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LA PALMA

RGO/LA PALMA TECHNICAL NOTE No. 35

Spectrum of the Copper-Neon lamp on the IDS

Plots of the spectrum cover the wavelength region 320-970nm from IPCS and CCD data taken on the INT on La Palma.

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Richard Collyns

28 November 1985

SPECTRUM OF THE COPPER-NEON LAMP ON THE IDS

1. INTRODUCTION

This series of maps of the copper-neon spectrum is provided in the same spirit as the earlier copper-argon maps, that is, spectra at some representative integration time are shown, with a standard slit width and useful filter combinations. It is not so complete as the copper-argon survey, in that only a few gratings have been used and all data is from the 235 camera. However, this is probably adequate in view of the nature of this particular spectrum.

2. THE SPECTRUM

The neon spectrum is dominated by very strong features due to NeI in the wavelength range 5850-7440 \AA , and continues redwards with prominent lines between 7500 and 9600 \AA , although these are as much as ten times fainter than those in the former range. Bluewards, neon lines, both NeI and NeII, appear only weakly, and the copper spectrum provides the strongest features. Even so, the strong copper line (CuI) at 5105 \AA is barely one-twentieth the intensity of the principal NeI lines described earlier.

This implies that for low dispersion work the copper-neon lamp is only really useful for wavelength calibrations in the red, whilst the red features are so spaced that at high dispersions very few lines may be present in certain wavelength intervals. In the blue, at intermediate and high dispersions, it may be useful to supplement the argon spectrum in regions where this is a bit sparse e.g. 3300-4000 \AA etc., but significantly longer integrations would be necessary to bring up the few NeII lines below 3250 \AA .

3. THE DATA

The data used for these arc maps were obtained by Paul Murdin (IPCS) and Charles Jenkins (CCD).

The IPCS data covers 3200-8900 using the 1200g/mm and R400V gratings on the 235 camera. It is perhaps a little misleading in the far red because of the inherent intensity of the neon spectrum compensating for fall-off in image tube sensitivity, but is a useful identification guide.

To supplement this, some data from a few CCD frames have been added to extend the material to 9700Å, taken with the R632V grating on the 235 camera.

According to wavelength tables, it should be possible to reach a few weaker lines redward of this, in particular in the 1.0-1.1 μ m region. Careful second-order filtering would be required, and probably considerable patience, but CCD frames in this wavelength range need to be calibrated if work is to be carried out, say, on the 10830Å line.

4. THE PLOTS

While direct pictures of the lamp spectrum are reproduced for both IPCS and CCD data in the first 3 figures, plots are only given for the IPCS. Full details of the instrument configuration are given with each plot, which shows 2000 channels as the mean of five central cross sections in histogram form, together with a representation of the spectrum as if on the 1311 display.

Less data are available for the CCD frames, and these are offered only as an extension to the line identification sequence.

On the plots, a few of the principal lines have been labelled, but the sources of reference for detailed identification are:

5. PRINCIPAL REFERENCES

New lines "The Iron-Neon Hollow-Cathode Spectrum". H M Crosswhite.
J. Res. N.B.S. 79A, 17, 1975.

Copper lines Tables of Spectral line intensities arranged by Elements.
W F Meggers, C H Corliss, B F Scribner.
N.B.S. Mono. 32., 61, 1961.

Both supplemented by data from

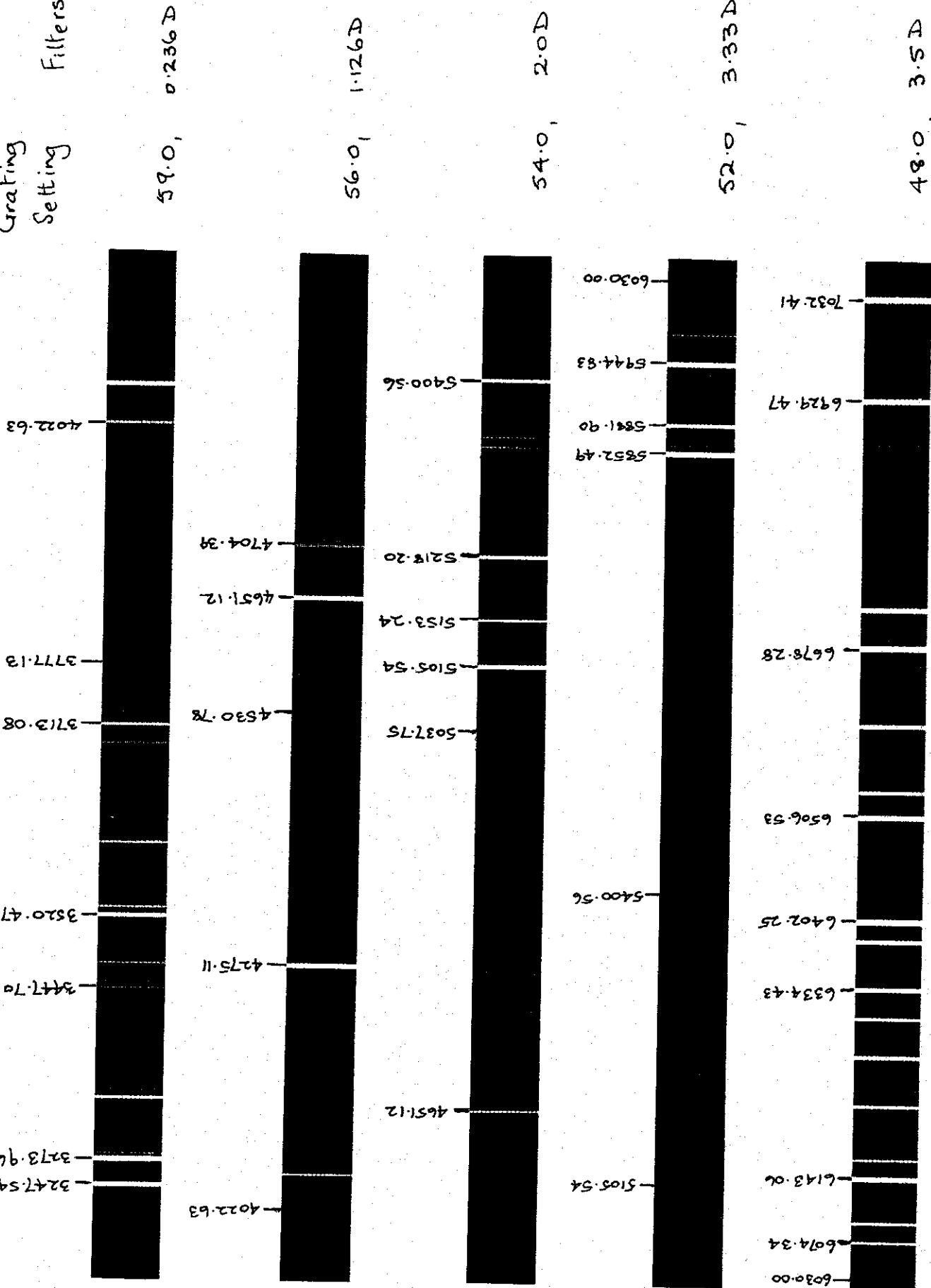
M. I. T. Wavelength Tables, G R Harrison (Wiley, 1939)
Revised Multiplet Tables, Moore, C E
N.B.S. Tech. note 36 (1959).

6. ACKNOWLEDGEMENTS

This note is offered in response to many requests for information on the copper-neon spectrum. With the instrumentation less available for detailed work of this nature, we are obliged to coerce observers into providing self-consistent sets of data for us to present to users. We are therefore most grateful to Paul Murdin for his painstaking trip through the necessary routines to provide the IPCS material, and we thank Charles Jenkins for permitting us to use data from a few of his CCD test frames to extend the wavelength coverage. Richard Worth patiently produced the photographs from the Starlink display monitors, and we appreciate his contribution.

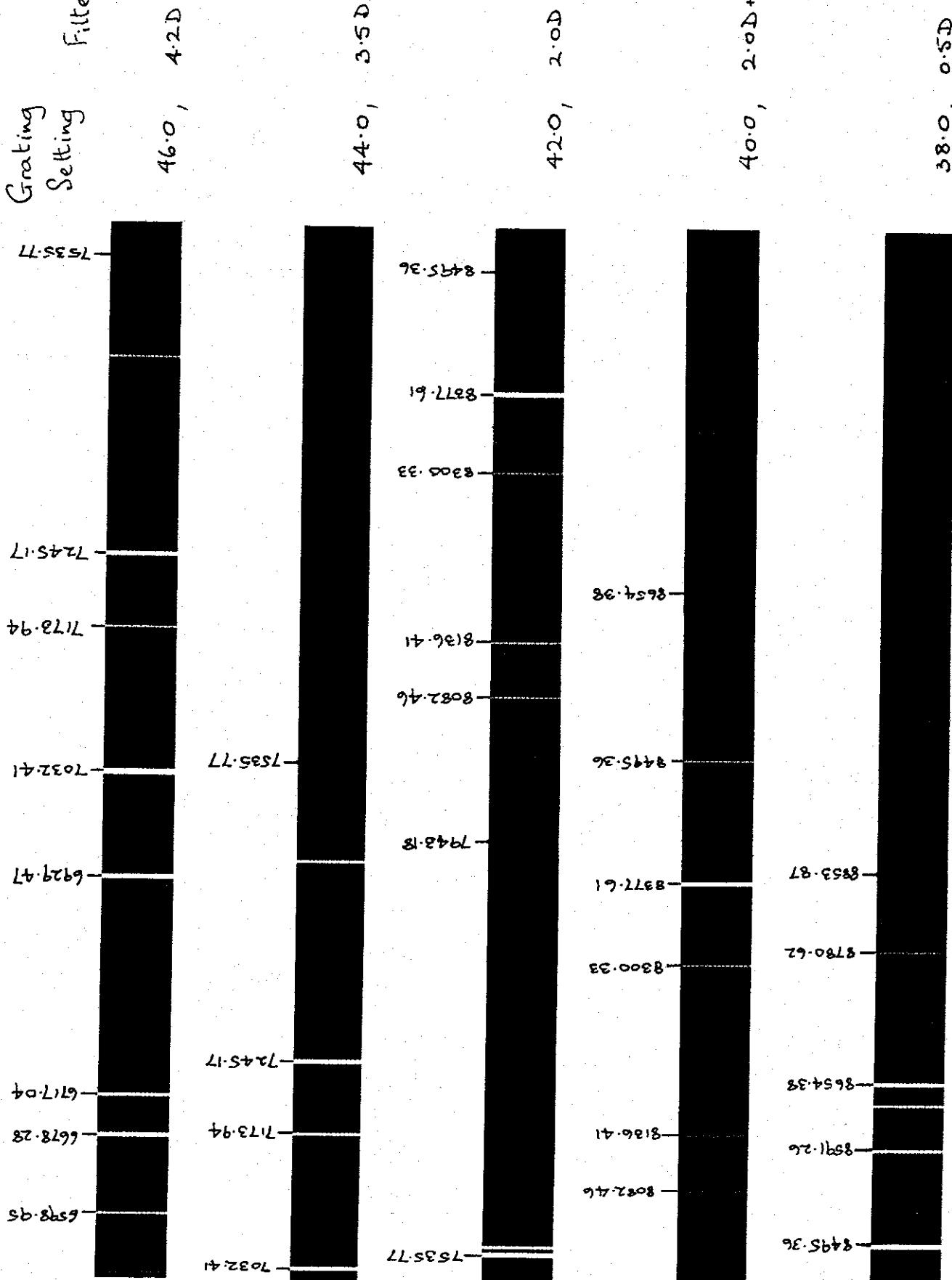
Finally, we would be grateful if any errors could be brought to the attention of Bill Martin, who cares for the general documentation.

