

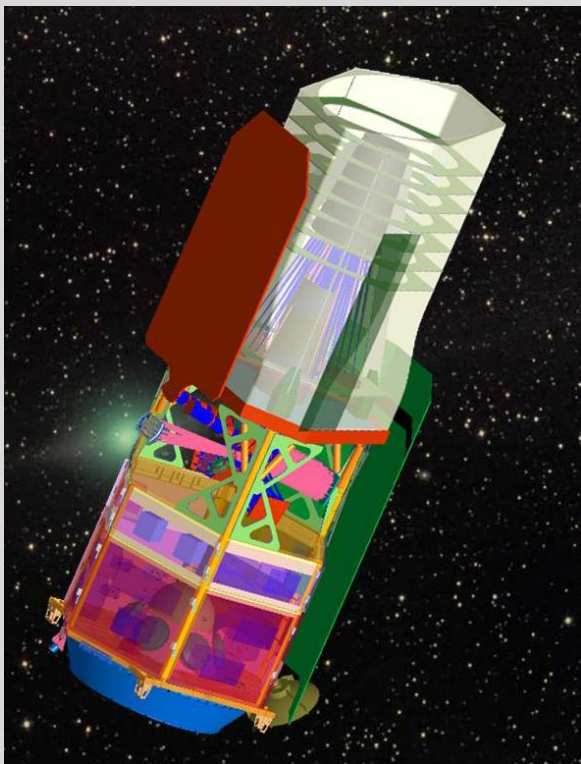
# NASA's Wide Field Infrared Survey Telescope (WFIRST)

Jason Rhodes (JPL/Caltech)

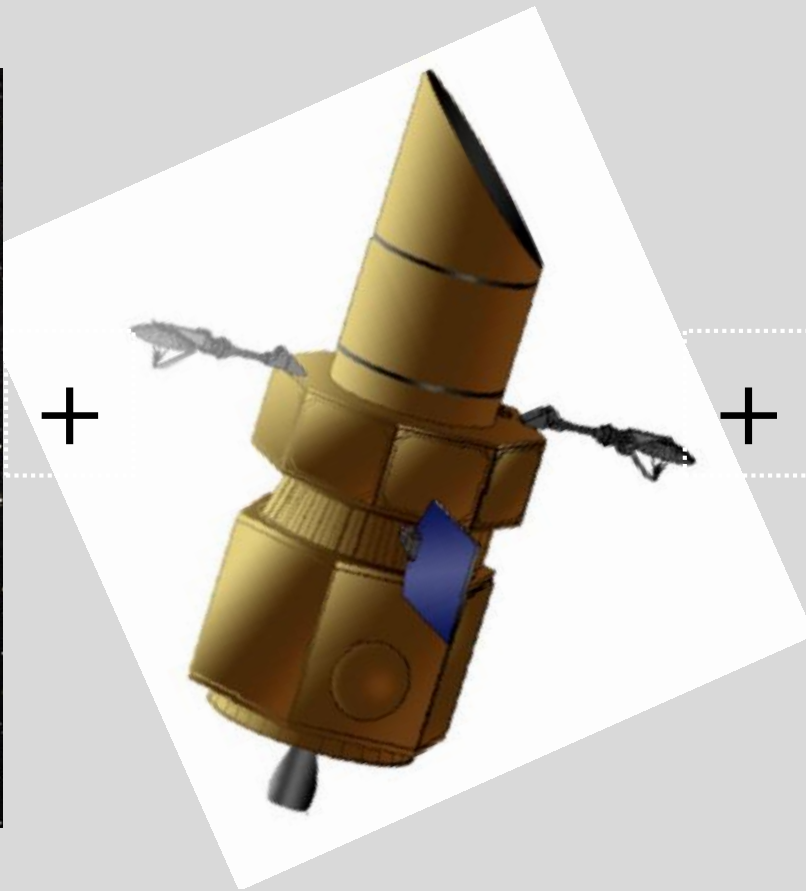
Multi-Object Spectroscopy in the Next Decade

