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The WEAVE MOS and IFU spectrograph will produce millions of spectra

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allow end users to access the data reliably and easily for their pipelines CPS and APS. The underlying database structure specific research projects. We envisage here a Science Archive and performance will be guaranteed with a redundant composed of (i) a repository including every WEAVE data network of disks and CPUs, and robust open-source software product and (ii) a database which intelligently indexes the for which troubleshooting support is available for a high content for subset selection and data retrieval. A GUI will fraction of the time. grant public or privileged access to the archive, giving the

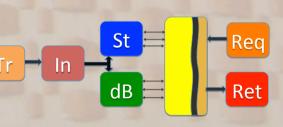
possibility of data mining, selection and retrieval of the data from different configurations of the WEAVE instrument, i.e. This big-data product needs a robust high-performance tool to the spectra or IFU meta-images produced by the WEAVE data

WAS inside: the data flow in its subsystems.

System) through a suitable and secure research. transportation system [Tr] which may A layer of organic processes will have

Data arrive to WAS (WEAVE Archive order suitable for any astrophysical

make use of VPNs or more sophisticated access to storage and databases and will techniques like GlusterFS in order to allow other (external) processes to send simplify the management on the requests (Req) and get results (Ret). producers' side (OCS, CPS and APS). The Among plug-in programs we foresee a web subsequent ingestion process [In] drives based common GUI interface, a prerelease all FITS file and tables into a redundant organizer used before public releases and repository storage [St] and populates the the possibility to develop ad-hoc analysis series of databases [dB] which will which need to access the whole amount of organize data and metadata in an indexed data.



Architecture down to the foundations.

The WEAVE archive architecture will sit atop of the On the other hand, the GUI will be built on web scalability.

the WEAVE standard. The ingestion process continues scientist data. and sends the files into the database.

latest technologies in work load distribution and data technology for the maximum flexibility to the scientist management in order to minimize the points of failure community. Web technologies such as HTML5 and and at the same time, increasing its reliability and Ajax will be part of the design to increase of the user experience on the browser. Initially, the web front end The archive will be able to ingest (according to the will be used as a search engine to display the target initial estimations) more than 50GB of data for and IFU features. Later on, the GUI could also give around 10,000 targets on daily basis. The raw FITS access to the database Analysis adding new features files and their sub-products, will be stored on file- such as Data Mining, Graphs and Machine Learning system with a previous quality check that conforms to which are an extra bonus to capacity of the existing

nds for WHT Enhanced Area Velocity Explorer. The name also reflects the challenge of positioning correctly in a 2-dimensional plane a large number of fibres. to happens to be a long-standing traditional handloraft on the island of La Palmai The background image comes from the Silk Museum in El Paso, on the

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Le fil rouge from telescope to users.

Data coming from the WEAVE metadata at any level are sent to the (Advanced Processing Software) project board. turns the spectra into physical observable quantities. All the

instrument are managed and WAS archive for storage and packaged by the Observing Control organization. Users can access the Software (OCS) and then sent to the wealth of data (from raw to final archive (WAS) and to the Core astrophysical quantities and IFU Processing Software (CPS) in order maps) through the Web interface to for spectra to be extracted and the archive. Regular, public releases calibrated. In the next phase the APS of data will be overseen by the WEAVE

OCS

CPS

APS

Users' GUI access to the archive.

A fairly complex gui is required to display phase(target, area, IFU, search for data obtained with the observing log) which will offer the different observing modes (fibers, second filtering selection before the IFUs). The process involves first a retrieval of data or frames. search by e.g. area or object, then a