Grating
Setting
Filters

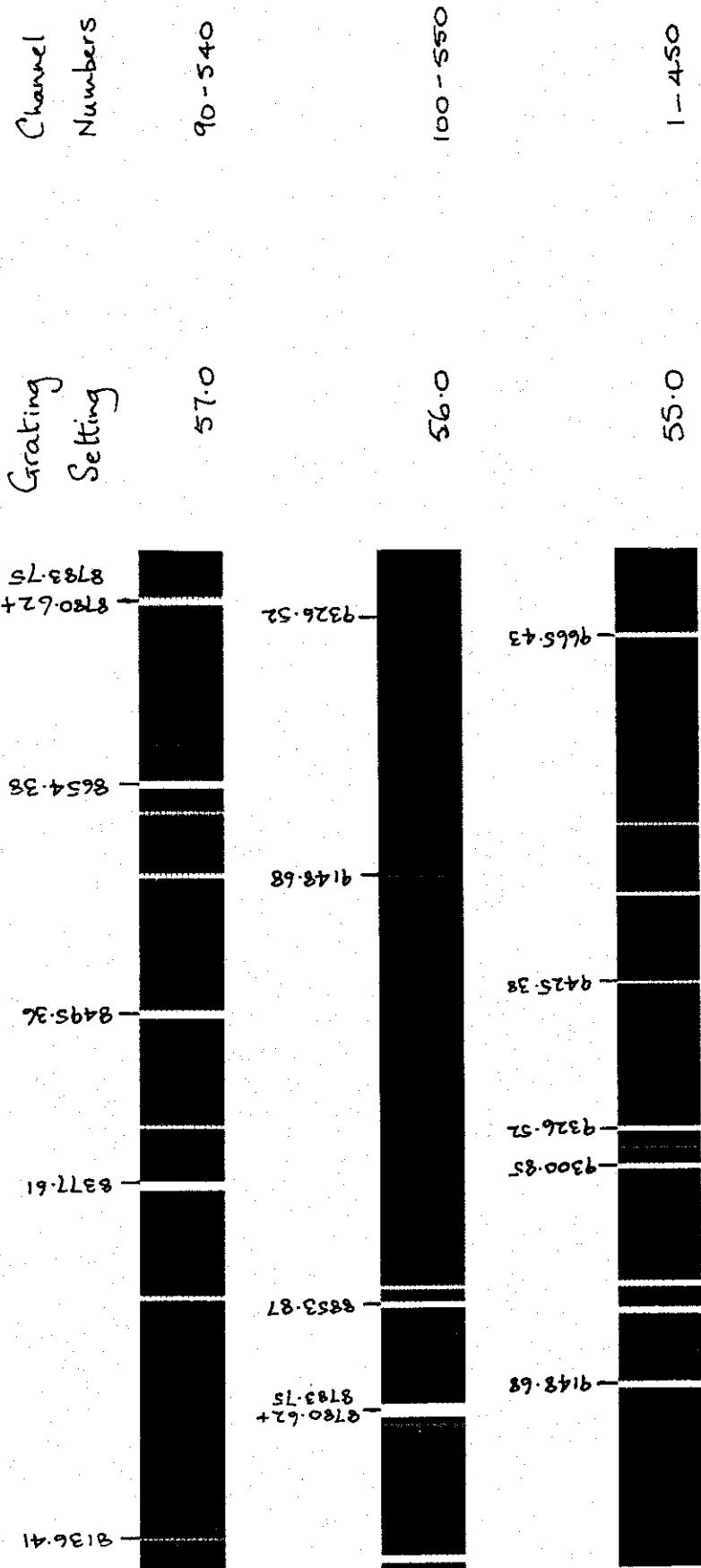


Copper-Neon lamp, 235 Camera, Grating R1200B, 1PCS

Grating
Setting Filters



Copper-Neon lamp, 235 Camera, Grating R1200R, IPCS

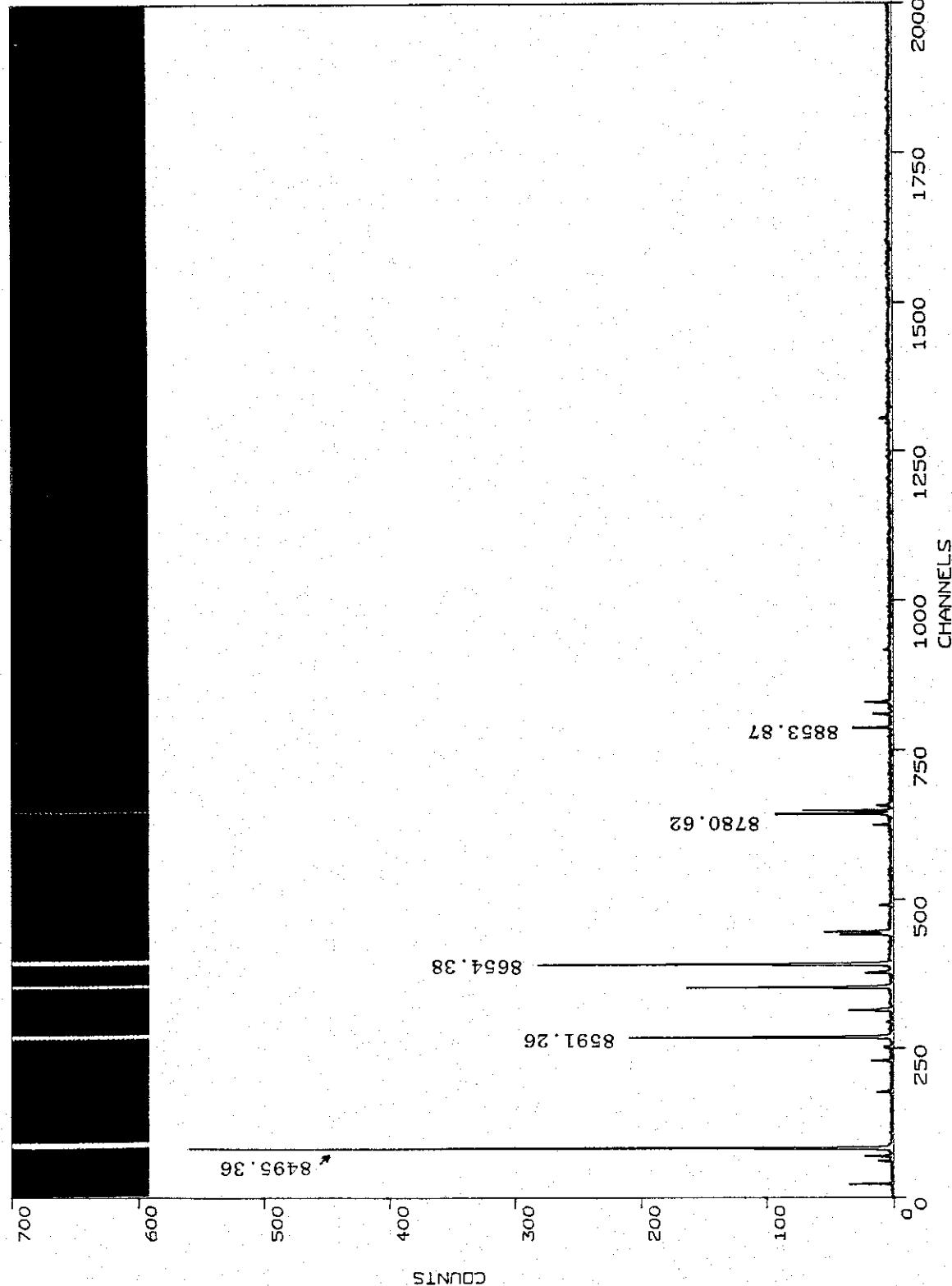


Copper-Neon lamp, 235 Camera, Grating R600I, GEC CCD.

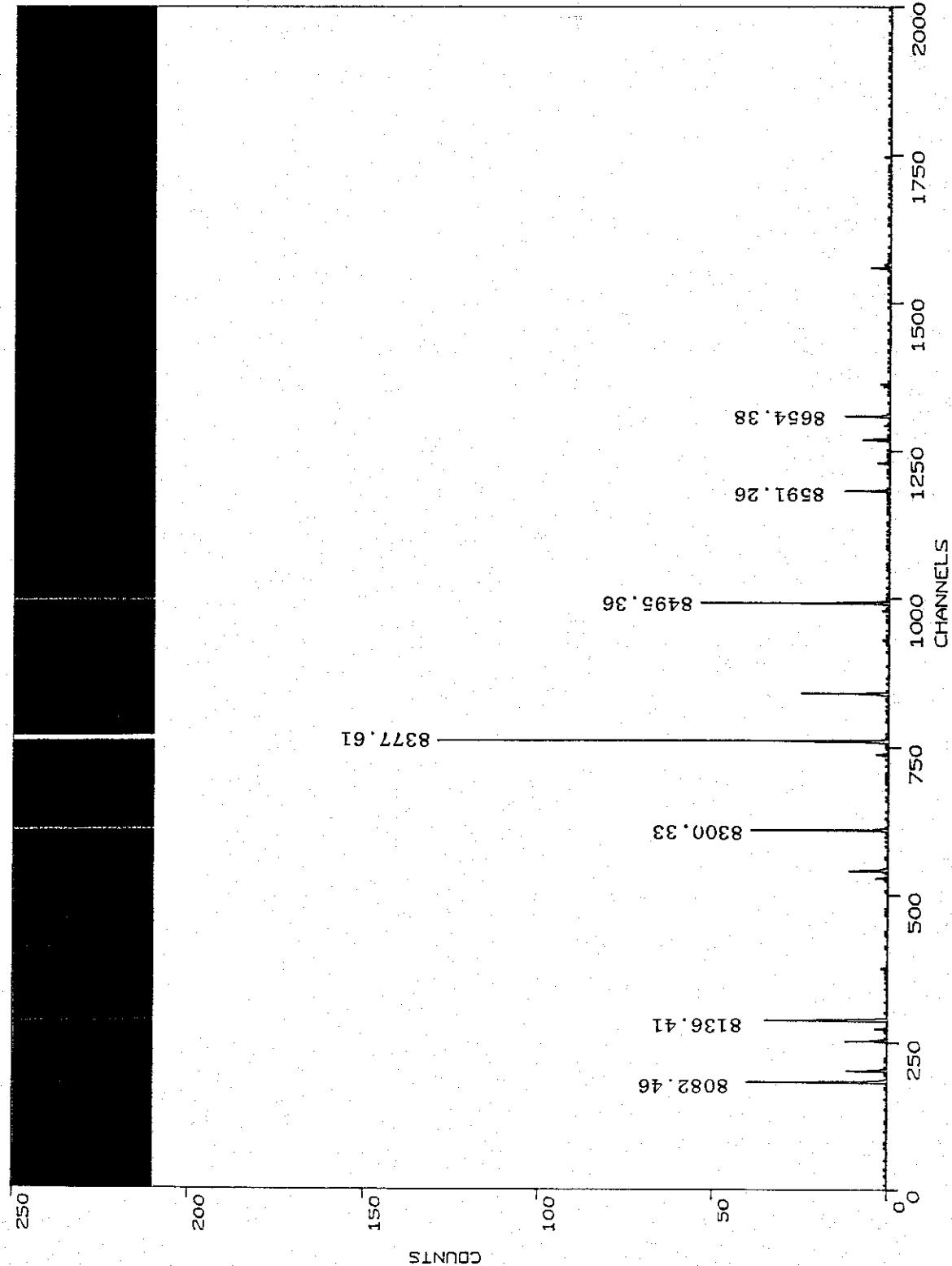
HIGH DISPERSION CU-NE ARC MAPS 8350-3250

