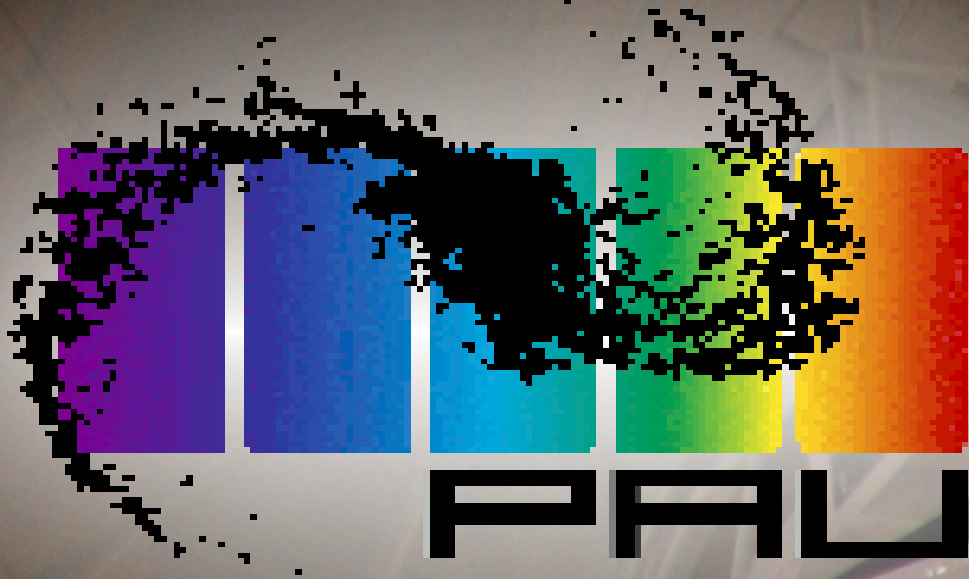
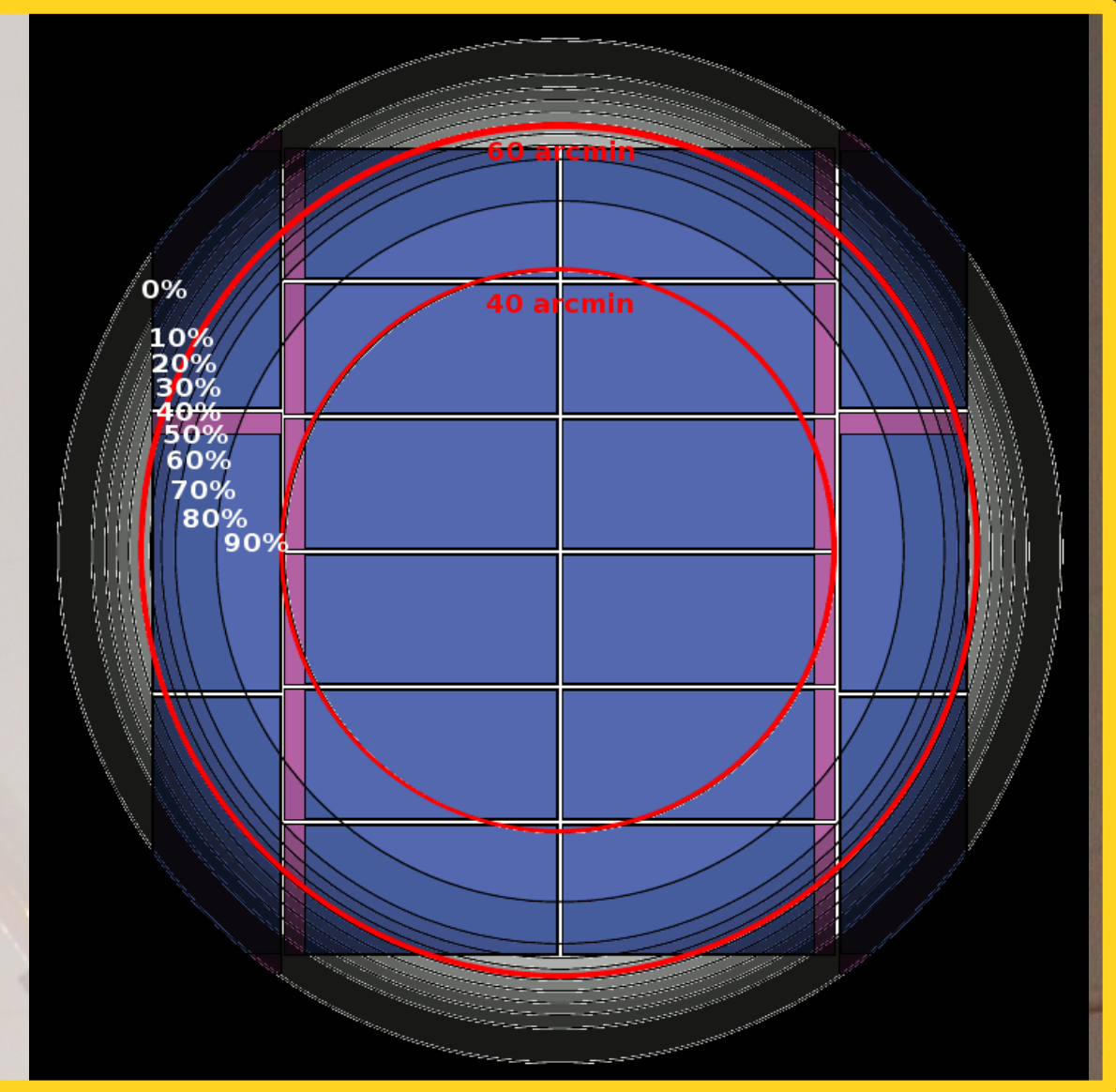
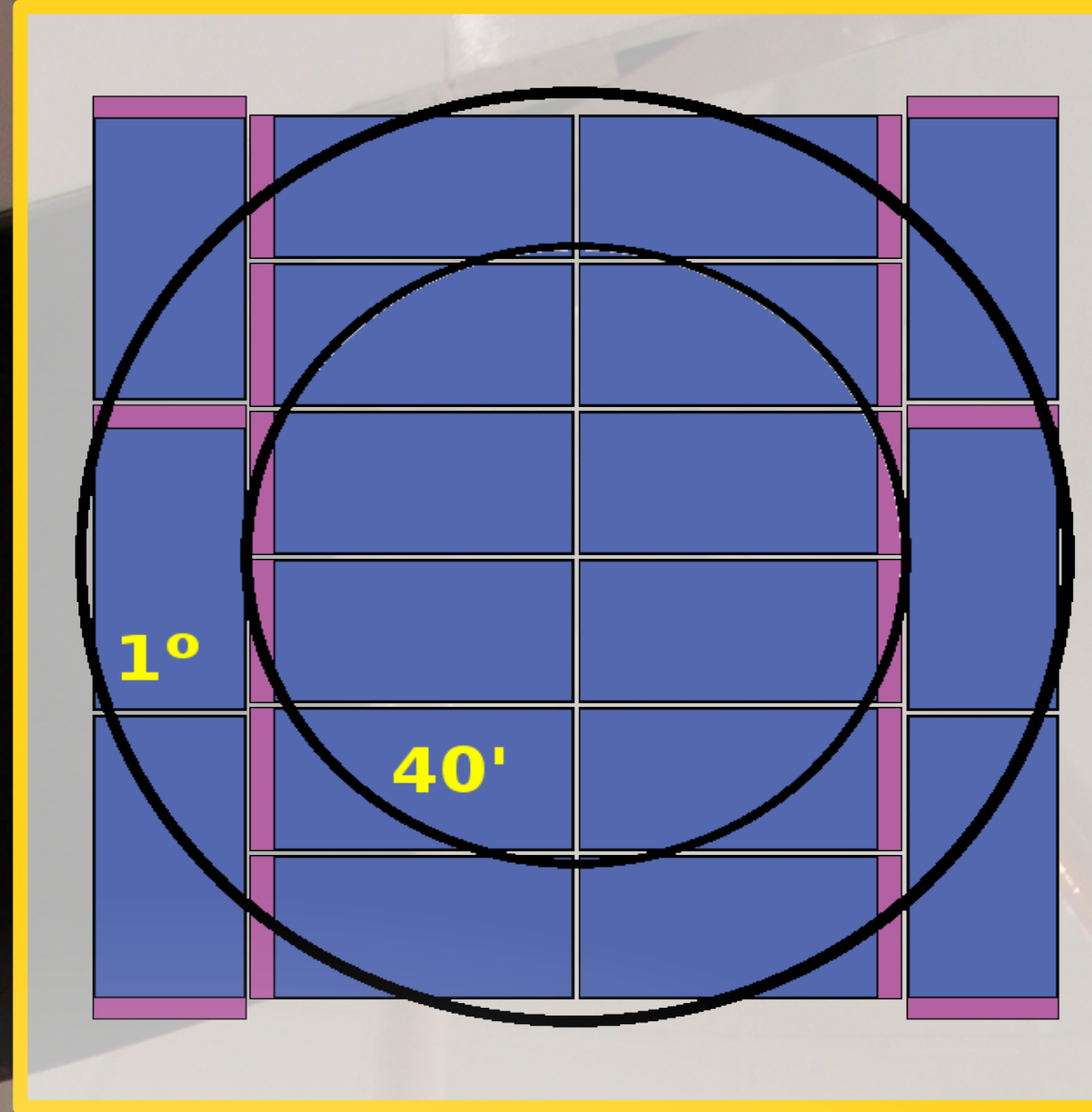
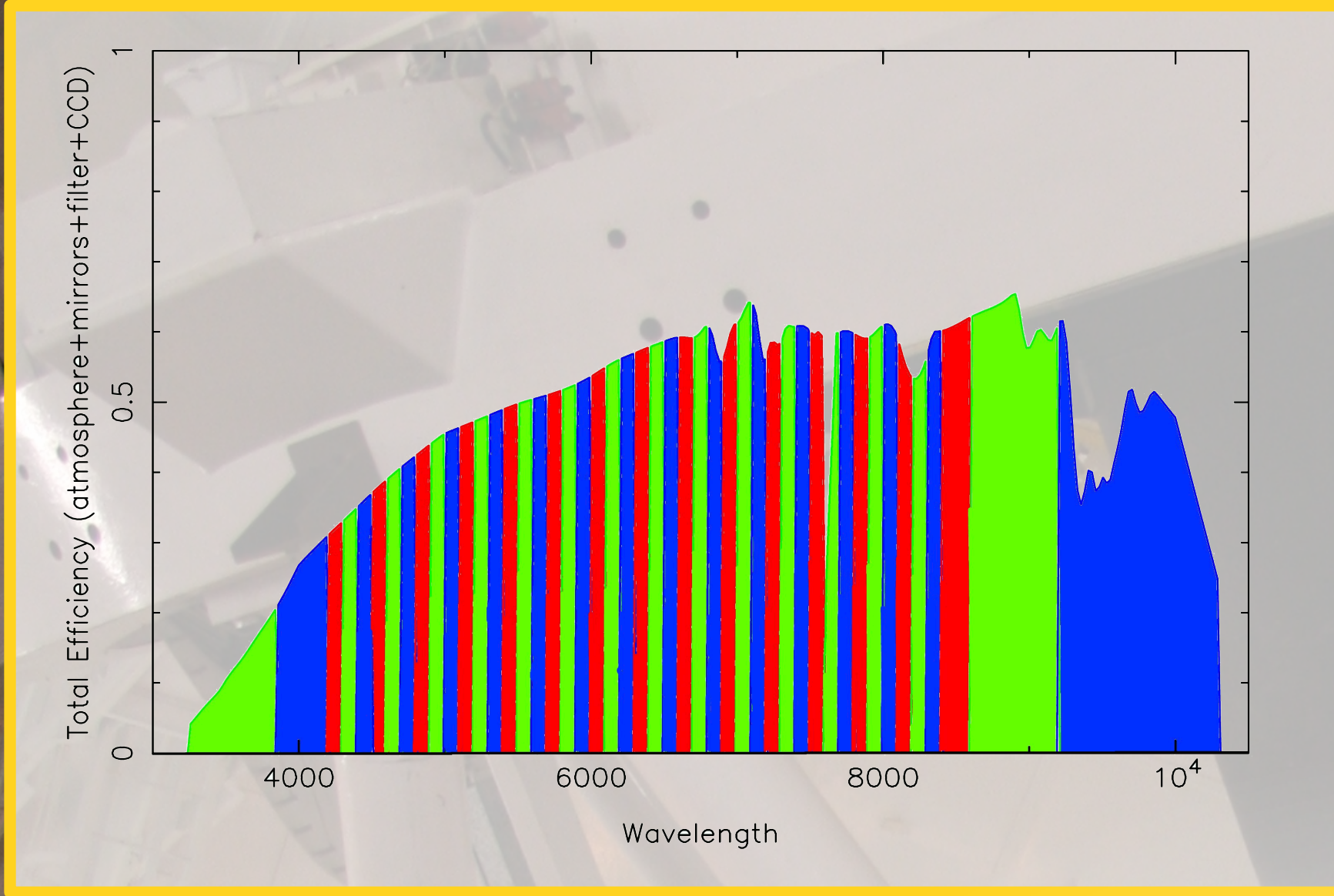


# The PAU Camera

F. J. Castander(1), O. Ballester(2), L. Cardiel(2), J. Carretero(1), R. Casas(1), J. Castilla(3), M. Croce(1), J. de Vicente(3), M. Delfino(4), E. Fernández(2), P. Fosalba(1), J. García-Bellido(5), E. Gaztañaga(1), F. Grañera(2), F. Madrid(1), P. Martí(2), R. Miquel(2), Ch. Neissner(4), E. Sánchez(3), S. Serrano(1), I. Sevilla(3), R. Ponce(3)  
 (1)ICE (IEEC-CSIC), (2)IFAE, (3)CIEMAT, (4)PIC, (5) UAM



The PAU (Physics of the Accelerating Universe) collaboration is building an instrument, intended for the William Herschel Telescope prime focus, designed to perform a large area survey for cosmological studies.

