

# The EURONEAR Moving Object Detection System

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# Content

- ≈ Introduction - EURONEAR
- ≈ EURONEAR Discoveries
- ≈ The EURONEAR Moving Object Detection System
- ≈ Problems encountered
- ≈ Future works
- ≈ Q&A

# Introduction - EURONEAR

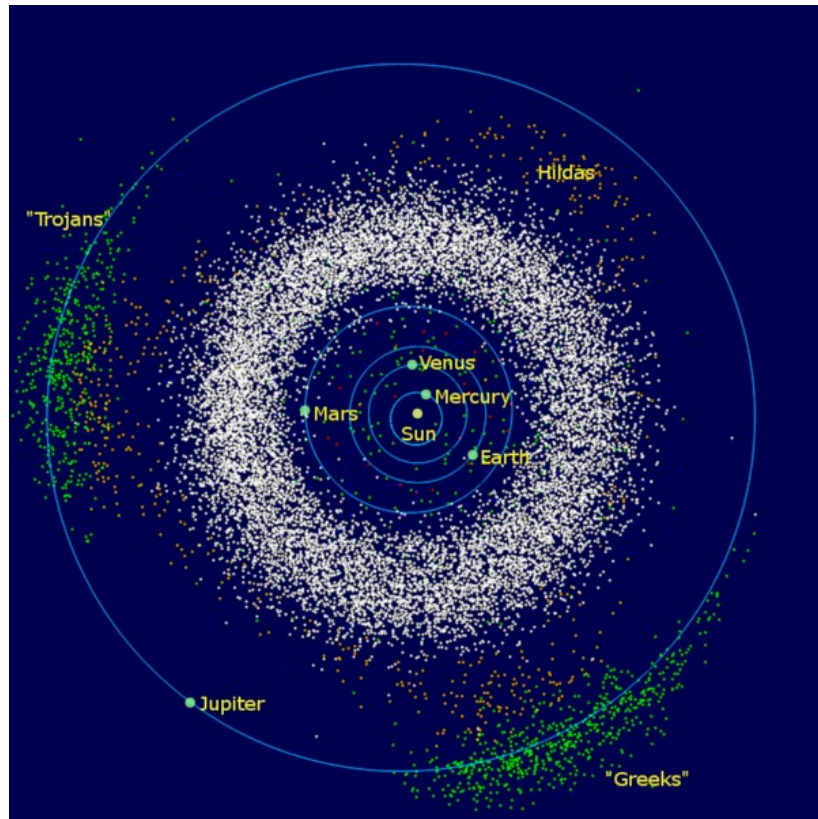


## The **EURO**pean **Near Earth Asteroid Research**

- ≈ A project aiming to increase the European contribution in the Near Earth Asteroids (NEA) research;
- ≈ Born in 2006 at IMCCE Paris (by O. Vaduvescu and M. Birlan);
- ≈ Including people from 14 institutions in 8 European countries plus Chile;
- ≈ Proposing to improve orbital and physical properties of NEAs using a network of existing telescopes or maybe in the future some dedicated infrastructure (understood our logo?)
- ≈ A collaborative project contributing to education and public outreach, involving amateurs and students collaborators as reducers, discoverers and paper co-authors.

# Introduction - EURONEAR

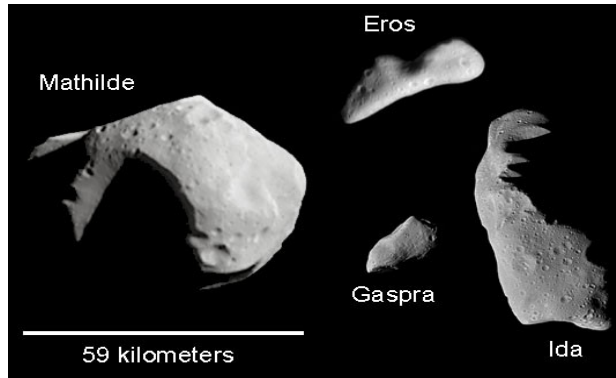
## The inner Solar System and the asteroids



- ≈ January 2017:
  - ≈ More than ...
    - ≈ 750,000 Main Belt Asteroids (MBAs)
    - ≈ 15,000 Near Earth Asteroids (NEAs)
    - ≈ 1,750 PHAs (Potentially Hazardous)
    - ≈ 500-600 Virtual Impactors (VIs)
  - ≈ ... known objects (Minor Planet Center, MPC)

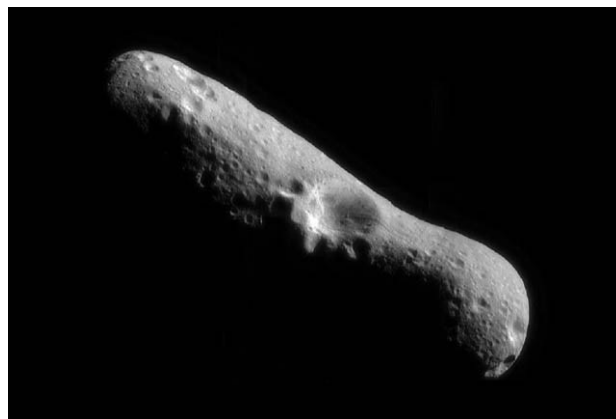
# Introduction - EURONEAR

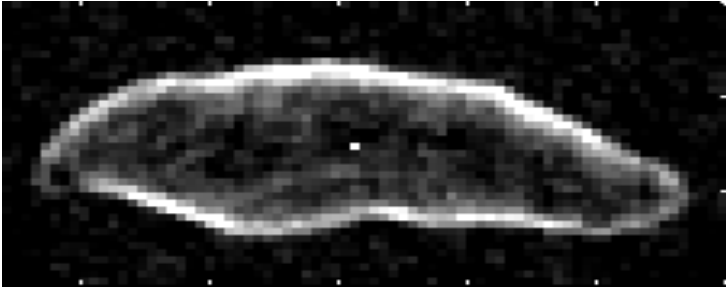
Only few asteroids & NEAs were visited by space missions:



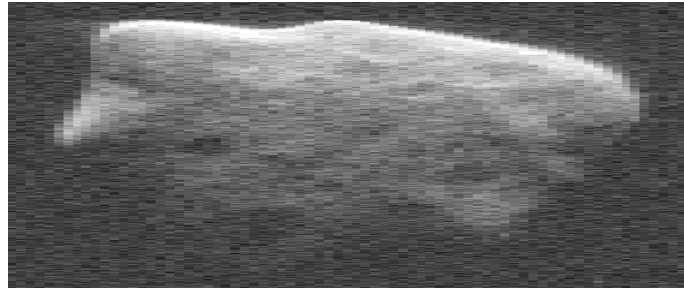
(951) Gaspra – Galileo 1991  
(243) Ida – Galileo 1993  
(253) Mathilde – NEAR 1997

NEA (433) Eros – NEAR 2000



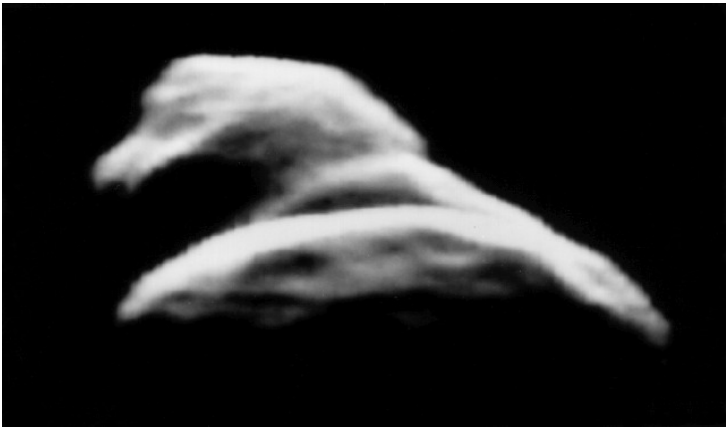


NEA (1620) Geographos – Goldstone 1994

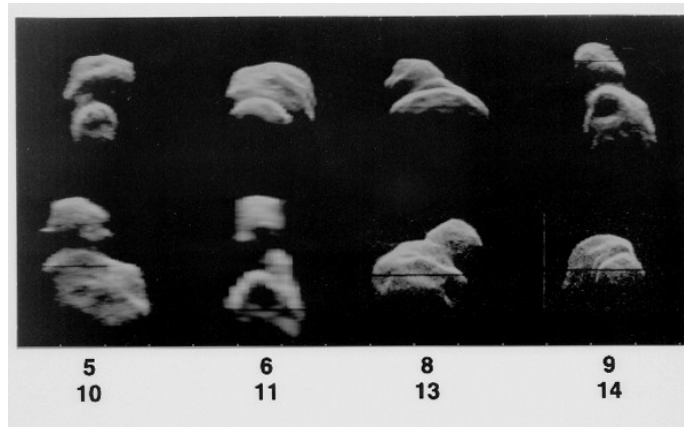


PHA (53319) 1999 JM8 – Goldstone 1999

And a few hundred NEAs were observed by radar



PHA (4179) Toutatis – Goldstone 1992 (~5km size, one of the most dangerous)



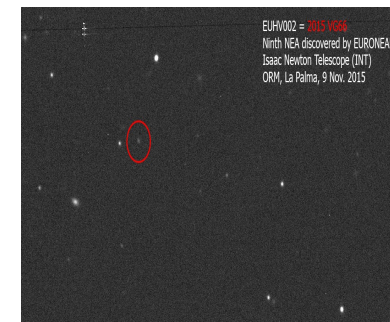
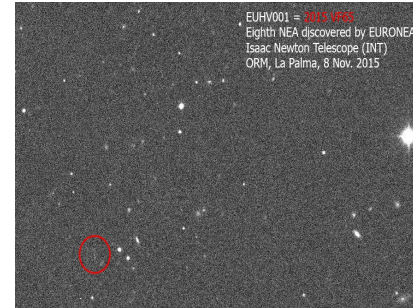
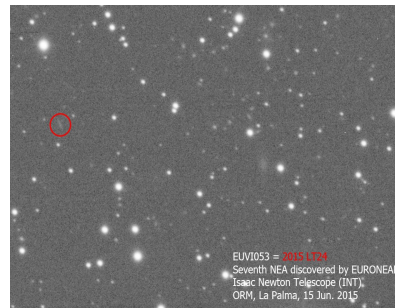
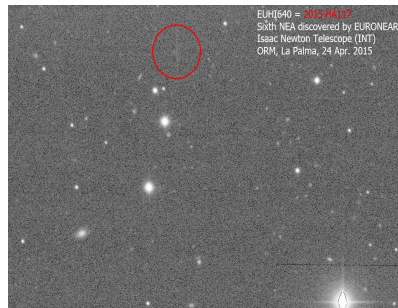
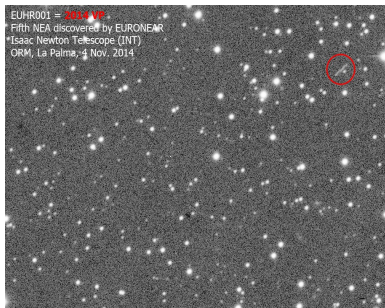
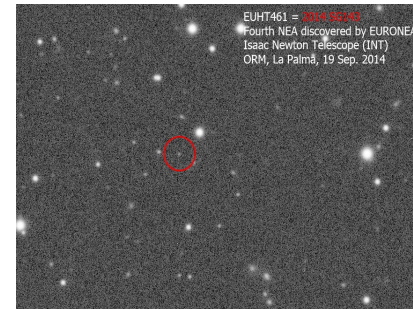
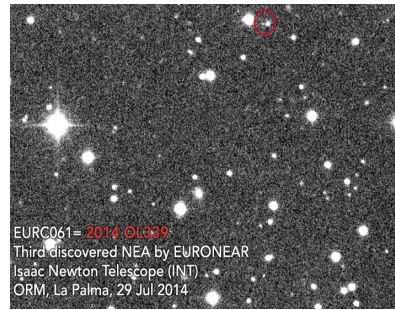
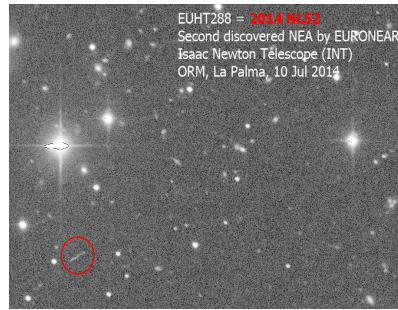
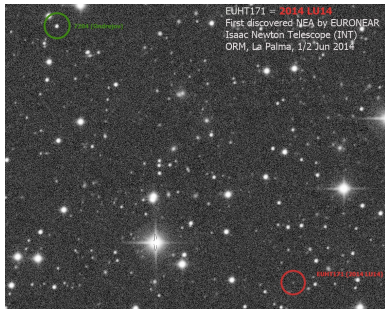