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IPCS
235
R 1200R**

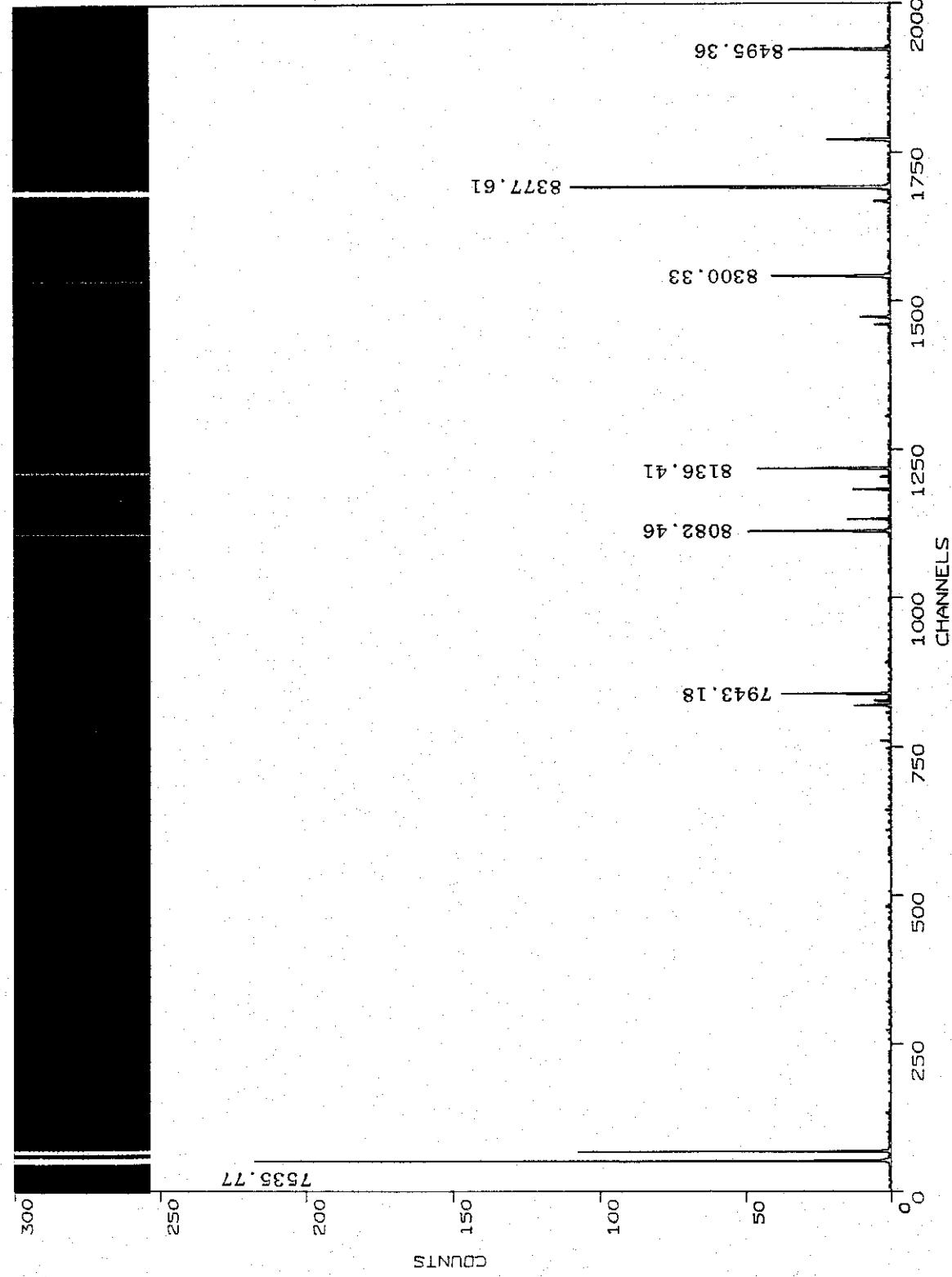
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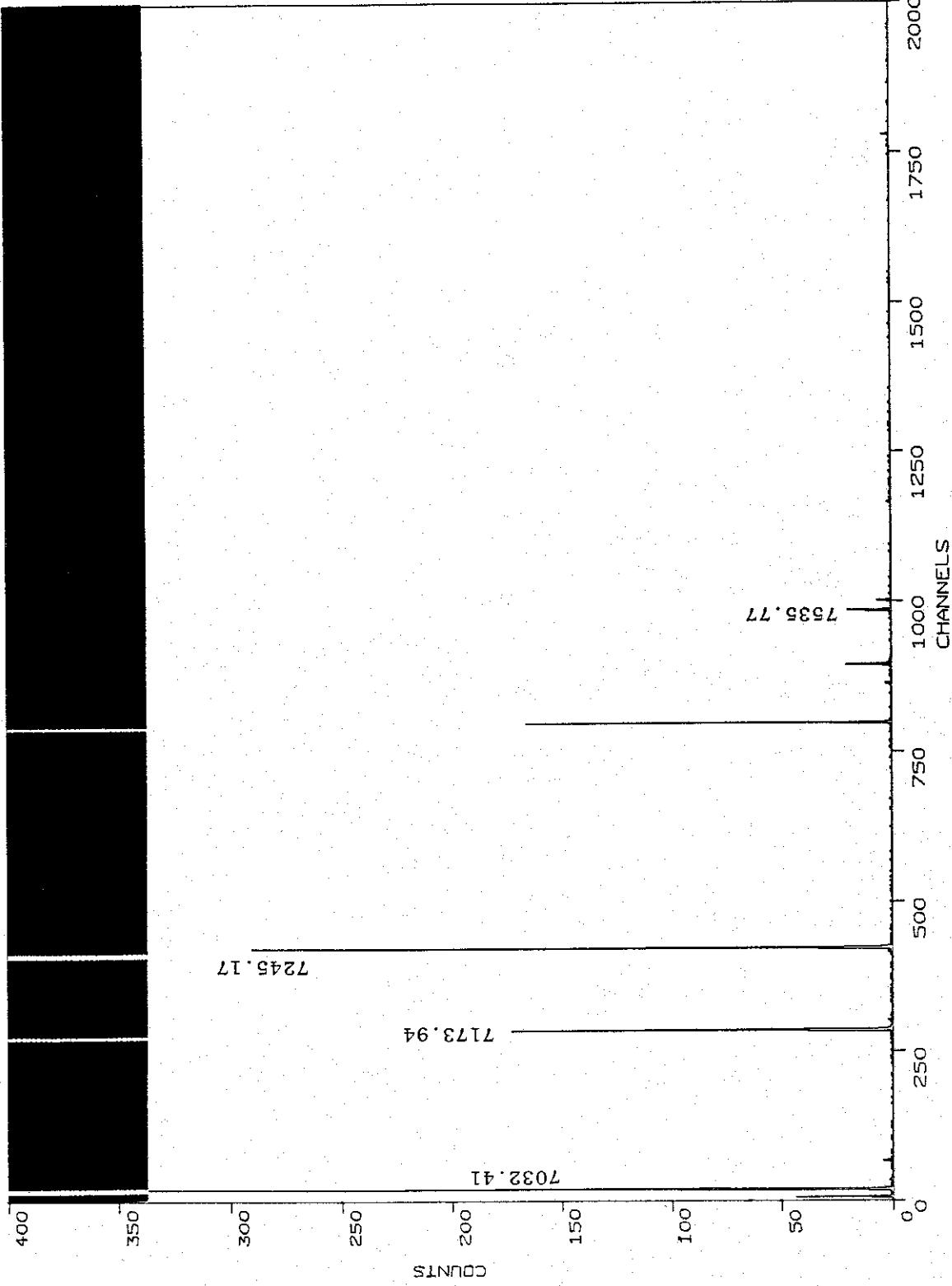
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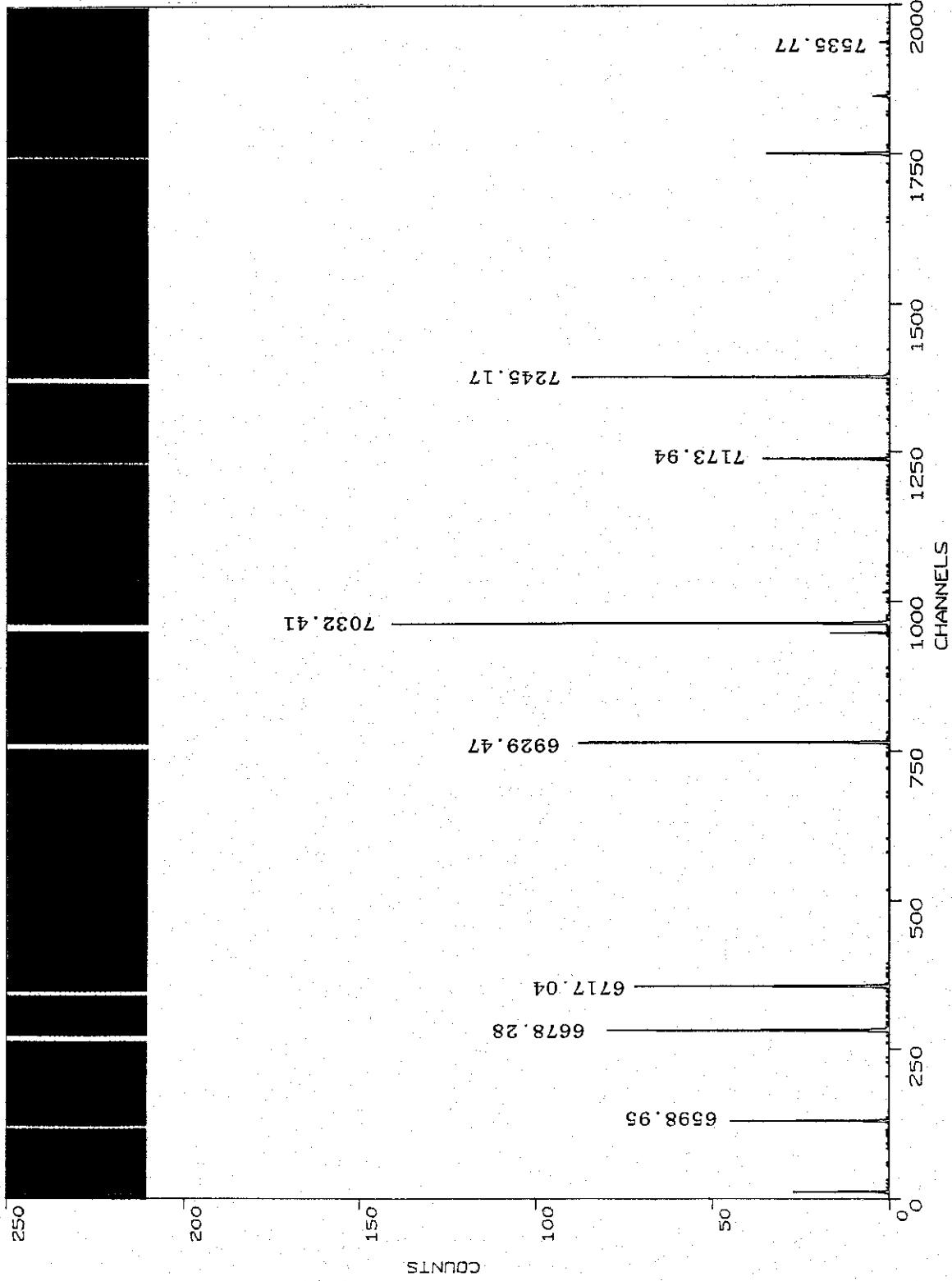
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235 CAMERA WITH TIPCS



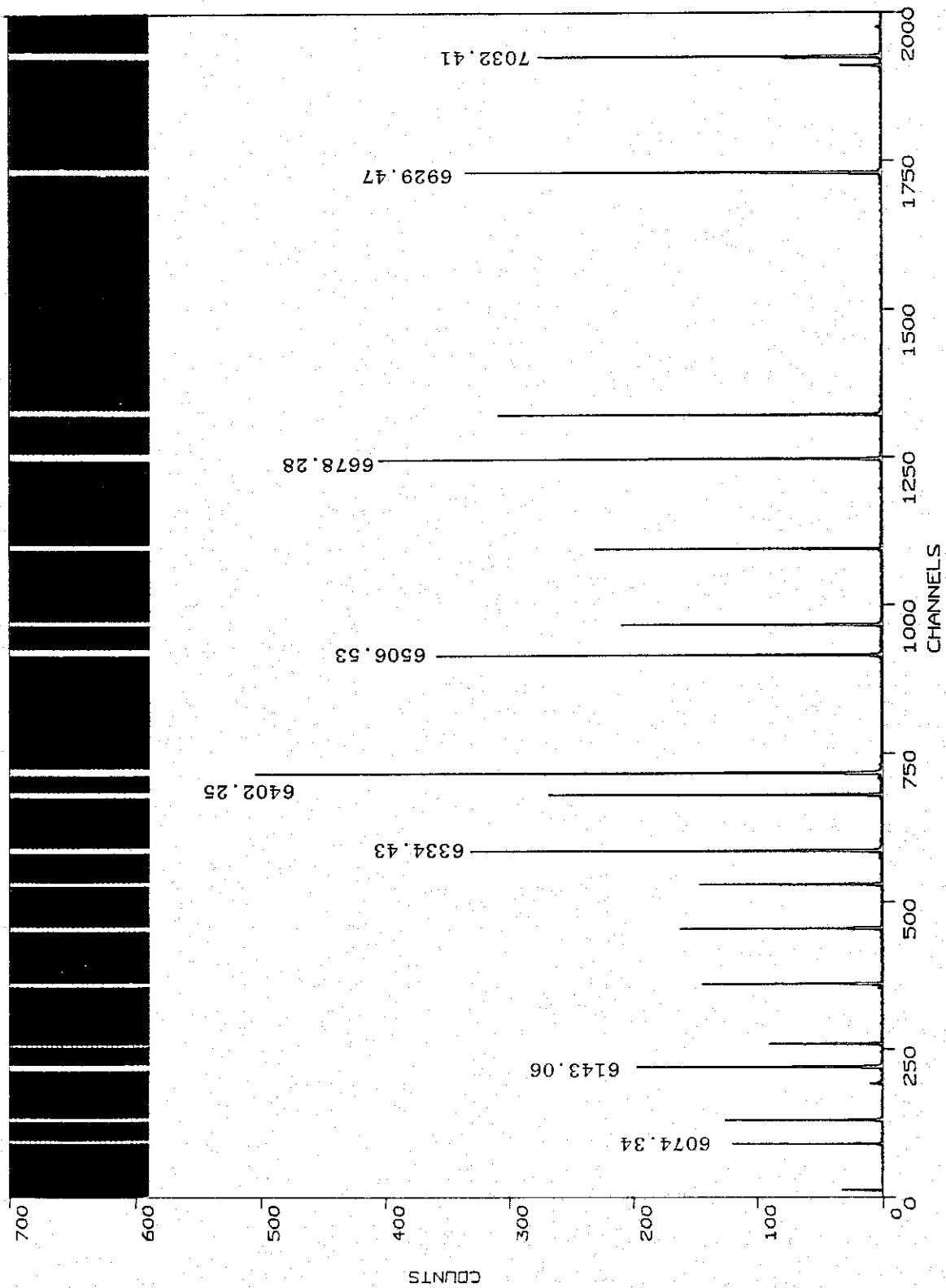
2.35 CAMERA WITH TIPCS



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..... (A=5, B=4)
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SOURCE CUNE