March 5, 2015

# WFIRST =



JDEM-Ω

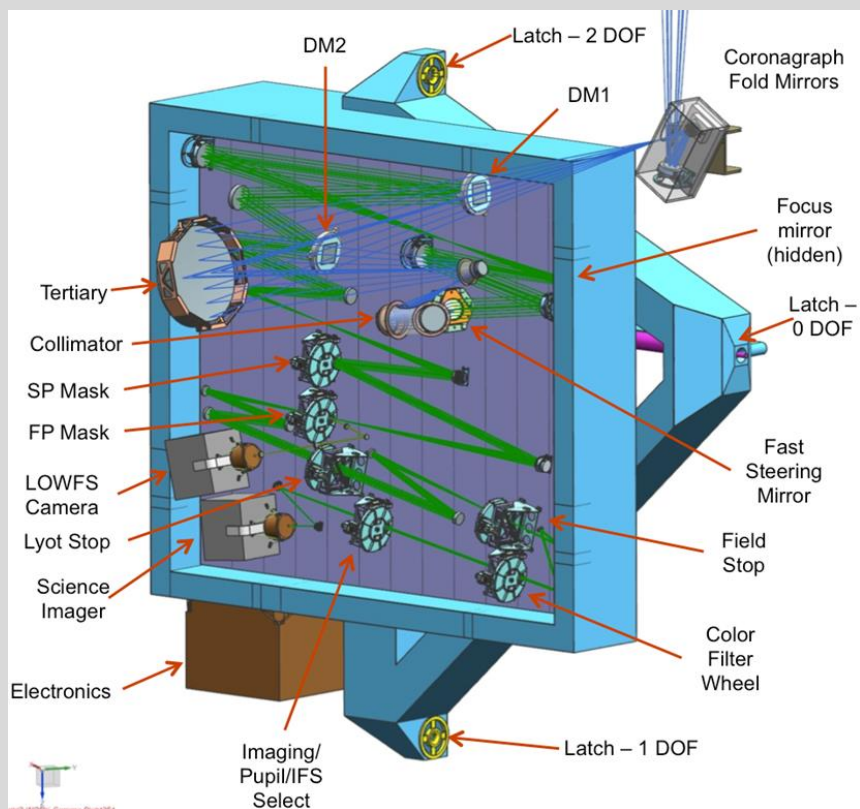
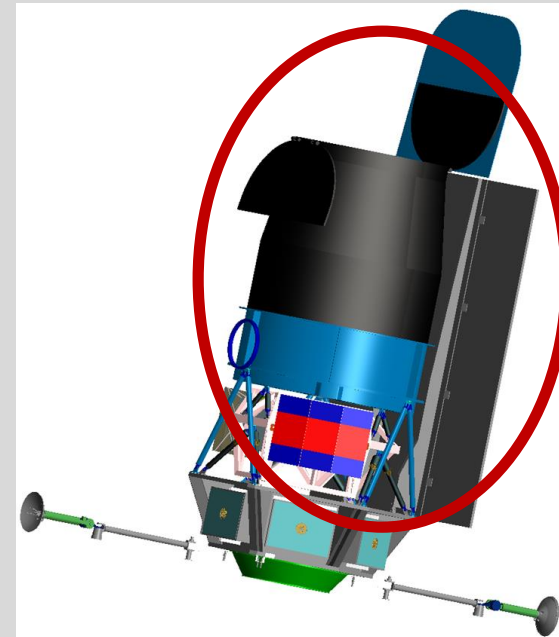


MPF

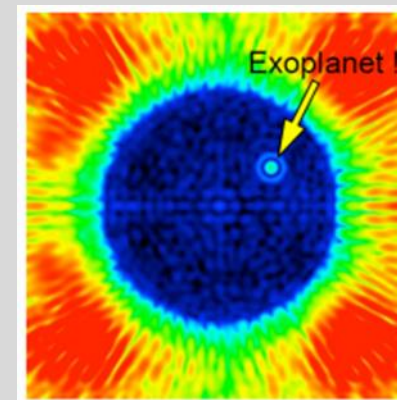


NIRSS

- Latest Design Reference Mission is **WFIRST– AFTA** (Astrophysics Focused Telescope Asset)
- AFTA is a repurposed **2.4 m** telescope from the US National Reconnaissance office (NRO)
- The AFTA telescope is already built, and sitting in a storage facility



