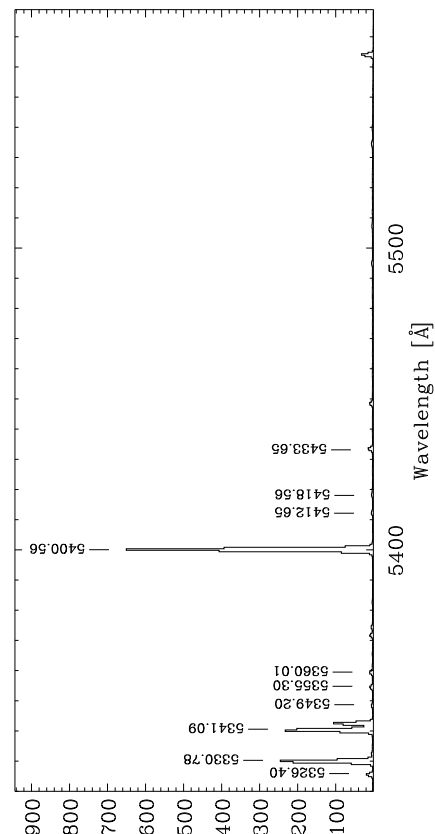
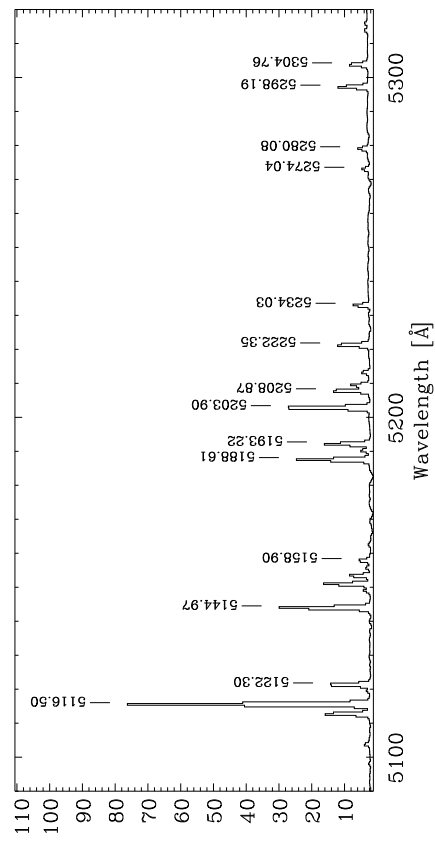
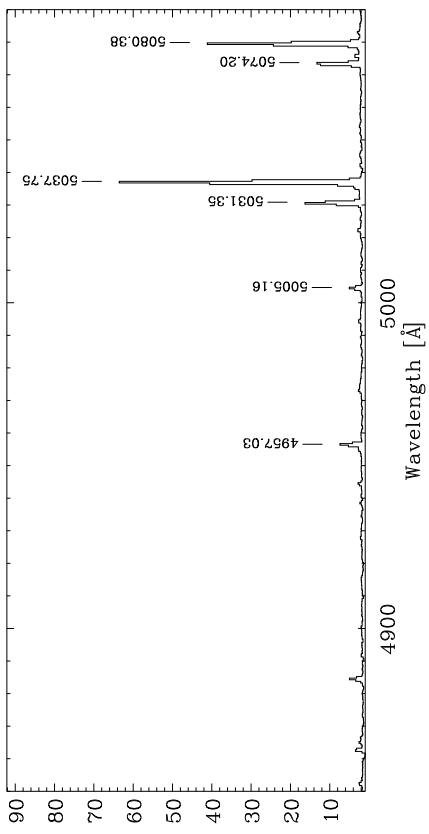
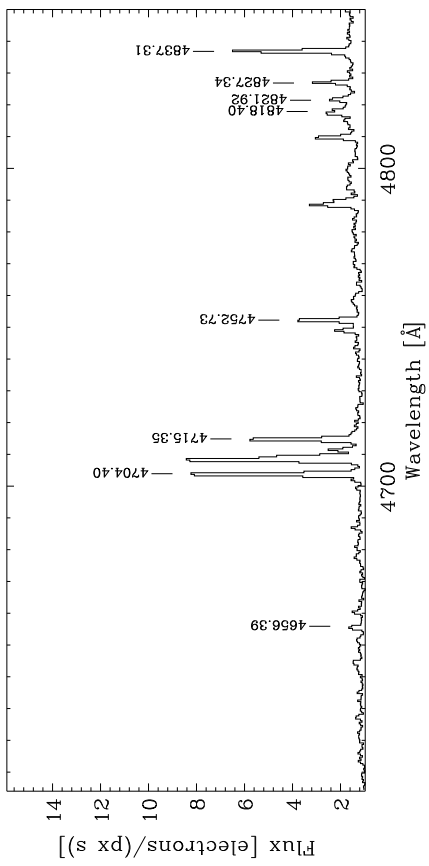
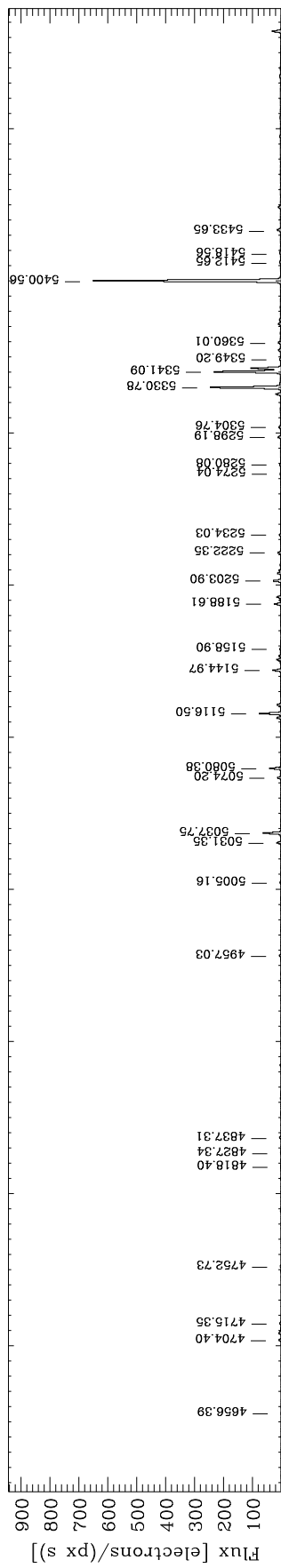
H1800V $\lambda_{\text{cen}} = 5120\text{\AA}$ He



H1800V

$\lambda_{\text{cen}} = 5120\text{\AA}$

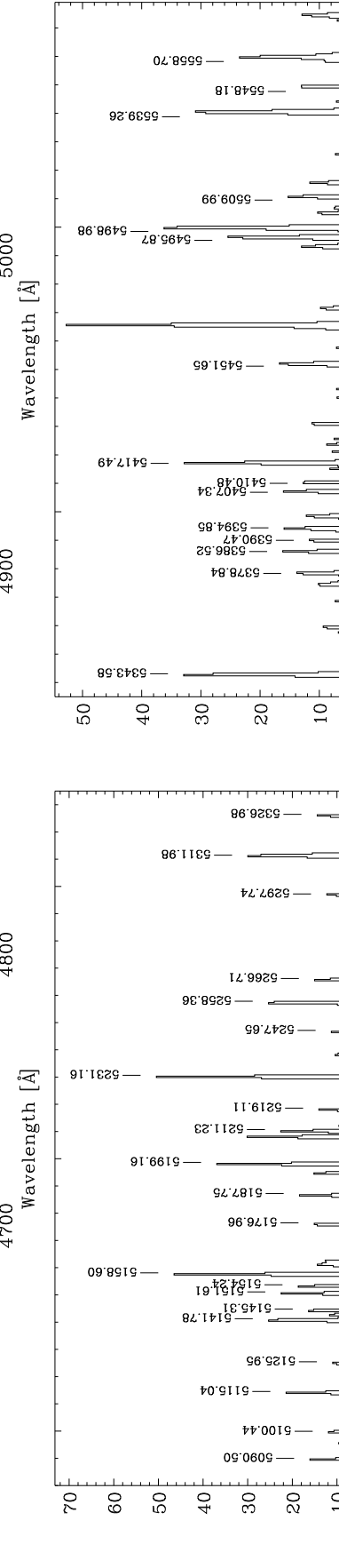
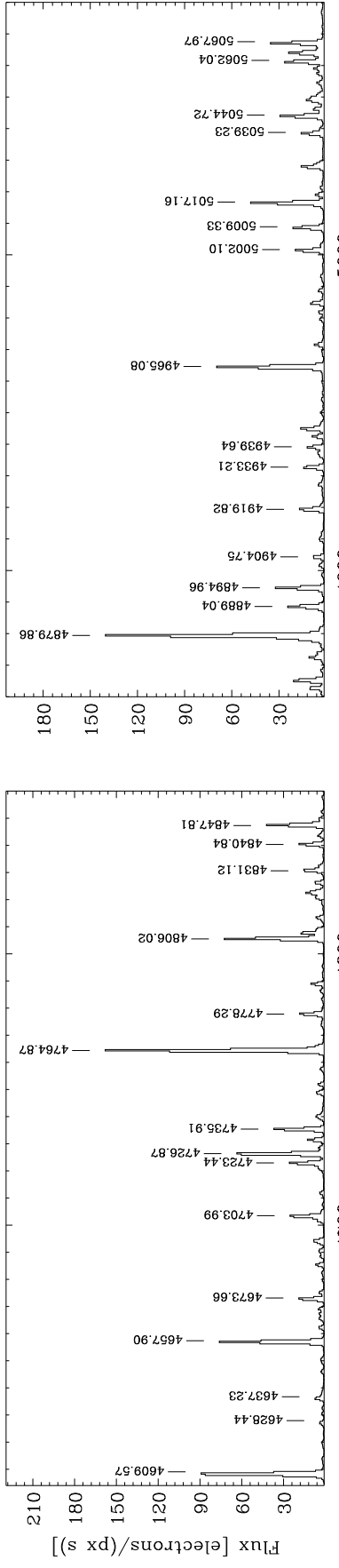
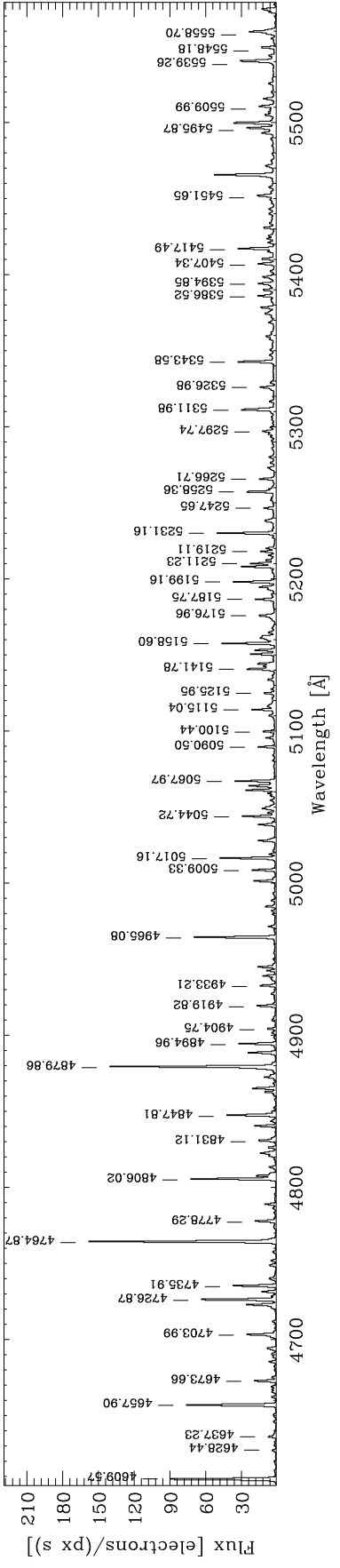
Ne