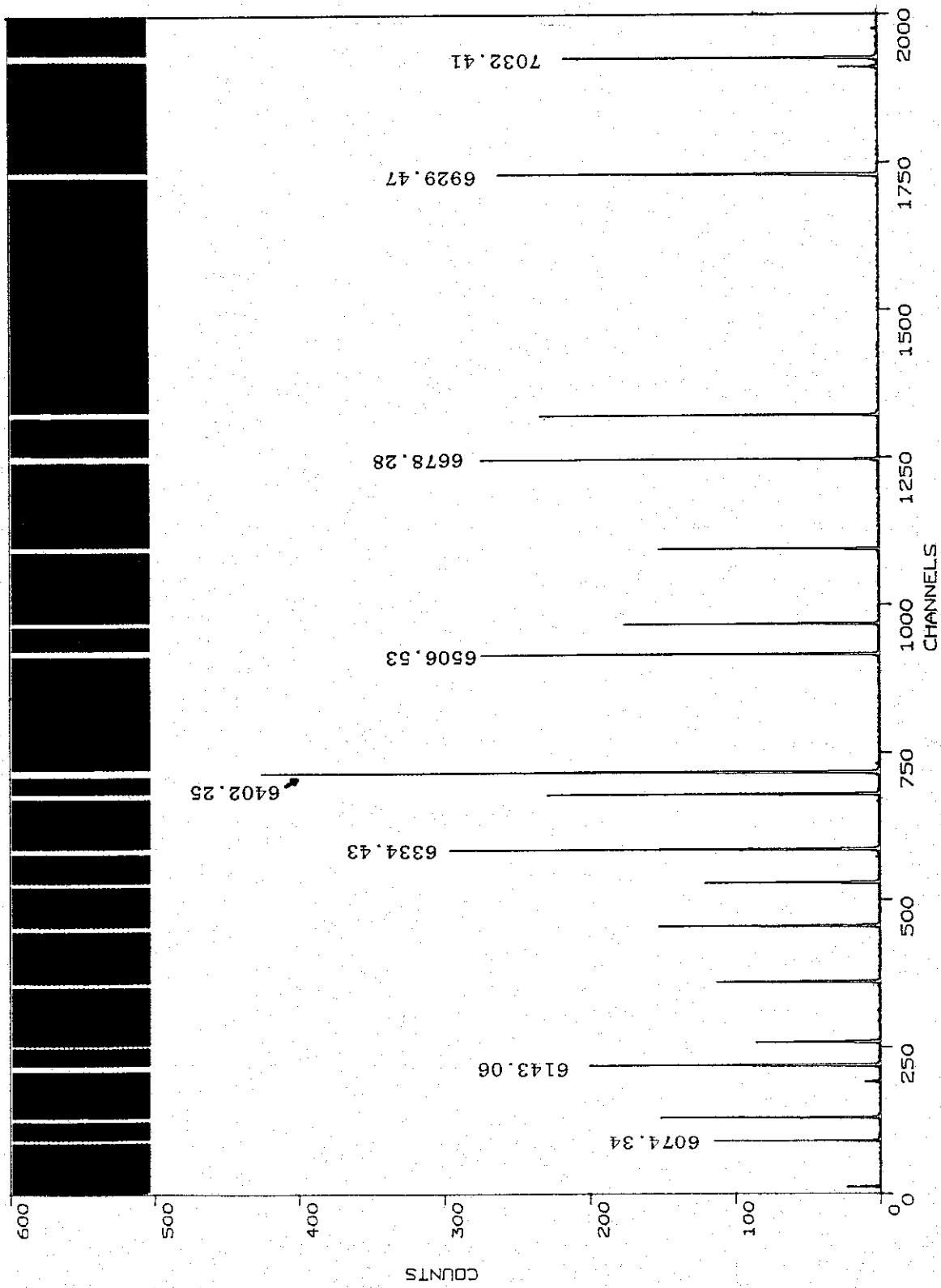
235 CAMERA WITH I PCCS



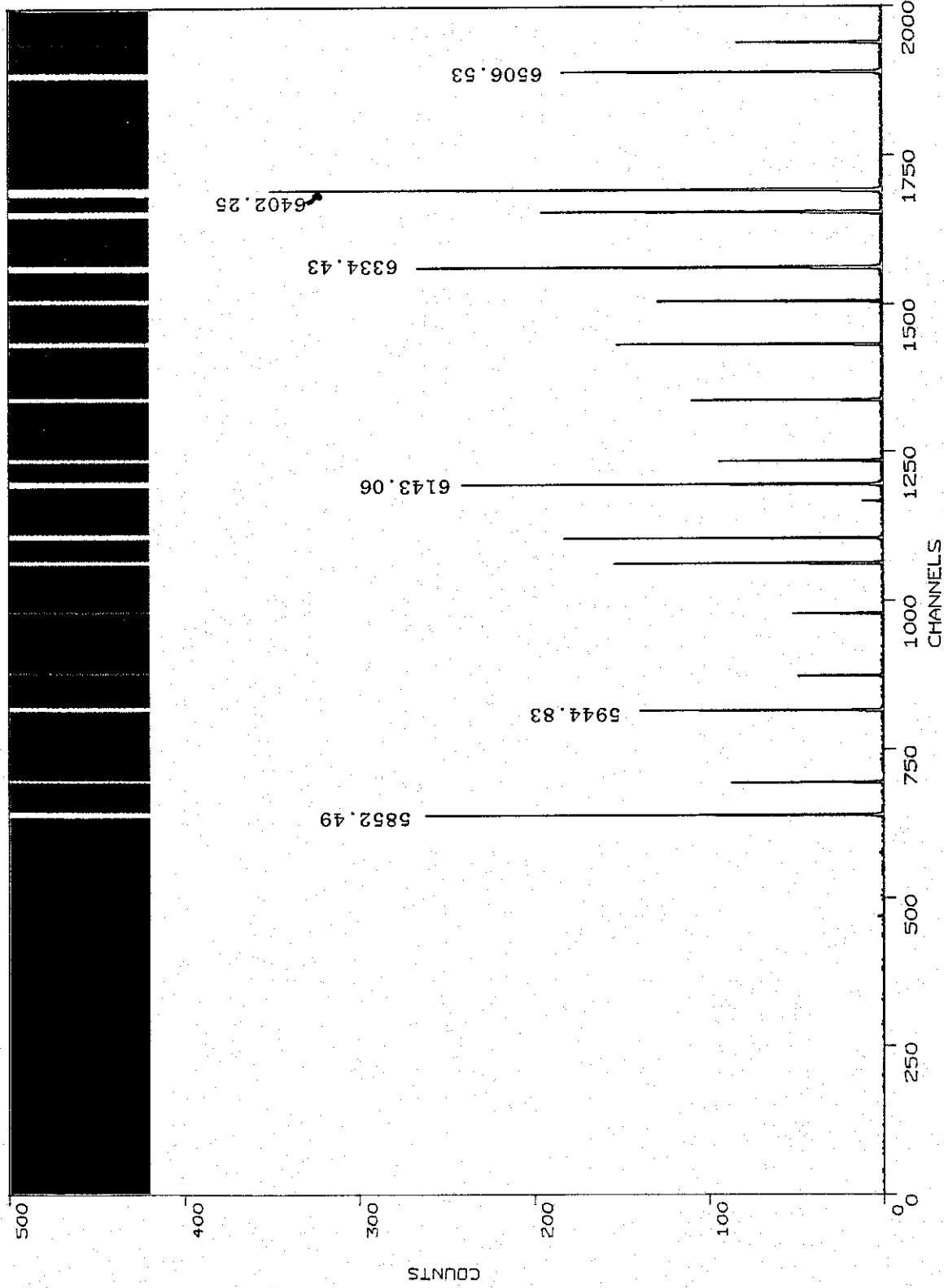
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EXPOSURE 200S
SOURCE CUNE