## Earth Impact Database

- ≈ (Univ. New Brunswick, Canada <http://www.passc.net/EarthImpactDatabase>)
  - ≈ We know about 160 craters due to cosmic collisions!



# EURONEAR Discoveries (INT & ESO/MPG)

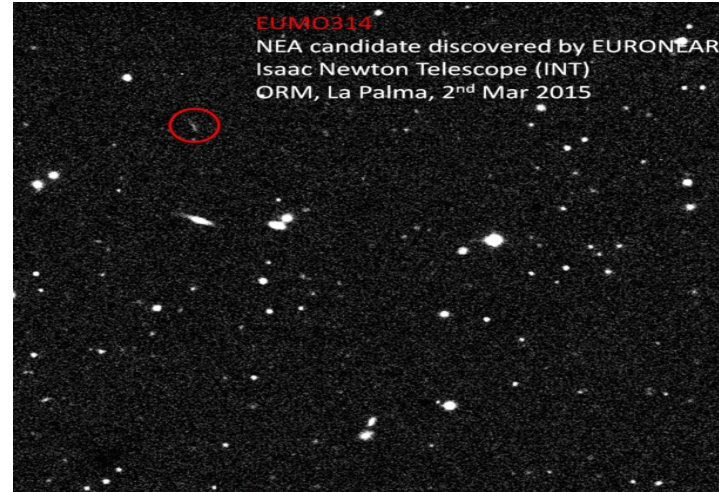
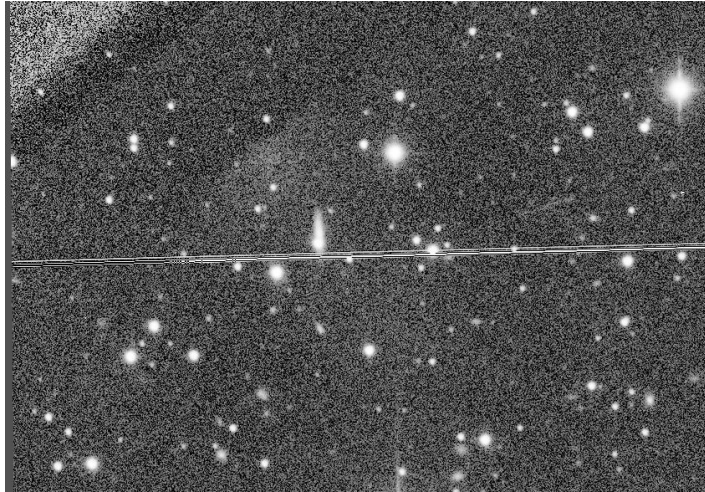


- ≈ Few hundreds MBAs (two mini-surveys);
- ≈ 9 secured serendipitous NEAs (the first in La Palma, during ToO time to recover known NEAs)
- ≈ Few lost NEAs (no follow-up time or other telescopes).



