Terra Hunting Experiment HARPS-3 @ INT

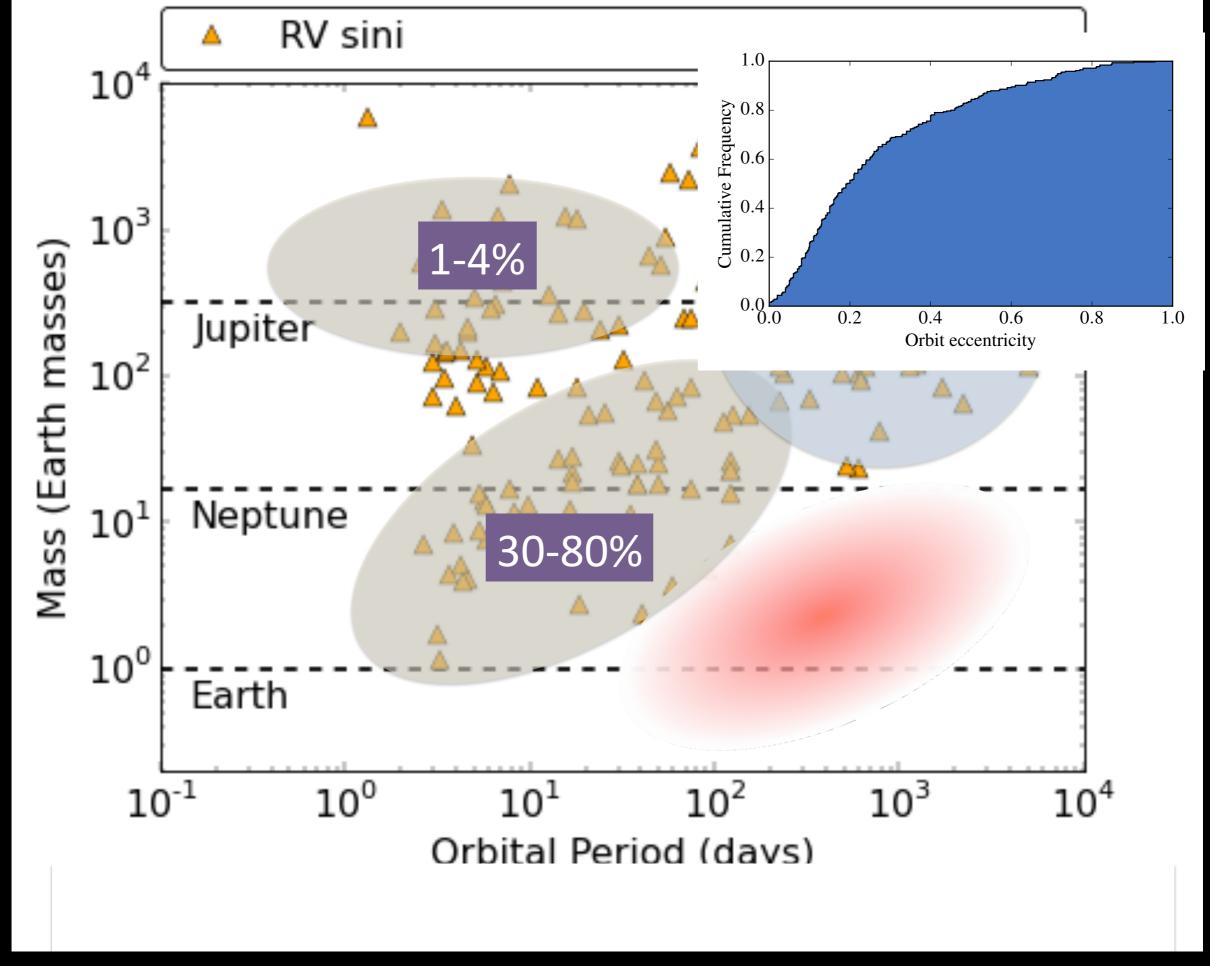
Cambridge, Exeter, NOVA, IAC, Stockholm, Geneva, U. Queen Belfast, Oxford, Princeton, Flatiron