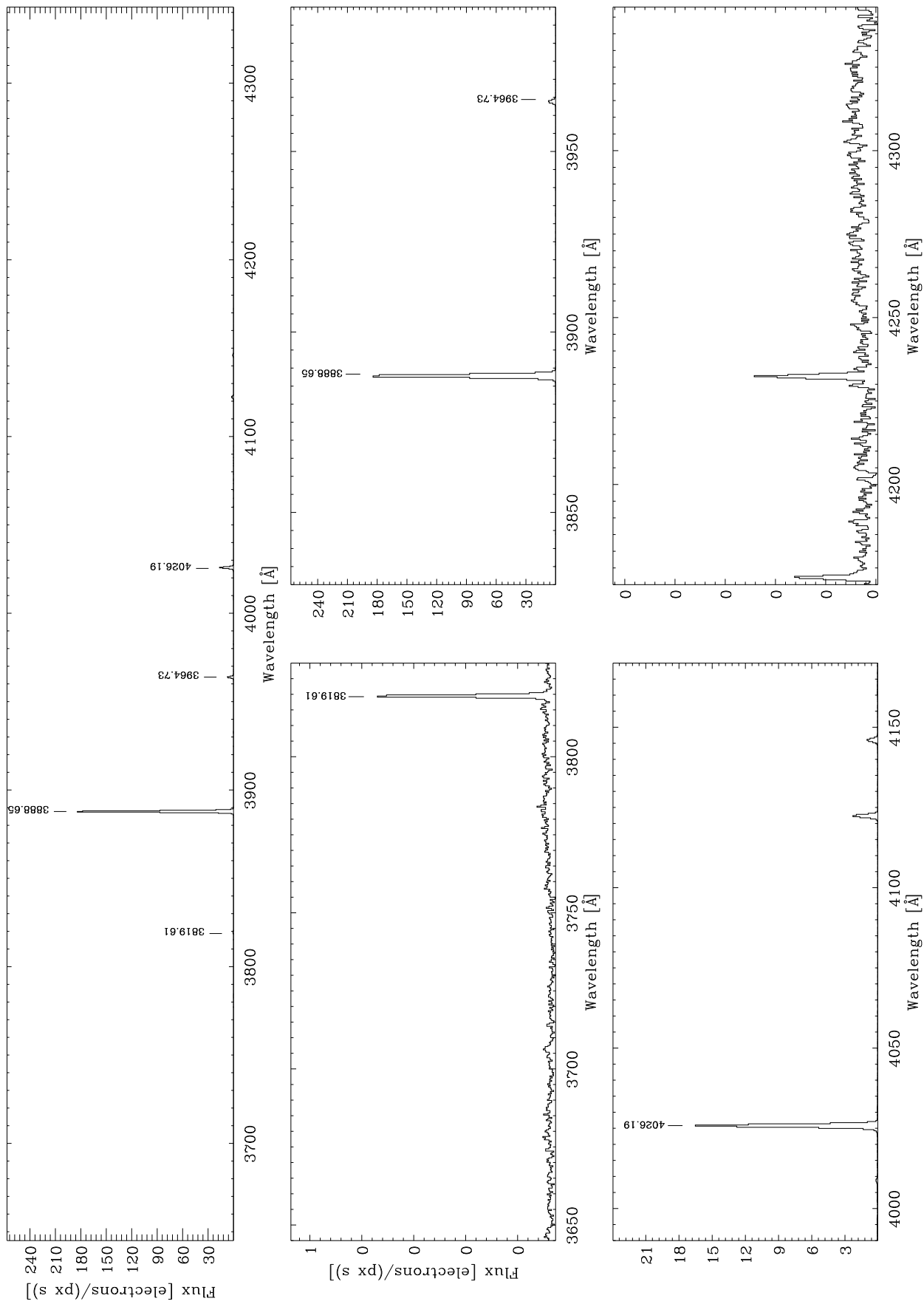
H1800V

$\lambda_{\text{cen}} = 5120\text{\AA}$

ThAr



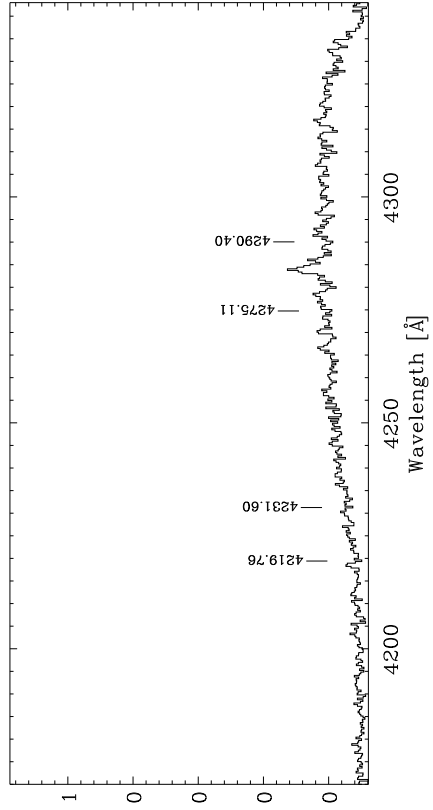
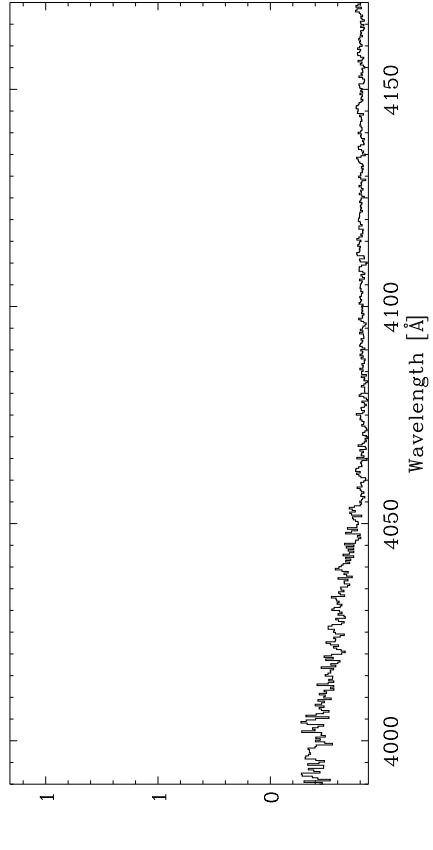
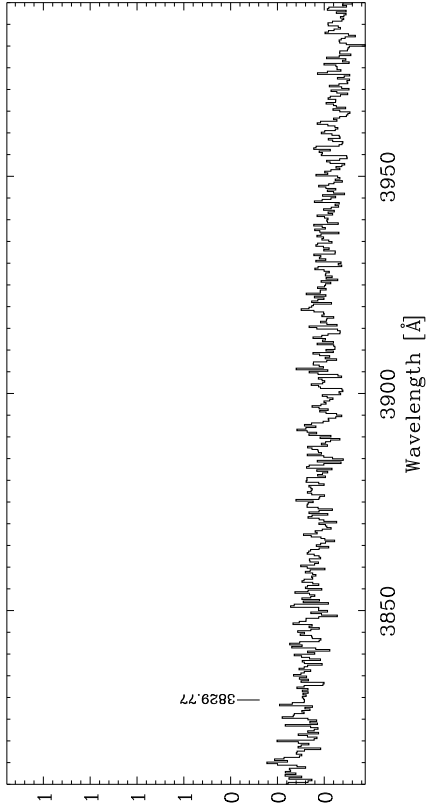
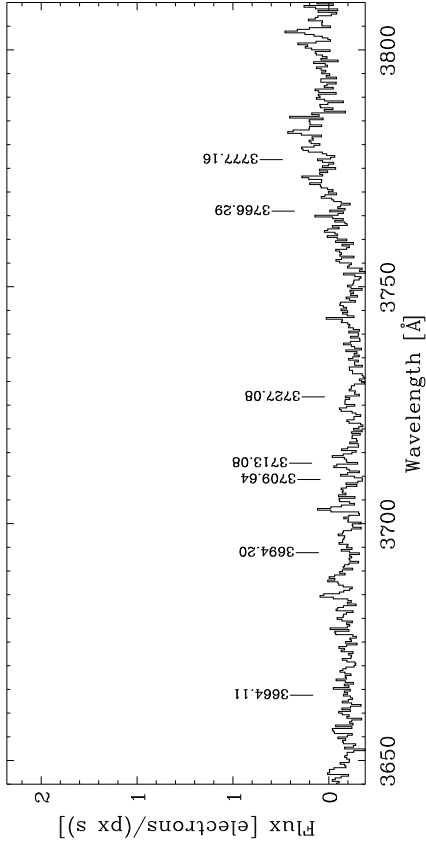
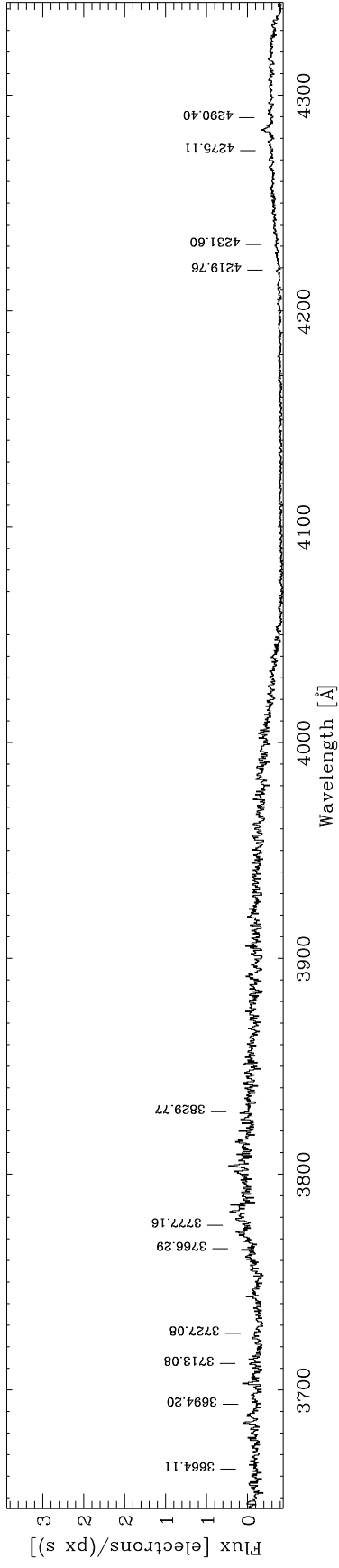
H2400B $\lambda_{\text{cen}} = 4000\text{\AA}$ He



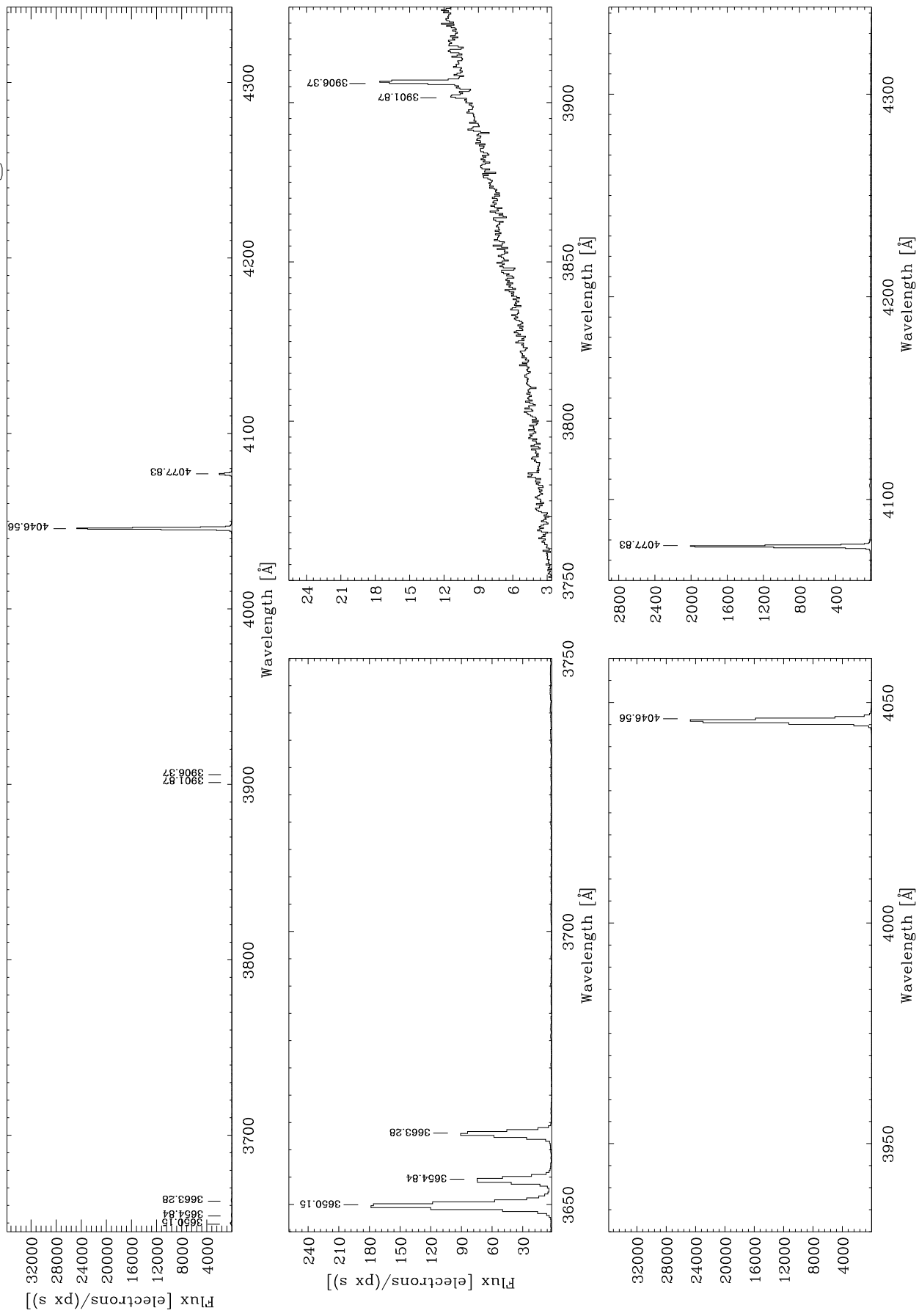
H2400B

$\lambda_{\text{cen}} = 4000\text{\AA}$

Ne



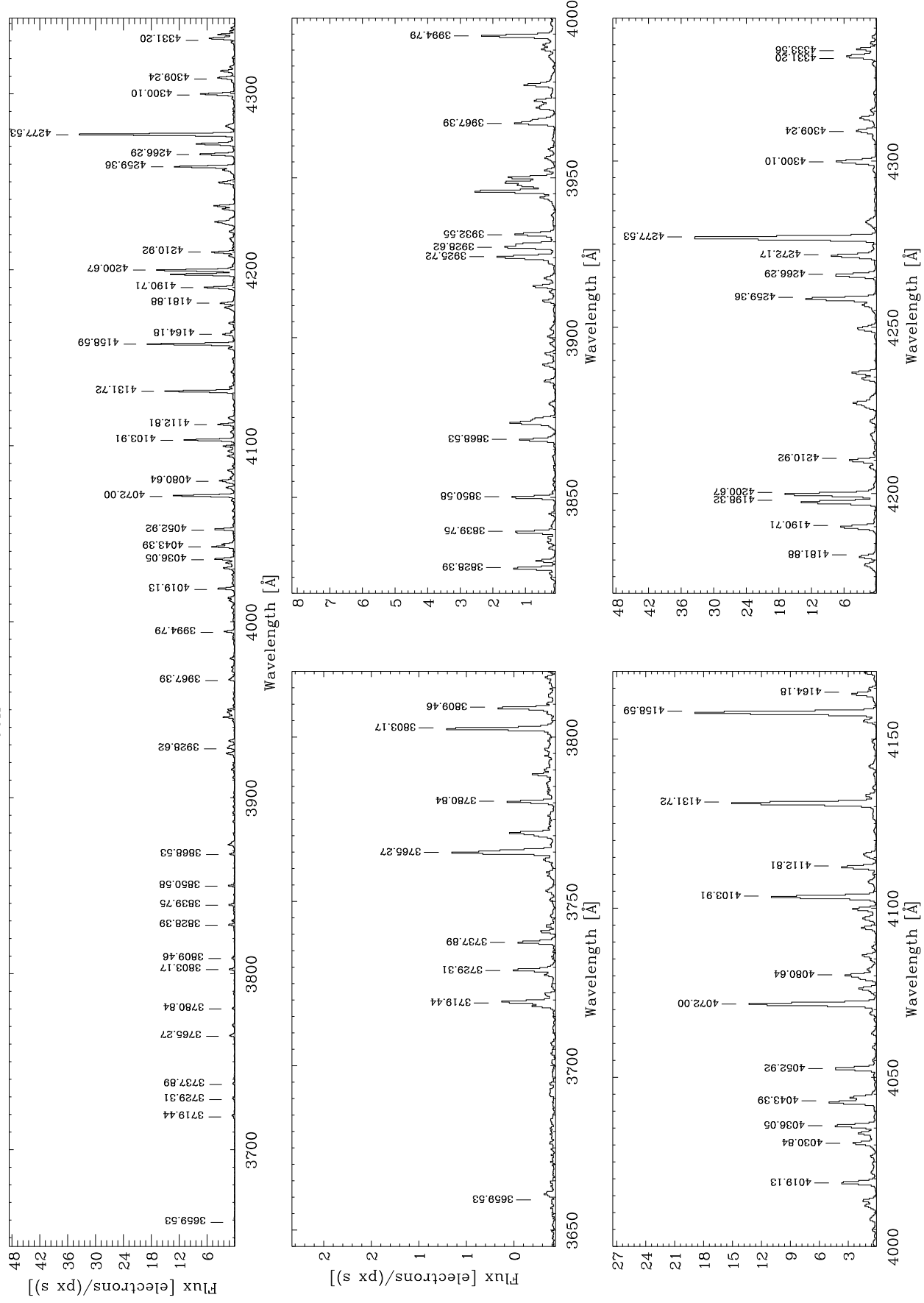
H2400B $\lambda_{\text{cen}} = 4000\text{\AA}$ H γ