- WFIRST-AFTA includes a coronagraph to image exoplanets
- This was not envisaged by the decadal survey
- Enabled by the 2.4 meter mirror

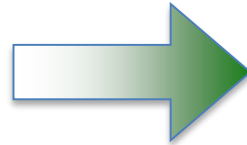
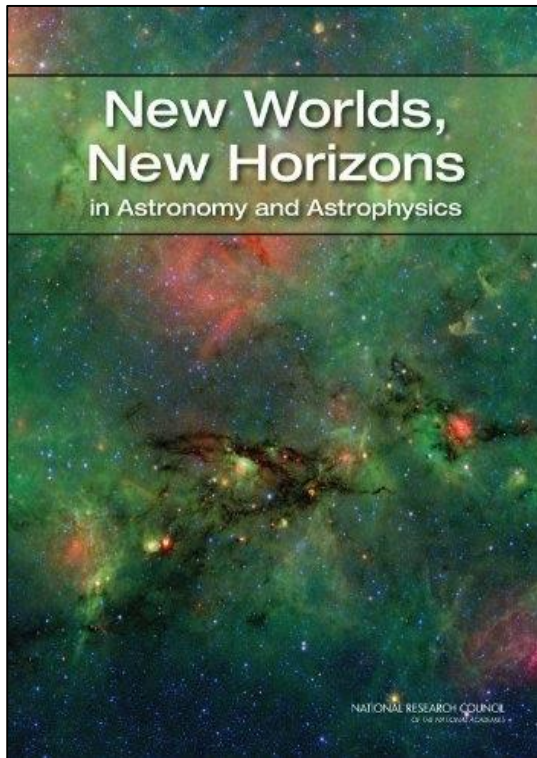


## Highest Priority

#1 Large-Scale Priority - Dark Energy, Exoplanets

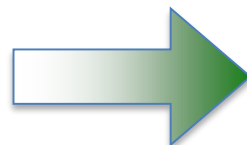
#1 Medium-Scale Priority - New Worlds Tech. Development  
(prepare for 2020s planet imaging mission)

## WFIRST covers many other NWNH science goals



### 5 Discovery Science Areas

- ID & Characterize Nearby Habitable Exoplanets ✓
- Time-Domain Astronomy ✓
- Astrometry ✓
- Epoch of Reionization ✓
- Gravitational Wave Astrometry

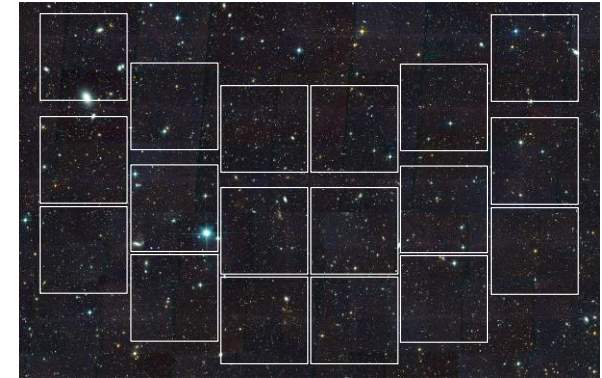


### 20 Key Science Questions

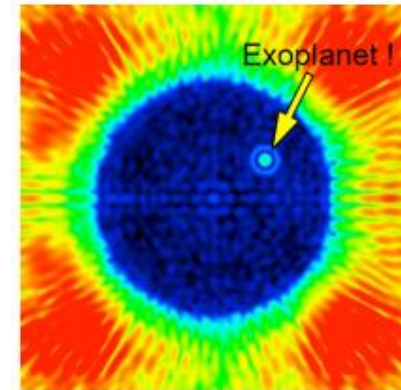
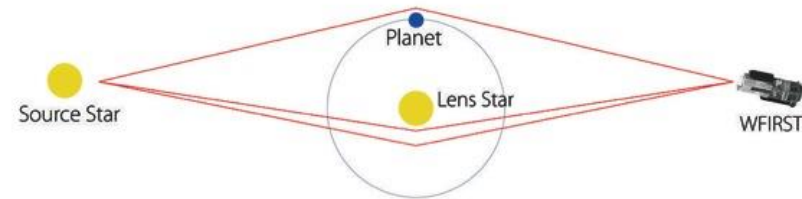
- Origins (**7/7 key areas**)
- Understanding the Cosmic Order (**6/10 key areas**)
- Frontiers of Knowledge (**3/4 key areas**)