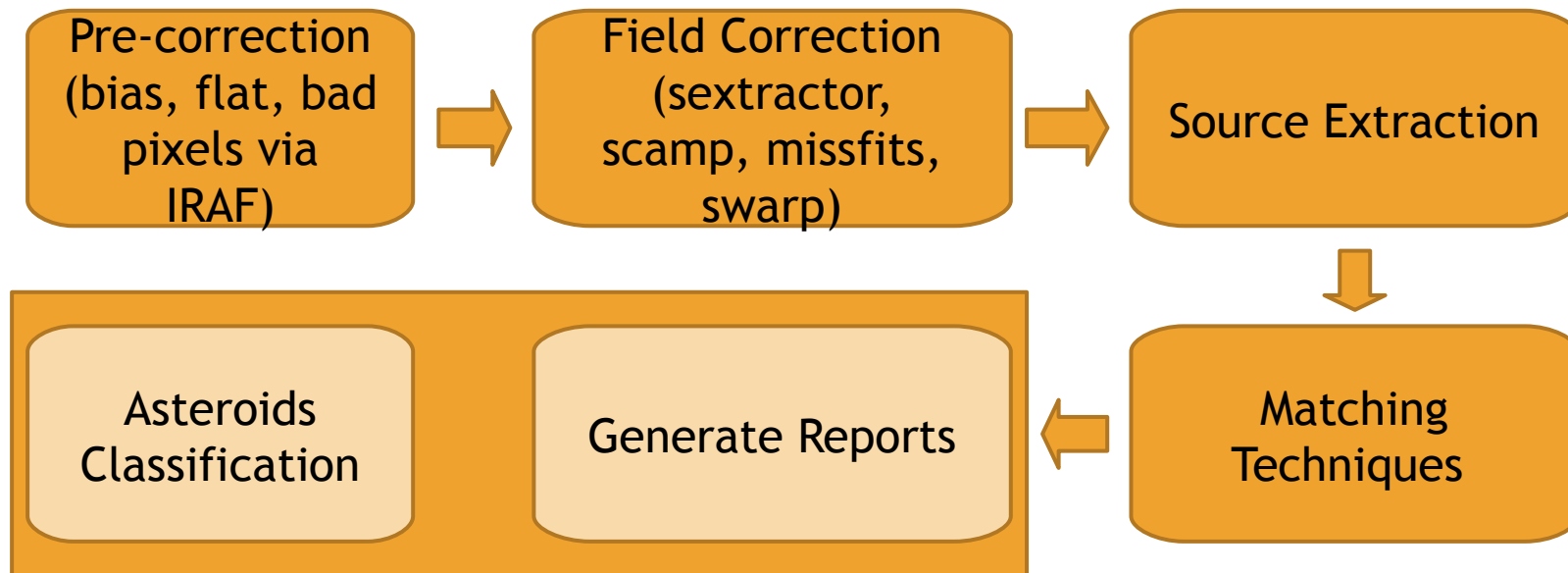
# EURONEAR Discoveries

## Found but Lost



# The EURONEAR Moving Object Detection System

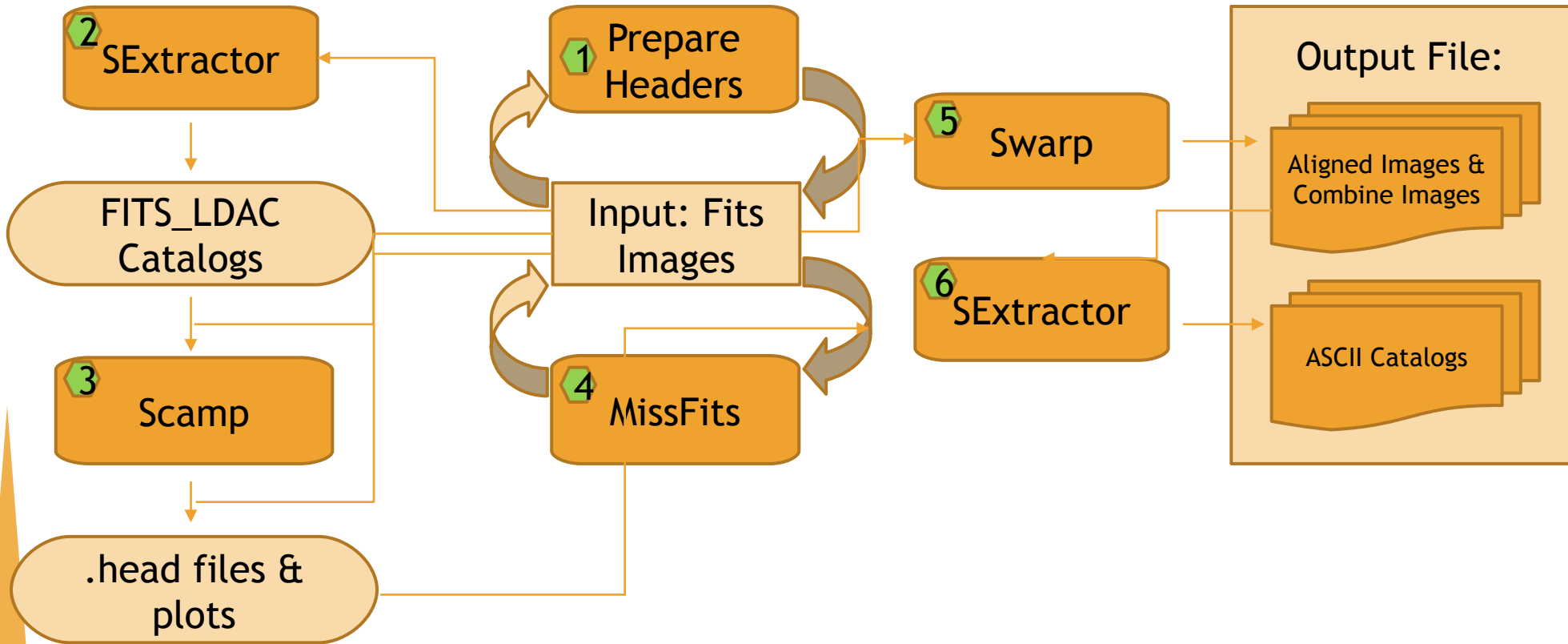
≈ a prototype pipeline written mainly in Python calling 3rdParty libraries for image pre-correction and correction



# Pre-correction Module

- ≈ Python calls to IRAF in order to apply:
  - ≈ Bad pixels
  - ≈ Bias
  - ≈ Flat field
- ≈ Input: Raw FITS Images
- ≈ Output: Reduced FITS Images