H2400B

$\lambda_{\text{cen}} = 4000\text{\AA}$

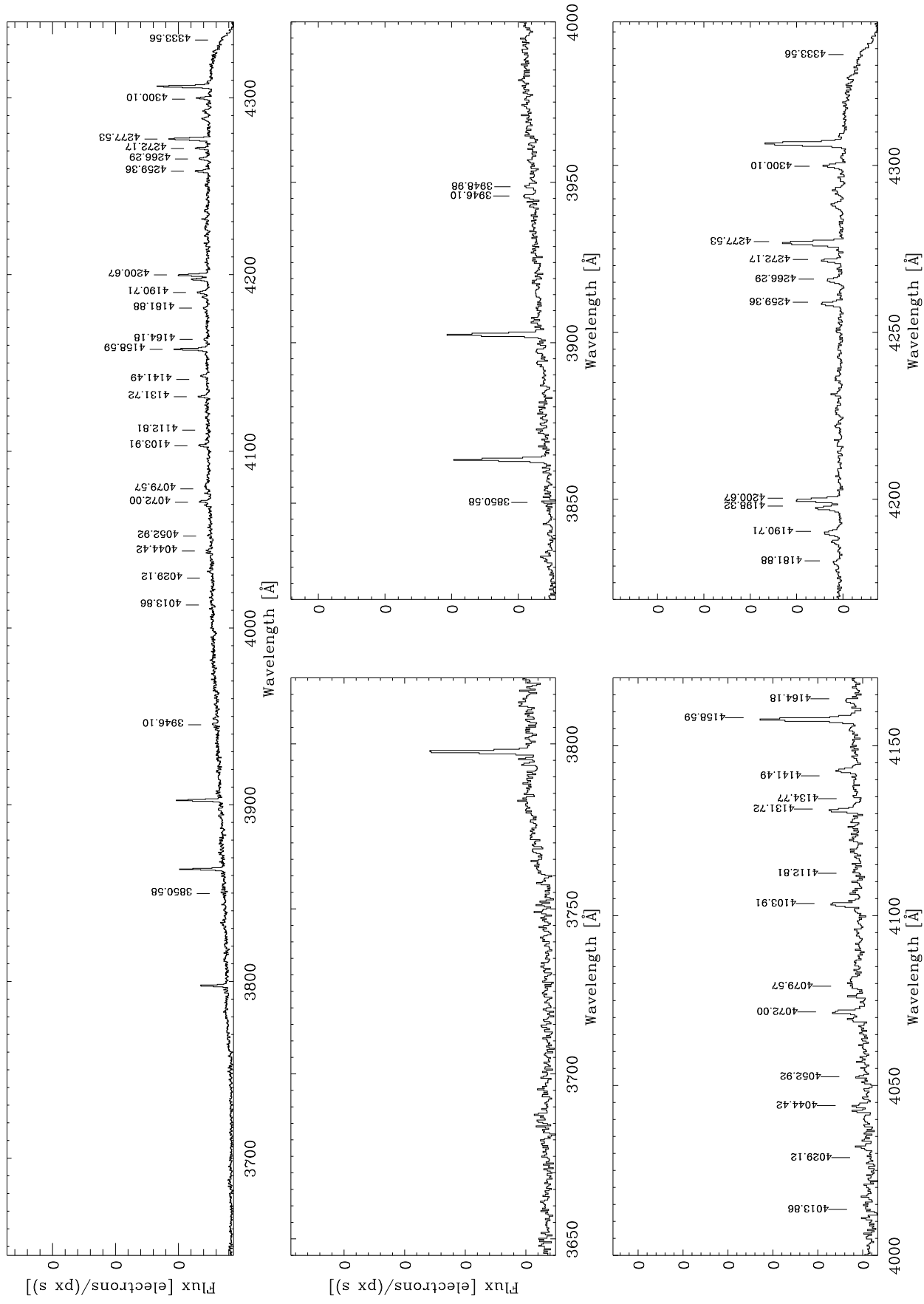
ThAr



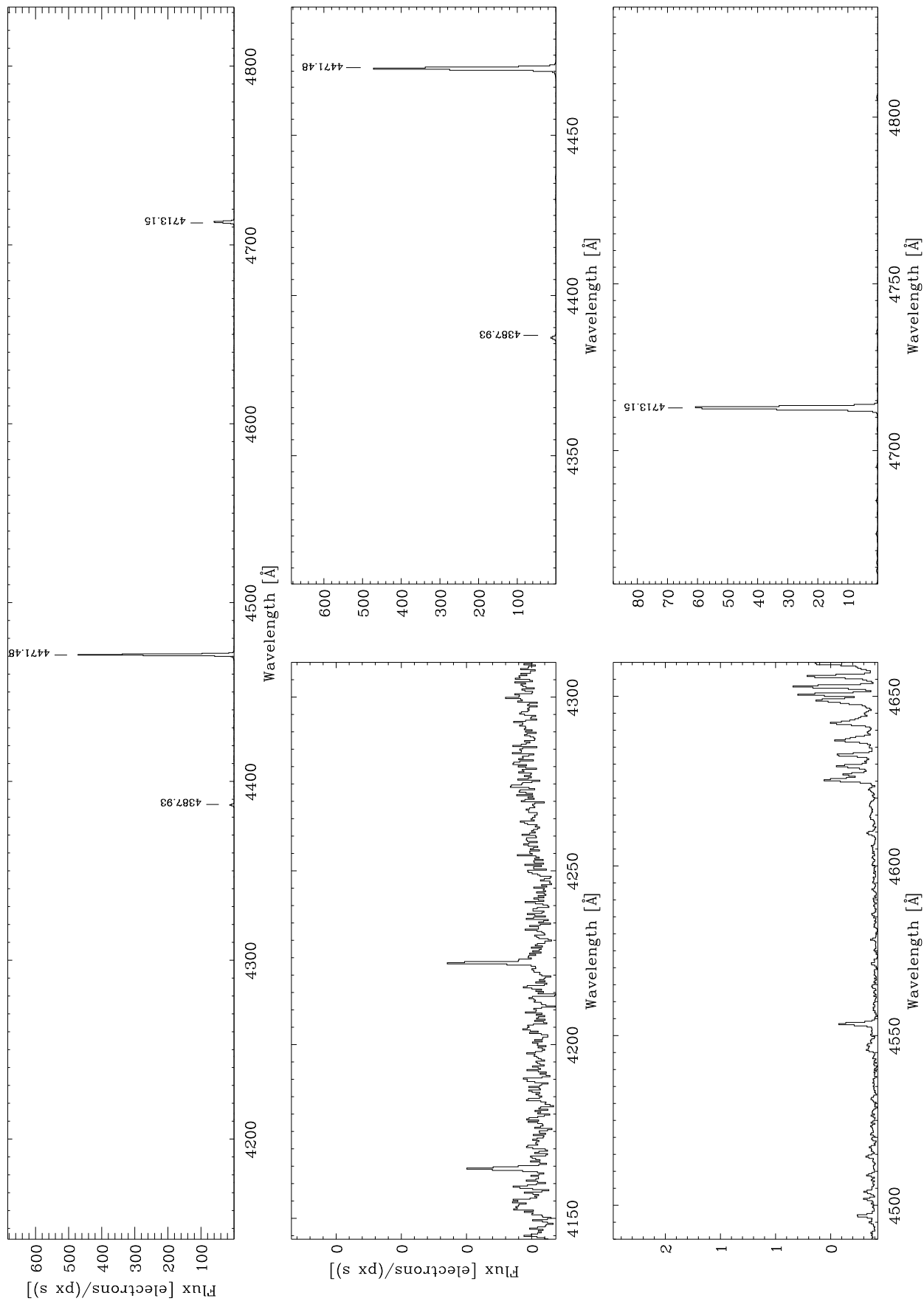
H2400B

$\lambda_{\text{cen}} = 4000\text{\AA}$

Cd



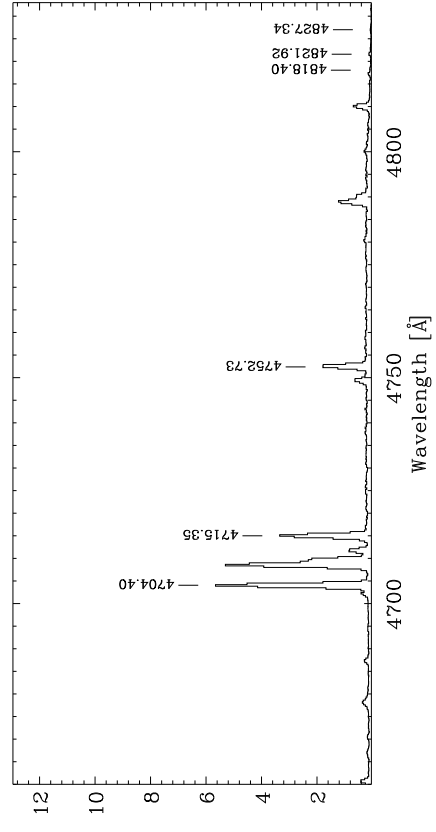
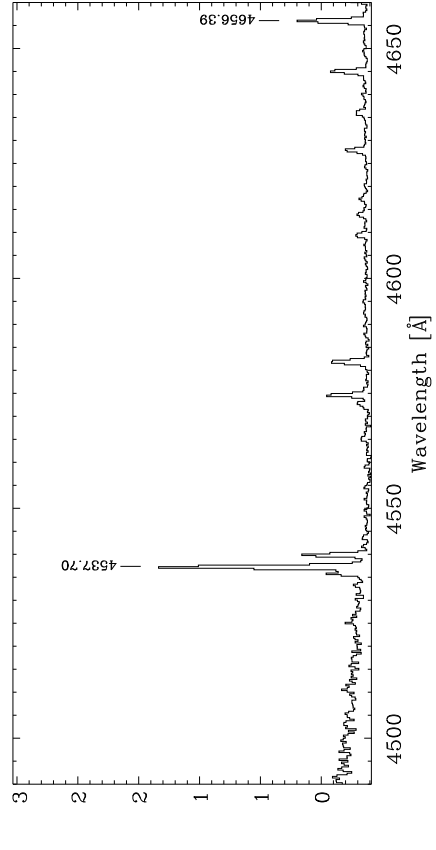
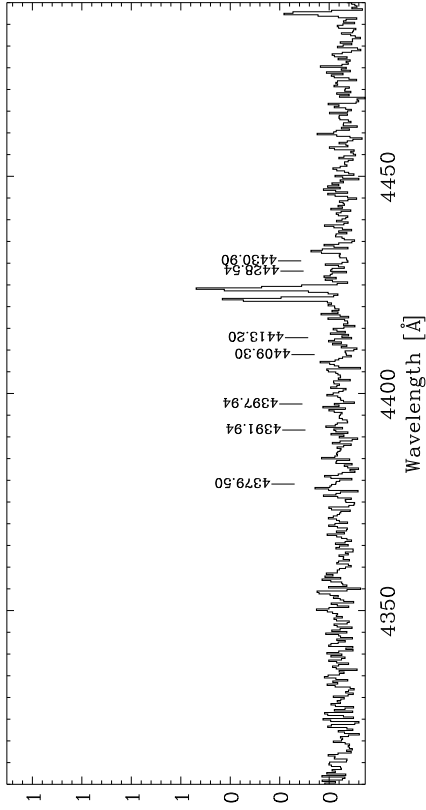
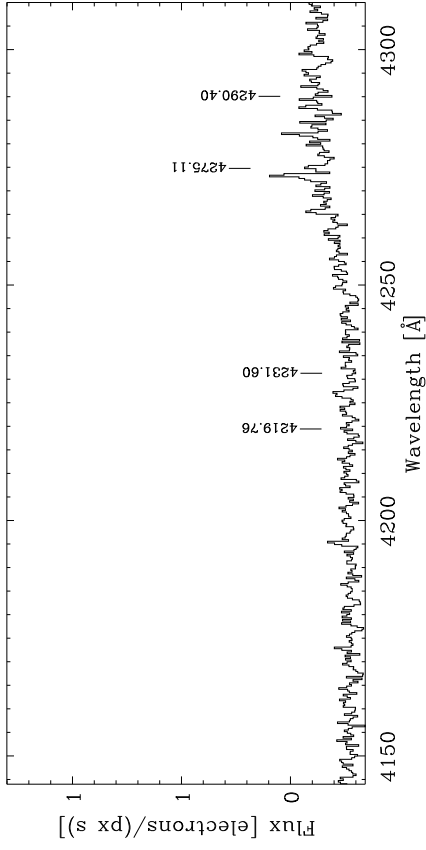
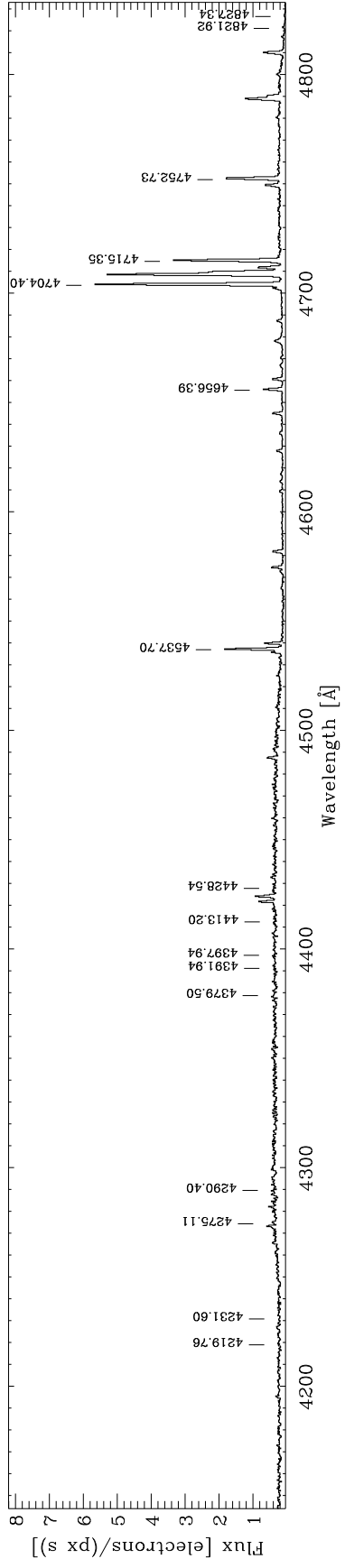
H2400B $\lambda_{\text{cen}} = 4500\text{\AA}$ He