- WFIRST is the highest ranked NWNH large space mission.
  - Determine the nature of the dark energy that is driving the current accelerating expansion of the universe
  - Perform statistical census of planetary systems through microlensing survey
  - Survey the NIR sky
  - Provide the community with a wide field telescope for pointed wide observations
  
- Coronagraph characterizes planets and disks, broadens science program and brings humanity closer to imaging Earths.
  
- WFIRST gives Hubble-quality and depth imaging over thousands of square degrees
  
- The WFIRST-AFTA Design Reference Mission has
  - 2.4 m telescope (already exists)
  - NIR instrument with 18 H4RG detectors
  - Baseline exoplanet coronagraph
  - 6 year lifetime



HST/ACS    HST/WFC3    JWST/NIRCAM



*complements  
Euclid*

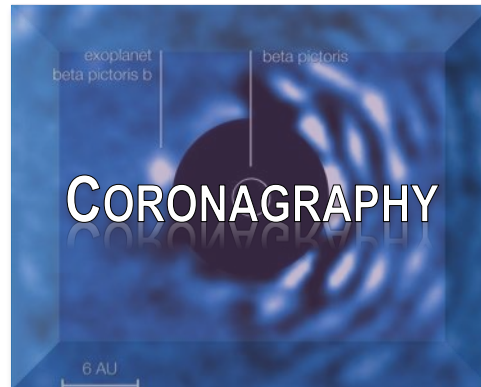
**BARYON ACOUSTIC  
OSCILLATIONS**

**GRAVITATIONAL  
LENSING**

**LEGACY SCIENCE  
WITH SURVEYS**

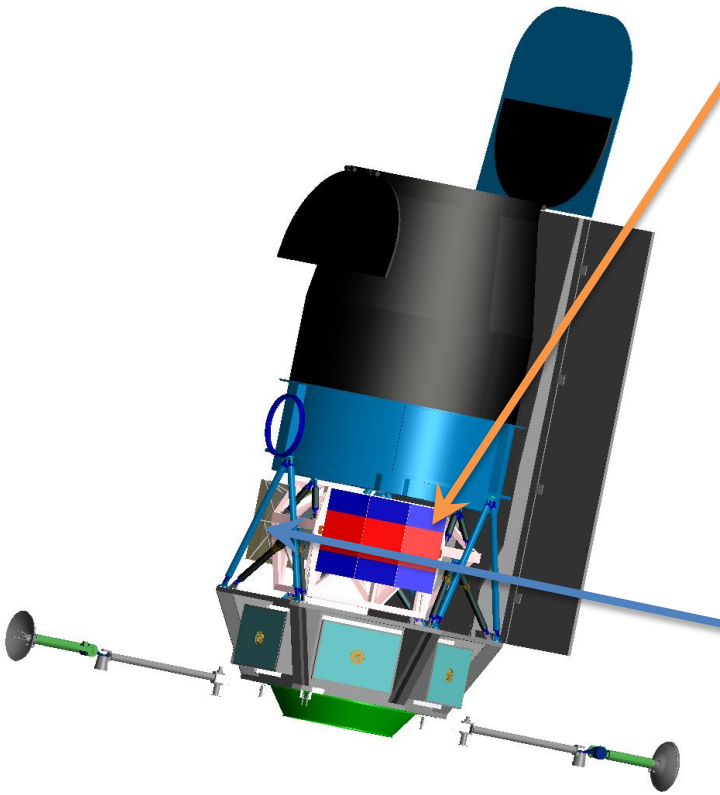
*complements  
LSST  
complements  
Kepler*