# Field Correction Module

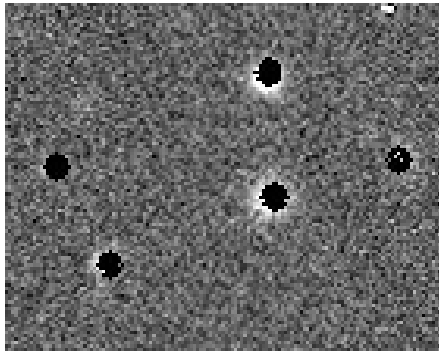


# Source Extraction Module

- ≡ Pixel Based

- ≡ (using star subtraction)

- ≡ Using IRAF - not good enough



- ≡ Can be improved and used if the operation is done with tools like OpenCV (dedicated tool for image processing)

- ≡ Space Objects Based

- ≡ (using sources in catalogs)

- ≡ Remove all fixed objects from individual catalogs that match sources from the combined catalog (using the combined image)

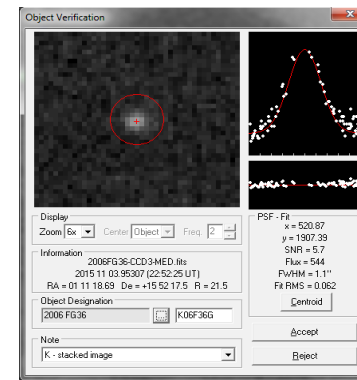
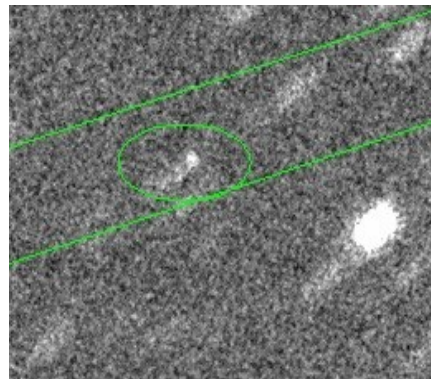
- ≡ The remaining objects are asteroids or noise, and should be paired in the individual images



# Matching Module

## Two Detection Techniques

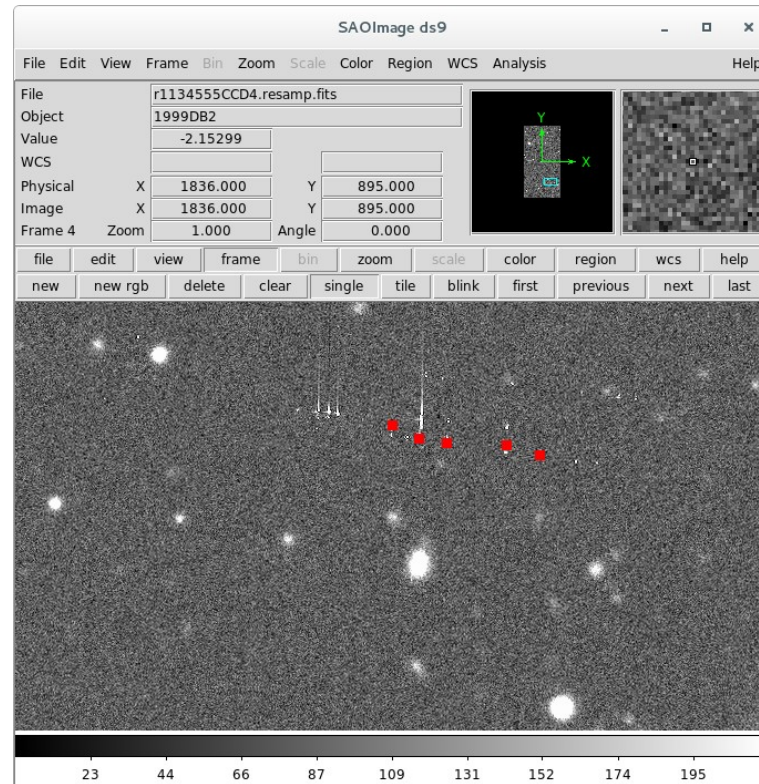
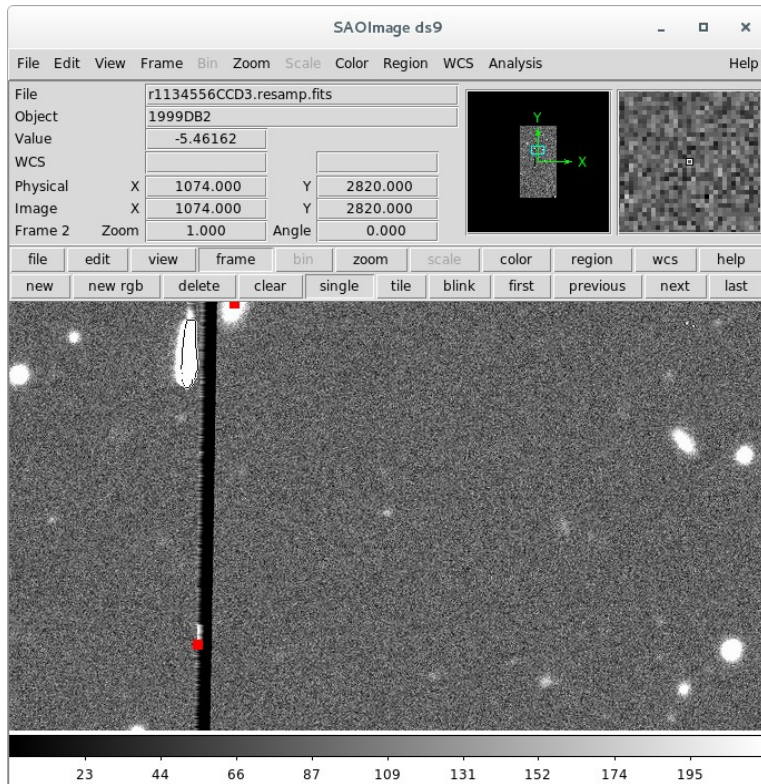
- ≈ Improved “Blink“ Algorithm
- ≈ in series of a few (4-5) images, pairing (at least 3) sources moving linearly and rejecting noise
- ≈ “Synthetic (digital) tracking”
- ≈ using series of many (dozens) of images, in the search of unknown objects moving in any direction and with any proper motion