H2400B

$\lambda_{\text{cen}} = 4500\text{\AA}$

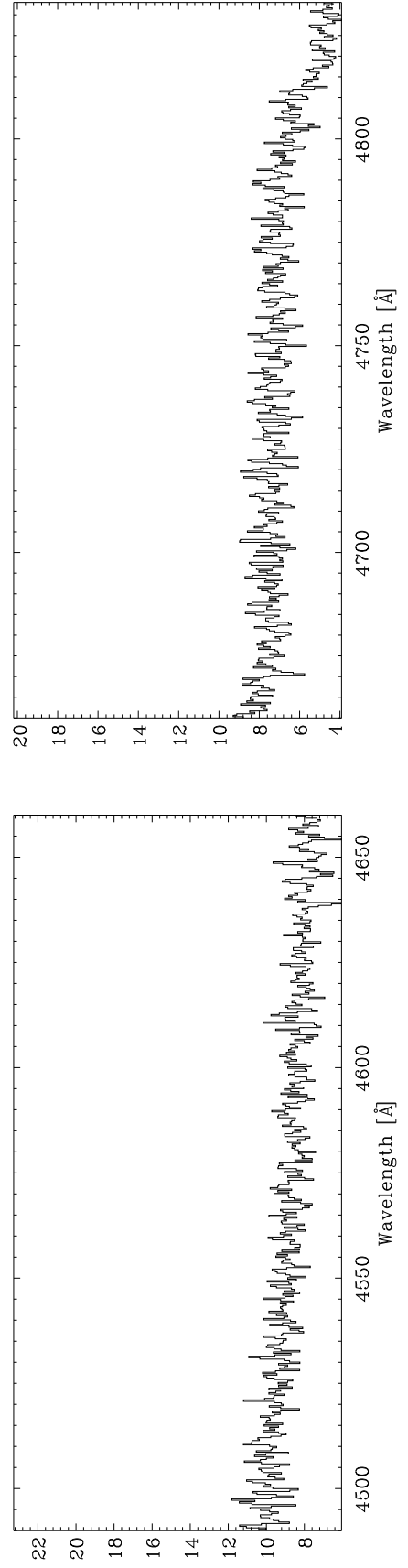
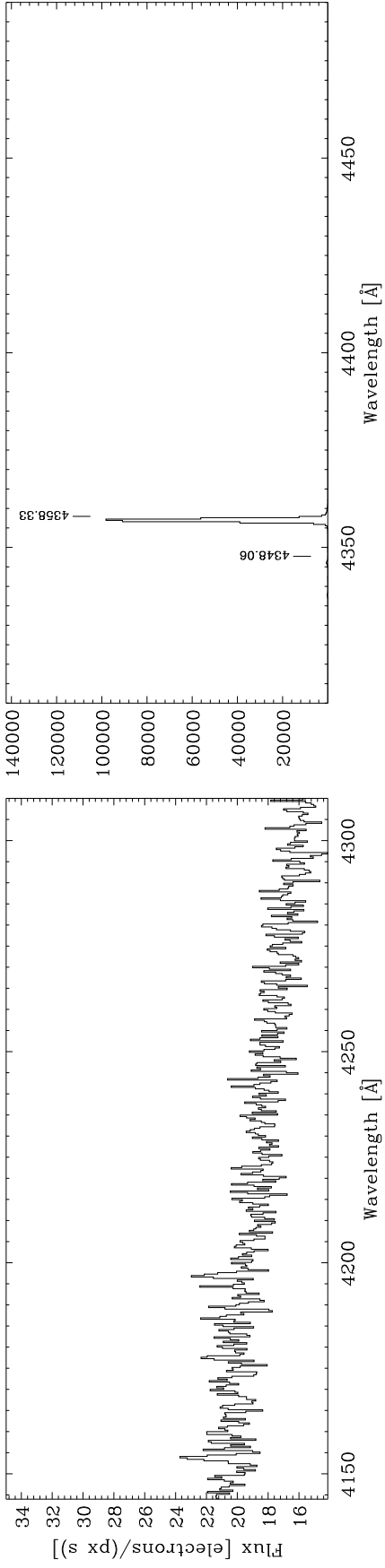
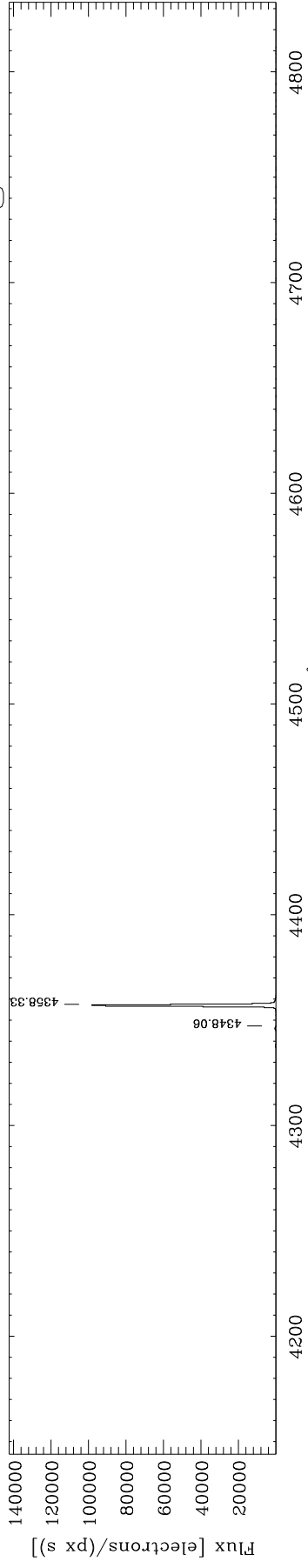
Ne



H2400B

$\lambda_{\text{cen}} = 4500\text{\AA}$

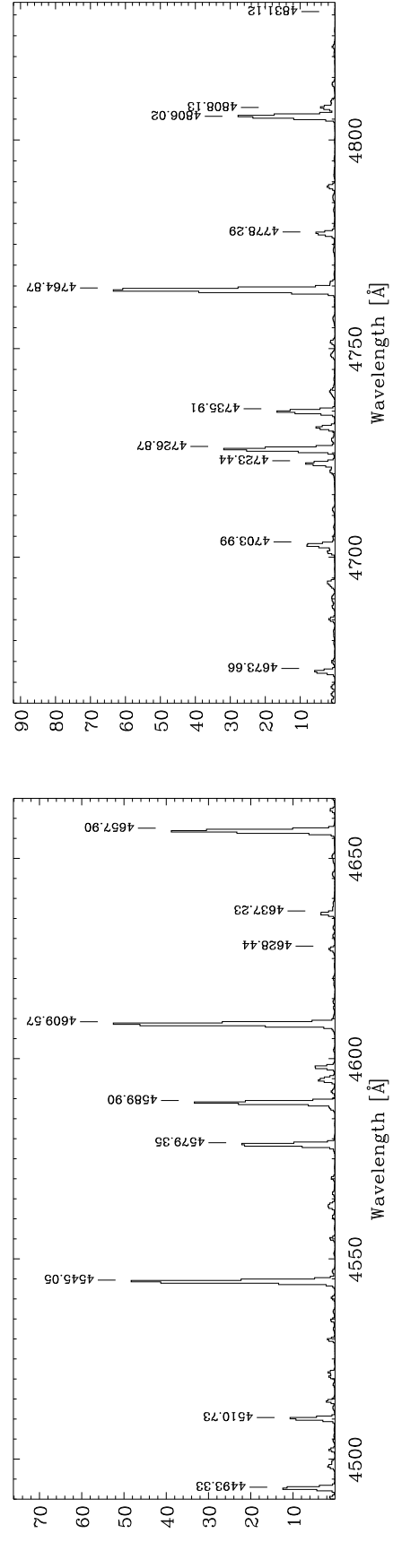
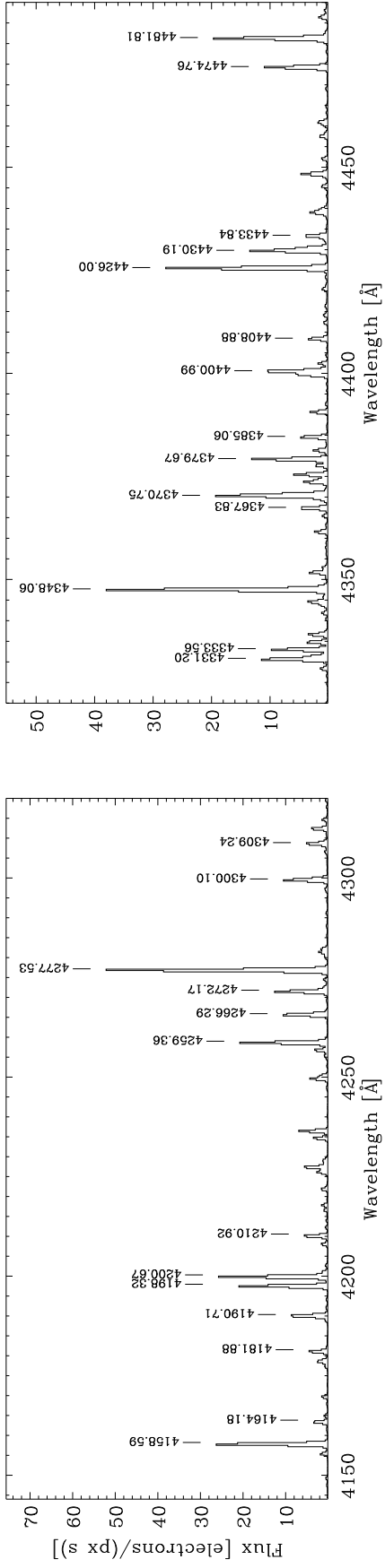
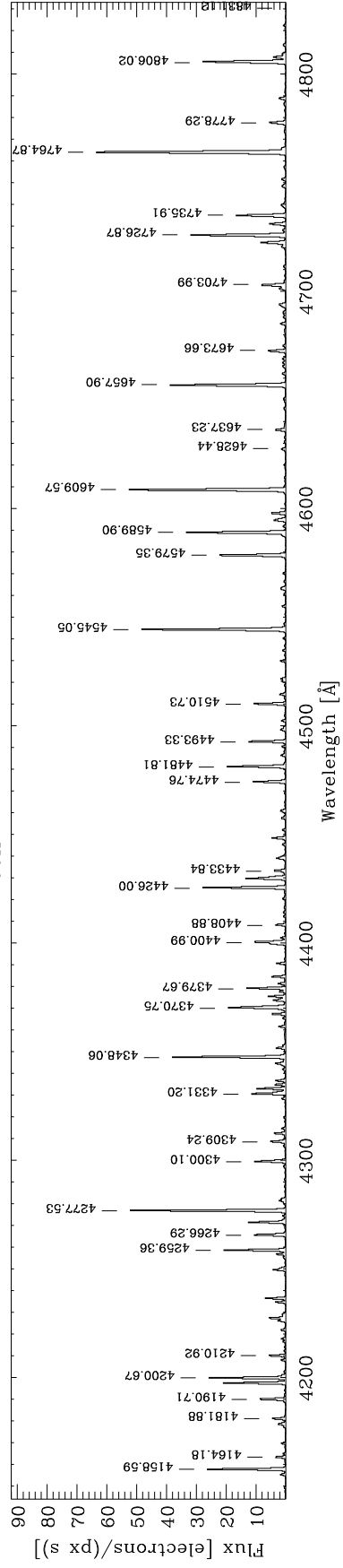
Hg



H2400B

$\lambda_{\text{cen}} = 4500\text{\AA}$

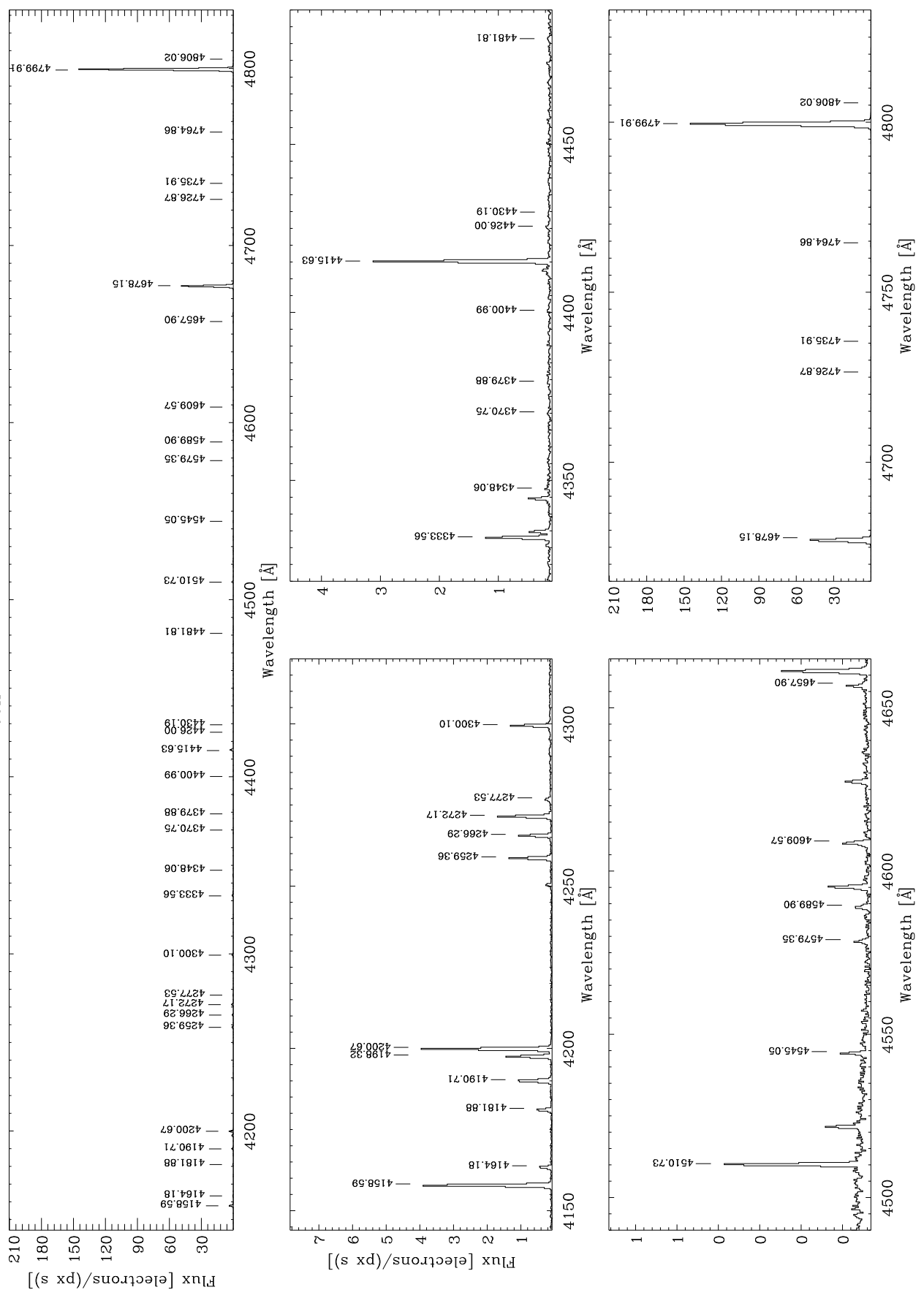
ThAr



H2400B

$\lambda_{\text{cen}} = 4500\text{\AA}$

Cd



Appendix 1: Laboratory spectral lines

For Hg and ThAr the lines were taken from the SALT Longslit Line Atlas⁷. The laboratory lines for He, Ne, and Cd were extracted from the NIST⁸ database. Lines with low relative Intensity were removed from the NIST data.