**SUPERNOVAE**



*continues  
Great  
Observatory  
legacy*

# WFIRST-AFTA Instruments



## Wide-Field Instrument

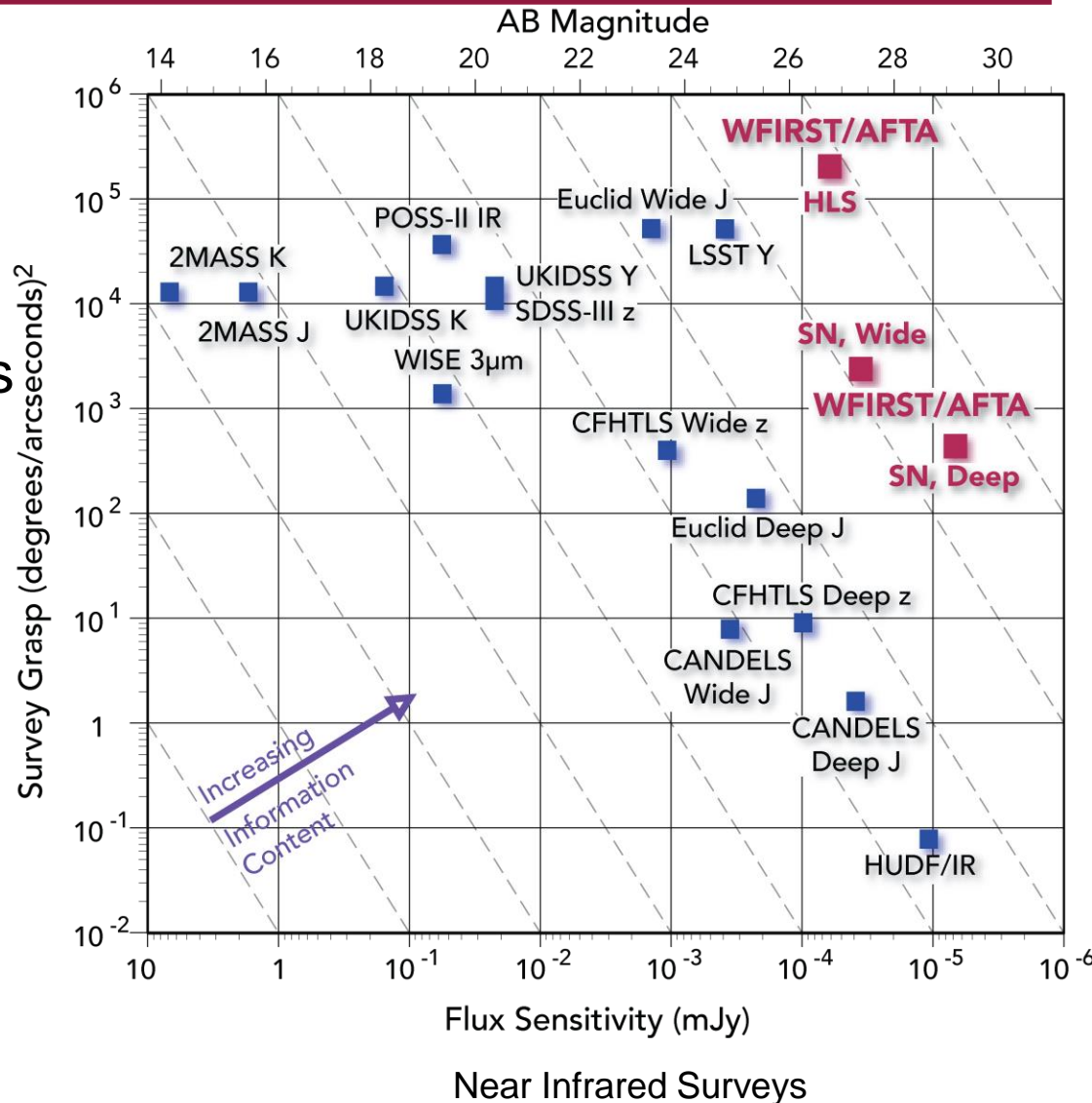
- *Imaging & spectroscopy over 1000s of sq. deg.*
- *Monitoring of SN and microlensing fields*
- $0.7 - 2.0 \mu\text{m}$  (imaging) &  $1.35-1.89 \mu\text{m}$  (spec.)
- $0.28 \text{ deg}^2$  FoV (100x JWST FoV)
- 18 H4RG detectors (288 Mpixels)
- 6 filter imaging, grism + IFU spectroscopy