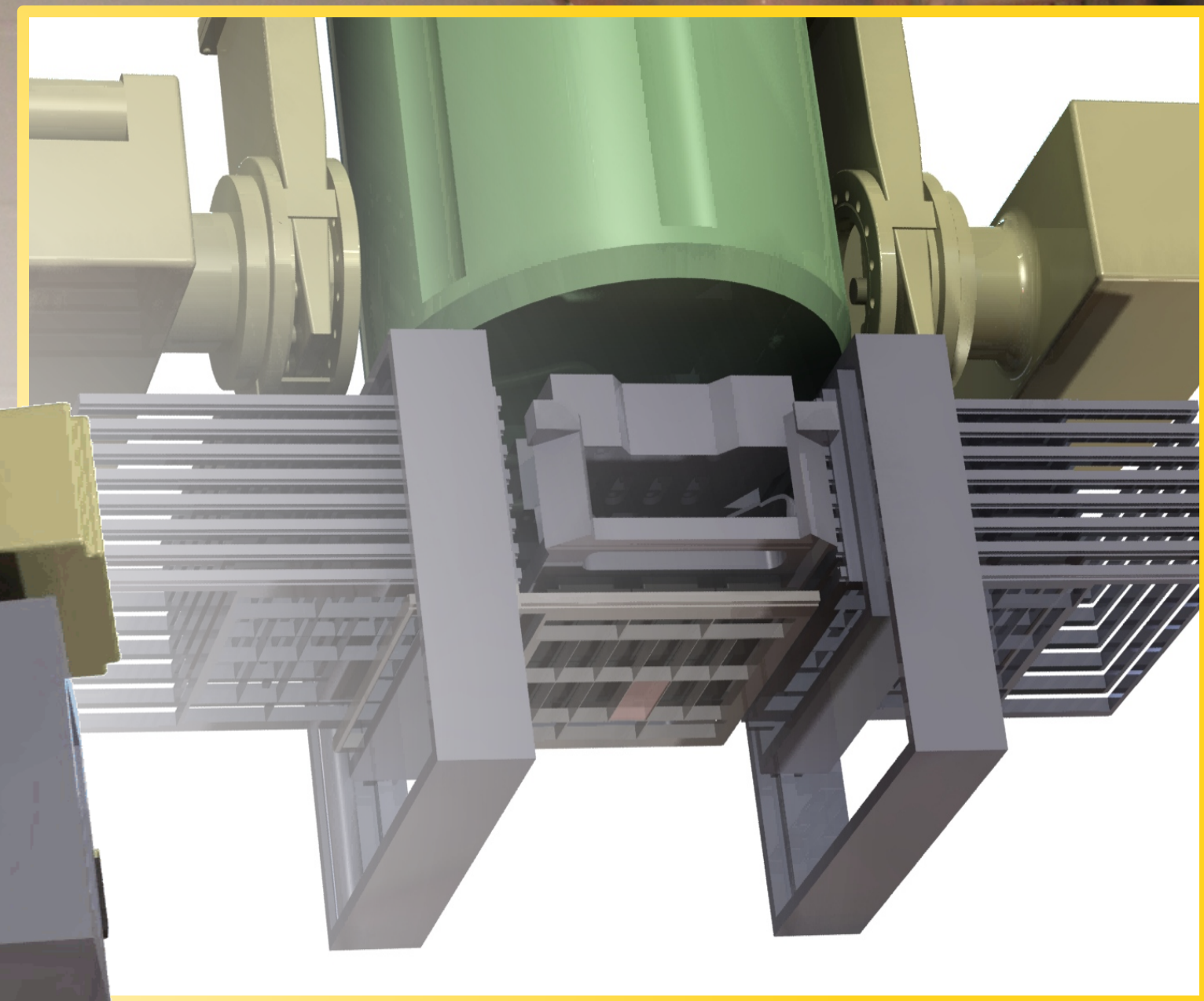
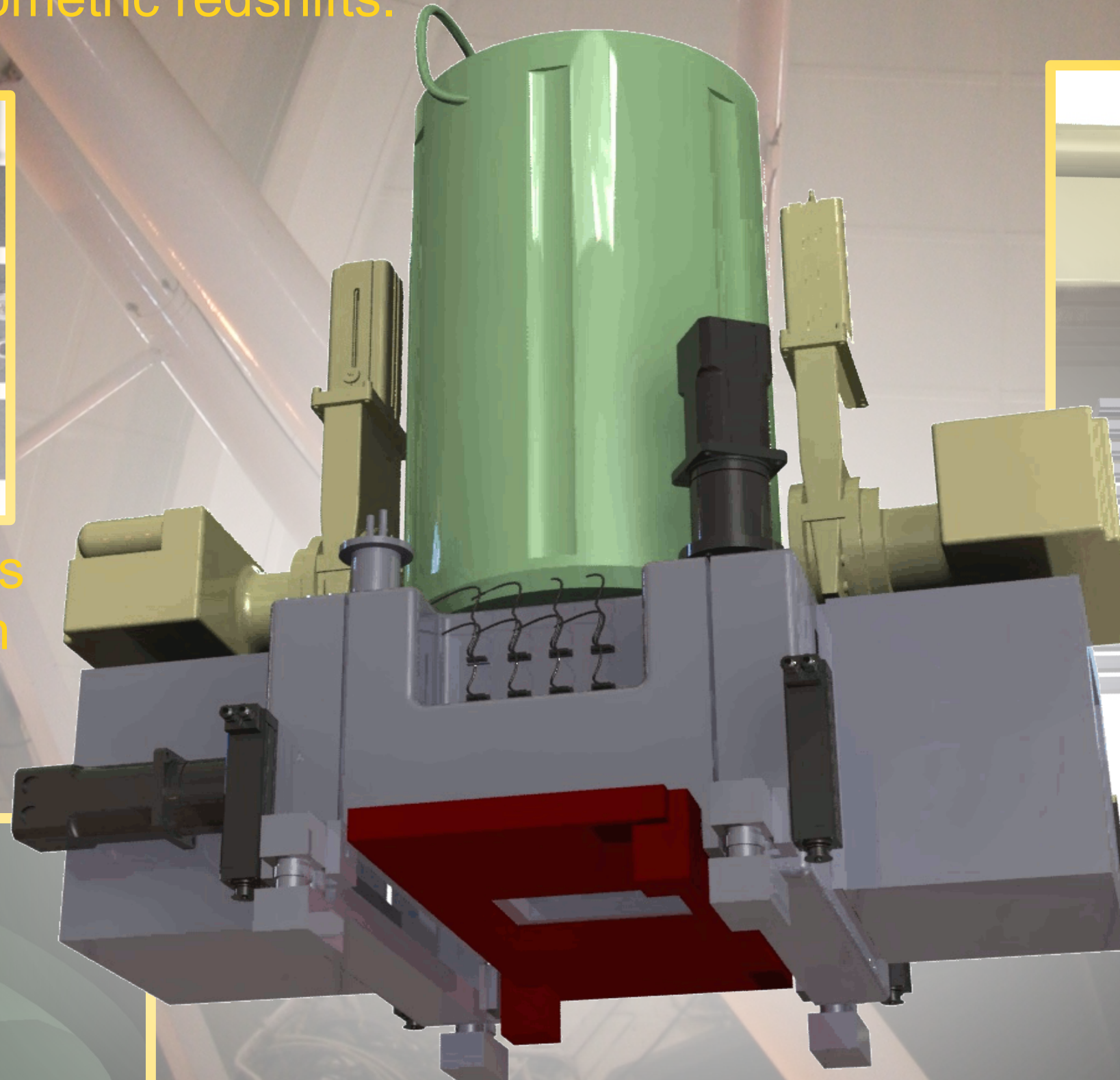


Narrow band imaging: SEDs are sampled using ~40 filters ~100Å wide spanning the optical wavelength range to accurately measure photometric redshifts.