**CU-NE
IPCS
235
R 1200B**

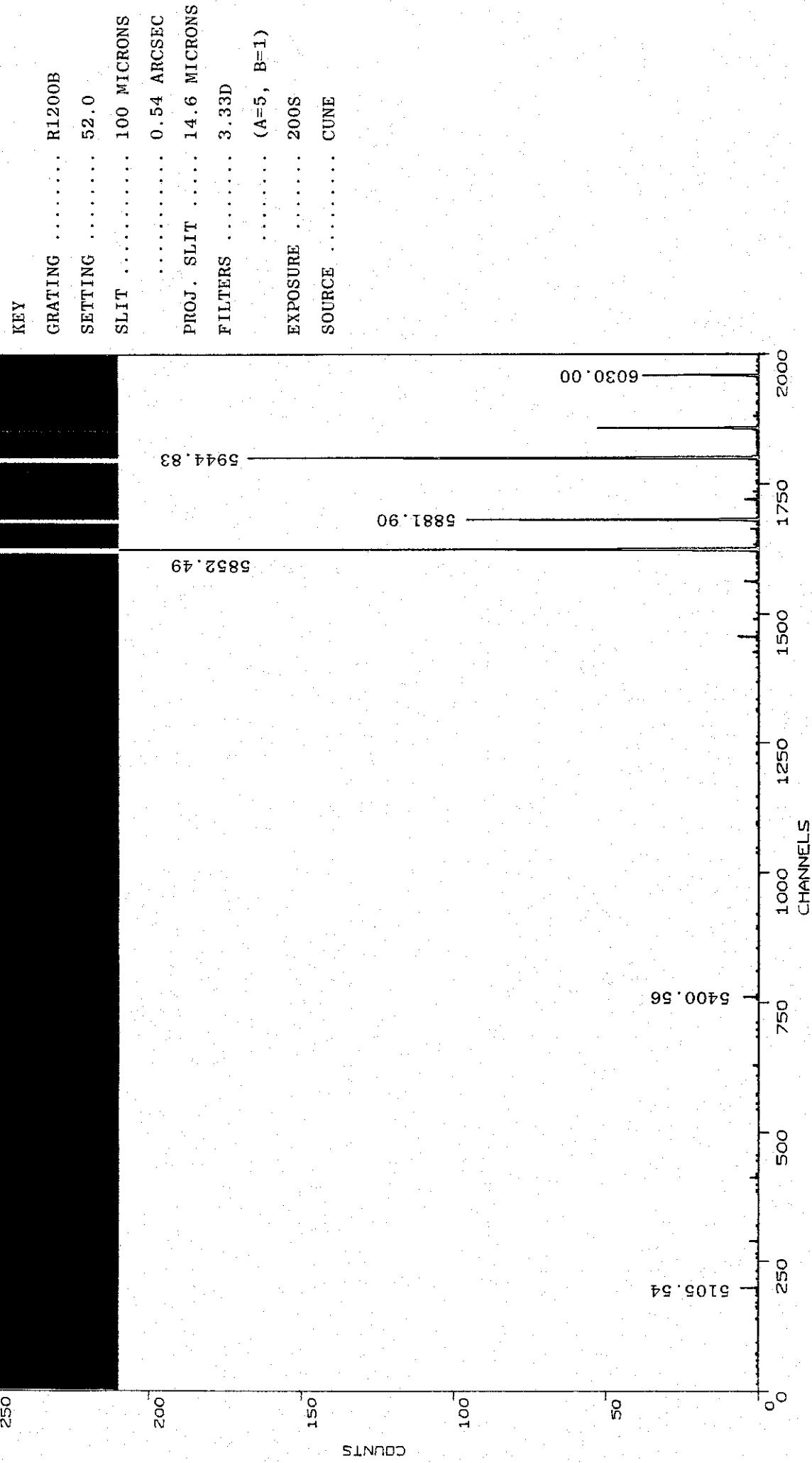
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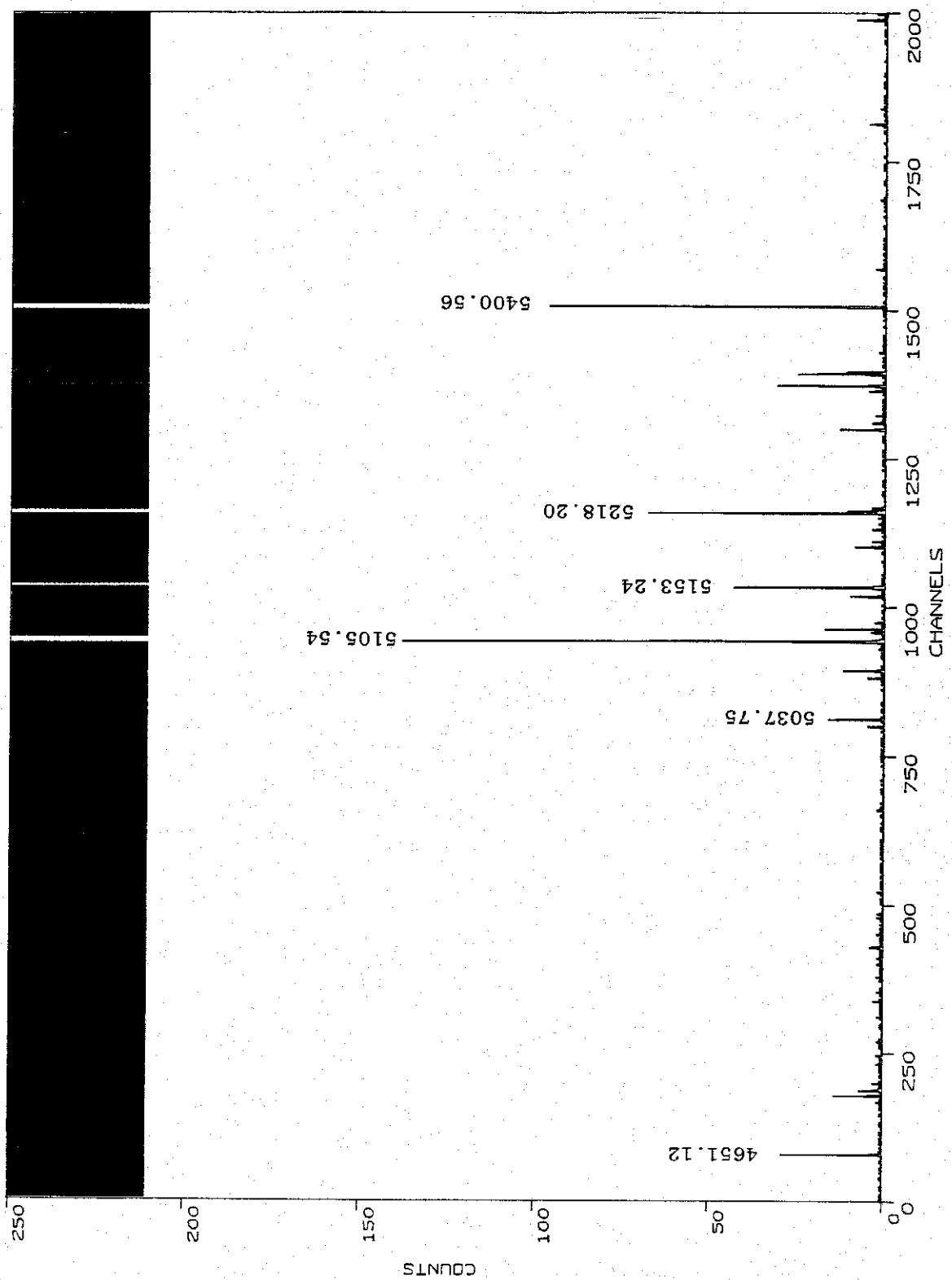
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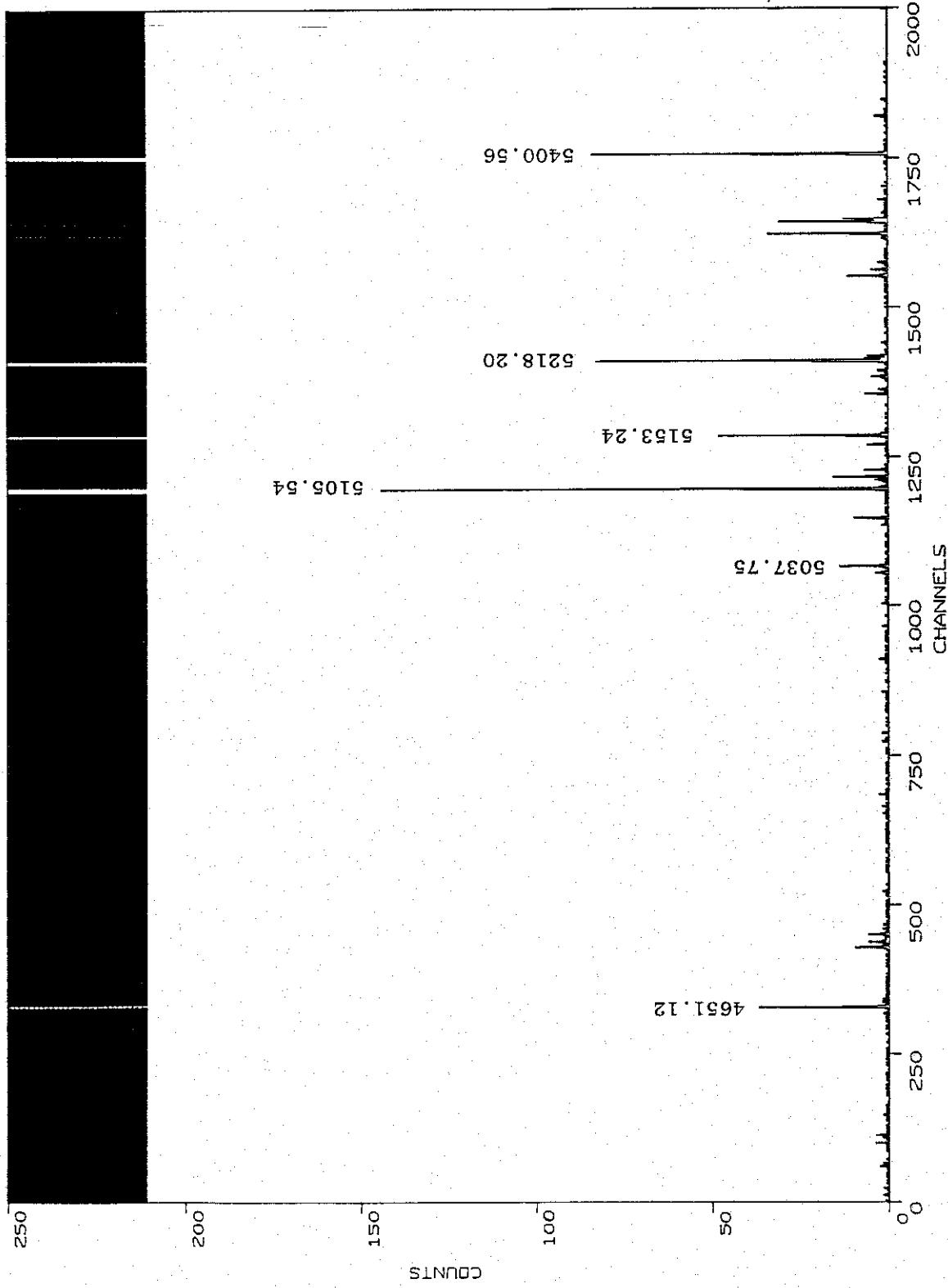
235 CAMERA WITH TIPCS