## Coronagraph

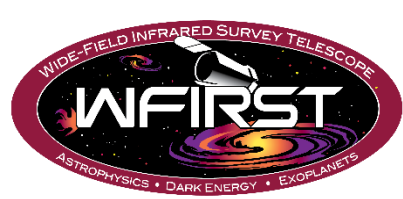
- *Image and spectra of exoplanets from super-Earths to giants*
- *Images of debris disks*
- $430 - 970 \text{ nm}$  (imaging) &  $600 - 970 \text{ nm}$  (spec.)
- Final contrast of  $10^{-9}$  or better
- Exoplanet images from 0.1 to 1.0 arcsec



- Multiple surveys:
  - High-Latitude Survey
    - Imaging, spectroscopy, supernova monitoring
  - Repeated Observations of Bulge Fields for microlensing
  - 25% Guest Observer Program
  - Coronagraph Observations
- Flexibility to choose optimal approach







# Capabilities



## WFI:

Imager **0.76-2.0 microns** 0.28° FoV, 0.11" pixel scale

Filters: z (0.76 - 0.98), Y (0.93-1.19), J (1.13-1.45), H(1.38-1.77),  
F184 (1.68-2.0), W149 (0.93-2.00)

Grism: **1.35-1.89 microns** 0.28° FoV, R=461λ, 0.11" pixel scale

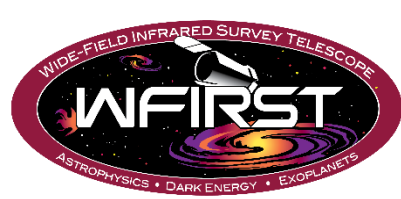
IFU: **0.6-2.0 microns** 3" & 6" FoV, R~100, 0.075" pixel scale

## Coronagraph:

Imager: **0.43-0.97 microns** 1.63" FoV (radius), 0.01" pixel scale, 1k x 1k EMCCD, 10<sup>-9</sup> final contrast, 100-200 mas inner working angle

IFS: **0.60-0.97 microns** 0.82" FoV (radius), R~70

Field of Regard: 54° - 126° 60% of sky



# Design Reference Mission Yields



## Attributes

## WFIRST-AFTA Yields

Imaging survey

J ~ 27 AB over 2200 sq deg

Slitless

This is in a 6 year mission with:

- 2 years High Latitude Survey (HLS) imaging and spectroscopy

Number

- ~6 months SN search and follow-up with IFU

Number

- ~1 year for coronagraph

Number

- ~1 year for microlensing planet search

Number

- ~1.5 years (25%) dedicated to Guest Observers (GO)

Number

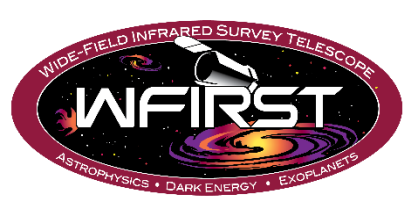
**An extended mission (10+ years) would consist of an expanded GO program**

Number of microlens exoplanets

2600

Number of imaged exoplanets

10s



# WFIRST-AFTA Status

- Significant WFIRST-AFTA funding added to the NASA budget by Congress for FY14 and FY15 for a total of \$106.5M.
- Funding is being used for pre-Phase A work to prepare for a rapid start and allow a shortened development time
  - Detector array development
  - Coronagraph technology
  - Science simulations and
  - Observatory design work
- ROSES "Preparatory Science" proposals funded at ~\$15 million
- NASA HQ charge for telescope coronagraph is "not driven by the fastest, cheapest implementation of the mission" but forward
- Community engagement: PAGs, conferences and outreach
  - Special sessions held at January & June 2014 and January 2015 AAS conferences
  - Wide-Field Infrared Surveys conference held in November 17-22, 2014 in Pasadena
- **Foreign contributions being actively pursued (Europe, Canada, Japan, Korea)**