He

Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion
3819.607	HeI	4026.191	HeI	4713.145	HeI	5047.738	HeI	7065.188	HeI
3888.646	HeI	4387.928	HeI	4921.931	HeI	5875.621	HeI	7281.349	HeI
3964.729	HeI	4471.479	HeI	5015.68	HeI	6678.151	HeI		

Ne

Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion
3323.75	NeII	4537.704	NeI	5355.3	NeI	6532.882	NeI	8544.7	NeI
3334.87	NeII	4656.394	NeI	5360.012	NeI	6598.953	NeI	8571.353	NeI
3369.87	NeI	4704.395	NeI	5400.562	NeI	6652.09	NeI	8582.91	NeI
3378.28	NeII	4715.347	NeI	5412.649	NeI	6678.276	NeI	8591.259	NeI
3417.903	NeI	4752.732	NeI	5418.558	NeI	6717.043	NeI	8634.647	NeI
3447.703	NeI	4818.4	NeI	5433.651	NeI	6929.467	NeI	8647.04	NeI
3454.195	NeI	4821.924	NeI	5656.659	NeI	7024.05	NeI	8654.384	NeI
3466.579	NeI	4827.344	NeI	5662.549	NeI	7032.413	NeI	8681.922	NeI
3472.571	NeI	4837.312	NeI	5689.816	NeI	7059.107	NeI	8704.113	NeI
3501.216	NeI	4957.033	NeI	5719.225	NeI	7173.939	NeI	8771.659	NeI
3520.472	NeI	5005.159	NeI	5748.298	NeI	7245.167	NeI	8780.622	NeI
3568.53	NeII	5031.35	NeI	5764.419	NeI	7438.899	NeI	8783.754	NeI
3593.56	NeI	5037.751	NeI	5804.098	NeI	7472.438	NeI	8853.867	NeI
3600.169	NeI	5074.2	NeI	5820.156	NeI	7488.871	NeI	8865.306	NeI
3664.112	NeII	5080.385	NeI	5852.488	NeI	7535.774	NeI	8919.499	NeI
3694.197	NeII	5116.503	NeI	5881.895	NeI	7544.044	NeI	8988.58	NeI
3709.64	NeII	5122.3	NeI	5902.462	NeI	7839.06	NeI	9148.68	NeI
3713.084	NeII	5144.97	NeI	5944.834	NeI	7927.11	NeI	9201.76	NeI
3727.08	NeII	5158.902	NeI	5965.471	NeI	7937.01	NeI	9220.05	NeI
3766.29	NeII	5188.612	NeI	5975.534	NeI	7943.181	NeI	9226.67	NeI
3777.16	NeII	5193.224	NeI	6029.997	NeI	8082.458	NeI	9275.53	NeI
3829.77	NeII	5203.895	NeI	6074.338	NeI	8118.549	NeI	9300.85	NeI
4219.76	NeII	5208.865	NeI	6096.163	NeI	8128.908	NeI	9310.58	NeI
4231.6	NeII	5222.351	NeI	6128.451	NeI	8136.406	NeI	9313.98	NeI
4275.11	NeI	5234.028	NeI	6143.063	NeI	8259.379	NeI	9326.52	NeI
4290.4	NeII	5274.039	NeI	6163.594	NeI	8266.079	NeI	9373.28	NeI
4379.5	NeII	5280.085	NeI	6217.281	NeI	8300.326	NeI	9425.38	NeI
4391.94	NeII	5298.19	NeI	6266.495	NeI	8365.749	NeI	9459.21	NeI
4397.94	NeII	5304.758	NeI	6304.789	NeI	8377.606	NeI	9486.68	NeI
4409.3	NeII	5326.396	NeI	6334.428	NeI	8418.427	NeI	9534.167	NeI
4413.2	NeII	5330.778	NeI	6382.992	NeI	8463.37	NeI	9547.4	NeI
4428.54	NeII	5341.094	NeI	6402.246	NeI	8484.45	NeI	9665.424	NeI
4430.9	NeII	5349.204	NeI	6506.528	NeI	8495.359	NeI		

Hg

Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion
3125.67	HgI	4916.07	HgI	6598.684	ArI	7272.935	ArI	8103.693	ArII
3131.55	HgI	5017.162	ArII	6604.853	ArI	7300.306		8115.311	ArI
3341.48	HgI	5460.74	HgI	6677.278	ArI	7353.293	ArI	8264.523	ArI
3606.522	ArI	5769.6	HgI	6716.423	HgI	7372.118	ArI	8408.21	ArI
3649.832	ArI	5790.66	HgI	6752.83	ArI	7383.98	ArI	8424.647	ArI
3650.15	HgI	5882.623	ArI	6871.286	ArI	7503.868	ArI	8521.442	ArI
3654.84	HgI	5942.667	ArI	6907.521	HgI	7514.651	ArI	8667.944	ArI

⁷ <http://pysalt.salt.ac.za/lineatlas/lineatlas.html>

⁸ http://physics.nist.gov/PhysRefData/ASD/lines_form.html

3663.28	HgI	6032.124	ArI	6965.426	ArI	7635.106	ArI	8716.654	
3901.87	HgI	6072.72	HgI	7030.252	ArI	7723.76	ArI	9122.967	ArI
3906.37	HgI	6098.8	ArI	7067.217	ArI	7944.66	HgII	9224.499	ArI
4046.56	HgI	6145.439	ArI	7068.73	ArI	7948.176	ArI		
4077.83	HgI	6234.393	HgI	7081.894	HgI	8006.157	ArI		
4348.063	ArII	6384.714	ArI	7091.861	HgI	8014.786	ArI		
4358.33	HgI	6416.304	ArI	7147.041	ArI	8093.126			

Cd

Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion
3080.822	CdI	4259.362	ArI	5162.285	ArI	6483.083	ArII	7635.106	ArI
3133.167	CdI	4266.286	ArI	5187.746	ArI	6538.112	ArI	7723.761	ArI
3250.33	CdII	4272.169	ArI	5268.01	CdII	6567.65	CdII	7724.207	ArI
3252.524	CdI	4277.528	ArII	5271.6	CdII	6604.853	ArI	7891.075	ArI
3261.055	CdI	4300.101	ArI	5337.48	CdII	6638.22	ArII	7948.176	ArI
3403.652	CdI	4333.561	ArI	5378.13	CdII	6643.697	ArII	8006.157	ArI
3466.2	CdI	4348.064	ArII	5381.89	CdII	6666.359	ArII	8014.786	ArI
3467.655	CdI	4370.753	ArII	5451.652	ArI	6677.282	ArI	8053.308	ArI
3491.536	ArII	4379.667	ArII	5495.874	ArI	6684.292	ArII	8066.99	CdII
3509.778	ArII	4379.881	ArII	5558.702	ArI	6725.78	CdII	8103.693	ArI
3514.387	ArII	4400.986	ArII	5572.541	ArI	6752.834	ArI	8103.693	ArII
3545.595	ArII	4415.63	CdII	5606.733	ArI	6759.19	CdII	8115.311	ArI
3554.306	ArI	4426.001	ArII	5650.704	ArI	6766.612	ArI	8200.309	CdI
3559.508	ArII	4430.189	ArII	5739.52	ArI	6778.116	CdI	8264.522	ArI
3561.03	ArII	4481.81	ArII	5843.3	CdII	6861.269	ArII	8392.27	ArI
3576.615	ArII	4510.733	ArI	5860.31	ArI	6871.289	ArI	8408.21	ArI
3588.44	ArII	4545.052	ArII	5880.22	CdII	6888.174	ArI	8424.648	ArI
3606.522	ArI	4579.349	ArII	5882.624	ArI	6937.664	ArI	8521.442	ArI
3610.508	CdI	4589.898	ArII	5888.584	ArI	6965.431	ArI	8605.776	ArI
3612.873	CdI	4609.567	ArII	5912.085	ArI	7030.251	ArI	8667.944	ArI
3614.453	CdI	4657.901	ArII	5928.813	ArI	7067.218	ArI	8771.86	ArII
3850.581	ArII	4678.149	CdI	6032.127	ArI	7068.736	ArI	8849.91	ArI
3946.097	ArII	4726.868	ArII	6043.223	ArI	7107.478	ArI	9017.592	ArII
3948.979	ArI	4735.905	ArII	6052.723	ArI	7125.82	ArI	9075.394	ArI
4013.856	ArII	4764.864	ArII	6059.372	ArI	7147.042	ArI	9106.562	ArII
4029.12	CdII	4799.912	CdI	6099.142	CdI	7158.839	ArI	9122.967	ArI
4042.893	ArII	4806.02	ArII	6111.495	CdI	7206.98	ArI	9194.638	ArI
4044.418	ArI	4847.81	ArII	6114.923	ArII	7237.01	CdII	9224.499	ArI
4052.921	ArII	4879.863	ArII	6145.441	ArI	7265.172	ArI	9291.531	ArI
4072.004	ArII	4881.72	CdII	6172.278	ArII	7272.936	ArI	9292.0	CdI
4079.573	ArII	4889.042	ArII	6173.096	ArI	7284.38	CdII	9354.22	ArI
4103.912	ArII	4904.751	ArII	6212.503	ArI	7311.716	ArI	9601.925	ArII
4112.815	ArII	4933.209	ArII	6296.872	ArI	7316.005	ArI	9657.786	ArI
4131.723	ArII	4965.079	ArII	6307.657	ArI	7345.67	CdI	9758.637	ArII
4134.77	CdII	4972.16	ArII	6325.166	CdI	7353.293	ArI	9773.567	ArII
4141.49	CdII	5009.334	ArII	6330.013	CdI	7372.118	ArI	9783.079	ArII
4158.59	ArI	5017.163	ArII	6354.72	CdII	7383.98	ArI	9784.503	ArI
4164.18	ArI	5025.5	CdII	6359.98	CdII	7392.98	ArI	9803.69	ArII
4181.884	ArI	5062.037	ArII	6384.717	ArI	7412.337	ArI	9849.458	ArII
4190.713	ArI	5085.822	CdI	6416.307	ArI	7435.368	ArI	9854.061	ArII
4198.317	ArI	5141.783	ArII	6438.47	CdI	7503.869	ArI	9993.863	ArII
4200.674	ArI	5145.308	ArII	6464.94	CdII	7514.652	ArI		

ThAr

Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion	Line [Å]	Ion
3281.701	ArII	4433.838	ArII	5343.581	ThI	6212.503	ArI	7107.48	ArI
3382.132	ArII	4474.759	ArII	5378.836	ThI	6215.939	ArI	7125.82	ArI
3491.536	ArII	4481.81	ArII	5386.525	ArII	6224.528	ThI	7147.04	ArI
3545.843	ArII	4493.334	ThI	5390.466	ThII	6234.856	ThI	7168.9	ThI

3559.508	ArII	4510.73	ArI	5394.852		6243.12	ArII	7206.98	ArI
3561.03	ArII	4545.052	ArII	5407.345	ArII	6261.418	ThI	7218.054	ThI
3576.557	ThI	4579.349	ThI	5410.475	ArI	6307.657	ArI	7272.94	ArI
3588.44	ArII	4589.898	ArII	5417.486	ThI	6327.278	ThI	7311.716	ArI
3612.427	ThI	4609.567	ArII	5451.652	ArI	6342.86	ThI	7315.067	ThI
3659.529	ArI	4628.441	ArI	5495.874	ArI	6348.625		7353.29	ArI
3719.435	ThI	4637.232	ArII	5498.985		6355.911	ThI	7372.12	ArI
3729.309	ArII	4657.901	ArII	5509.994	ThI	6369.575	ArI	7383.98	ArI
3737.889	ArII	4673.661	ThI	5539.262	ThI	6376.931	ThI	7435.37	ArI
3765.27	ArII	4703.99	ThI	5548.176	ThI	6384.72	ArI	7471.16	ArI
3780.84	ArII	4723.438	ThI	5558.702	ArI	6411.9	ThI	7484.33	ArI
3803.172	ArII	4726.868	ArII	5587.026	ThI	6416.307	ArI	7503.869	ArI
3809.456	ArII	4735.906	ArII	5595.064	ThI	6457.28	ThI	7514.651	ArI
3828.385	ThI	4764.865	ArII	5606.733	ArI	6462.614	ThI	7567.74	ThI
3839.746	ThII	4778.294	ThI	5615.32	ThI	6466.553	ArI	7585.792	ThI
3850.58	ArII	4806.021	ArII	5639.746	ThII	6483.083	ArII	7635.11	ArI
3868.528	ArII	4808.134	ThI	5650.704	ArI	6490.738	ThI	7724.206	ArI
3925.718	ArII	4831.121	ThI	5700.918	ThII	6512.364	ThI	7788.937	ThI
3928.623	ArII	4840.843	ThI	5707.103	ThII	6531.34	ThI	7817.771	ThI
3932.546	ArII	4847.81	ArII	5720.183	ThI	6538.112	ArI	7847.54	ThI
3967.392	ThI	4879.864	ArII	5760.551	ThI	6554.16	ThI	7891.08	ArI
3994.792	ArII	4889.042	ArII	5789.645	ThI	6577.215	ThI	7948.176	ArI
4019.129	ThII	4894.955	ThI	5804.141	ThI	6583.906	ThI	7978.97	ThI
4030.842	ThI	4904.751	ArII	5834.264	ArI	6588.54	ThI	8006.157	ArI
4036.047	ThI	4919.816	ThII	5852.774		6593.94	ThI	8014.786	ArI
4043.39	ThI	4933.208	ArII	5860.31	ArI	6604.85	ArI	8053.31	ArI
4052.921	ArII	4939.642	ThI	5882.624	ArI	6638.22	ArII	8103.693	ArI
4072.005	ArII	4965.08	ArII	5888.584	ArI	6639.74	ArII	8115.311	ArI
4080.636	ArII	5002.097	ThI	5891.451	ThI	6643.7	ArII	8143.139	ThI
4103.912	ArII	5009.334	ArII	5912.09	ArI	6662.27	ThI	8186.914	ThI
4112.815	ArII	5017.163	ArII	5925.893	ThII	6666.359	ArII	8264.523	ArI
4131.724	ArII	5039.23	ThI	5928.812	ArI	6677.28	ArI	8330.45	ThI
4158.591	ArI	5044.719	ThI	5973.665	ThI	6684.29	ArII	8408.21	ArI
4164.18	ArI	5062.037	ArII	5989.044	ThII	6719.219	ArI	8424.648	ArI
4181.884	ArI	5067.974	ThI	5994.129	ThI	6727.459	ThI	8446.509	ThI
4190.713	ArI	5090.495	ArII	6005.725	ArI	6752.833	ArI	8478.36	ThI
4198.317	ArI	5100.442	ArII	6007.072	ThI	6756.453	ThI	8521.44	ArI
4200.675	ArI	5115.044	ThI	6021.036	ThI	6766.611	ArI	8573.122	ThI
4210.923	ThI	5125.95	ThI	6025.15	ArI	6778.313	ThI	8605.78	ArI
4259.36	ArI	5141.783	ArII	6032.127	ArI	6780.413	ThI	8620.46	ArI
4266.286	ArI	5145.308	ArII	6037.698	ThI	6829.036	ThI	8667.94	ArI
4272.169	ArI	5151.612	ThI	6043.223	ArI	6834.925	ThI	8709.236	ThI
4277.53	ArII	5154.243	ThI	6049.075	ArII	6861.269	ArII	8748.033	ThI
4300.1	ArI	5158.604	ThI	6098.803	ArI	6871.29	ArI	8761.686	ArI
4309.239	ArII	5176.961	ThI	6105.635	ArI	6874.754	ThI	8967.641	ThI
4331.199	ArII	5187.746	ArI	6114.92	ArII	6879.583	ArI	9048.252	ThI
4333.561	ArI	5199.164	ThI	6121.879	ArI	6888.174	ArI	9122.967	ArI
4348.064	ArII	5211.23	ThI	6145.441	ArI	6911.23	ThI	9194.638	ArI
4367.831	ArII	5219.11	ThI	6151.993	ThI	6937.66	ArI	9218.986	ArII
4370.75	ArII	5231.16	ThI	6155.239	ArI	6943.61	ThI	9224.498	ArI
4379.666	ArII	5247.654	ThII	6169.822	ThI	6965.43	ArI	9291.532	ArI
4385.057	ArII	5258.36	ThI	6172.28	ArII	6989.66	ThI	9354.218	ArI
4400.99	ArII	5266.71	ThI	6182.622	ThI	7000.806	ThI	9657.786	ArI
4408.882	ThI	5297.743	ThI	6191.906	ThI	7018.57	ThI		
4426.001	ArII	5311.982	ArII	6198.223	ThI	7030.25	ArI		
4430.189	ArII	5326.976	ThI	6203.493	ThI	7067.218	ArI		