### Track & Stack

Detection of very faint known NEA with INT from 6 images

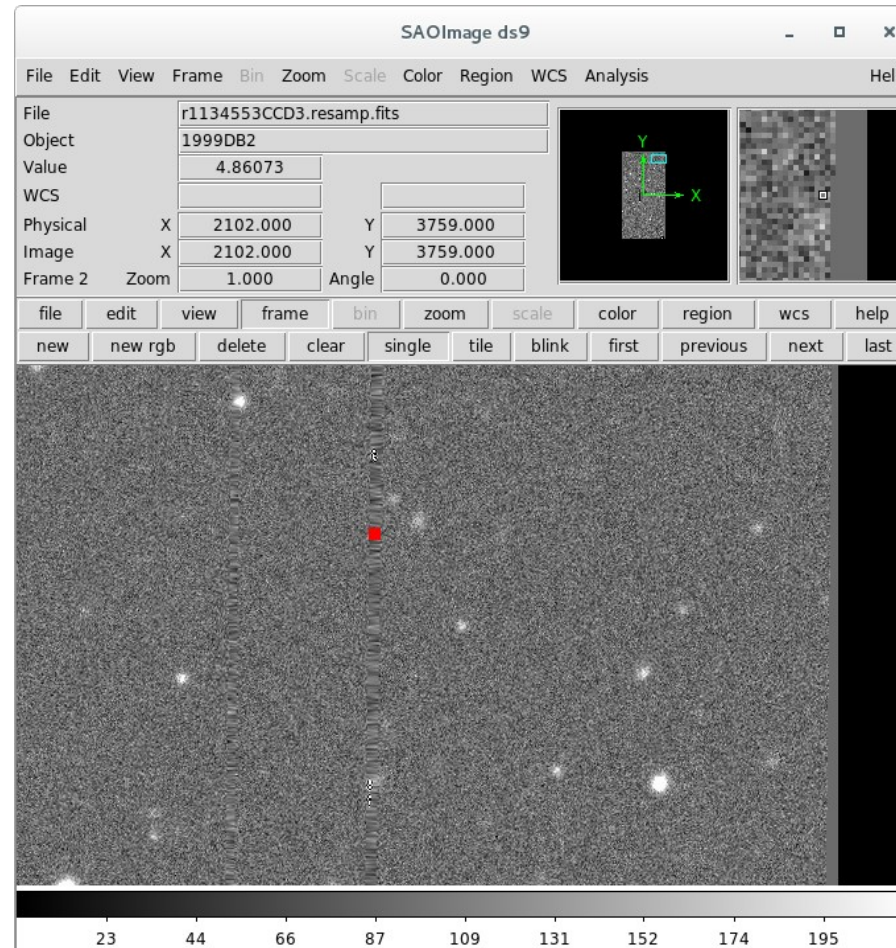
# Problems encountered



Caused by bad pixels

# Problems encountered

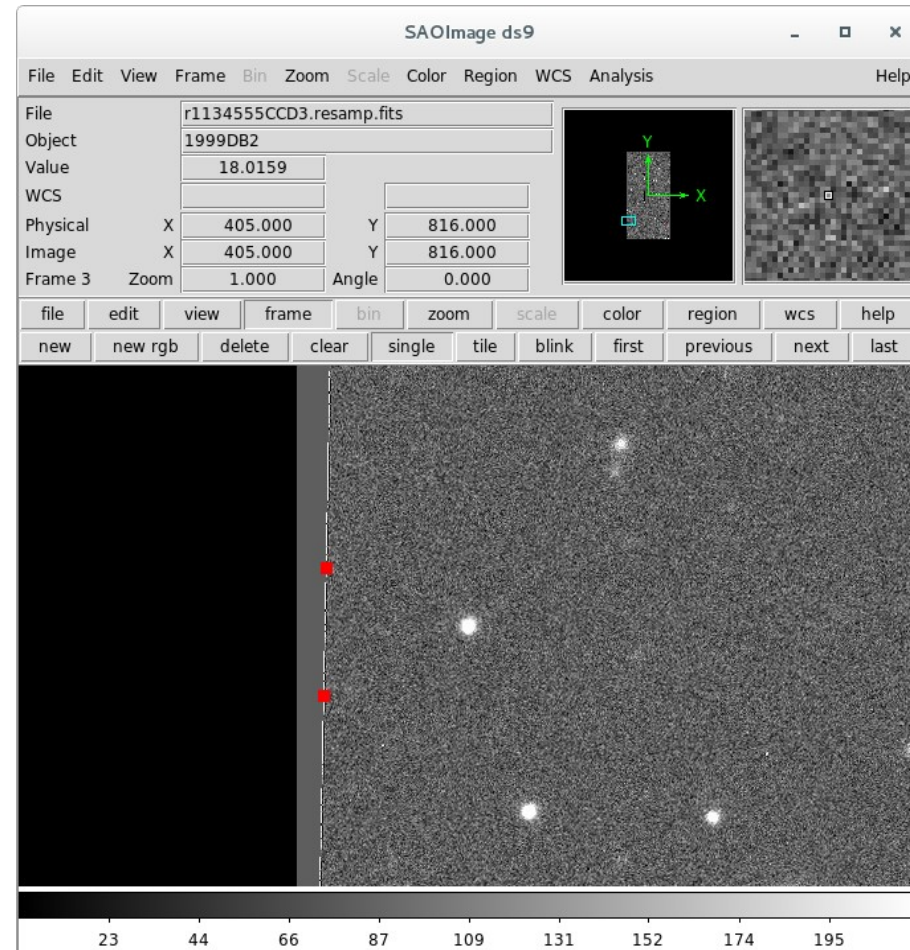
Caused by interpolation





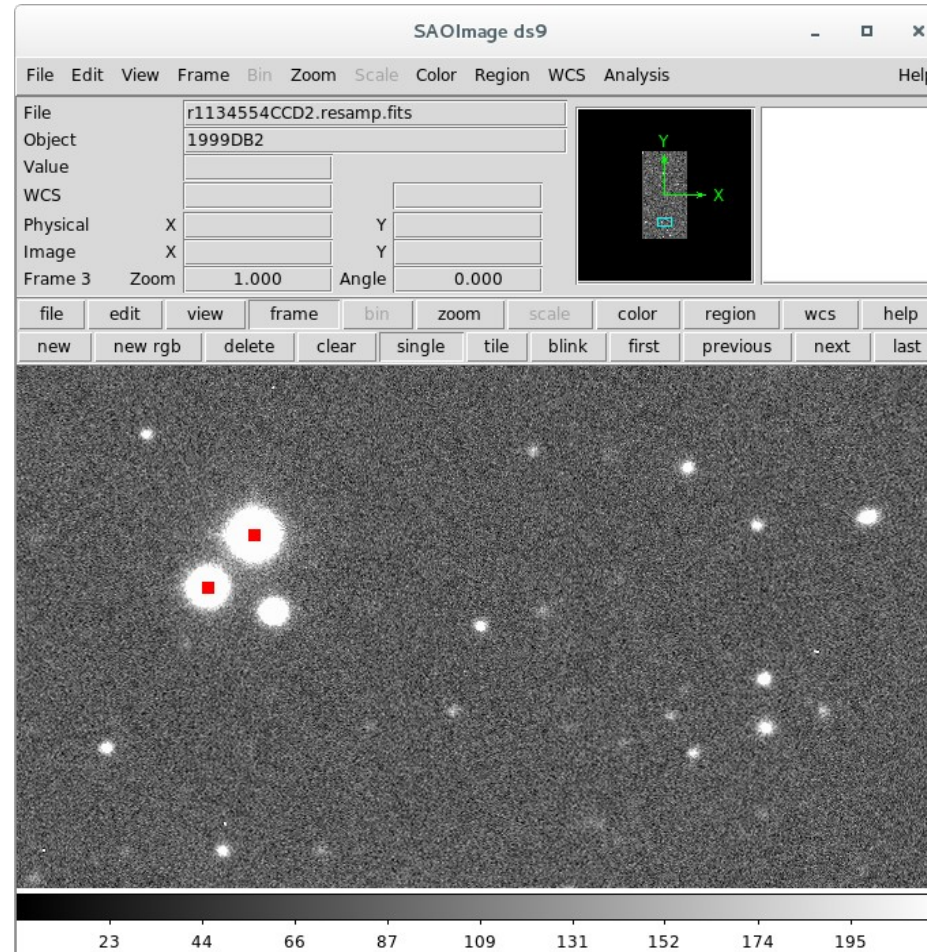
# Problems encountered

Caused by image registration



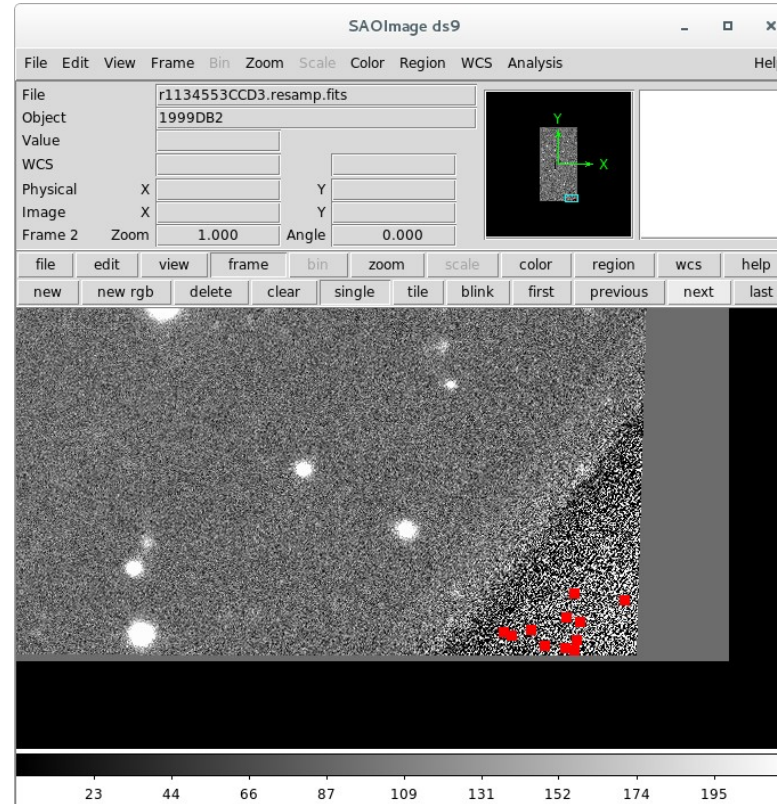
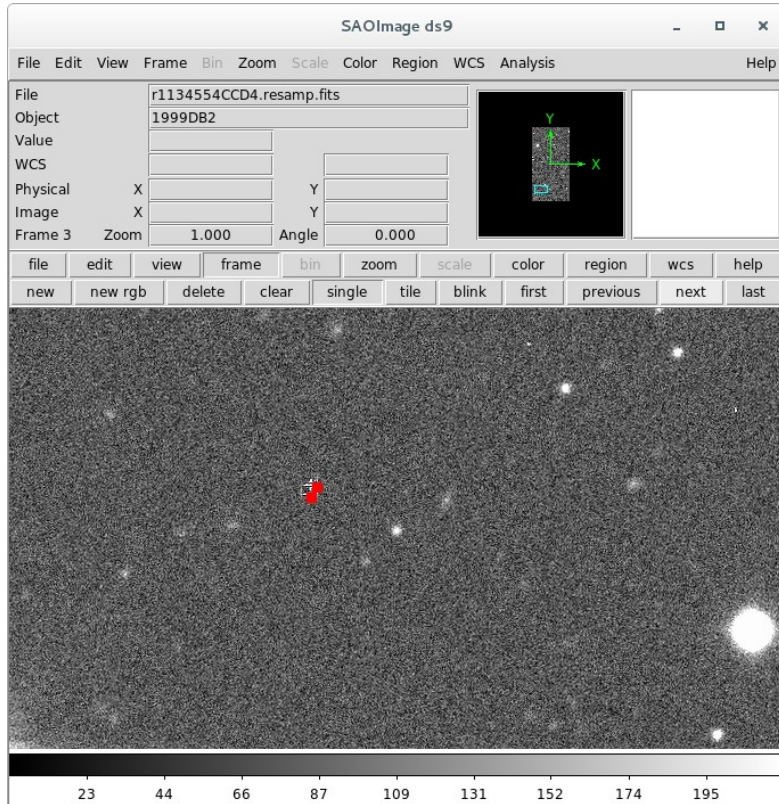
# Problems encountered

Caused by star saturation





# Problems encountered

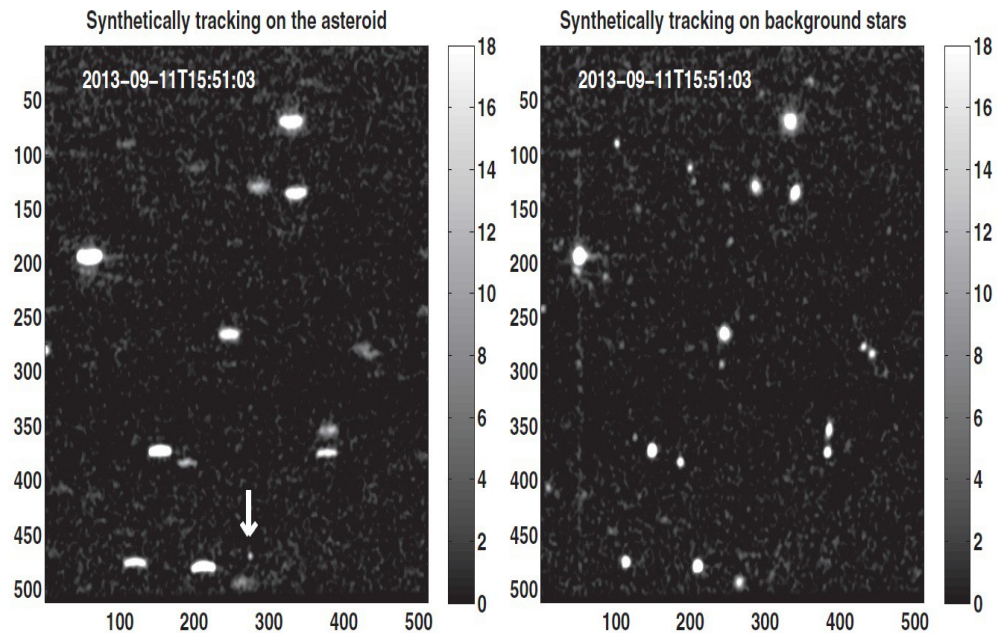


Caused by noise

# Future works

## ≡ Astronomy:

- ≡ Complete the latest “blink” modules