Jason's best guess:  
2017 full start for  
2024 launch with a 10  
year mission

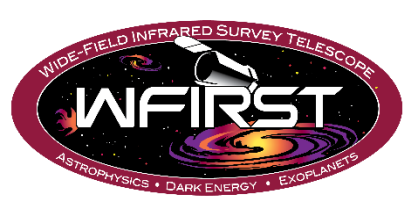




# Additional Material

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# WFIRST-AFTA SDT



## Co-Chairs

- David Spergel, Princeton University
- Neil Gehrels, NASA GSFC

## Members

- Charles Baltay, Yale University
- Dave Bennett, University of Notre Dame
- James Breckinridge, California Institute of Technology
- Megan Donahue, Michigan State University
- Alan Dressler, Carnegie Institution for Science
- Scott Gaudi, Ohio State University
- Tom Greene, NASA ARC
- Olivier Guyon, Steward Observatory
- Chris Hirata, Ohio State University
- Jason Kalirai, Space Telescope Science Institute
- Jeremy Kasdin, Princeton University
- Bruce Macintosh, Stanford University
- Warren Moos, Johns Hopkins University

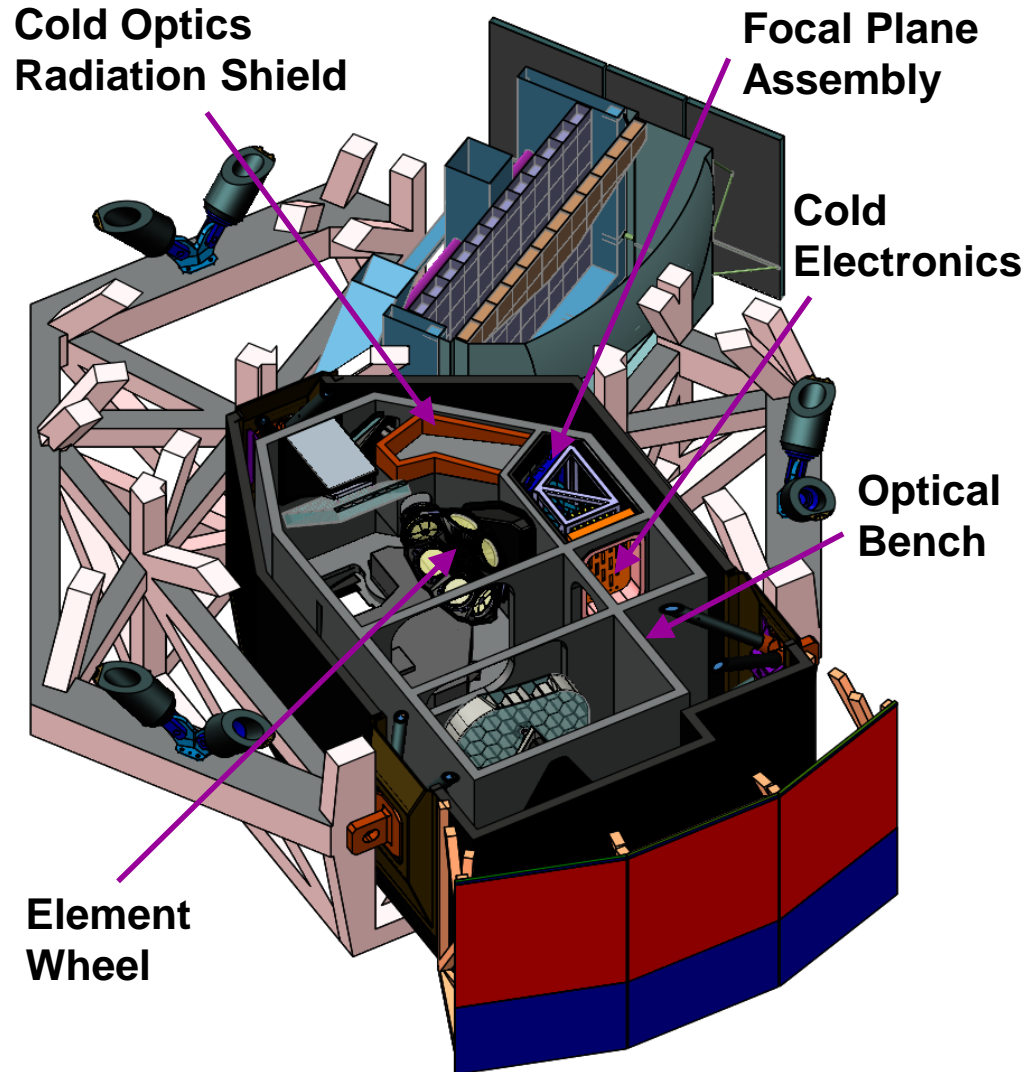
- Saul Perlmutter, University of California Berkeley
- Marc Postman, Space Telescope Science Institute
- Bernie Rauscher, NASA GSFC
- Jason Rhodes, NASA JPL
- Yun Wang, IPAC/Cal Tech
- David Weinberg, Ohio State University