Appendix 2: Atlas plotting code

This appendix gives the IDL code for plotting the arc spectra with the labels and ticks on some identified lines. The inputs are two files, one containing the wavelengths and the fluxes, and another one with the wavelengths of the lines to be marked. The output is a formatted pdf file showing five plots: one covers the entire spectrum and the other four zoom in four splits, in order to facilitate the line identification.

```

1 ;*****
2 ; idl_plot.pro procedure
3 ; July 2015
4 ; Edited for WYFOS: Ronny Errmann
5 ; original Authors: Hassan Fathivavari, Javier Mendez and Liam Hardy:
6 ; http://www.ing.iac.es/astromy/observing/manuals/ps/tech_notes/tn133.pdf
7 ;*****
8 ; It requires files <grating>-<cenwave>-<arc>-<extension>.txt
9 ; and ticksall-<grating>-<cenwave>-<arc>-<extension>.txt in the
10 ; same directory
11 ;*****
12 ; <grating>-<cenwave>-<arc>-<extension>.txt is the 1-D ascii file
13 ; containing two columns: wavelength and flux.
14 ;*****
15 ; ticksall_* contains all the lines to mark. Example:
16 ; 5100 He II
17 ; 5105.54 Hg+Ar
18 ; 5330.78 (text is optional and not used)
19 ;*****
20 ; Execute as follows:
21 ; IDL> idl_plot, grating, cenwave, arc, arctext, exptime, extension, 5, wavelengths, line
    offset
22 ; the wavelength are: begin, 3x splitting, end
23 ; IDL> idl_plot, 'R900V', '6500', 'hgar', 'Hg+Ar', 10, 'p1234578
    ', 3500, 5800, 6600, 8500, 10500, 0
24 ;*****
25 pro idl_plot, gra, cenw, arc, arctext, exptime, extension, x1, x2, x3, x4, x5, offset
26
27 only_first_plot=0 ;to create ps files with only the first plot use
    only_first_plot=1
28 set_plot, 'ps'
29 device, filename=gra + '_' + cenw + '_' + arc + '.ps', /color, bits_per_pixel=8, /
    landscape
30 ;device, XSIZE=29.7-5.6, YSIZE=21.0-4.0, /cm, filename=gra + '_' + cenw + '_' + arc +
    '.ps', /color, bits_per_pixel=8, /landscape ;for a4
31 loadct, 0
32
33 ticksfile='ticksall_' + gra + '_' + cenw + '_' + arc + '_' + extension + '.txt'
34 filename=gra + '_' + cenw + '_' + arc + '_' + extension + '.txt'
35 xran=[fix(x1), fix(x2), fix(x3), fix(x4), fix(x5)]
36
37 readcol, ticksfile, ticksdata
38 n=file_lines(filename)
39 data=fltarr(2,n)
40 get_lun, lun
41 openr, lun, filename
42 readf, lun, data
43 free_lun, lun
44 wave=data(0,*)
45 flux=data(1,*)/exptime
46
47 nf = n_elements(flux)
48 nt = n_elements(ticksdata)
49
50 ytitl='Flux-[electrons/(px-s)]'
51 ;***** Subwindow Settings *****
52 for xk = 0,4*(1-only_first_plot) do begin ;4 normally, otherwise only first plot
53 if xk eq 0 then begin
54 !Y.MARGIN=[4,6] ;distance to paper edges
55 !p.multi=[0,1,3]
56 xb=xran(0) ;global min
57 xe=xran(4) ;global max
58 endif else begin
59 !Y.MARGIN=[4,2]
60
61 endif
62 if xk eq 1 then begin
63 !p.multi=[4,2,3]
64 xb=xran(0)
65 xe=xran(1) ;first subplot
66
67 endif
68 if xk eq 2 then begin
69 !p.multi=[3,2,3]
70 xb=xran(1)

```

```

69     xe=xran(2);second subplot
70     ytitl='',
71 endif
72 if xk eq 3 then begin
73     !p.multi=[2,2,3]
74     xb=xran(2)
75     xe=xran(3);third subplot
76 endif
77 if xk eq 4 then begin
78     !p.multi=[1,2,3]
79     xb=xran(3)
80     xe=xran(4) ;fourth subplot
81     ytitl='',
82 endif
83
84 charsize=0.5
85 xycharsize=1.5
86
87 if only_first-plot eq 1 then begin
88     !p.multi=0
89     !Y.MARGIN=[4,3]
90     DEVICE, YSIZE=7.7, /cm
91     ytitl='Flux_[electrons/px]'
92     charsize=0.6
93     xycharsize=1.0
94 endif
95
96 ;***** Ticks Mark *****
97 dx = xe - xb ;wavelength range
98 ;xx = 0.003666141 * dx ;minimum distance of two lines in plot
99 if xk eq 0 then begin
100     xx = 0.006 * dx ;plotting offset
101     if strmatch(arc, 'thar') EQ 1 then xxn = 0.008 * dx else xxn = 0.006 * dx ;for ThAr
102         wider sep till next line
103 end else begin
104     xx = 0.01 * dx ;plotting offset
105     xxn = 0.01 * dx ;minimum distance of the text of two lines
106 end
107
108 t = fltarr(nt,3) ;empty nt entry array
109 for k=0,nt-1 do begin
110     if (ticksdata(k) lt xb) or (ticksdata(k) gt xe) then continue ;only in range
111     msb = value_locate(wave, ticksdata(k)+offset)
112     if msb eq -1 then continue ;if outside of the flux range
113     t(k,0)=total(flux(max([msb-2,0]):min([msb+2,nf-1]))) ;flux around the line
114     t(k,1)=ticksdata(k)
115     t(k,2)=max(flux(max([msb-5,0]):min([msb+5,nf-1])))
116 endfor
117 tsort=reverse(sort(t(*,0))) ;tsort contains index starting with the highest flux
118 ticks = fltarr(nt)
119 m = 1
120 ticks(0)=t(tsort(0),1) ;wavelength with highest flux
121 maxi=t(tsort(0),2) ;highest flux
122 for k=1,nt-1 do begin ;adding the other wavelengthes, starting with second
123     highest flux and only, if no line close to it is already added
124     if t(tsort(k),1) lt 1 then break ;stop, if no more catalog data is available
125     add=1
126     for l=0,m-1 do begin ;search the already added lines
127         if abs(t(tsort(k),1)-ticks(l)) LE xxn then begin ;found a close line
128             add=0
129             break ;-> don't use
130         endif
131     endfor
132     if add eq 1 then begin ;add line
133         ticks(m) = t(tsort(k),1) ;add the line to the ticks
134         maxi=max([maxi,t(tsort(k),2)])
135         m = m + 1
136     endif
137 endfor
138 ticks = ticks(sort(ticks(0:m-1))) ;sort by wavelength and ignore the zeros
139 print, ticks
140
141 ;***** Borders *****
142 msb0=value_locate(wave, xb+1)
143 msb1=value_locate(wave, xe-1)
144 print, msb0
145 print, msb1
146 maxi=max([maxi,max(flux(msb0:msb1))/1.4]) ;if highest line has no tick it
147     creates unused space
148 mini=min(flux(msb0:msb1))
149 std=stddev(flux(msb0:msb1))
150 ;print, std, mini, maxi, maxi-mini

```

```

148     if (maxi-mini lt 15*std) and (maxi-mini lt 10) then maxi=mini+min([10,15*std])
149
150 ;***** Ticks *****
151 xtickint=1000
152 if xe - xb lt 2000 then xtickint=500
153 if xe - xb lt 1000 then xtickint=200
154 if xe - xb lt 500 then xtickint=100
155 if xe - xb lt 200 then xtickint=50
156 if xe - xb lt 100 then xtickint=20
157 if xk eq 0 then begin
158     xtickint=1000
159     if xe - xb lt 4000 then xtickint=500
160     if xe - xb lt 2000 then xtickint=200
161     if xe - xb lt 1000 then xtickint=100
162     if xe - xb lt 400 then xtickint=50
163 endif
164
165 ytickint=(maxi-mini)*1.45/8 ;unrounded
166 rounded=10^round(alog10(ytickint)-0.3)
167 ytickint = round(ytickint/(rounded+0.0))*rounded ;rounded
168 ;ytickint = 30000000
169
170 ;***** Plot *****
171 plot, wave, flux, psym=10, xrange=[xb, xe],/xstyle, yrange=[mini, maxi*1.45],/ystyle
, thick=0.65*(1+only_first_plot), ytickinterval=ytickint, ytickformat='(I8)',
ytickklen=0.01, yminor=5, ycharsize=ycharsize, ytitle=ytitl, xtickinterval=
xtickint, xtickklen=0.02, xminor=10, xcharsize=xcharsize, xtitle='Wavelength_
[!6!sA!r!u!9_%!6!n]'
172
173 if ticks(0) gt 1 then begin ;only if catalog data available
174     for i=0,m-1 do begin
175         msb = value_locate(wave, ticks(i)+offset) -0
176         y = max(flux(max([msb-5,0]):min([msb+5,nf-1]))) ;maximum flux around line to
plot tick marks right
177         plots, [ticks(i)-0.2*xx+offset, ticks(i)-0.2*xx+offset], [y+0.07*maxi, y+0.15*
maxi], thick=0.8 ;tick marks
178         xyouts, ticks(i)+0.5*xx+offset, y+0.08*maxi, ticks(i), charsize=charsize,
alignment=0.2, orientation=90, charthick=0.8, /data ;tick text
179     endfor
180 endif
181
182 endfor ;subplots
183
184 if only_first_plot eq 1 then begin
185 xyouts,[2000, 10000, 20000], [7000, 7000, 7000], [gra, '!7k!X!Icen!N='+cenw+'!6!sA!r!u
!9_%!6!n', arctext], CHARSIZE = 2, /DEVICE ;Title text
186 endif else begin
187 xyouts,[2000, 10000, 20000], [17000, 17000, 17000], [gra, '!7k!X!Icen!N='+cenw+'!6!sA!r
!u!9_%!6!n', arctext], CHARSIZE = 2, /DEVICE ;Title text
188 endelse
189
190 device, /close
191
192 if only_first_plot ne 1 then begin
193 spawn, 'rm_f_out.pdf'
194 spawn, 'ps2pdf13_' + gra + '_' + cenw + '_' + arc + '.ps_out.pdf'
195 ;spawn, 'ps2pdf13 -sPAPERSIZE=a4 ' + gra + '_' + cenw + '_' + arc + '.ps out.pdf'
;leaves a too big border on the left
196 spawn, 'rm_f_' + gra + '_' + cenw + '_' + arc + '_' + extension + '.pdf'
197 spawn, 'pdftk_out.pdf_cat_1-endN_output_' + gra + '_' + cenw + '_' + arc + '_' +
extension + '.pdf' ;rotates the pdf by 0deg
198 spawn, 'mv_' + gra + '_' + cenw + '_' + arc + '.ps_' + gra + '_' + cenw + '_' + arc
+ '_' + extension + '.ps'
199 print, 'Output_saved_to_' + gra + '_' + cenw + '_' + arc + '_' + extension + '.pdf'
200 endif
201
202 end

```

Appendix 3: Preparation for the atlas plotting

This code combines the flux from fibers by running a median on the “iraf_*.fit” files, produced by the WYFFOS pipeline. As each fiber has a different throughput the fibers were scaled and a 2σ clipping was applied before combining them. To determine the scaling factor, the scaling ratio against the median flux for each pixel was determined. These factors were combined by using the weighted average with the median flux as weights.