- ≡ Test the pipeline using INT-WFC archive images and a future mini-survey;
- ≡ Implementing the “Digital (or Synthetic) Tracking” method, assisted by modern computer infrastructure at two Romanian universities.



## References

- Tyson, J. A., Guhathakurta, P., Bernstein, G. M., & Hut, P. (1992, Septembrie). Limits on the Surface Density of Faint Kuiper Belt Objects. *American Astronomical Society*.
- Gladman, B., & Kavelaars, J. J. (1997, Ianuarie). Kuiper Belt searches from the Palomar 5-m telescope. *Astronomy and Astrophysics*.
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- Zhai et al., (2014) - Detection of a Faint Fast-moving Near-Earth Asteroid Using the Synthetic Tracking Technique, *Apj* 792, p. 60
- Heinze, A. N., Metchev, S., & Trollo, J. (2015). Digital Tracking Observations Can Discover Asteroids Ten Times Fainter than Conventional Searches. *The Astronomical Journal*, 51.

# Future works

## ≈ Computer Science:

- ≈ Migrate modules to cloud as Docker Containers or as Virtual Machine
- ≈ Replace some of the modules with OpenCV operations/algorithms
  - ≈ After that it can become even more accelerated with the GPU and using hardware accelerators tools such as CUDA or OpenCL (provided by NVIDIA)

# Q&A

☞ Thank you for your attention!!!