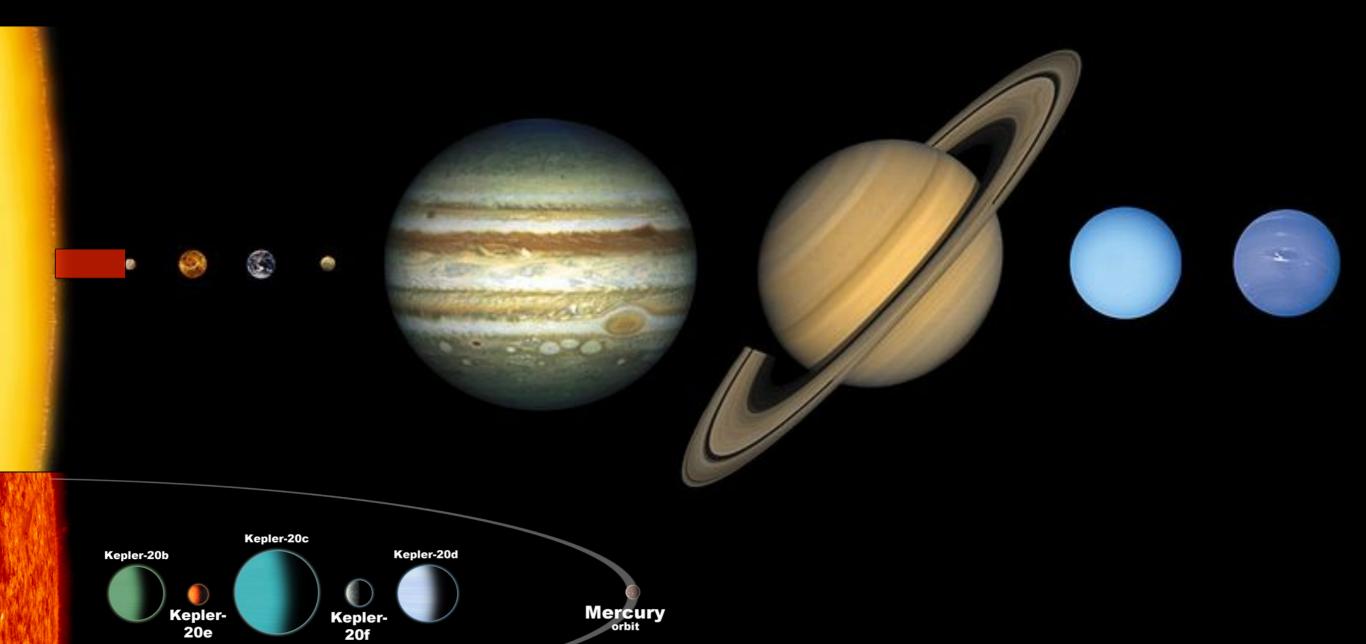
- How diverse are planetary systems?

 Our Solar system in perspectives

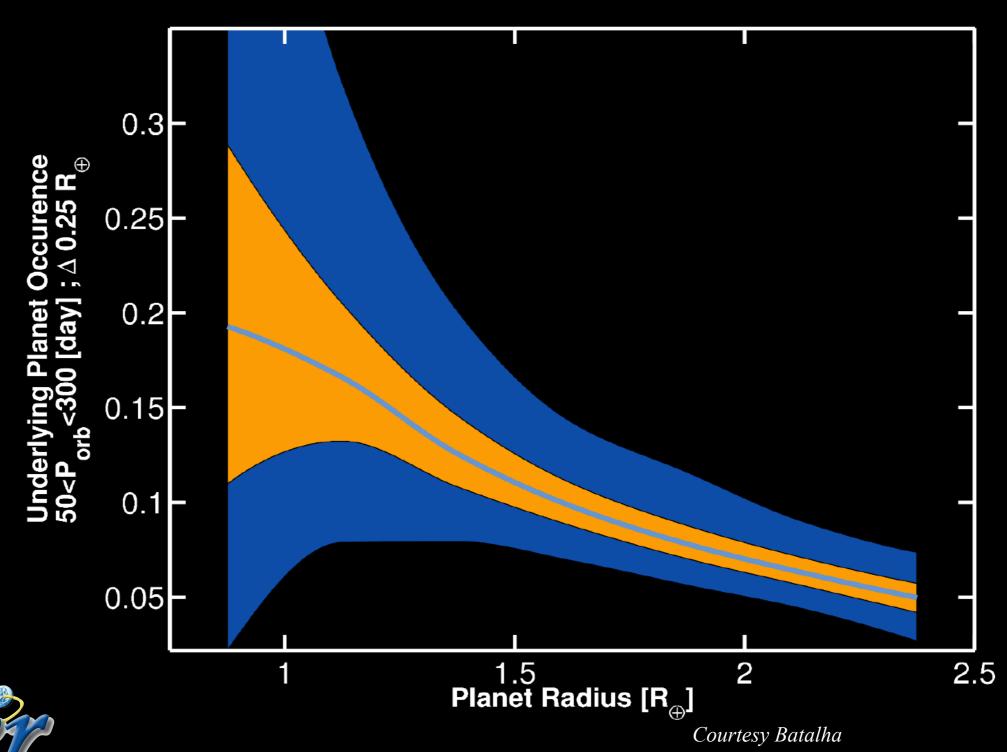
 Towards an universal model for planetary systems
- Is there life elsewhere?