## Ex Officio

- Dominic Benford, NASA HQ
- Mike Hudson, Canadian Space Agency
- Woong-Seob Jeong, Korea Astronomy and Space Science Institute
- Yannick Mellier, European Space Agency
- Wes Traub, NASA JPL
- Toru Yamada, Japan Aerospace Exploration Agency

## Key Features

- Wide field channel instrument for both imaging and spectroscopy
  - 3 mirrors, 1 powered
  - 18 4k x 4k HgCdTe detectors cover 0.76 - 2.0  $\mu\text{m}$
  - 0.11 arc-sec plate scale
  - Single element wheel for filters and grism
  - Grism used for GRS survey covers 1.35 – 1.89  $\mu\text{m}$  with  $R = 461\lambda$  (~620 – 870)
- IFU channel for SNe spectra, single HgCdTe detector covers 0.6 – 2.0  $\mu\text{m}$  with  $R$  between 80-120







# Wide Field Channel Description & Modes



- The wide field channel's only routinely moving part is the element wheel (EW)
- 8 positions: 6 filters, blank, grism (galaxy redshift survey)
- Table shows how measurement modes and observations align

#	Min $\lambda$ (mm)	Max $\lambda$ (mm)	R $\lambda$	SN Detect		SN Spec	HLS		Microlensing		Available for GO
				Shallow	Med/Deep		Image	Spec	Monitor	Color	
Z087	0.760	0.977	4.0							2X Daily	All
Y106	0.927	1.192	4.0	X			Photo-z				
J129	1.131	1.454	4.0	X	X						
H158	1.380	1.774	4.0		X		Photo-z & Shapes				
F184	1.683	2.000	5.81								
W149	0.927	2.000	1.442						15 min cadence		
GRS	1.35	1.95	461*					X			
IFU	0.600	2.000	75			X					

## Key Features

- **Telescope:** 2.4 m aperture primary mirror
- **Instruments**
  - Wide Field Imager/Spectrometer & Integral Field Unit
  - Internal Coronagraph with Integral Field Spectrometer
- **Overall Dry Mass:** 4059 kg (CBE)
- **Structure:** high stiffness composites; modular packaging for avionics
- **GN&C/Propulsion:** inertial pointing, 3-axis stabilized, mono-prop system for stationkeeping & momentum unloading
- **Data Downlink Rate:** Continuous 600 Mbps Ka-band to dedicated ground station
- **C&DH:** low rate bus for housekeeping and spacecraft control, high speed bus for science data
- **Power:** ~2400 W average power (CBE)
- **GEO orbit**
- **Launch Vehicle:** Delta IV Heavy
- **GSFC:** leads mission, wide field instrument, spacecraft
- **JPL:** leads telescope, coronagraph