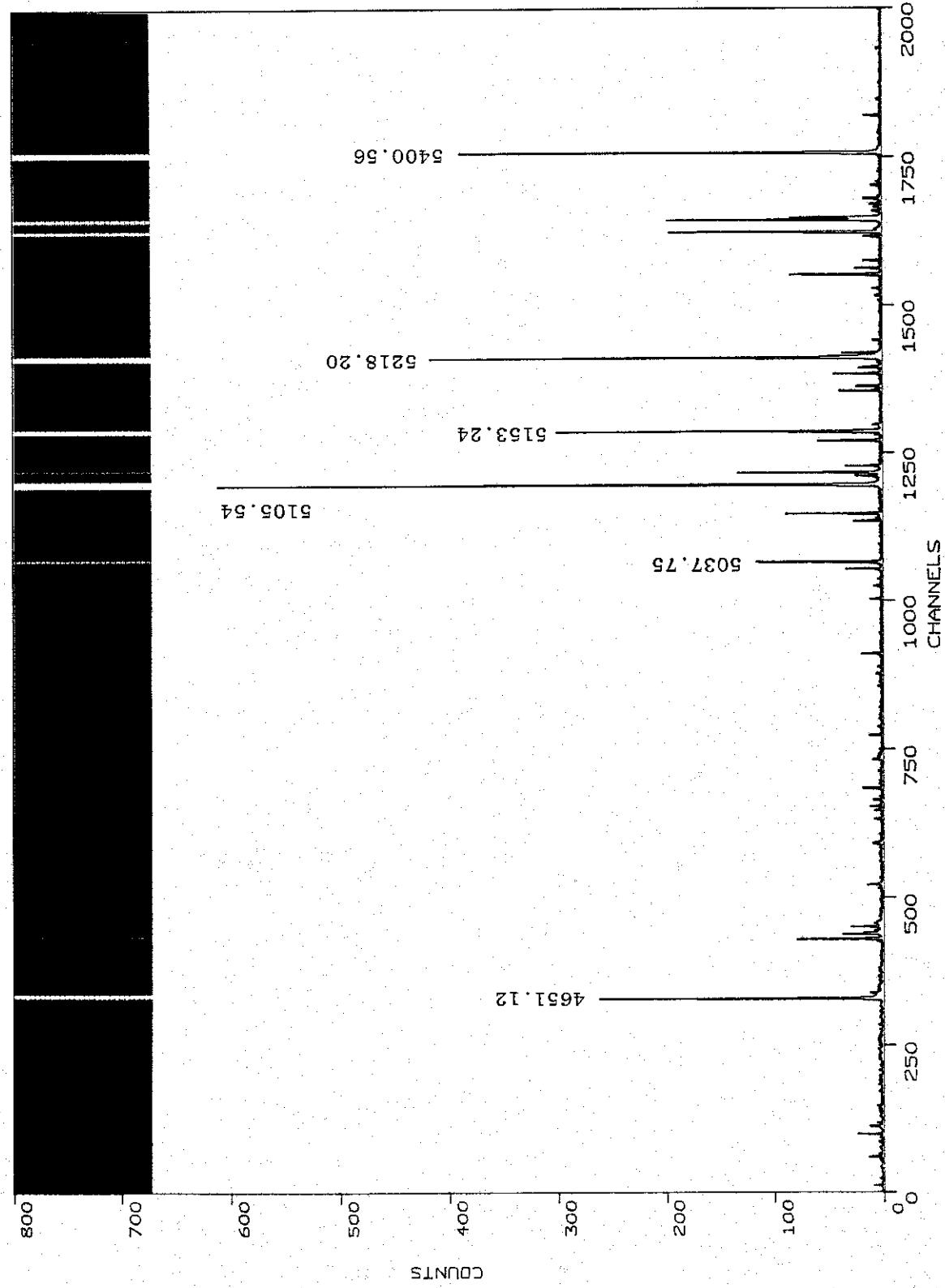
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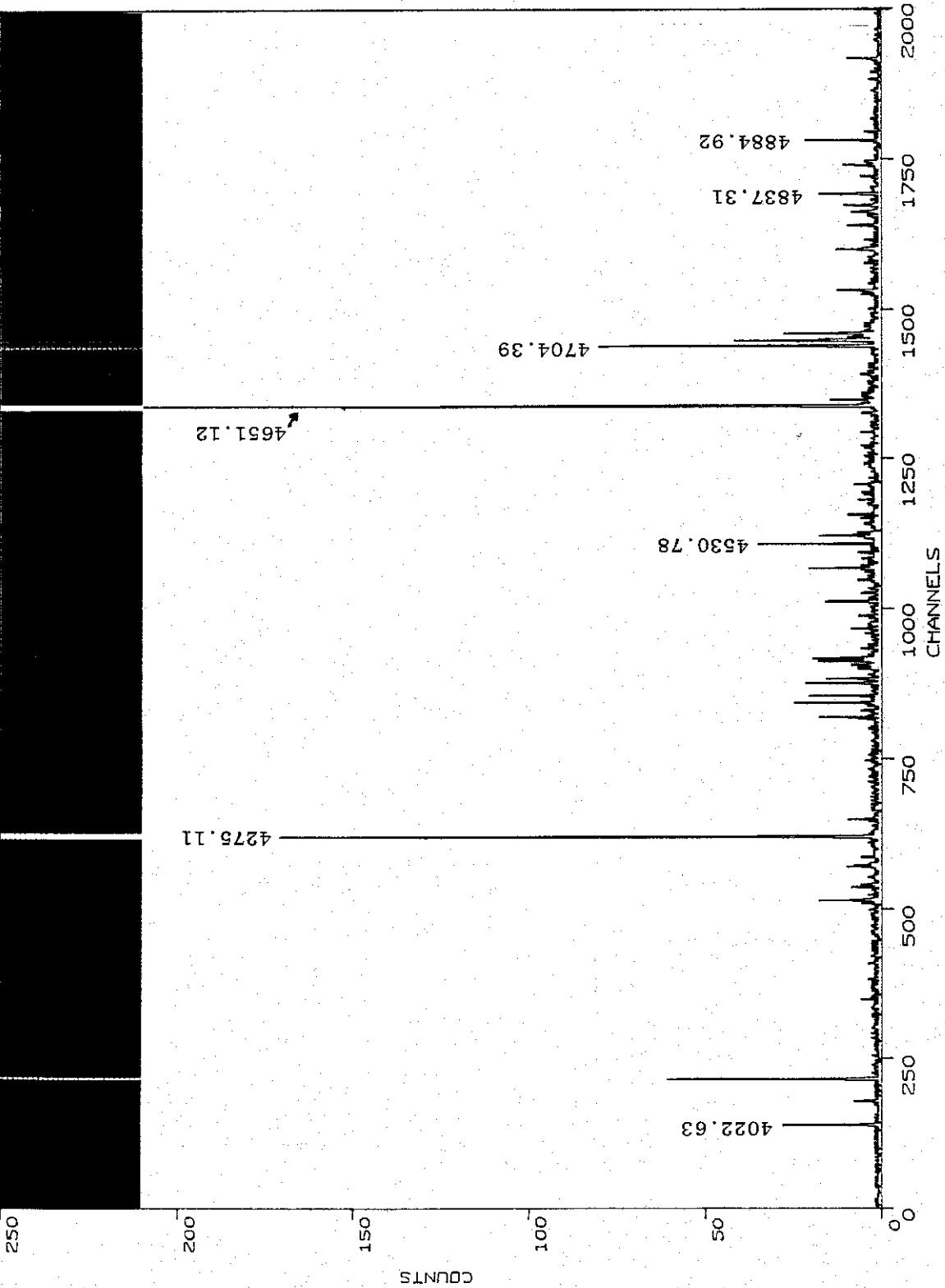
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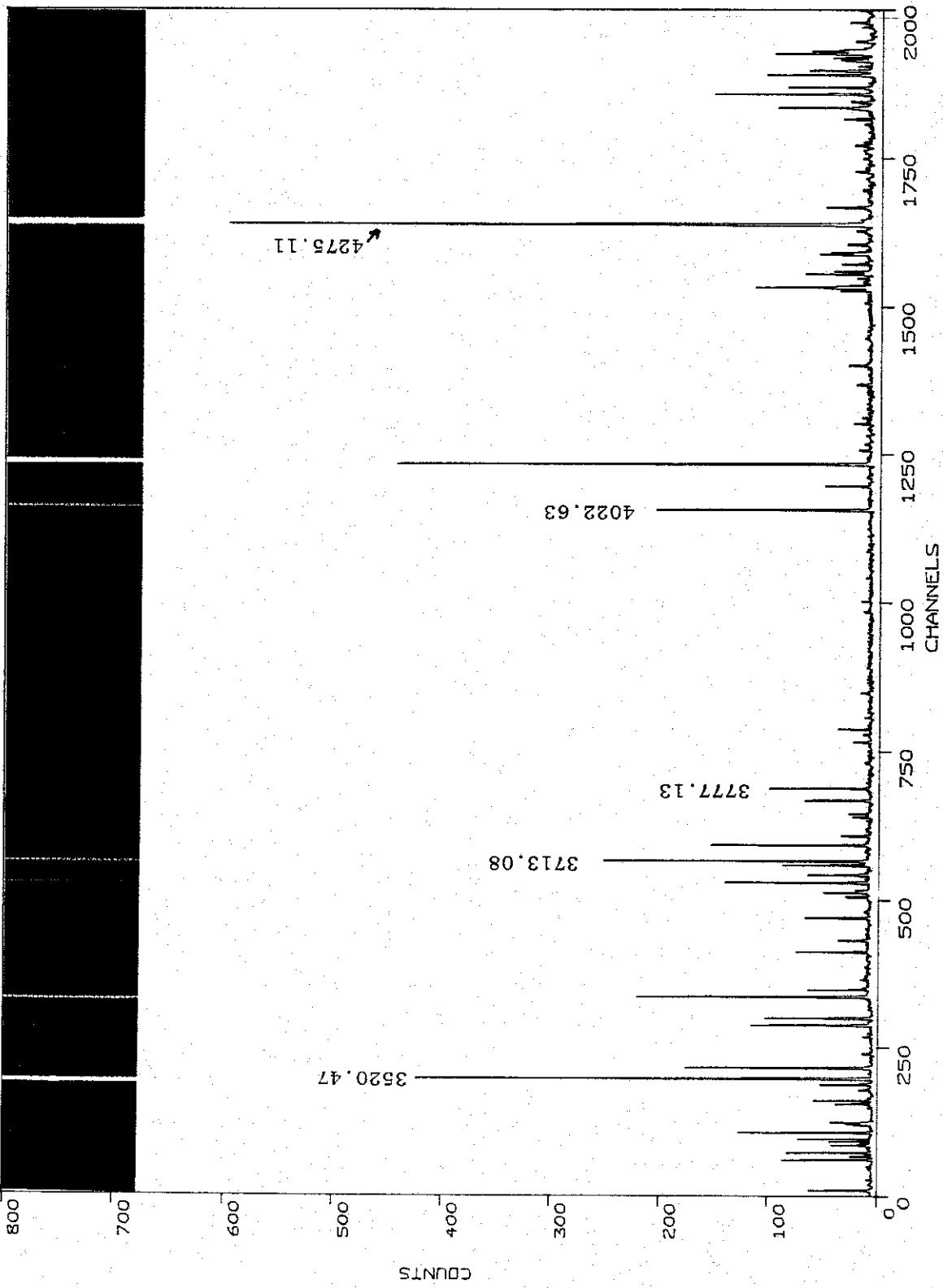
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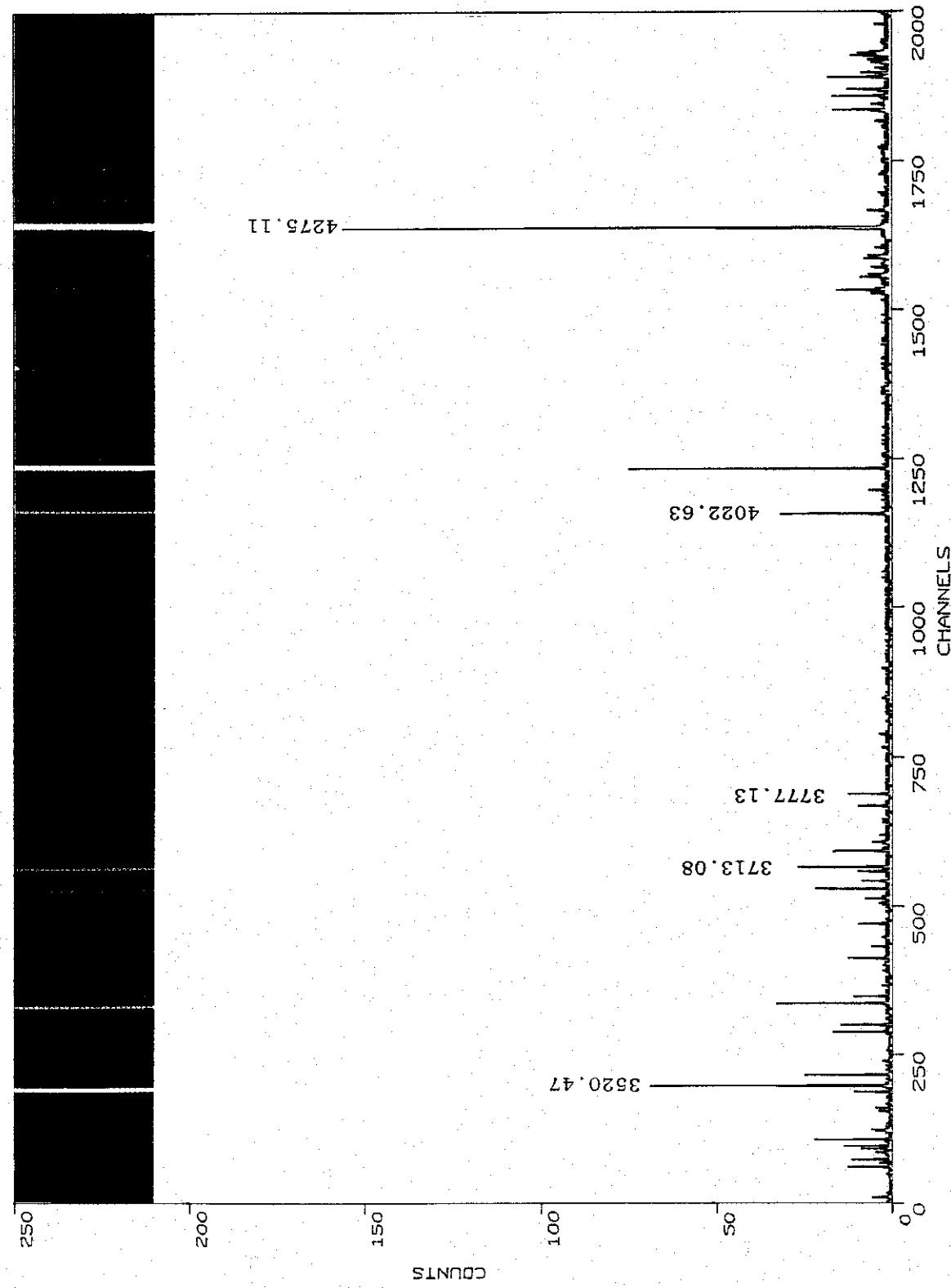
235 CAMERA WITH TIPCS