 Understanding the origin, prevalence and nature of life in the Universe.

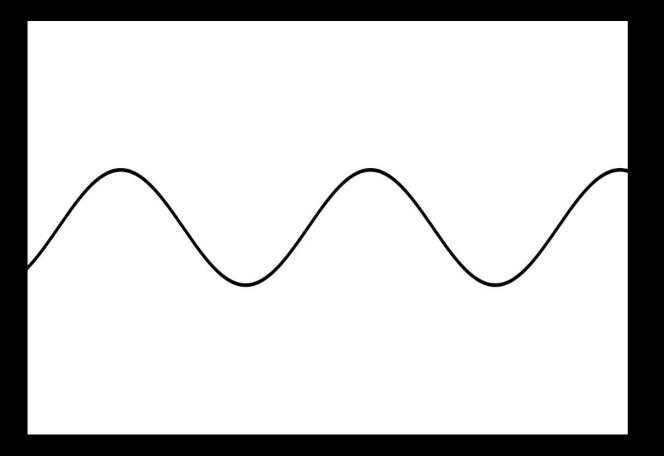


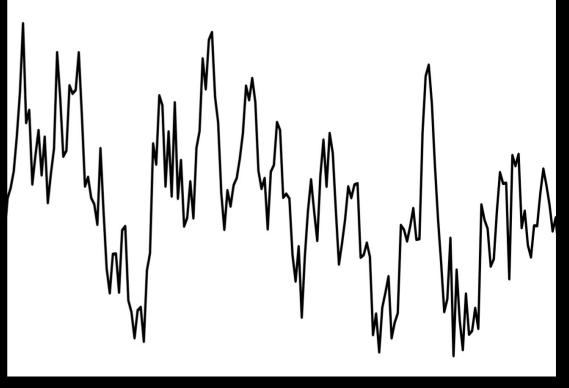


Kepler, Eta Earth measurement...