Additionally files necessary to create this document are also prepared: (1) The catalog with

the line labels for each individual setup is created. (2) Header information are collected and put together in a table (`table.tex`), which can be imported into a `tex` file. (3) Furthermore the lists for all lines plotted in the graphs are written in a text file (`arc_catalogs.txt`). (4) A script to be run in `idl` is created and executed (`run_in_idl.pro`).

```

1  import os
2  import sys
3  import time
4  import numpy
5  import pyfits
6  from operator import itemgetter
7  #copy data into plot folder
8  #find /data/ronny/data-af2-proj/ -name "iraf_*.fit" -exec cp -puv {} . \;
9  #find /obsdata/outgoing/ldp/LAMPS/ -name "iraf_*.fit" -exec cp -puv {} . \;
10 #run: python ../combine_fibers.py iraf_*
11
12 fibers=[1,161] #[77,85]          #which fiber range should be used to determine the
    median flux for the setup
13 comborders=True                #False does not create the single order files for idl_plot (
    again). Only use it to save computing time
14 sort_out3=False                #True doesn't plot graphs with less than 3 lines
15 sort_out_not_usable=True       #True doesn't plot graphs from the not_usable list
16 use_ignore_files=True          #True means the list "ignore_files" is applied and the files
    in the list are not used
17 borders=True                  #True means that the min and max wavelength from dispersions
    is applied to have all lamps of a configuration equal
18 min_flux=1                     #if median flux in a fiber is lower, the fiber is not used
19 lampfolder='/home/ronny/Programs/WYFFOS_pipeline.v3.0.3/control_standards/' #Where
    are the catalog files located?
20 #as ND filter is missing in header create a file 'ndfilter.lst' with the columns
    runnumber and header keyword, e.g. "2282280 WYC-clear1"
21
22 keywords=[]
23 keywords.append('EXPTIME')      #IDS,WYFFOS
24 keywords.append('WYFCRAT')     #WYFFOS
25 keywords.append('CENWAVE')     #IDS,WYFFOS
26 keywords.append('WYFORDER')    #WYFFOS
27 keywords.append('WYFCLAMP')    #WYFFOS
28 keywords.append('LAMP')        #WYFFOS
29
30 dispersions=[] #grating, cenwave, dispersion, plot borders (dispersion for 2x2
    binning from webpage for Red+4, or interpolated to write into tex table)
31 dispersions.append(['R158B', '3600', '6.4', '', '']) # '3.2')#
    binning 1x1
32 dispersions.append(['R158B', '4500', '6.4']) # '3.2')
33 dispersions.append(['R158B', '4900', '6.4']) # '3.2')
34 dispersions.append(['R158R', '6500', '6.4']) # '3.2')
35 dispersions.append(['R158R', '7500', '6.4']) # '3.2')
36 dispersions.append(['R300B', '4000', '3.6']) # '1.8')
37 dispersions.append(['R300B', '4500', '3.6', '3000', '8305']) # '1.8')
38 dispersions.append(['R300B', '5300', '3.6']) # '1.8')
39 dispersions.append(['R316R', '6500', '3.4']) # '1.7')
40 dispersions.append(['R316R', '7500', '3.4', '4418', '10000']) # '1.7')
41 dispersions.append(['R600B', '3900', '1.8']) # '0.89')
42 dispersions.append(['R600B', '4000', '1.8']) # '0.89')
43 dispersions.append(['R600B', '5000', '1.8', '3351', '6682']) # '0.89')
44 dispersions.append(['R600R', '6500', '1.8', '4834', '8177']) # '0.89')
45 dispersions.append(['R600R', '7000', '1.8']) # '0.89')
46 dispersions.append(['R600R', '8000', '1.8', '6371', '9668']) # '0.89')
47 dispersions.append(['R1200B', '4000', '0.86', '3177', '4815']) # '0.43')
48 dispersions.append(['R1200B', '4500', '0.84', '3678', '5298']) # '0.42')
49 dispersions.append(['R1200B', '5000', '0.81', '4193', '5804']) # '0.41')
50 dispersions.append(['R1200B', '5500', '0.79']) # '0.39')
51 dispersions.append(['R1200R', '5500', '0.79', '4665', '6292']) # '0.39')
52 dispersions.append(['R1200R', '6000', '0.76', '5169', '6746']) # '0.38')
53 dispersions.append(['R1200R', '6500', '0.74', '5681', '7266']) # '0.37')
54 dispersions.append(['R1200R', '7000', '0.71', '6195', '7748']) # '0.36')
55 dispersions.append(['R1200R', '7200', '0.70']) # '0.35')
56 dispersions.append(['R1200R', '7500', '0.69', '6711', '8229']) # '0.34')
57 dispersions.append(['R1200R', '8000', '0.66', '7208', '8708']) # '0.33')
58 dispersions.append(['R1200R', '8200', '0.65']) # '0.33')
59 dispersions.append(['R1200R', '8500', '0.64', '7724', '9184']) # '0.32')
60 dispersions.append(['R1200R', '9000', '0.61']) # '0.31')
61 dispersions.append(['R1200R', '9350', '0.59', '8601', '10000']) # '0.30')
62 dispersions.append(['H1800V', '5120', '0.44']) # '0.22')
63 dispersions.append(['H1800V', '5500', '0.44']) # '0.22')
64 dispersions.append(['H2400B', '4000', '0.38', '3645', '4343']) # '0.19')
65 dispersions.append(['H2400B', '4500', '0.38', '4144', '4833']) # '0.19')
66 dispersions.append(['', '', '']) #if no matching dispersion fill dispersion in tex
    table with empty entry
67
68 replace_cenwave=[]            #the cenwave in the header is not exact, with this list values

```

```

        are replaced
69 replace_cenwave.append(['3499','3500'])
70 replace_cenwave.append(['3501','3500'])
71 replace_cenwave.append(['9004','9000'])
72 replace_cenwave.append(['9005','9000'])
73 replace_cenwave.append(['9301','9300'])
74 replace_cenwave.append(['9351','9350'])
75 replace_cenwave.append(['9352','9350'])
76
77 ignore_files=[]
78 ignore_files.append('p2279722') #'H2400B','4500','cd','Cd','120'          saturated
79 ignore_files.append('p2279719')
80 ignore_files.append('p2271410') #'R600B','4000','he','He','60'
81 ignore_files.append('p1710557') #'R158B','4900','hgar','Hg','2' replace by new
    dedector
82
83 overrideSplittings=[] #grat,cenwave,arc,wave1,wave2,wave3,wave4,wave5      #'
    for wave uses auto
84 overrideSplittings.append(['H2400B','4000','he','','3830','','',''])
85 overrideSplittings.append(['H2400B','4000','hgar','','3750','3930','4060','',''])
86 overrideSplittings.append(['R1200R','9350','ne','','','9290','',''])
87 overrideSplittings.append(['R1200R','9350','hgar','','','9680','',''])
88 overrideSplittings.append(['R158B','4500','he','3000','4200','5400','7200','10000'
    ])
89 overrideSplittings.append(['R158B','4500','ne','3000','5500','6800','7800','10000'
    ])
90 overrideSplittings.append(['R158B','4500','hgar','3000','5100','5850','7850','10000'
    ])
91 overrideSplittings.append(['R600R','8000','hgar','','7200','8050','',''])
92 overrideSplittings.append(['R600R','8000','thar','','7245','8030','',''])
93 overrideSplittings.append(['R600R','8000','cd','','7190','8020','8820','',''])
94
95 manual_shift=[] #move the lines by how many angstrom as small offset occurs
96 manual_shift.append(['ECHELLE-3','8685','hgar','0.7'])
97 manual_shift.append(['ECHELLE-5','5675','ne','-0.4'])
98 manual_shift.append(['ECHELLE-5','5400','ne','-0.5'])
99
100 not_usable=[] #settings which should not be used for the final plot
101 not_usable.append(['H2400V','4500','ne']) #no flux
102 not_usable.append(['H2400V','4000','he']) #bad solution
103
104 fibers=range(fibers[0],fibers[1]+1)
105 names=[]
106 for j in range(1,len(sys.argv)):
107     name=sys.argv[j].replace('.fits','')
108     name=name.replace('.fit','')
109     if name==sys.argv[j]: #no change -> no fits
110         continue
111     names.append([name,sys.argv[j]])
112 if comborders==True: #will make the spectra to text, read them and combine the
    fibers into one text-file
113     for name in names:
114         data=[]
115         header_table = pyfits.getheader(name[1])
116         data_cube = pyfits.getdata(name[1], 0) #for good files: 0==2 contains
            flux, 1==3 contains errors?, 4?, 5?, otherwise data_cube.shape
            ===(2,*) for 0
117         if len(data_cube)==2:
118             data.append(data_cube[0])
119             print "extraction_was_done_for_only_one_fiber",name[0]
120         else:
121             print name[0],":\t",
122             for j in [fibers,range(1,162)]: #all fibers in case
                the selected ones contain no data
123                 for i in j:
124                     keyword='F%3.3i%i+TYPE'
125                     if keyword in header_table.keys():
126                         if header_table[keyword][0]<>'X':
127                             #X means 0 flux, as fiber
                                has a problem
128                                 if numpy.median(data_cube[i
                                    -1])>min_flux: #only
                                        if enough flux
129                                     data.append(data_cube[
                                        i-1])
130
131             if data<<[]:
132                 break
133             print "no_data_in_the_selected_fibers:",j,"_try_with_
                all_fibers","\t\t",
134             print len(data),"fibers_used"
wavelength=[header_table['CRVAL1']] #prepare wavelength list
step=header_table['CDELTA1']

```

```

135     for i in range(1, len(data[0])):           #create wavelength list
136         wavelength.append(wavelength[i-1]+step)
137     data=numpy.array(data)
138
139     if len(data)>3:                             #scale the flux between fibers by determining
140         the median flux and determining the factor
141         data_sub=[]
142         for i in range(len(wavelength)):         #for every pixel
143             if numpy.median(data[:,i])>min_flux: #if enough
144                 flux
145                 data_sub.append(data[:,i])
146         data_sub=numpy.array(data_sub)
147         data_sub=numpy.transpose(data_sub)      #
148         subarray of data which contains flux
149         med_flux=[]                             #
150         median flux in each pixel
151         for i in range(len(data_sub[1,:])):     #for every pixel
152             med_flux.append(numpy.median(data_sub[:,i])) #
153             determine median flux
154         med_flux=numpy.array(med_flux)         #median flux
155         factor=[]                               #scaling factors for
156         each fiber
157         for j in range(len(data_sub)):         #for every fiber
158             factors=numpy.array(data_sub[j,:]/med_flux) #
159             factors for each pixel towards median_flux
160             factors=numpy.where(factors>0.33, factors, 10) #
161             replace factors below 1/3 by 10
162             temp=numpy.where(factors<3)        #get
163             indexes where factor between 1/3 and 3 to use only
164             these ones.
165             factor.append(numpy.average(factors[temp[0]], weights=
166             med_flux[temp[0]**2)) #get weighted average
167             of factors for each pixel by weighting them with
168             the median flux of that pixel
169         factor/=numpy.median(factor)           #normalize to one by
170         dividing through median
171         std_factor=2*numpy.std(factor, ddof=1) #2 times standard
172         deviation
173         i=0
174         for j in range(len(data))[:, -1]:     #for every
175             fiber backwards
176             if abs(factor[j]-1)<std_factor:     #median(factor
177                 )==1
178                 data[j,:]/=factor[j]          #if in 2*std,
179                 then scale data of this fiber
180             else:
181                 data=numpy.delete(data,(j), axis=0) #
182                 otherwise delete fiber
183                 i+=1
184         if i>0:
185             print "deleted", i, "bad_fiber(s)"
186             file=open(name[0].replace('iraf_', '')+'.txt', 'w')
187             for i in range(len(wavelength)):
188                 file.write('%5.3f'%wavelength[i]+' \t %4.2f'%(numpy.median(data
189                [:,i]))+' \n')
190             file.close()
191     ndfilters=[]
192     if os.path.isfile('ndfilter.lst')==True:
193         file=open('ndfilter.lst')
194         for line in file:
195             line=line.split()
196             if len(line)<2:
197                 continue
198             #line[1]=line[1].replace(' ', '')
199             line[1]=line[1].replace('WYC_clear1', '')
200             line[1]=line[1].replace('WYC_', '')
201             line[1]=line[1].replace('0', '')
202             ndfilters.append(line)
203         file.close()
204     if ndfilters==[]:
205         print "no_or_empty_file:_ndfilter.lst"
206
207     result=[]
208     arc_catalogs=[]
209     log_lines=[]
210     for name in names: #does the important stuff to each file
211         shortname=name[0].replace('iraf_', '')
212         if shortname in ignore_files and use_ignore_files==True:
213             continue
214         spectrum=[]
215         if os.path.isfile(shortname+'.txt')==False:
216             print shortname+'.txt_does_not_exist, _please_set_ "comborder" _to_"

```

```

    True" '
197     exit(1)
198     file=open(shortname+'.txt','r')
199     for line in file:
200         line=line.split()
201         spectrum.append([float(line[0]),float(line[1]),len(spectrum)])
202     file.close()
203     keyresult=[]
204     header_table = pyfits.getheader(name[1])
205     for keyword in keywords:
206         if keyword in header_table.keys():
207             temp=str(header_table[keyword])
208             if temp.find('_')>-1:
209                 keyresult.append(header_table[keyword].replace('_', ''))
210
211         else:
212             keyresult.append(header_table[keyword])
213
214     else:
215         keyresult.append('')
216
217     #get data for wavelength in plot
218     spectrum=numpy.array(spectrum)
219     spectrum=sorted(spectrum, key=itemgetter(0)) #to have smallest and highest
220     wavelength at beginning and end, respectively
221     spectrum=numpy.array(spectrum)
222     minwave=spectrum[0,0]
223     maxwave=spectrum[-1][0]
224     exptime='%1.1i'%float(keyresult[0])
225     if exptime=='0':
226         exptime='1'
227     grat=keyresult[1]
228     if grat=='ECHELLE':
229         grat=grat+'-'+keyresult[3]
230     cenwave=str(keyresult[2])
231     for line in replace.cenwave: #replace not even numbers in cenwave
232         if cenwave==line[0]:
233             cenwave=line[1]
234         break
235     arc=keyresult[4]
236     if arc=='':
237         arc=keyresult[5]
238     if arc=='Hg':
239         if name[0].find('p18')>-1: #for p18 the lamp in the header is
240             wrong
241             arc='ThAr'
242         else:
243             arc='HgAr'
244
245     compwaves=[]
246     file=open(lampfolder+'/'+arc+'.txt','r')
247     for line in file:
248         if len(line)<5: #empty lines might be a problem
249             continue
250         line=line.split()
251         if float(line[0])>minwave and float(line[0])<maxwave:
252             temp=[arc, line[0]]
253             if len(line)>1:
254                 compwaves.append(line[0]+'\\t'+line[1]+'\\n')
255                 temp.append(line[1].replace('_', ''))
256             else:
257                 compwaves.append(line[0]+'\\n')
258                 temp.append('')
259             if temp not in arc_catalogs:
260                 arc_catalogs.append(temp)
261     file.close()
262     if sort_out3==True:
263         cont=False
264         if len(compwaves)<3: #at least 3 lines should be available for
265             configuration
266             print 'not_enough_lines:',shortname,minwave,maxwave,grat,
267                 cenwave,arc
268             continue
269     if sort_out_not_usable==True:
270         cont=False
271         for line in not_usable: #doesn't plot the configurations in not_usable
272             if line[0]==grat and line[1]==cenwave and (line[2]==arc.lower
273                 () or line[2]=='*'):
274                 print 'not_used:',shortname,minwave,maxwave,grat,
275                     cenwave,arc
276                 cont=True
277         if cont==True:
278             continue
279     if grat.find("ECHELLE")<>-1: #doesn't plot ECHELLE as not complete
280     print 'all_ECHELLE_modes_are_sorted_out:',shortname,minwave,maxwave,