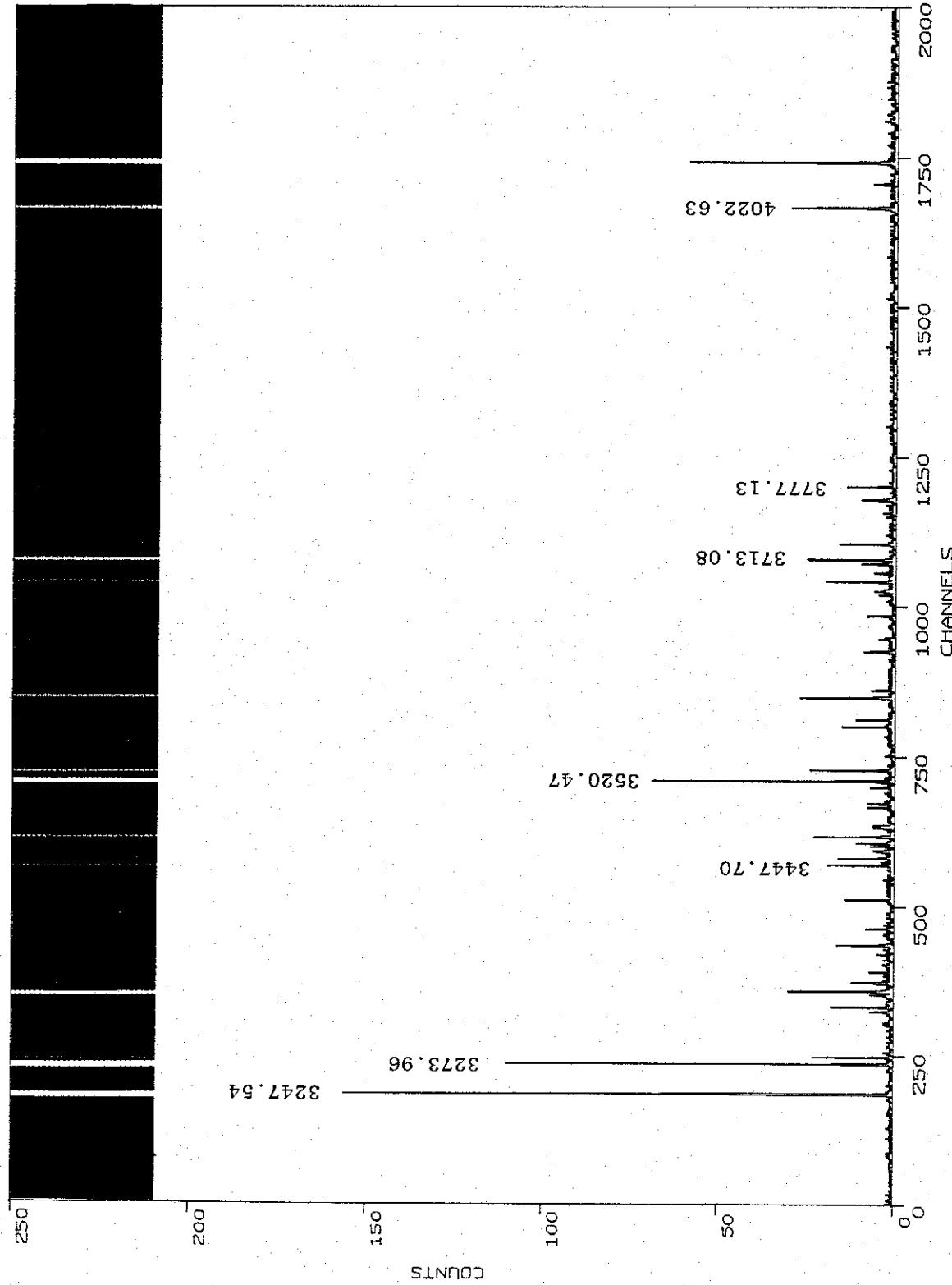
235. CAMERA WITH LIPCS



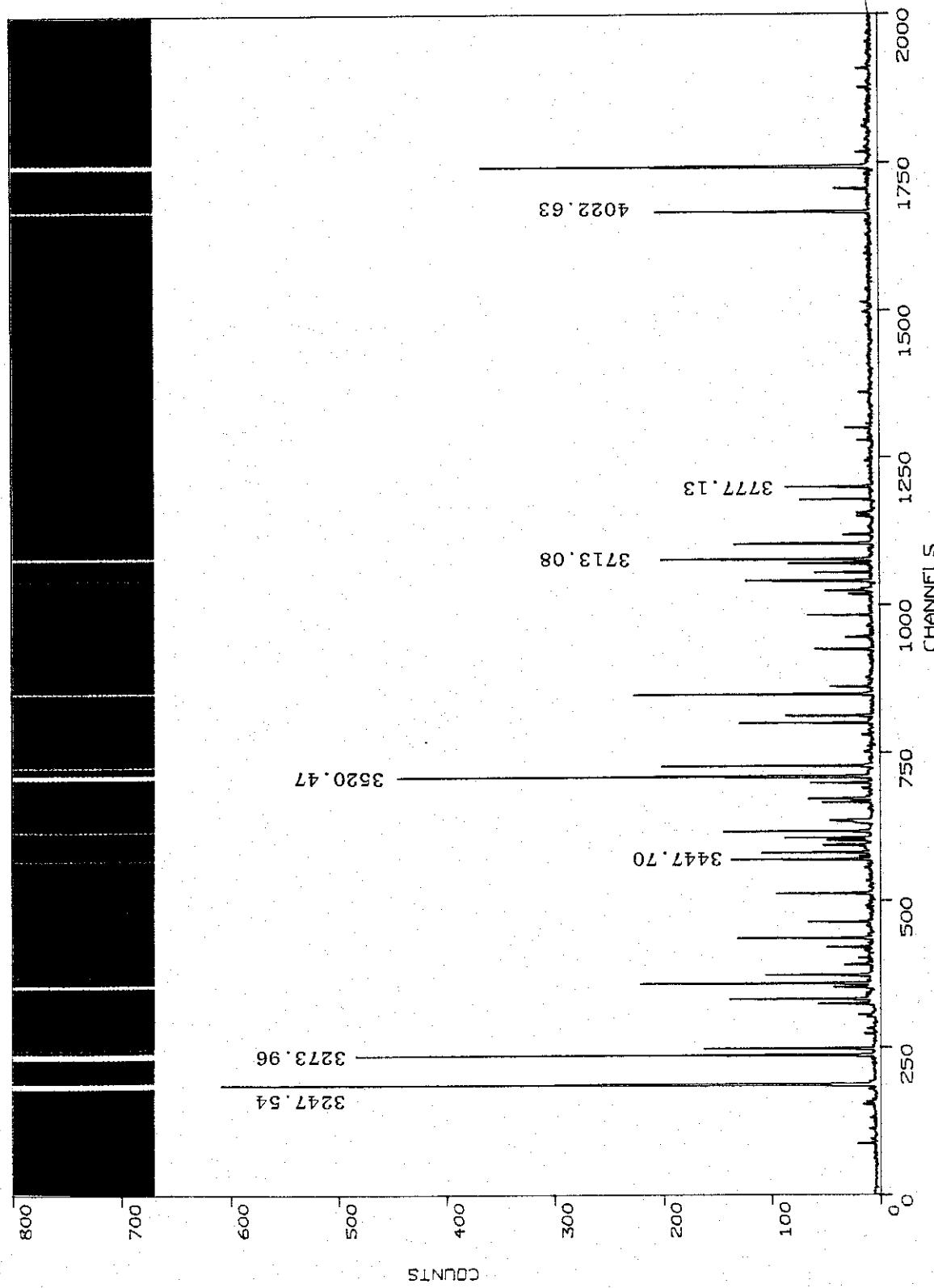
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2 3 5 C A M E R A W I T H I P C S

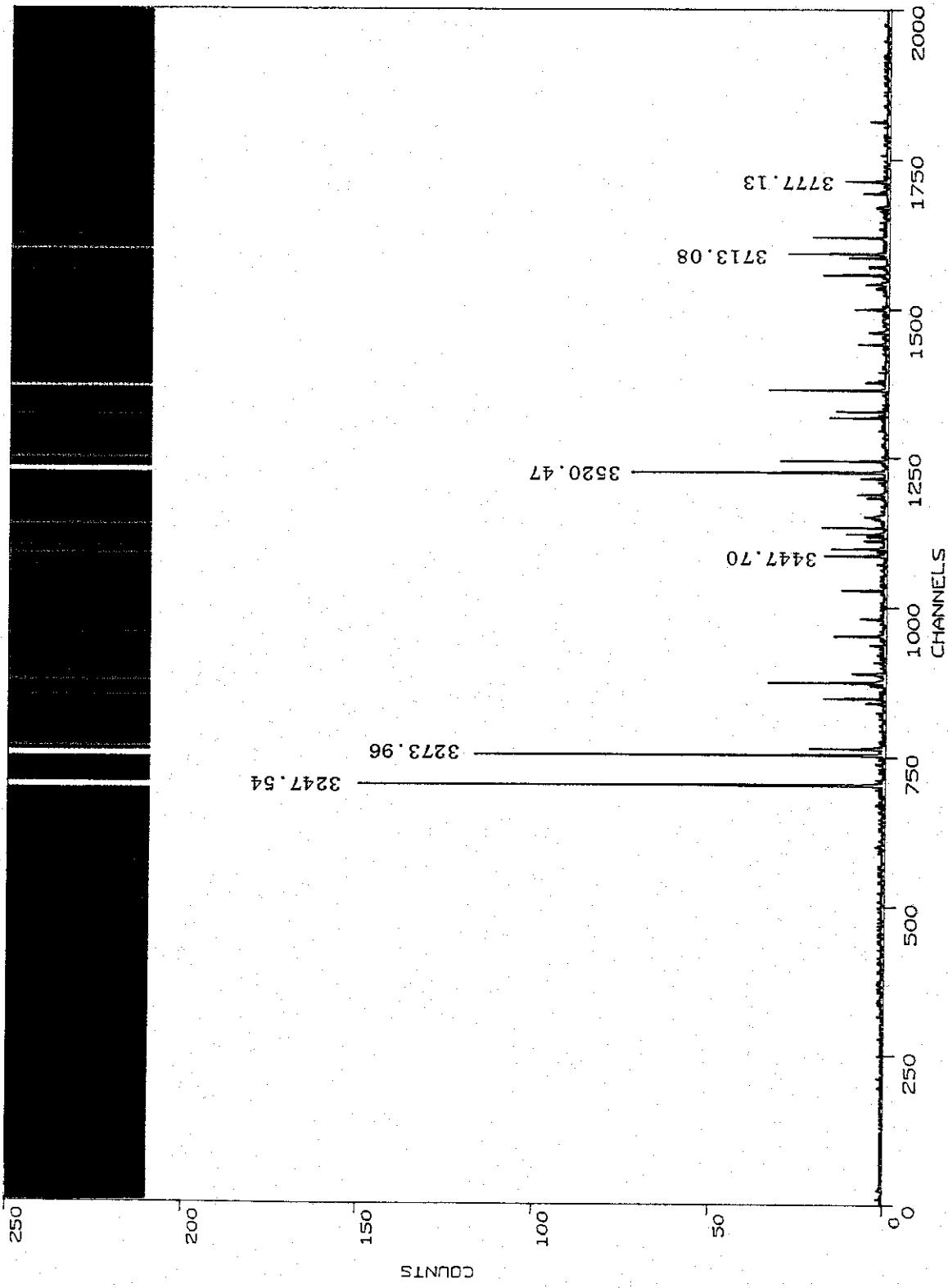


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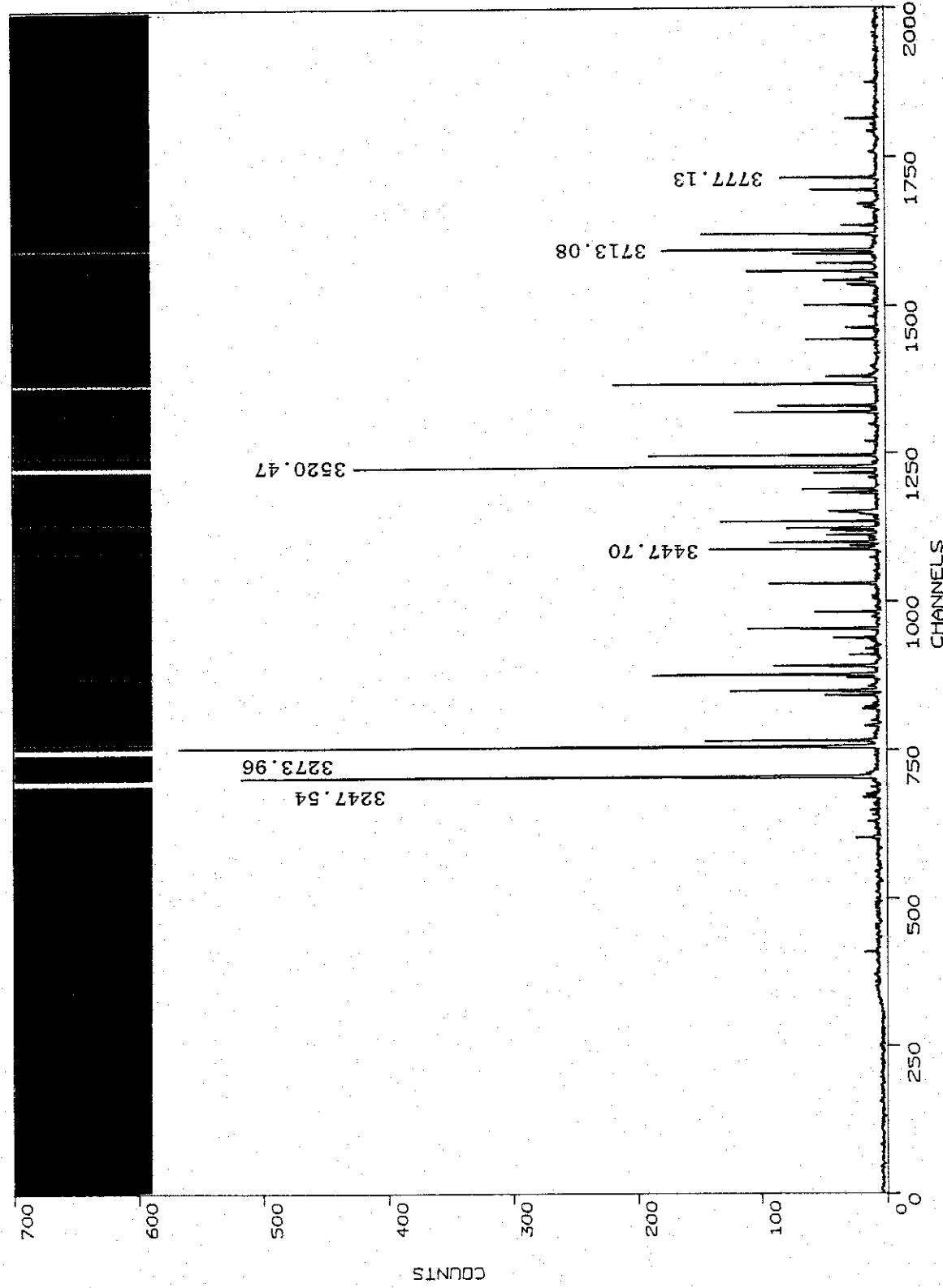