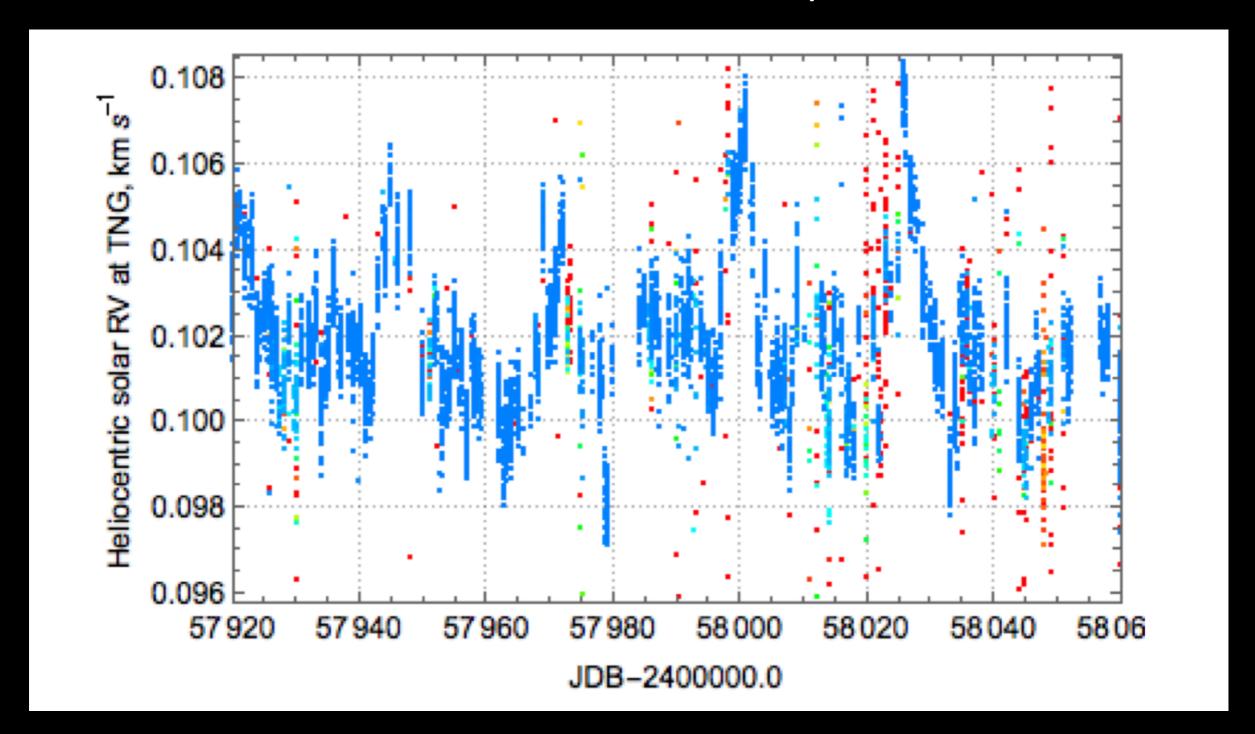




Planet RV

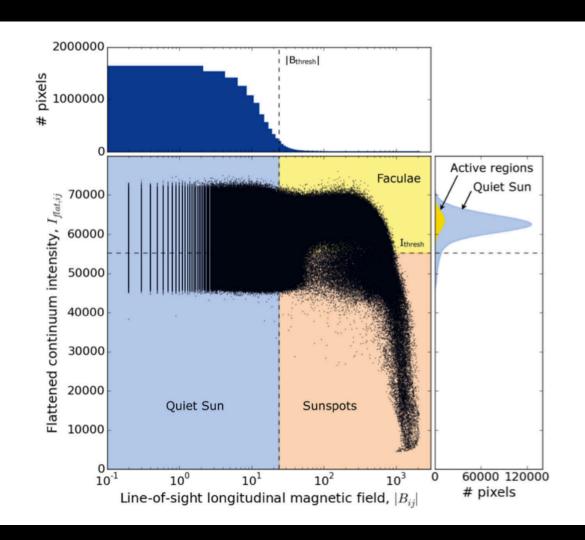
Stellar RV

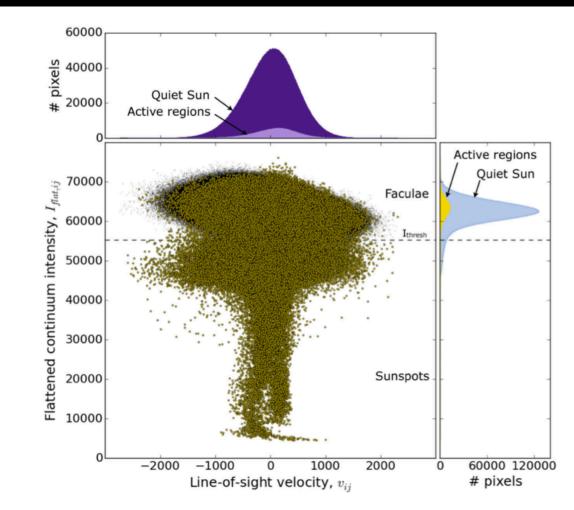
HARPS-N solar telescope



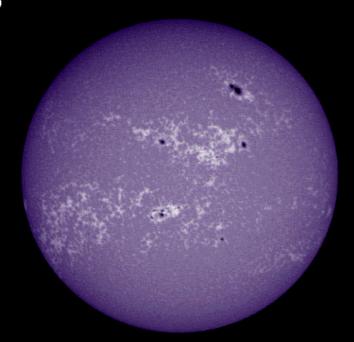
Cameron et al. 2019

Haywood et al. 2017

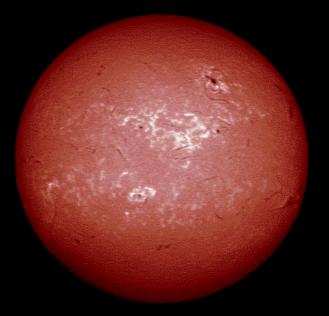




Ca II 3934 Å

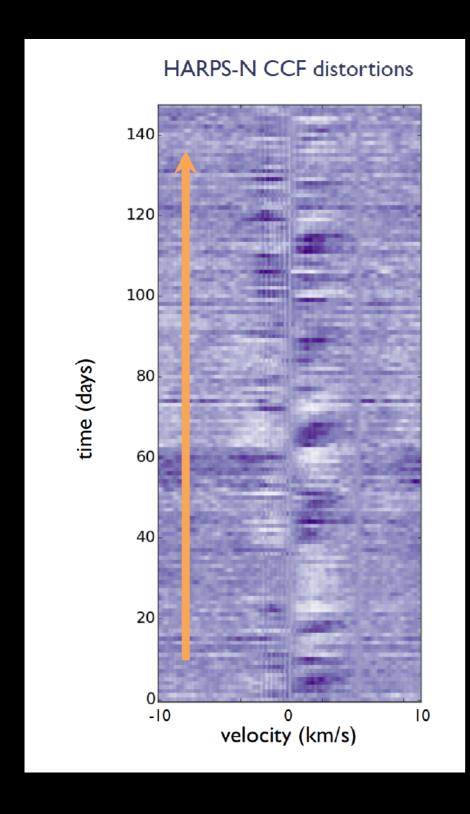


HI 6563 Å BBSO

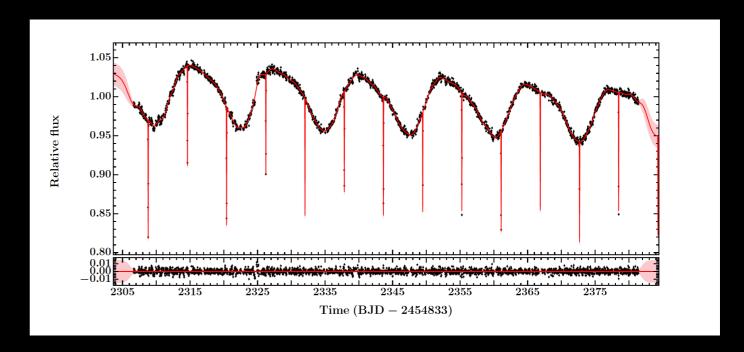


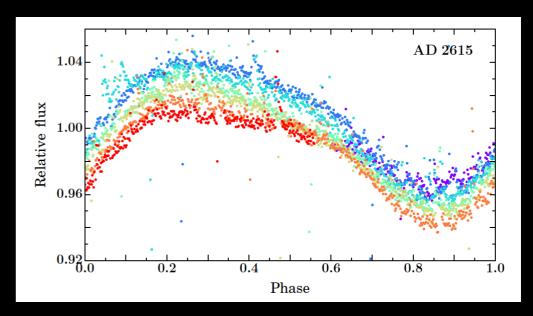
1991 May 13 1991 May 13

Our Sun as a star



A eclipsing binary star in Praesepe





Gillen et al 2017

Haywood et al 2018

Science Goal:

Detecting 1-4 Earth mass, with period from 50d to 300d on star brighter than 8.5th (RV amplitude 10-100cm/s)

Accuracy Requirements:

- 48 hours calibration <10 cm/s (differential reference)
- Yearly "systematics" 20 cm/s (absolute reference)
- 30 cm/s photon noise per RV measurements $(SN_{r.e}=300)$

Survey model

(essentially based on R. Hall et al. 2019)

- Season Long uninterrupted "weather permit" nightly series of RV
- (never less than 4h/night, every day*)
- Season is between 6-9 months
- 150 RV/year x (up to 10 years)
- Stellar jitter "residual pink noise" less than 1m/s (peak-to-peak)

^{*} Excl. Eng. nights

Uninterrupted nightly series is ING magic wand!