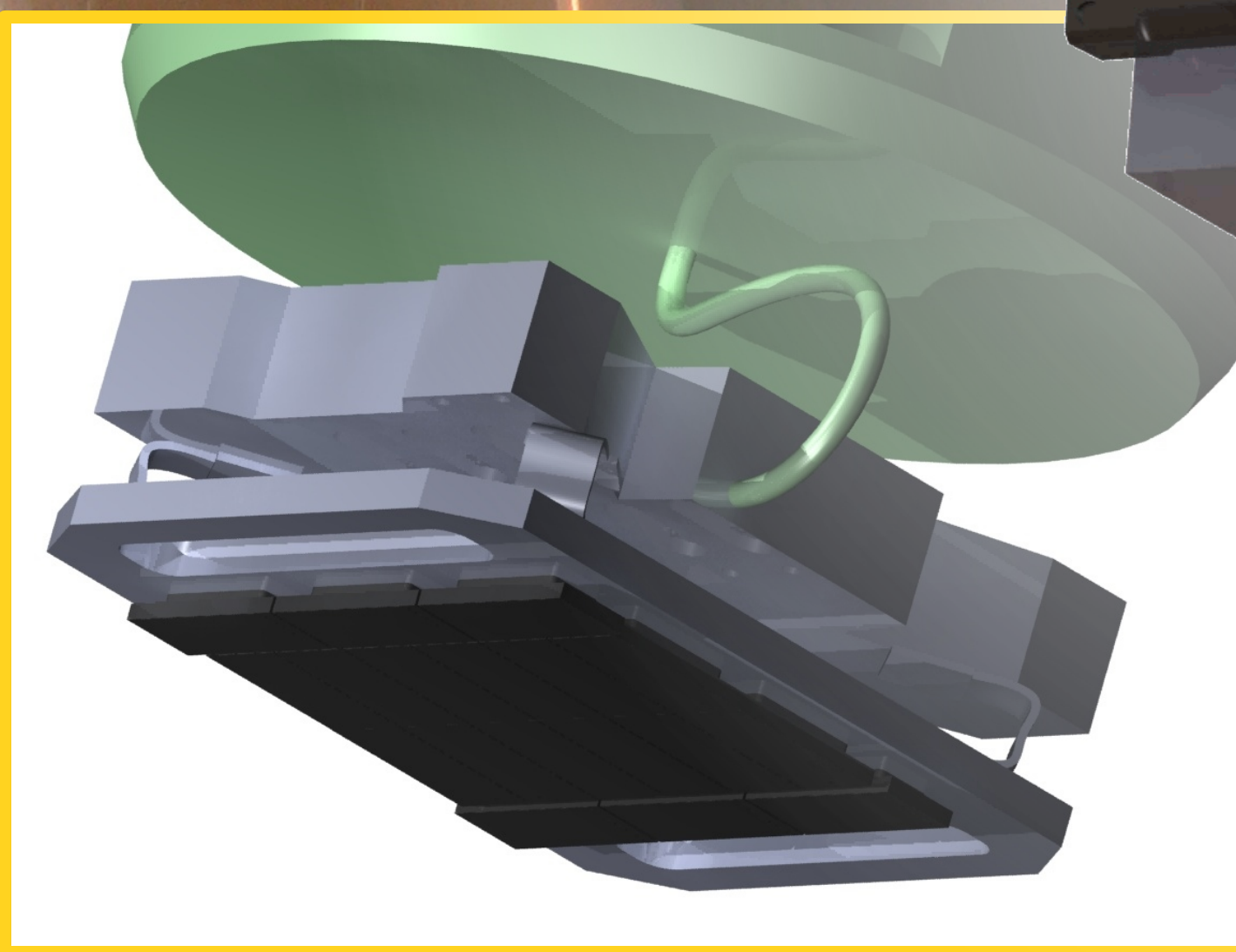
The FoV is 1 degree, densely populated by 18 state of the art CCDs. The maximum vignetting is 50% of the focal plane.



Low read-out noise of the CCDs is being prepared using the Monsoon architecture developed by NOAO.



The filters, placed as close as possible to the CCD detector on segmented exchangeable filter trays. An innovative jukebox-like exchanging mechanism inside the cryostat based on rolling hybrid bearings technology with tungsten disulfide as solid lubricant has been chosen to hold the movements inside the cryostat.



A pressurized liquid nitrogen tank feeding a boiler inside the cryostat. CCD's will be kept at 170K. Temperature homogeneity achieved with a coldplate: heat spreader and thermal capacitance smoothing the temperature ramps.

Two independent servomotors with absolute encoders outside the cryostat through magnetic feedthroughs will motion the system.



Testing setup for CCD characterization.