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 SOURCE CUNE

2 3 5 C A M E R A W I T H I P C S



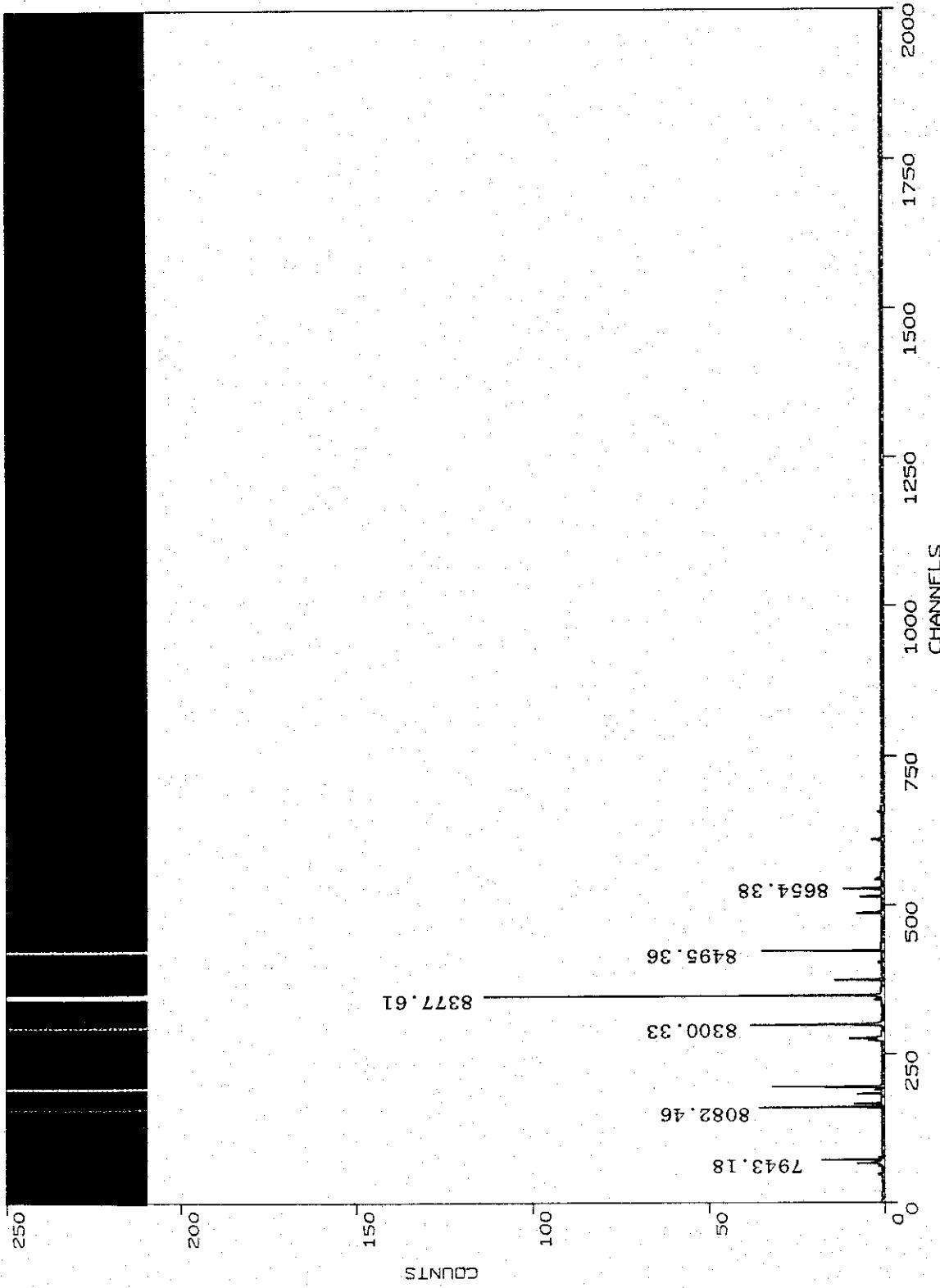
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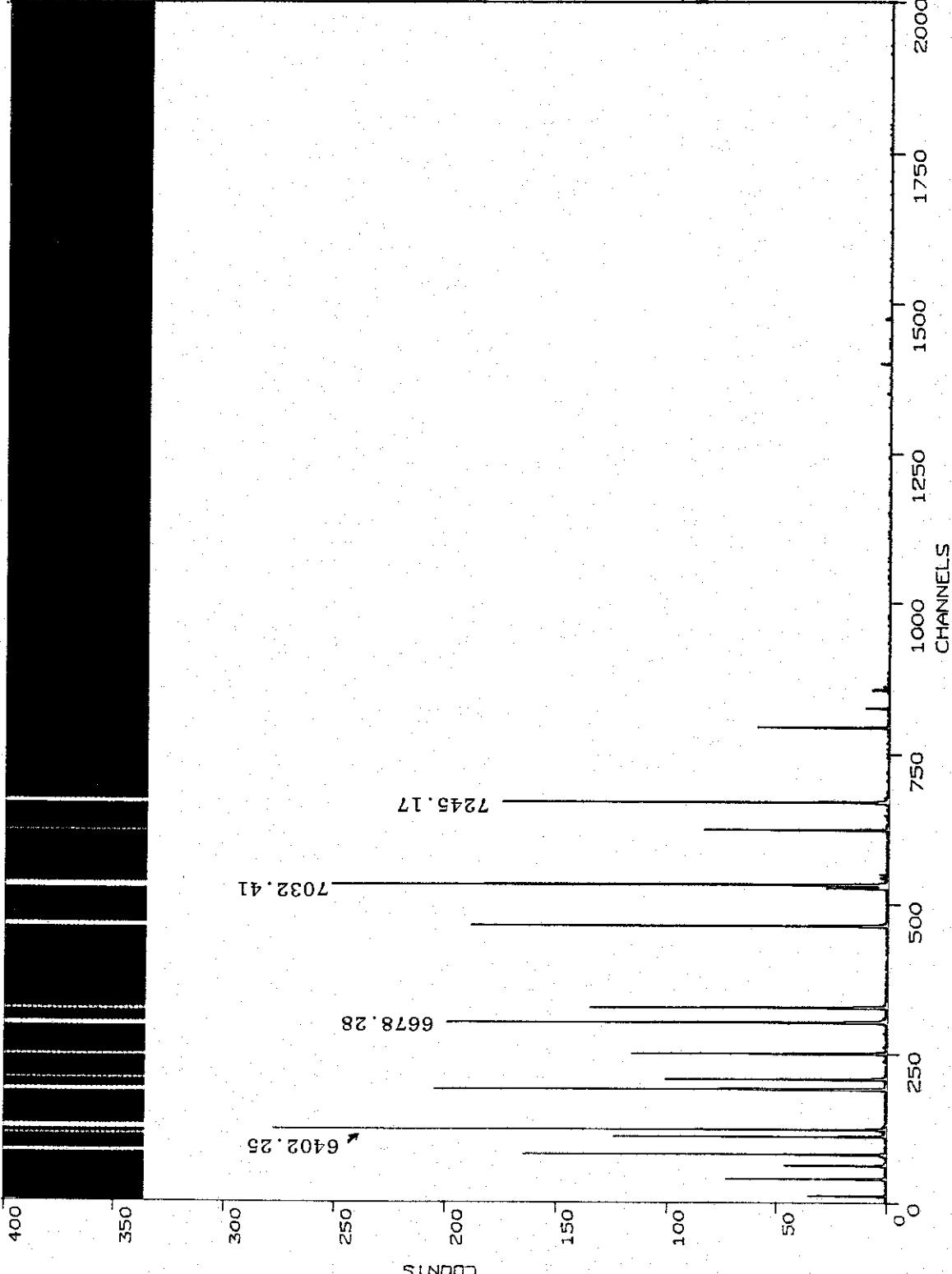
LOW DISPERSION CU-NE ARC MAPS 8650-4000

**CU-NE
IPCS
235
R 400V**

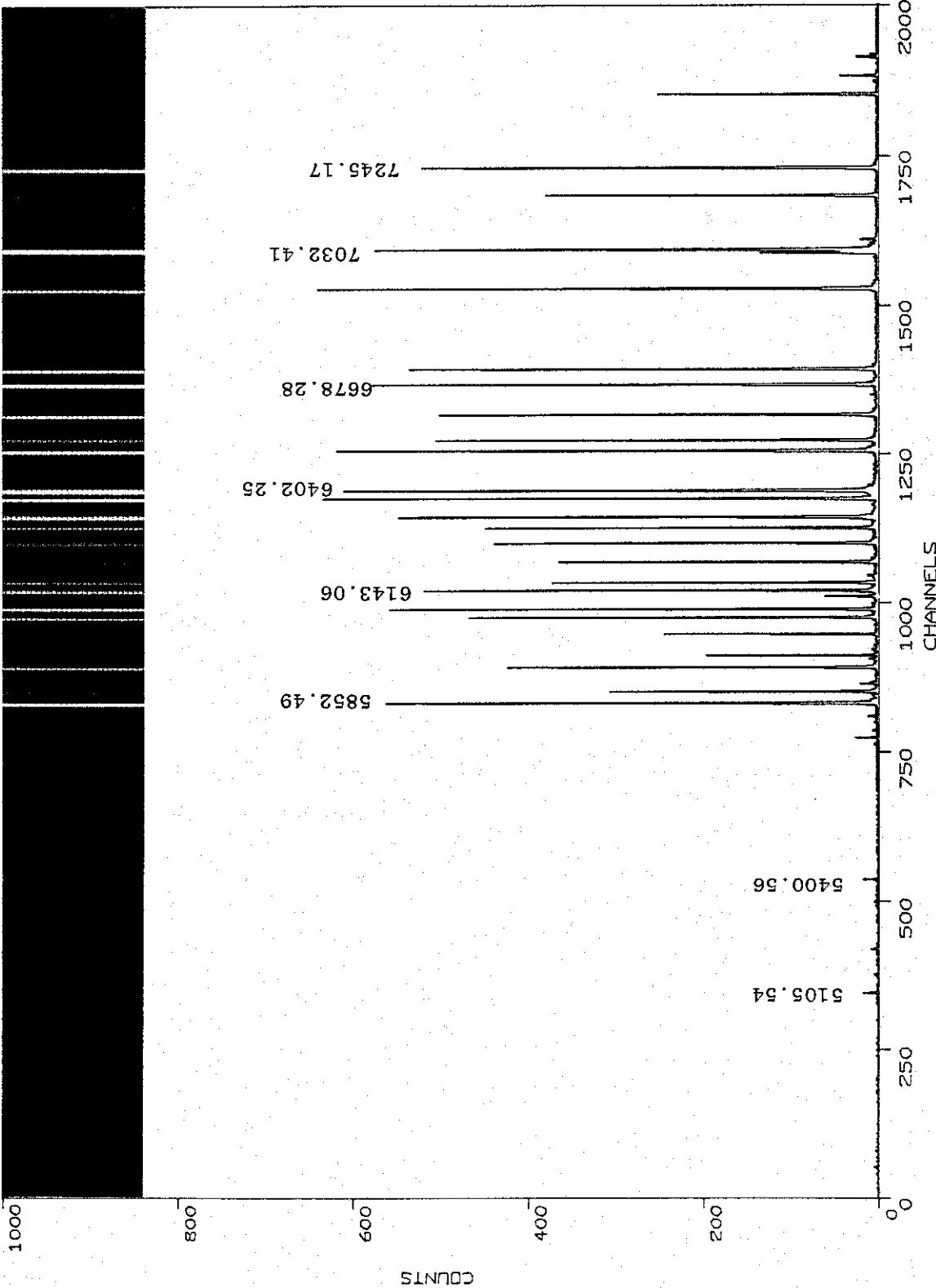
235 CAMERA WITH IPCCS



235 CAMERA WITH IPCS



2 3 5 C A M E R A W I T H I P C S



235 CAMERA WITH TIPCS

250