```



```

271         grat , cenwave , arc
272     continue
273 if compwaves==[]:
274     print 'no_template_data_from_'+lampfolder+'/'+'+arc+'.txt_for ',
275     compwaves.append("0\n")
276 temp=grat+'_'+cenwave+'_'+arc.lower()+ '_'+shortname+'.txt'
277 if os.path.isfile(temp)==True:
278     os.system('rm_'+temp)
279 os.system('ln_s_'+shortname+'.txt_'+temp) #for idl_plot
280 file=open('ticksall_'+grat+'_'+cenwave+'_'+arc.lower()+ '_'+shortname+'.txt', 'w')
281 #for idl_plot
282 for line in compwaves:
283     file.write(line)
284     file.close()
285 print shortname, minwave, maxwave, grat , cenwave , arc , #file , min(wavelength)
286 , max(wavelength) , header-key-result
287 deltawave=(maxwave-minwave)/4
288 wave=[',',',',',',',',',',']
289 wave[1]=str(int(round(minwave+deltawave,-1)))
290 wave[2]=str(int(round(minwave+2*deltawave,-1)))
291 wave[3]=str(int(round(minwave+3*deltawave,-1)))
292 wave[0]=str(int(round(minwave,0))) #1 makes problems in the idl_plot , as
293 the exact wavelength is necessary
294 wave[4]=str(int(round(maxwave,0)))
295 for line in overrideSplittings:
296     if line[0]==grat and line[1]==cenwave and line[2]==arc.lower():
297         for i in range(5):
298             if line[i+3]<>',' :
299                 wave[i]=line[i+3]
300     print "->_with_manual_settings",
301 if borders==True:
302     for dispersion in dispersions:
303         if len(dispersion)>=5: #only dispersions with border
304             information
305             if dispersion[0]==grat and dispersion[1]==cenwave:
306                 #dispersion and cenwave are matching
307                 wave[0]=dispersion[3]
308                 wave[4]=dispersion[4]
309                 break
310 offset='0' #offset
311 for line in manual_shift: #check, if manual offset has to be applied
312     if line[0]==grat and line[1]==cenwave and line[2]==arc.lower():
313         offset=line[3]
314 #spectrum=sorted(spectrum, key=itemgetter(1,2)) #sort by flux and pixel
315 #spectrumori=spectrum
316 for i in range(len(compwaves)):
317     line=compwaves[i].split()
318     compwaves[i]=float(line[0]) #replace ['wavelength',line] by float(
319 wavelength)
320 compwaves.sort()#sort by wavelength
321 dlambd=1.0*(maxwave-minwave)/len(spectrum)
322 if dlambd<1.5:
323     step=[2,4,6,10] #search-range for [maximum, close peak, too high peak
324 too close, not real maximum]
325 else:
326     step=[2,3,5,7]
327 maxflux=[]
328 posis=[]
329 quarter=0
330 for i in range(len(spectrum)): #get max(flux) in each of the four subplots
331     if spectrum[i,0]>int(wave[quarter+1]):
332         maxflux.append(max(spectrum[posis,1])/100.)
333         posis=[]
334         quarter+=1
335         if quarter>3:
336             break
337     posis.append(i)
338 if len(maxflux)<4:
339     maxflux.append(max(spectrum[posis,1])/100.)
340 maxflux.append(max(maxflux)/10)
341 if len(maxflux)<>5:
342     exit(100)
343
344 quarter=0
345 usedlines=[]
346 goodlines=[]
347 verygoodlines=[]
348 todelete=[]
349 for compwave in compwaves:
350     text=grat+'_'+cenwave+'_'+arc.lower()+'\t'+shortname+'_'+str(compwave)
351    )+'_%5.2f'%dlambd+'\t'
352     for i in range(quarter,4):

```

```

344         if compwave>int(wave[quarter+1]):           #next subplot
345             quarter+=1
346         else:
347             break
348         temp=abs(spectrum[:,0]-compwave)           #distance to compwave
349         mini=min(temp)                             #smallest distance to compwave
350         posi1=numpy.where(temp==mini)[0]           #index of smallest distance to
351             compwave
352         posi1=posi1[0]                             #make it number instead of array
353         if posi1<step[3] or posi1>len(spectrum)-step[3]: #only central
354             lines, these line is too far to the borders
355             log_lines.append("6_"+text+'border\t'+str(posi1))
356             continue
357         if mini/dlambda<0.3: #well centered
358             posi1=[posi1]
359         elif temp[posi1-1]<temp[posi1+1]: #lower pixel is also close to
360             catalog line
361             posi1=[posi1-1,posi1]
362         else: #higher pixel is also close to catalog
363             line
364             posi1=[posi1, posi1+1]
365         maxi=max(spectrum[posi1[0]-step[0]:posi1[-1]+step[0]+1,1]) #
366             maximum of column 2 (flux)
367         if maxi<maxflux[quarter] or maxi<maxflux[4]: #otherwise the
368             lines are not visible in the subplot
369             log_lines.append("5_"+text+'brightness\t'+str([maxi,maxflux[
370                 quarter],maxflux[4]]))
371             continue
372         posi2=numpy.where(spectrum[posi1[0]-step[0]:posi1[-1]+step[0]+1,1]==
373             maxi)[0] #where is maximum of the line
374         posi2=posi2[0]+posi1[0]-step[0] #make it
375             number instead of array
376         text=text+str(spectrum[posi2,0])+'\t'+str([posi1, posi2])+'\t'+str(maxi
377             )+'\t'
378         if max(spectrum[posi2-step[1]:posi2+step[1]+1,1])>maxi:
379             #another peak too close
380             log_lines.append("1_"+text+'neighbor1\t'+str(max(spectrum[
381                 posi2-step[1]:posi2+step[1]+1,1])))
382             continue
383         if max(spectrum[posi2-step[2]:posi2+step[2]+1,1])*0.1>maxi:
384             #too high peak too close
385             log_lines.append("2_"+text+'neighbor2\t'+str(max(spectrum[
386                 posi2-step[2]:posi2+step[2]+1,1])))
387             continue
388         posis=range(posi2-step[3],posi2-step[0]+1)+range(posi2+step[0],posi2+
389             step[2]+1)
390         std=numpy.std(spectrum[posis,1], ddof=1)
391         med=numpy.median(spectrum[posis,1])
392         if maxi<med+3.5*std: #not significant enough
393             log_lines.append("3_"+text+'significance\t'+str(numpy.round([
394                 med,std],1)))
395             continue
396         if posi2 in usedlines:
397             if goodlines not in todelete:
398                 todelete.append(goodlines[-1])
399             log_lines.append("4_"+text+'same_line\t'+str(goodlines[-1]))
400         log_lines.append("0_"+text+'good_line\t'+str(numpy.round([med,std],1)
401             ))
402         usedlines.append(posi2)
403         goodlines.append(compwave)
404         if maxi>10*maxflux[quarter]: #if more than 10% of brightest
405             line in quarter
406             verygoodlines.append(compwave)
407     for i in range(len(goodlines))[:-1]:
408         if goodlines[i] in todelete:
409             del goodlines[i]
410     for i in range(len(verygoodlines))[:-1]:
411         if verygoodlines[i] in todelete:
412             del verygoodlines[i]
413     file=open('goodlines_'+grat+'_'+cenwave+'_'+arc.lower()+'.txt','w')
414     for line in goodlines:
415         if line in verygoodlines:
416             file.write(str(line)+'\t'+str(line)+'\n')
417         else:
418             file.write(str(line)+'\n')
419     file.close()
420     #print goodlines
421     ndfilter='-1'
422     for line in ndfilters: #add, if neutral density filter was used
423         if line[0].find(shortname)<>-1 or shortname.find(line[0])<>-1:
424             ndfilter=line[1]
425     break

```

```

408     if ndfilter=='-1':
409         print "not_in_list_of_ndfilters",
410     print ""
411
412     #line="idl_plot,'" + grat + "','"+cenwave + "','"+arc.lower() + "','"+arc + "','"+
413     #exptime + "','"+shortname + "','"+wave[0] + "','"+wave[1] + "','"+wave[2] + "','"+wave
414     # [3] + "','"+wave[4]
415     index=0 #for nice sorting
416     index+=len(grat)*1E6
417     if grat.find('ECHELLE')==0:
418         index+=(9-int(grat[-1]))*1E4 #start with ECHELLE-6 and end with
419         ECHELLE-3
420     else:
421         index+=int(grat[1:3])*1E4 #start with low resolutions
422         if grat[-1]=='B': #to distinguish between same
423             gratings with B
424             index+=0*1E4
425         elif grat[-1]=='R': #... and R
426             index+=1*1E4
427         index+=int(cenwave[: -1])*1E1 #only 3 digits
428         if arc.lower()=='he':
429             index+=1
430         if arc.lower()=='ne':
431             index+=2
432         if arc.lower()=='hgar':
433             index+=3
434         if arc.lower()=='thar':
435             index+=4
436         if arc.lower()=='cd':
437             index+=5
438         arctext=arc
439         if arc=='HgAr':
440             arctext='Hg'
441         #if arc<>'Cd':
442         # continue
443         line=[index, grat, cenwave, arc.lower(), arctext, exptime, shortname, wave[0], wave
444             [1], wave[2], wave[3], wave[4], offset, str(len(goodlines)), str(len(
445             verygoodlines)), ndfilter]
446         result.append(line)
447
448     result.sort()
449     text=''
450     file=open('run_in_idl.pro','w')
451     #file.write('.RESET_SESSION\n\n')
452     for line in result:
453         text=text+'_' +line[1]+'_' +line[2]+'_' +line[3].lower()+'_' +line[6]+' .pdf'
454         plot=True
455         if os.path.isfile(line[1]+'_' +line[2]+'_' +line[3].lower()+'_' +line[6]+' .pdf')
456             ==True:
457             if os.path.getmtime(line[1]+'_' +line[2]+'_' +line[3].lower()+'_' +line
458             [6]+' .pdf')>os.path.getmtime(line[1]+'_' +line[2]+'_' +line[3].lower
459             ()+'_' +line[6]+' .txt'):
460                 plot=False
461                 print line[1]+'_' +line[2]+'_' +line[3].lower()+'_' +line[6]+' .
462                 pdf'," is newer than",line[1]+'_' +line[2]+'_' +line[3].lower
463                 ()+'_' +line[6]+' .txt'," and therefore not plotted"
464         if plot==True:
465             file.write("idl_plot,'" +line[1]+' ',''+line[2]+' ',''+line[3].lower()+'
466             ',''+line[4]+' ',''+line[5]+' ',''+line[6]+' ',''+line[7]+' ',''+line[8]+
467             ',''+line[9]+' ',''+line[10]+' ',''+line[11]+' ',''+line[12]+' \n')
468             #grat          cenwave          arc
469             extension      arctext          exptime          wavelength          offset
470         file.write("\nspawn,_'pdftk_' +text+'_cat_output_catalog.pdf'\n")
471         file.write("print,_'catalog.pdf_erstellt'\n")
472         file.close()
473
474     arc_catalogs.sort()
475     before=''
476     file=open('../arc_catalogs.txt','w')
477     for line in arc_catalogs:
478         if line[0]<>before:
479             before=line[0]
480             file.write('\n'+line[0]+' \n')
481         file.write(str(round(float(line[1]),3)).ljust(10)+'_&L'+line[2].ljust(6)+'_&L
482         '\\\-\n')
483     file.close()
484
485     file=open('logfile_usable_lines','w')
486     for line in log_lines:
487         file.write(line+'\n')

```

```

473 file.close()
474
475 before=-1000
476 file=open('../table.tex','w')
477 file.write('\begin{center}\n')
478 file.write('\begin{longtable}{lccccclrrcr}\n')
479 file.write('\label{tab:allconfigs}\n')
480 file.write('\bf_Grating&\bf_\$lambda_{\mathrm{cen}}\$&\bf_\$lambda_{\mathrm{
min}}-\lambda_{\mathrm{max}}\$&\bf_Dispersion&\bf_Lamp&\bf_Exp.time&\bf
_Lines&\bf_ND\\n')
481 file.write('\bf_AA&\bf_AA&\bf_AA/px&\bf_s&\bf_
filter\\n')
482 file.write('(1)&(2)&(3)&(4)&(5)&(6)&(7)&(8)\\n')
483 file.write('\hline\n')
484 file.write('\endfirsthead\n\n')
485 file.write('\bf_Grating&\bf_\$lambda_{\mathrm{cen}}\$&\bf_\$lambda_{\mathrm{
min}}-\lambda_{\mathrm{max}}\$&\bf_Dispersion&\bf_Lamp&\bf_Exp.time&\bf
_Lines&\bf_ND\\n')
486 file.write('\bf_AA&\bf_AA&\bf_AA/px&\bf_s&\bf_
filter\\n')
487 file.write('\hline\n')
488 file.write('\endhead\n\n')
489 #file.write('\hline\n')
490 file.write('\endfoot\n\n')
491 file.write('\hline\n')
492 file.write('\endlastfoot\n\n')
493 for line in result:
494     for dispersion in dispersions:
495         if line[1]==dispersion[0] and line[2]==dispersion[1]: #dispersion
496             and cenwave are matching
497                 break
498         if line[0]-before>1E4:
499             file.write('\hline\n')
500             file.write(line[1]+'&'+line[2]+'&'+line[7]+'-'+line[11]+'$&'+
501                 dispersion[2]+'&'+line[4]+'&'+line[5]+'&'+line[13]+'-'+'line
502                 [14]+'&'+line[15]+'\\n')
503                 #grat          cenwave          min wave          max
504                 wave          dispersion          arctext          exptime good
505                 lines          ndfilter
506         elif line[0]-before>1E1:
507             file.write('\hline\n')
508             file.write('&'+line[2]+'&'+line[7]+'-'+line[11]+'$&'+
509                 dispersion[2]+'&'+line[4]+'&'+line[5]+'&'+line[13]+'-'+'line
510                 [14]+'&'+line[15]+'\\n')
511                 #no grat          cenwave min wave          max wave
512                 dispersion          arctext          exptime          good lines
513                 ndfilter
514         else:
515             file.write('&'+line[4]+'&'+line
516                 [5]+'&'+line[13]+'-'+'line[14]+'&'+line[15]+'\\n')
517             #file.write(' &'+line[7]+' &'+line[11]+' &'+line
518                 [4]+' &'+line[5]+' &'+old+' \\n')
519             #no gratcenwave min wave          max wave no dispersion
520             arctext          exptime
521         before=line[0]
522 file.write('\end{longtable}\n')
523 file.write('\end{center}\n')
524 file.close()
525
526 os.system('idl<-run.in.idl.pro')

```